

Near-Term Lunar Surface Gravimetry Science Opportunities

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Key Points

- We're developing a (nearly) new type of geophysical instrument: the VEGA space gravimeter
- VEGA enables a (nearly) new type of Lunar exploration
- VEGA's development is being supported by the Canadian Space Agency
- The instrument is a candidate for future CSA contribution to international planetary missions
- We are seeking exploration mission partners
 - Who can help arrange rides on Lunar rover missions



Relevance to This Meeting's Themes

Lunar Resources:

- Lunar surface gravimetry surveying could help identify ISRU ore-bodies
- E.g., water ice deposits
- Just as gravimetry surveys do here on Earth

The Dynamic Moon:

- Lunar surface gravimetry surveying can help delineate near-surface geological structures
- Some of which illustrate dynamic processes in the Moon's formation and evolution history
- E.g., lava tubes (volcanism)



Gedex, Terrestrial and Space Geophysics

- Located near Toronto, Canada
- Core business:
 - Airborne geophysics instrument development and exploration company
 - Developing world's most sensitive airborne gravity gradiometer
- Senior technical staff led development of Canada's first microsat missions
- Now also developing space geophysics instruments
- Bridging between terrestrial and space exploration communities









LSGG Study for Canadian Space Agency

- Lunar Surface Gravity Geophysics Science Definition Study
 - Funded by CSA's Exploration Core Program
 - March 2014 to March 2016
- Motivated in part by Gedex's development (underway) of a new space gravimeter instrument, VEGA
- Identify science that can be done on the Lunar surface with gravimeter and gravity gradiometer instruments
- Derive instrument requirements
- VEGA breadboard testing



LSGG Science Team Members / Co-authors

- David Hatch: Chief Geophysicist, Gedex
- Rebecca Ghent: Associate Professor, University of Toronto,
 Department of Earth Sciences
- Sabine Stanley: Professor, University of Toronto, Department of Physics
- Natasha Urbancic: Ph.D. candidate, University of British Columbia, Department of Earth, Ocean and Atmospheric Sciences
- Marie-Claude Williamson: **Geological Survey of Canada**
- W. Brent Garry: NASA Goddard Space Flight Center
- Manik Talwani: Schlumberger Chair, Rice University, Department of Earth Science



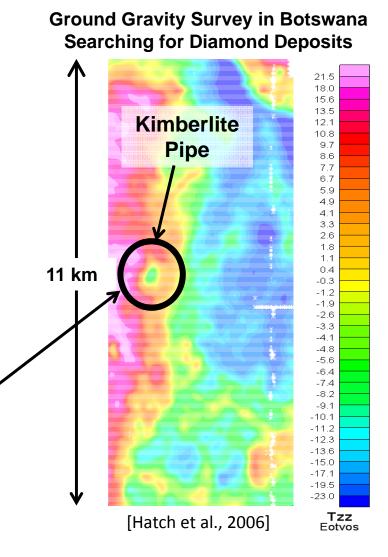
Lunar Surface Gravimetry

- Q: What could be detected by a gravimeter on the Moon's surface?
- A: The same sort of things as on Earth:
 - Subsurface density variations
 - Indicative of any local geology possessing anomalous density
- Example interesting Lunar subsurface structures that could cause such signals:
 - Stratigraphy: layers of varying density and thickness
 - Megaregolith: variations in depth
 - Bedrock: "topography" of top surface
 - Faults
 - Buried craters
 - Magmatic intrusions
 - Lave tubes
 - etc.



Gravimetry 101

- A standard geophysical technique used in exploring for resources on Earth
- Typically make measurements using a gravimeter at stations along a set of traverse lines covering a survey area
- Gravimeters respond to subsurface density variations
- Unlike with some other geophysical techniques, nothing blocks gravity
- The data reflect some of the compositional and structural variations in geological formations
- Hence some aspects of subsurface structure can be recovered from gravity maps



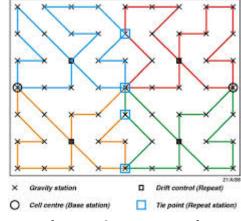


Existing Terrestrial Gravimeters

- The best field terrestrial gravimeters have drift-compensated repeatability within ~ 5-10 μGal (i.e., 5-10 milliG) over the course of a day
 - Relative gravimeters (not absolute)
 - Poor absolute accuracy
 - Arbitrary initial bias value
 - Bias drifts on the order of 1000 μGal/day
 - Bias drift rate varies significantly over 10-100 days
 - Very good *relative* accuracy achievable during surveys, but only by incorporating *daily loop-closing* in survey design
- Not suitable for the Lunar gravimetry application:
 - Can only function within +/- a few % of 1 G
 - Signal saturates outside that range
 - Daily loop-closing for a reasonably-sized survey would require a very fast rover



Scintrex CG-5 gravimeter



[Murray & Tracey, AGSO]



VEGA Instrument (Vector Gravimeter for Asteroids)

• Developed by Gedex

- Innovative (patented) design
- Spun off from airborne gravity gradiometer system technology

Measures absolute gravity vector

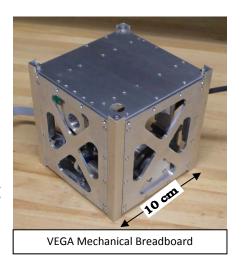
- Vector: Allows arbitrary lander orientation
- Absolute: Doesn't require looping-back to calibrate out drift — compatible with slow rovers

On the Moon:

- Better than 1 mGal (1 μG) absolute accuracy
- **Size:** 10x10x15 cm
- **Mass:** < 1.5 kg

Status:

- Canadian Space Agency is co-funding development
- Preliminary design done, breadboard testing and detailed design underway
- Test flight in LEO planned in 2017
- Seeking lunar rover (and asteroid lander/rover) missions to fly on!



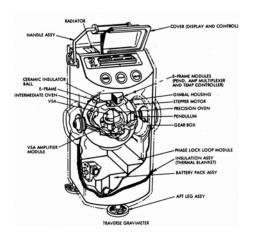


Past Lunar Surface Gravimetry

- Only one surface gravimetry survey has been conducted off-Earth, to date:
- Apollo 17 Traverse Gravimeter Experiment
 - Double-string accelerometer instrument
 - Built by Bosch-ARMA, then modified by MIT
- Measurements made at 9 stations during 3 EVAs by Eugene Cernan and Harrison Schmitt, spread across the 10 km width of the Taurus-Littrow valley floor
- Data analyzed and interpreted by Manik Talwani (P.I.)
- Results:
 - Valley floor composition determined to be basalt
 - Thickness of basalt layer determined (1 km)



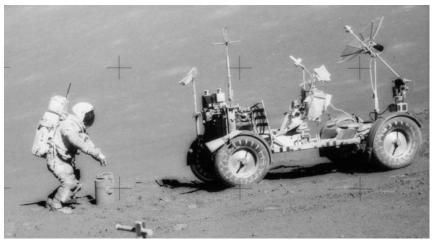
Traverse Gravimeter Experiment Taurus-Littrow Valley Survey





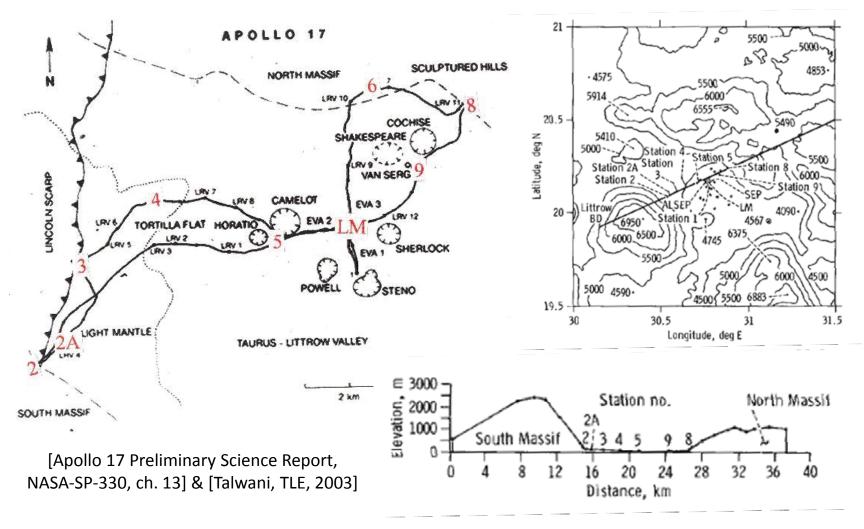






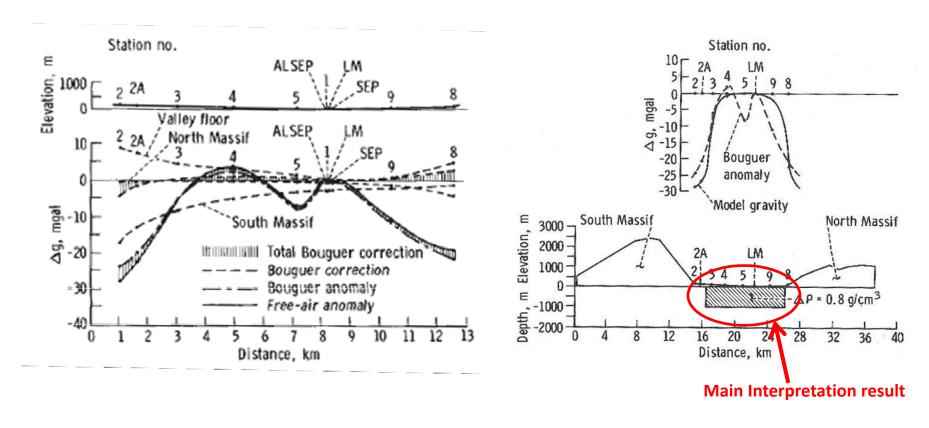


Taurus-Littrow Valley Survey Traverse Geography and Topography





Taurus-Littrow Valley Survey Data Processing and Interpretation



[Talwani et al., Apollo 17 Preliminary Science Report, NASA-SP-330, ch. 13] & [Talwani, TLE, 2003]



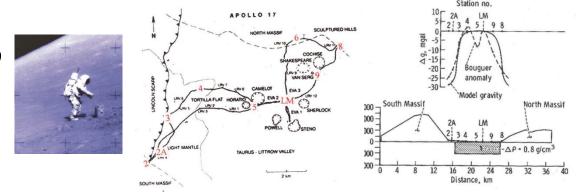
VEGA vis-à-vis TGE Gravimeter

- In the current era of Lunar exploration, VEGA provides an essentially new capability for Lunar geophysical exploration
- New for this era --- TGE was Then, VEGA is Now
 - TGE was a "one-off" instrument, long out of production
 - VEGA currently in development, soon ready for use in missions
- VEGA accuracy expected to be better than TGE accuracy
 - Opens up more geophysics applications
- VEGA is small enough to be carried on almost all Lunar rovers

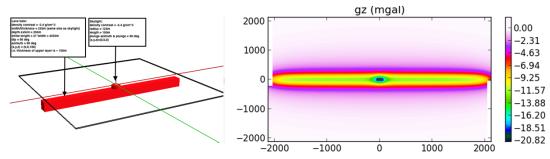


LSGG Science Applications Being Studied

(1) Follow up Apollo
 17 Lunar Traverse
 Gravimeter survey



- (2) Survey potential lava tubes from the surface
- (3) Survey for Lunar ice deposits in polar regions







Mission Opportunity #1: Part-Time Scientists GLXP Lunar Rover













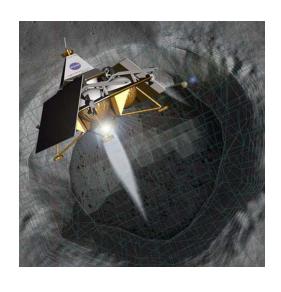
- German team, supported by Audi
- Planning to land 2 rovers near Apollo 17 landing site by end of 2017

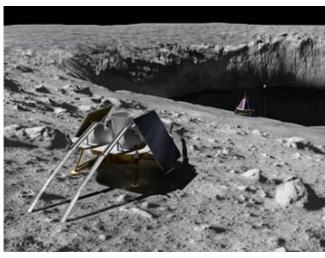
Gravimetry Survey Opportunity: Follow-Up Taurus-Littrow Gravimetry Survey

- Apollo 17 survey was quite limited in resolution
 - Due to EVA time limits
 - Only 9 gravimetry stations
 - Was able to "image" a large, deep geophysical target
- A follow-up survey could discern smaller, shallower features
- Rationale:
 - Extensive geological studies have been done of this area since 1972
 - Due in part to Harrison Schmitt's role in the mission, and his subsequent engagement as a geologist
 in studying results from Apollo 17 and subsequent Lunar missions
 - Those geological studies could guide and target next round of geophysical surveying, aiming to investigate/test current geological understanding of the area
 - Apollo 17 gravimetry survey has some intriguing features that need further data to investigate
- Gedex discussing with PTS the possibility of one of their rovers carrying a VEGA instrument, conducting a gravimetry survey around central part of Apollo 17 survey area
- Mission Concept:
 - CSA funds the VEGA instrument
 - Some combination of CSA and NASA funds the ride
 - Joint Canadian/US science team, based on LSGG team



Mission Opportunity #2: Astrobotic GLXP Lunar Rover

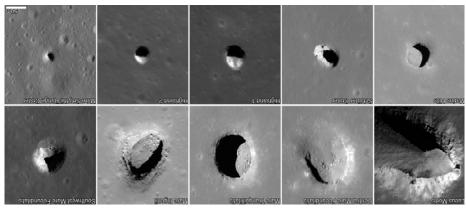






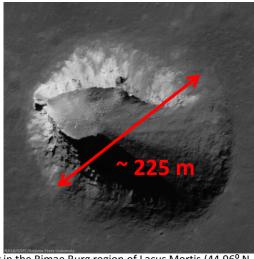
- Pittsburgh company, led by CMU's Red Whittaker
- Planning to land Griffin rover near a pit crater in Lacus Mortis by end of 2017
- Selling payload space commercially:
 - \$1.8M/kg on lander
 - \$2M/kg on rover

Gravimetry Survey Opportunity: Explore A Potential Lava Tube From Above



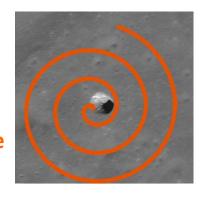
[Wagner & Robinson, (2014), NESF]

- Numerous pit craters have been found on the Moon
- Some Lunar pit craters appear to be skylights into subsurface void spaces
 - Localized voids, e.g., melt ponds?
 - Lava tubes?
- Lacus Mortis pit crater
 - Gravimetry survey around this could investigate presence/nature of subsurface void-space
 - Sooner/cheaper than other lava tube exploration concepts



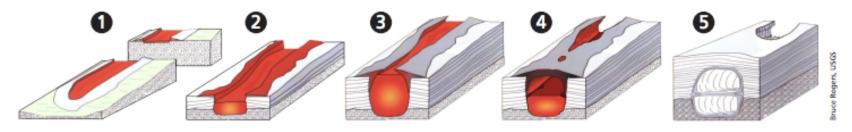
Pit crater in the Rimae Burg region of Lacus Mortis (44.96° N, 25.62° E), from LROC Narrow Angle Camera (NAC) observation M126759036L, orbit 3814, April 24, 2010; 49.4 angle of incidence, resolution 0.5 meters from 45.56 km [NASA/GSFC/Arizona State University, via http://lunarnetworks.blogspot.ca]

E.g., Spiral Survey Traverse





Possible Lava Tube Skylight?



[Bruce Rogers, USGS]



Lava tube skylight on Kīlauea, Hawaii, from http://commons.wikimedia.org/wiki/Kilauea

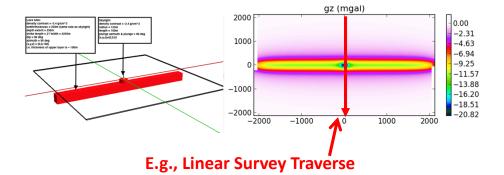


Valentine Cave in Lava Beds National Monument, California, from . http://en.wikipedia.org/wiki/Lava_tube

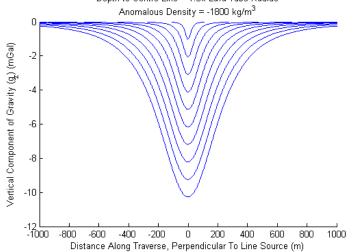


Surface Gravimetric Surveying Above a Lunar Lava Tube

- Gedex discussing with
 Astrobotic the possibility of
 their rover carrying a VEGA
 instrument, conducting a
 gravimetry survey around the
 pit crater --- could determine
 presence/size/shape of any
 associated subsurface void
- Mission Concept:
 - CSA funds the VEGA instrument
 - Some combination of CSA and NASA funds the ride
 - Joint Canadian/US science team, based on LSGG team



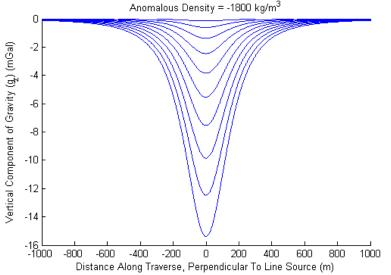
Vertical Component of Gravity From a Straight, Level Lava Tube of Circular Cross-Section
Lava Tube Radii From 20m to 200m in 20m increments
Depth to Centre-Line = 1.5x Lava Tube Radius





Traverse Results Indicative of Tube Diameter, Depth

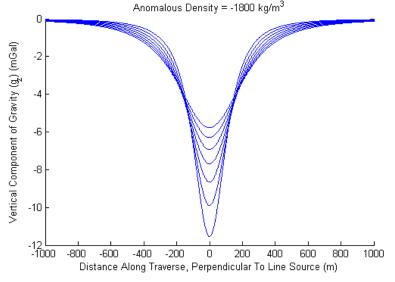
Vertical Component of Gravity From a Straight, Level Lava Tube of Circular Cross-Section Depth to Centre-Line of Lava Tube = 200m Lava Tube Radius From 20m to 200m in 20m_increments



Vertical Component of Gravity From a Straight, Level Lava Tube of Circular Cross-Section

Lava Tube Radius = 150m

Depth to Centre-Line From 150m to 300m in 25m increments



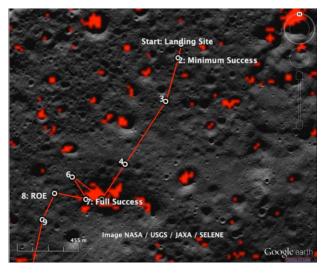
[Carroll et al., LPSC, 2015]

Mission Opportunity #3: NASA Resource Prospector Mission Rover

- 2020 launch, NASA
- Rover to Lunar polar region, to search for ice
- Ice deposits could cause density anomalies, make a signal detectable by VEGA
 - Solid ice specific gravity = 1
 - Dry regolith s.g. perhaps 2.5
 - Regolith with ice in voids s.g. > 2.5, perhaps3.0
- Could be an affordable Canadian contribution to RPM (<< \$10M)

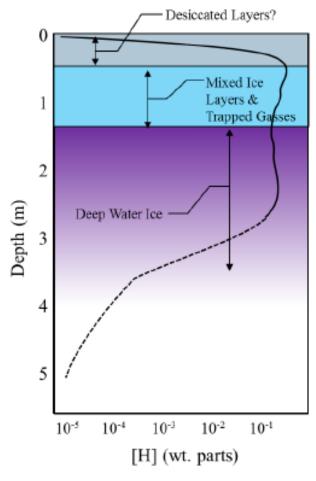








Nature of Lunar Ice Deposits Not Yet Known



[Colaprete et al., LEAG 2013]



[David Hollister, http://hop41.deviantart.com/art/Lunar-Ice-174785137]

- We don't *know* the nature of Lunar ice deposits
 --- depth, concentration, etc.
- Instruments used to date can only "see" down to ~ 1 m depth
- Nothing blocks gravity --- a gravimetry survey could detect signatures from deeper deposits



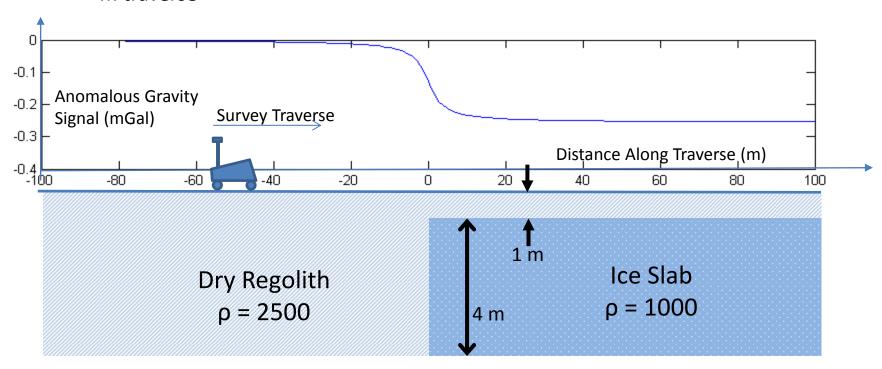
Gravimetric Detectability of Ice Deposits

- Assume large, horizontal deposit of uniform thickness t
- Anomalous gravity signal upper-bounded by Bouguer correction:
 - $-\Delta g_B = 2\pi G^* \Delta \rho^* t = 0.042^* \Delta \rho^* t$ [µGal]
 - Finite extent of deposit can reduce the peak signal
 - If there are defined edges they will also produce "edge effects" which can enhance detectability over a traverse
 - Sharpness of edge depends on the deposit's mean depth
- Two examples:
 - *Solid ice slab:* $\Delta \rho = 1000-2500 = -1500 \text{ kg/m}^3$
 - **Regolith with ice in voids:** $\Delta \rho$ might be as high as 500 kg/m³ (assuming 50% porosity)



Gravimetric Detectability of Ice Slab

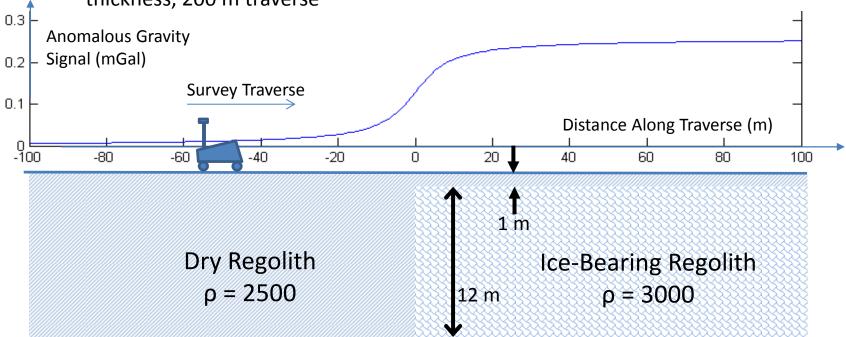
- For **solid ice**, $\Delta \rho = 1000-2500 = -1500 \text{ kg/m}^3$
 - Gravity low over the deposit
 - For t = 1 m, Δg_B = -62 μGal
 - To reach $\Delta g_B = -0.25$ mGal, t = 4 m
 - Plot (below) assumes 1 m depth to top of ice, 4 m ice slab thickness, 200 m traverse





Gravimetric Detectability of Ice In Regolith

- For *regolith with ice in voids*, $\Delta \rho$ might be as high as 500 kg/m³ (assuming 50% porosity)
 - Gravity high over the deposit
 - For t = 1 m, Δg_B = +21 μGal
 - To reach Δg_B = +0.25 mGal, t = 12 m
 - Plot (below) assumes 1 m depth to top of ice-bearing deposit, 12 m deposit thickness, 200 m traverse





Gravimetry Survey Opportunity: Look for Gravity Signature of Ice on RPM

Mission Concept:

- CSA funds the VEGA instrument
- CSA provides this as a Canadian contribution to the Resource Prospector Mission
- Joint Canadian/US science team, based on LSGG team
- RPM adds gravimetry measurements to its mission operations
 - Make a gravity measurement whenever the rover is stopped
 - Preferably (but not necessarily) when other rover payloads are quiescent
 - Baseline: analyze gravity data post-mission
 - Opportunistically: analyze data during RPM's traverse, use that to help guide drilling decisions



Conclusion

- VEGA gravimeter will soon be ready for use on Lunar rover missions
- VEGA is small enough to fit on even quite-small Lunar rovers
- Any Lunar rover planning to traverse much more than 10 m could potentially detect the signature of subsurface density variations
- There are numerous scientifically interesting subsurface structures to "look" for
- Including for three potential near-term Lunar rover missions

SEEKING EXPLORATION MISSION PARTNERS