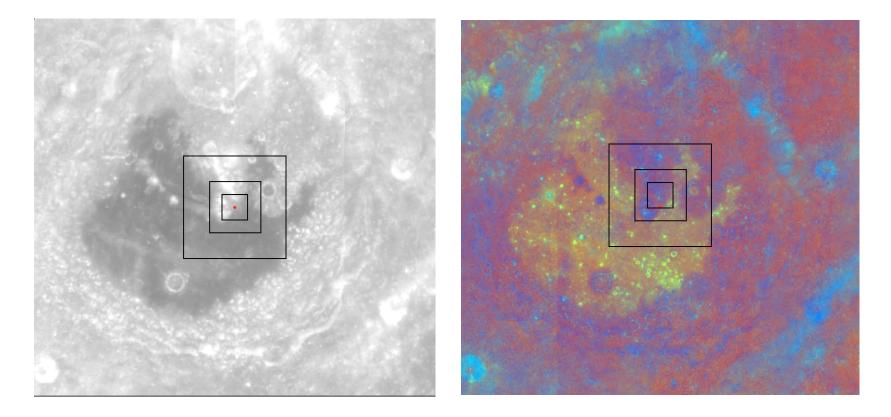
NASA Constellation Program Office Regions of Interest on the Moon:

An update for the 2009 Annual Meeting of the Lunar Exploration and Analysis Group





John E. Gruener Lunar Surface Systems Project Office NASA Johnson Space Center

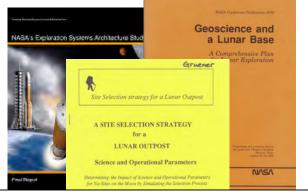
Constellation Program Office Regions of Interest on the Moon

The regions of interest identified by the Constellation Program Office (CxP):

- Illustrate the diversity of scientific and resource opportunities, and geographic terrains and locations
- As a set, they form a representative basis for scientific exploration, resource development, and mission operations

The CxP regions of interest <u>DO NOT</u> represent the initial step in a site selection process for future human missions to the Moon

CxP NAC Targets, ROIs and Products



Geoscience and a Lunar Base Workshop – 1988 Site Selection Strategy for a Lunar Outpost – 1990 Exploration Systems Architecture Study (ESAS) -2005

Data /

Products:

Leveraged heavily on past site selection work and recent Clementine and Lunar Prospector missions to the Moon

- Consideration of scientific, resource utilization, and operational merits and desire to span lunar terrain types
- CxP corrected & refined coordinates, organized into 50 sites grouped into two "tiers" (based on LROC PI recommendation) – 12/08
- Prepared CxP Target Catalog for eventual transfer to LRO 1/09, example entries to follow

Customers:

Surface Systems

Photo-mosaics & Digital Elevation Models	 Accurate <i>representative</i> site data for lander design and terrain relative navigation (TRN) & hazard avoidance system development. If targets are actual landing sites, pin-point landing accuracy through TRN Relevant ROI ~5 km radius 	 Accurate <i>representative</i> site data for trafficability assessments and mobility system designs. Relevant ROI ~ 10 km radius 		
Hazard Assessments	 Accurate <i>representative</i> lander-scale boulders, craters, slopes distributions for lander design and hazard avoidance system development. If targets are actual landing sites, pin-point landing site designation. Relevant ROI ~5 km radius 	 Statistical representation of surface mobility obstacles/hazards. If targets are actual landing sites, detailed operational traverse planning. Relevant ROI ~ 10 km radius 		

Altair/ALHAT

CxP Target Priority Strategy

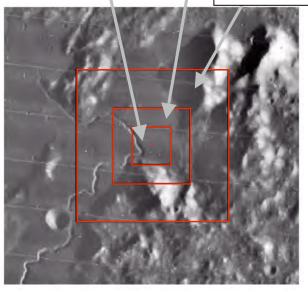
- CxP targets (50 total) are divided into two "tiers" of 25 each in event of mutual interference
- Region of Interest and observations necessary for specific data products prioritized to maximize total data return
- Targeting updates simple process, but acquisition becomes harder as mission progresses (interference, lack of time)
- CxP target coverage will extend into SMD mission transition/frozen orbit imaging possibilities

Priority 1: 10x10 km All Targets with "full observations*"

Priority 2: Other LROC Level 1 measurement requirements

Priority 3: 20x20 km All Targets "best effort" full observations + other Co-I or LRO science

Priority 4: 40x40 km All Targets "best effort*" nadir mosaics + outside science requests



<u>*The full set of observations include:</u>

- 1. Monoscopic image mosaic
- 2. Geometric stereo images
 - Two observations
 - One nadir, one at 20° off nadir (requires s/c slew)
 - Solar incidence angle 50-68° off vertical if possible
- 3. Photometric stereo images

Four observations with different solar incidence & azimuth angles All nadir

- 4. Hazards
 - Two Sets

One at solar incidence angle of 66-72° off vertical, one near 80° off vertical

Actual number of images required will depend on specific orbit

groundtracks, lighting, interference etc. and could be substantially more.

LEAG LROC SAT January-April 2009

Chair, Paul G. Lucey, U. Hawaii

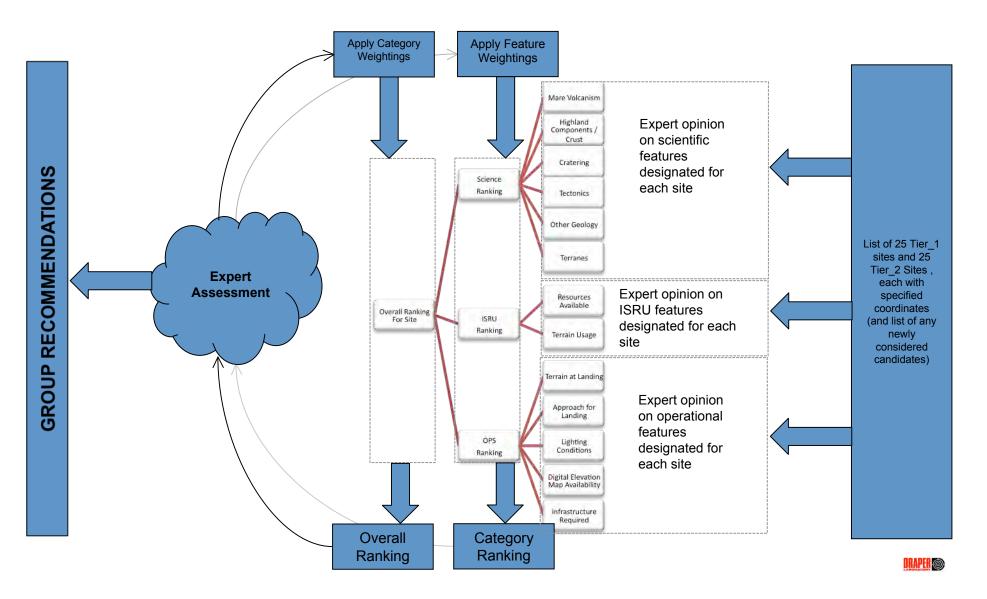
Science Subgroup

- Jeffrey Gillis-Davis, U. Hawaii
- B. Ray Hawke, U. Hawaii

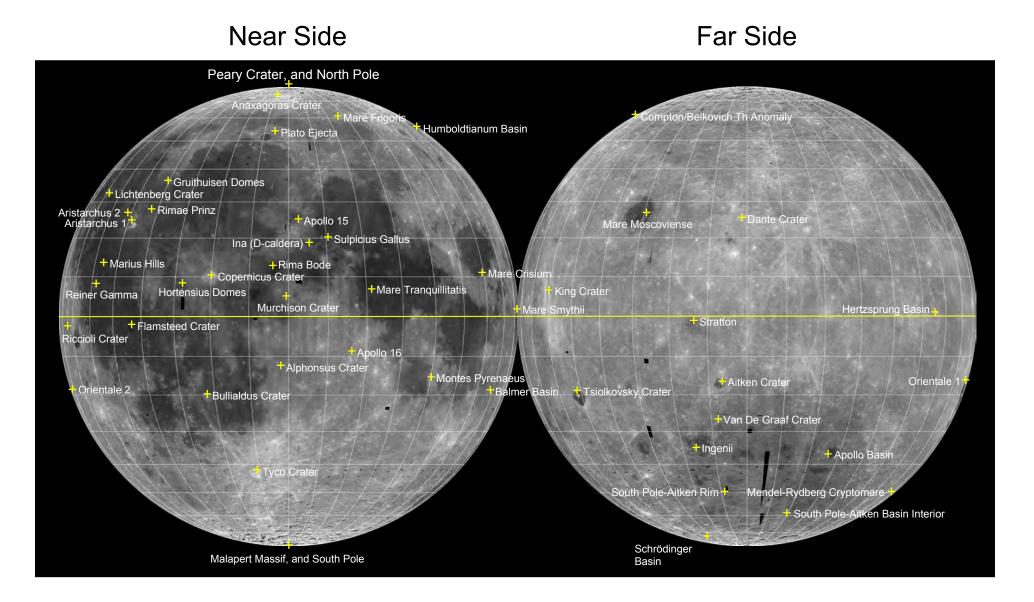
ISRU Subgroup

- Larry Taylor, U. Tenn, Knoxville
- Mike Duke, At-large lunar scientist
- Operations Subgroup
 - Tye Brady, Draper Lab
 - Todd Mosher, Sierra Nevada Corporation
- Observers
 - Mike Wargo, NASA
 - Steve Mackwell, LPI
 - Clive Neal, LEAG

LEAG LROC SAT Comparison Methodology Overview (Illustrated)

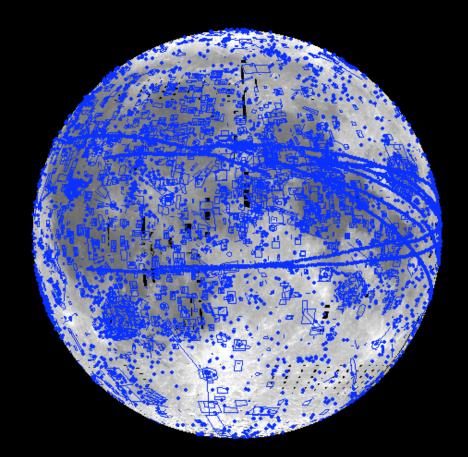


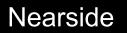
Constellation Program Office Regions of Interest



Constellation LRO NAC Targets											
Tier 1					Tier 2						
	Code	Name	Long	Lat			Code	Name	Long	Lat	
	ATK	Aitken Crater	173.48	-16.76		Tier 2	BAL	Balmer Basin	69.82	-18.69	
	ALP	Alphonsus Crater	-2.16	-12.56			СМР	Compton/Belkovich Th Anomaly	99.45	61.11	
Tier 1	ANX	Anaxagoras Crater	-9.30	73.48			DAN	Dante Crater	177.70	26.14	
	A15	Apollo 15	3.66	26.08			FLM	Flamsteed Crater	-43.22	-2.45	
	A16	Apollo 16	15.47	-9.00			HRT	Hortensius Domes	-27.67	7.48	
	APB	Apollo Basin	-153.72	-37.05			ним	Humboldtianum Basin	77.14	54.54	
	AR1	Aristarchus 1	-48.95	24.56			INA	Ina ('D-caldera')	5.29	18.65	
	AR2	Aristarchus 2	-52.40	27.70			ING	Ingenii	164.42	-35.48	
	BUL	Bullialdus Crater	-22.50	-20.70			LCT	Lichtenberg Crater	-67.23	31.65	
	COP	Copernicus Crater	-20.01	9.85			FRG	Mare Frigoris	26.10	59.80	
	GRT	Gruithuisen Domes	-40.14	36.03			MOS	Mare Moscoviense	150.47	26.19	
	HTZ	Hertzsprung	-125.56	0.09			SMT	Mare Smythii	85.33	2.15	
	KNG	King Crater	119.91	6.39			TRN	Mare Tranquillitatis	22.06	6.93	
	MAL	Malapert Massif	-2.93	-85.99			MAR	Marius Hills	-55.80	13.58	
	CRS	Mare Crisium	58.84	10.68			MEN	Mendel-Rydberg Cryptomare	-93.07	-51.14	
	MUR	Murchison Crater	-0.42	4.74			PYR	Montes Pyrenaeus	40.81	-15.91	
	NPO	North Pole	76.19	89.60			OR2	Orientale 2	-87.91	-18.04	
	OR1	Orientale 1	-95.38	-26.20			PLT	Plato Ejecta	-5.21	53.37	
	PRY	Peary Crater	30.00	88.50			RNG	Reiner Gamma	-58.56	7.53	
	RMB	Rima Bode	-3.80	12.90			RIC	Riccioli Crater	-74.28	-3.04	
	SPO	South Pole	-130.00	-89.30			RMP	Rimae Prinz	-41.72	27.41	
	SPA	South Pole-Aitken Basin Interior	-159.94	-60.00		SCH	Schrödinger	138.77	-75.40		
	STR	Stratton	166.88	-2.08			SPR	South Pole-Aitken Rim	170.92	-51.00	
	SPG	Sulpicius Gallus	10.37	19.87			TSK	Tsiolkovsky Crater	128.51	-19.35	
	TYC	Tycho Crater	-11.20	-42.99			VDG	Van De Graaf Crater	172.08	-26.92	

LROC Master List of targets as of Nov. 10, 2009 (16, 056 targets)







LEAG Workshop on Sustaining Lunar Exploration

Purpose and Scope:

- Learning from history is important for ensuring that a sustainable and longterm lunar return is possible
- The Apollo program was not sustainable
- Focus on the sustainability theme of the LEAG-coordinated Lunar Exploration Roadmap

- The remainder of this presentation will discuss those Cx regions of interest that support sustainability on the Moon:
 - Resources
 - Long-term habitation

Sustainability on the Moon A historical perspective

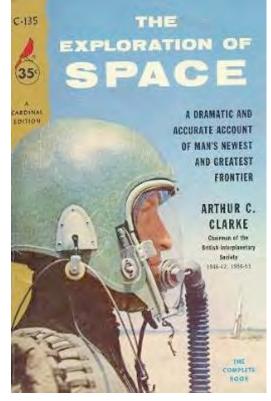
"The human race is remarkably fortunate in having so near at hand a full-sized world with which to experiment: before we aim at the planets, we will have had a chance of perfecting our astronautical techniques on our own satellite. . .the conquest of the Moon will be the necessary and inevitable prelude to remoter and still more ambitious projects."

"The crossing of interplanetary space, though a technical problem which will challenge all Man's ingenuity and resources, is not an end in itself but merely a beginning. There is no point in going to the planets unless we do something when we get there."

"Today we can no more predict what use mankind may make of the Moon than could Columbus have imagined the future of the continent he had discovered."

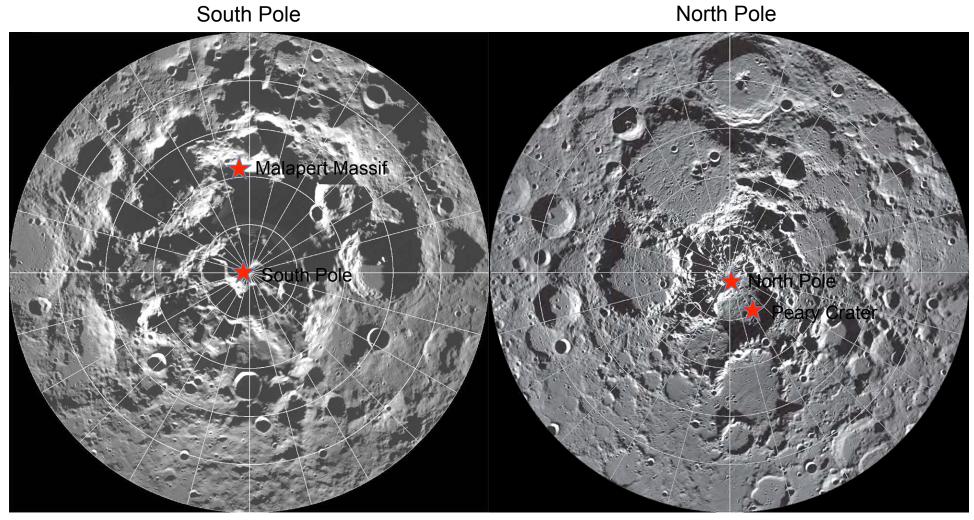
"The first lunar explorers will probably be mainly interested in the mineral resources of their new world, and upon these its future will very largely depend."

"For a considerable time all flights to the Moon would be directed to the same spot, so that material and stores could be accumulated where they would be most effective. There would be no scattering of resources over the Moon's twelve million square miles of surface."

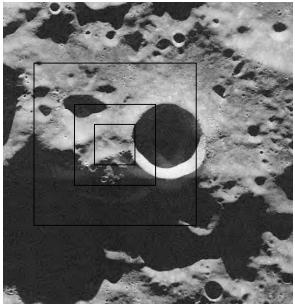


(1951)

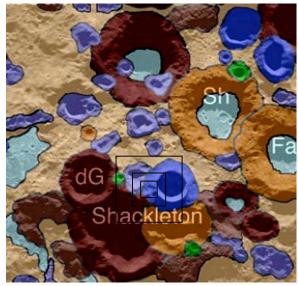
Constellation Polar Regions of Interest



from Paul Spudis, Clementine mosaics



Radar image from Margot et al., Science 284, 1658-1660 (1999)



Geologic map from Spudis et al., (2008)

South Pole

Location (longitude, latitude): -130, -89.3 (best estimate, see image to left)

Scientific Rationale:

South Pole-Aiken (SPA) basin geology Polar volatiles Impact process (e.g., Shackleton and other craters)

Resource Potential:

Highlands regolith Enhanced hydrogen in permanently shadowed polar craters (water ice?) Sunlight

Operational Perspective:

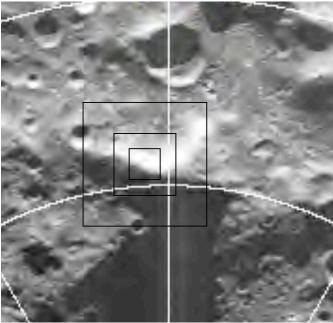
Highlands terrain Polar location Areas of permanent shadow Points of near-continuous sunlight

NASA References:

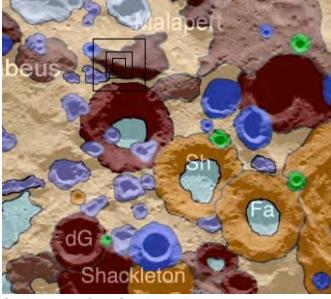
Exploration Systems Architecture Study (2005) Geoscience and a Lunar Base (1990)

Other References:

Spudis et al., GRL, 35, L14201, doi:10.1029/2008GL034468. Bussey et al., GRL, 26, no.9, 1187-1190 (1999)



Radar image from Margot et al., Science 284, 1658-1660 (1999)



Geologic map from Spudis et al., (2008)

Malapert Massif

Location (longitude, latitude): -2.93, -85.99 (best estimate, see image to left)

Scientific Rationale: South Pole-Aitken (SPA) basin rim? Basin geology Observatories

Resource Potential: Near-continuous sunlight (continuous?) Direct-to-Earth communication

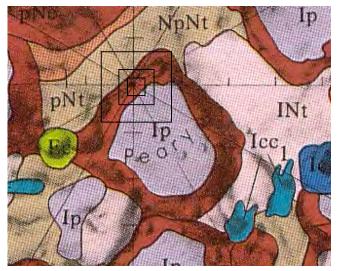
Operational Perspective:

Highlands terrain (e.g., massif) Polar location

NASA References:

Other References: Spudis et al., GRL, 35, L14201, doi:10.1029/2008GL034468





(Clementine uvvis color ratio image not available)

North Pole

Location (longitude, latitude): 76.19, 89.60 (best estimate, see image to left)

Scientific Rationale:

Polar volatiles Impact process (e.g., heavily cratered highlands) Distal Imbrium ejecta

Resource Potential:

Highlands regolith Enhanced hydrogen in nearby permanently shadowed polar craters (water ice?) Sunlight

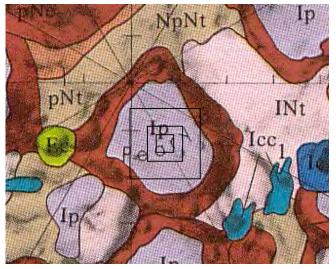
Operational Perspective:

Highlands terrain Polar location Nearby areas of permanent shadow Points of near-continuous sunlight

NASA References:

Exploration Systems Architecture Study (2005) Geoscience and a Lunar Base (1990)





(Clementine uvvis color ratio image not available)

Peary Crater

Location (longitude, latitude): 30.00, 88.50

Scientific Rationale: Polar volatiles Impact process

Resource Potential:

Highlands regolith Enhanced hydrogen in permanently shadowed polar craters (water ice?)

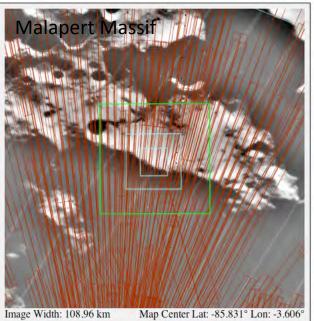
Operational Perspective:

Highlands terrain Polar location Areas of permanent shadow

NASA References:

Exploration Systems Architecture Study (2005) Geoscience and a Lunar Base (1990)

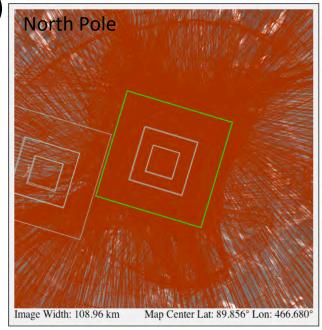
LROC NAC Coverage of Constellation Polar Regions of Interest

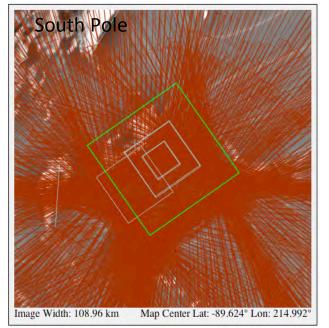


(as of November 10, 2009) North Pole

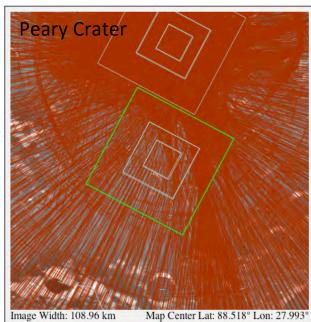
Potential resources:

- Sunlight
- H₂ in permanent shadow
- H₂⁻0 and solar wind H₂ in lunar regolith
- O₂ and Al in lunar regolith
- Regolith as construction material



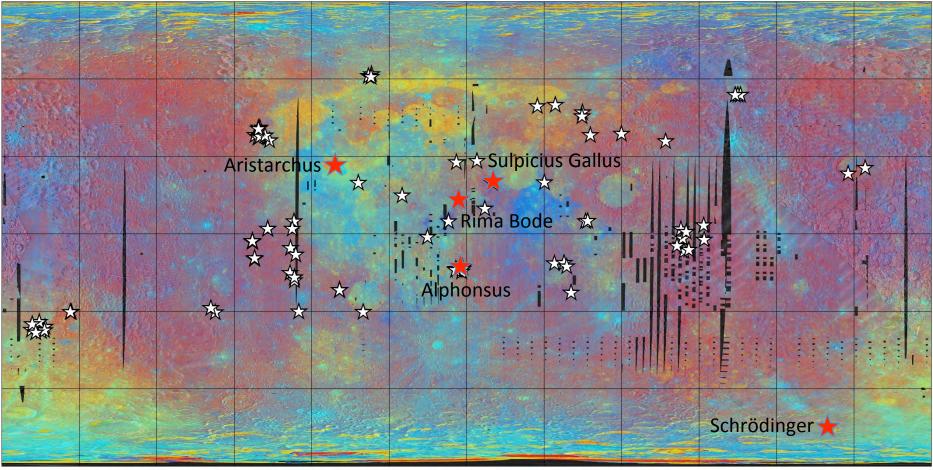


 See presentations from earlier in the workshop for further details

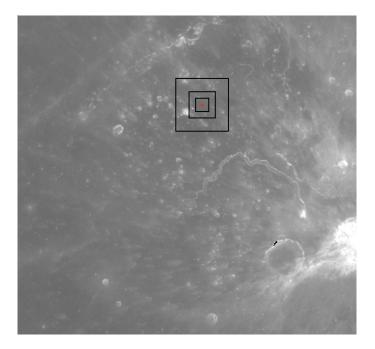


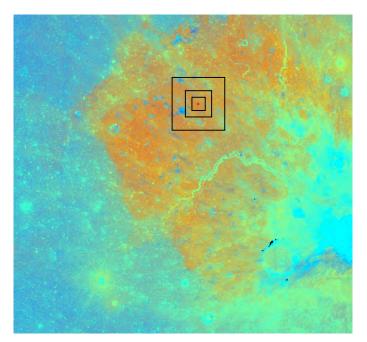
Location of pyroclastic deposits on the Moon

(**★** = Constellation Program region of interest)



from Lisa Gaddis, USGS Lunar Pyroclastic Volcanism Project





Aristarchus 2

Location (longitude, latitude): -52.40, 27.70

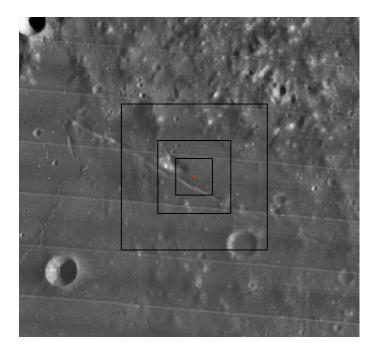
Scientific Rationale: Pyroclastic materials Nearby volcanic features

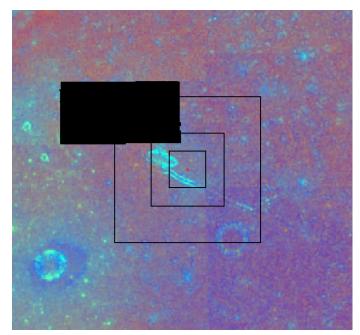
Resource Potential: Pyroclastic materials

Operational Perspective: Pyroclastic covered surface Near side location

NASA References:

Exploration Systems Architecture Study (2005) A Site Selection Strategy for a Lunar Outpost (1990) Geoscience and a Lunar Base (1990)





Rima Bode

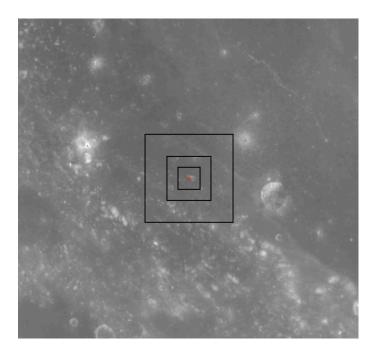
Location (longitude, latitude): -3.80, 12.90

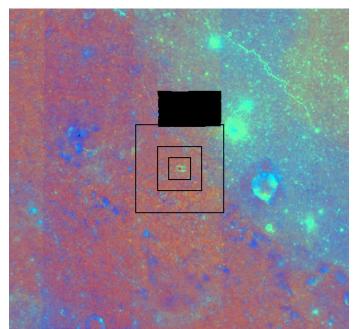
Scientific Rationale: High-Ti pryroclastic material Mantle xenoliths

Resource Potential: High-Ti pyroclastic material

Operational Perspective: Pyroclastic covered surface Highlands terrain Near side location

NASA References: Exploration Systems Architecture Study (2005) Geoscience and a Lunar Base (1990)





Sulpicius Gallus

Location (longitude, latitude): 10.37, 19.87

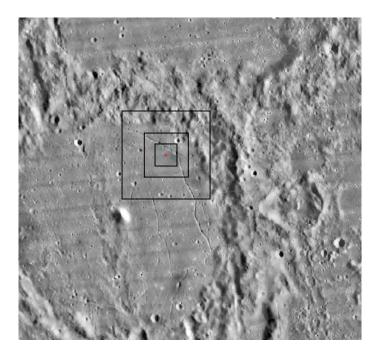
Scientific Rationale: Dark mantling material, pyroclastics Mantle xenoliths

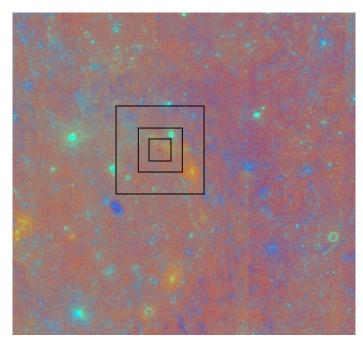
Resource Potential: Pyroclastic deposits

Operational Perspective: Smooth pyroclastic covered surface Mare terrain Near side location

NASA References: Geoscience and a Lunar Base (1990)

Other References: Lucchitta and Schmitt, 5th Lunar Conference (1974)





Alphonsus Crater

Location (longitude, latitude): -2.16, -12.56

Scientific Rationale:

Pryoclastic vents and materials Lunar transient events Alphonsus crater rim massifs Ranger 9 impact site

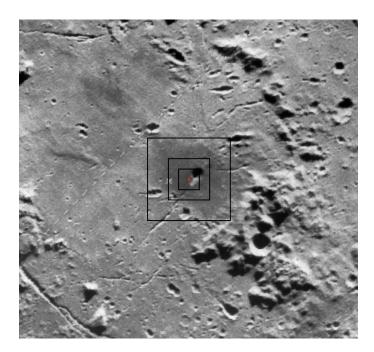
Resource Potential: Highlands regolith Pyroclastic materials

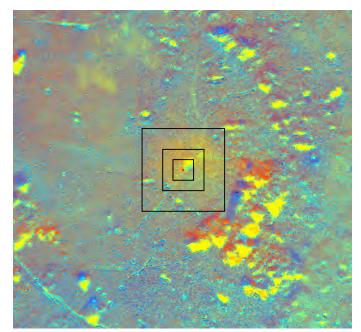
Operational Perspective: Highlands terrain

Pyroclastic covered surface Surface fracture

NASA References:

Optimizing Science and Exploration Working Group (OSEWG) Sortie Surface Scenario Workshop (2008) Geoscience and a Lunar Base (1990)





Schrödinger

Location (longitude, latitude): 138.77, -75.40

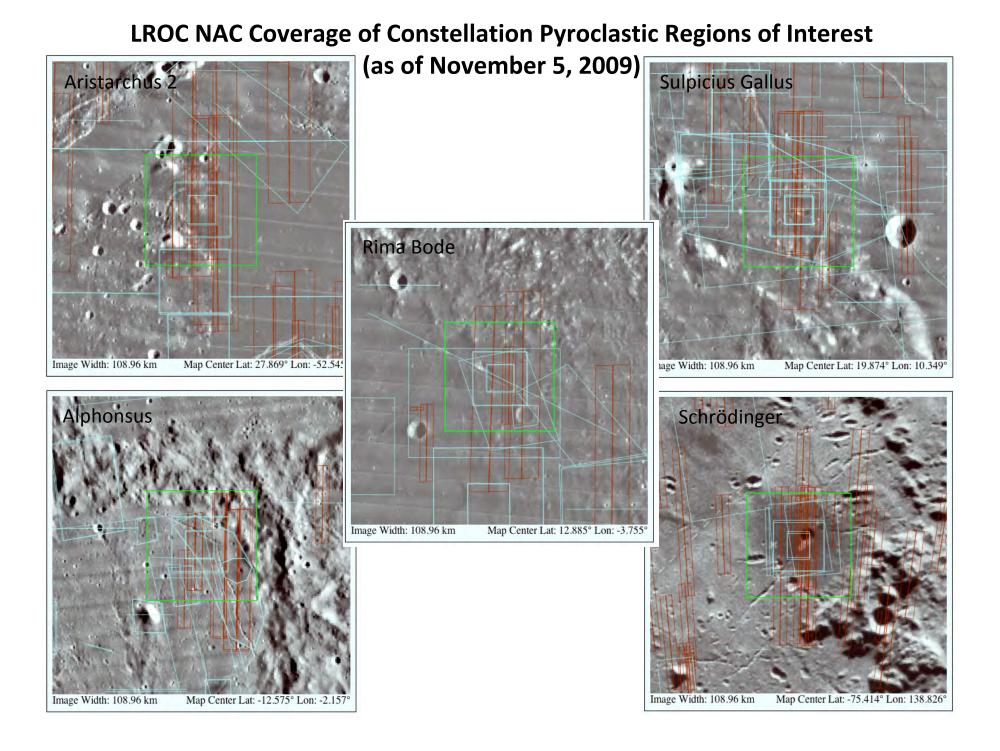
Scientific Rationale: Pyroclastic materials Mantle xenoliths Schrödinger basin impact melts and breccias

Resource Potential: Pyroclastic materials

Operational Perspective: Pyroclastic covered surface Far side location

NASA References: Geoscience and a Lunar Base (1990)

Other References: E.M. Shoemaker et al. (1994), Science Vol. 266, 1851-1854.



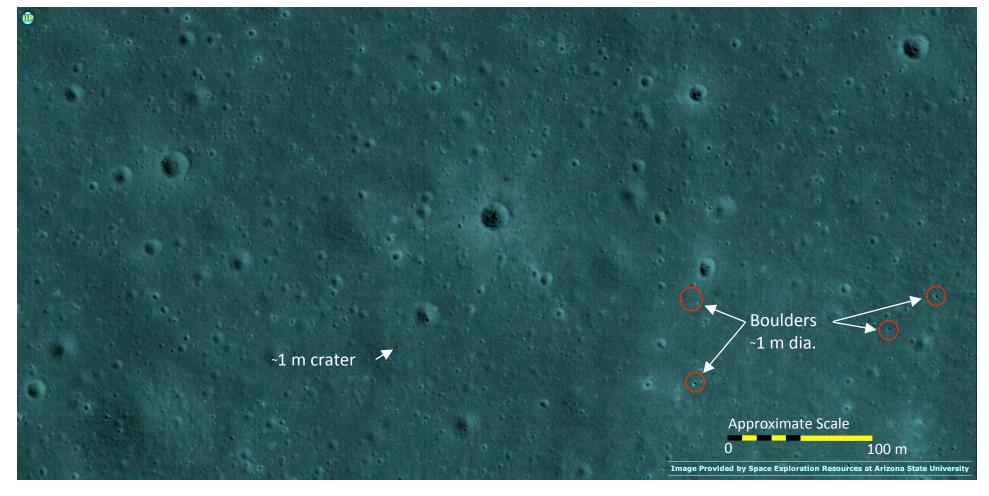


Cx ROI: Aristarchus 2

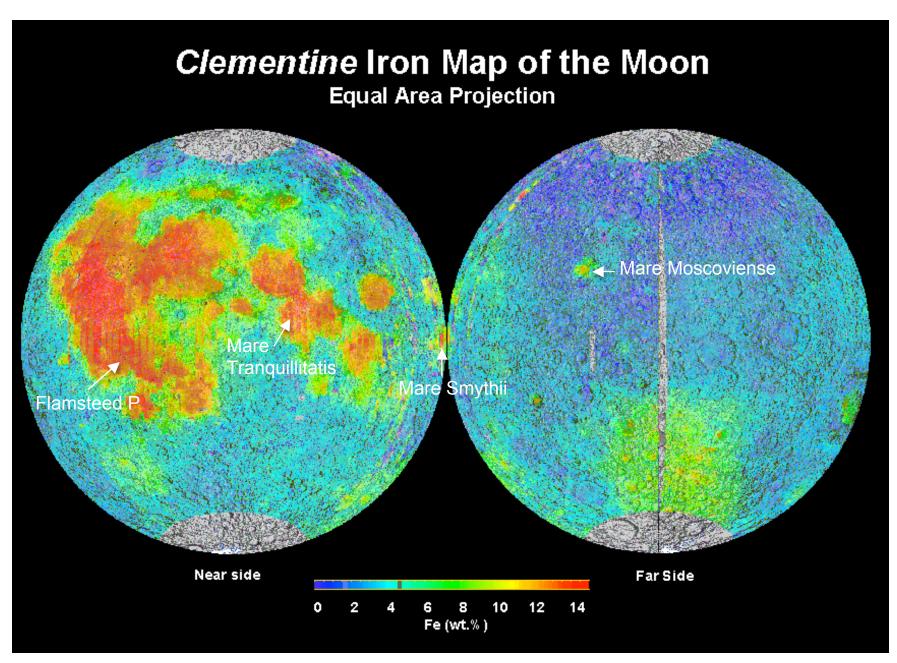
Simple Mosaic using NAC images: (left to right) M104862583L M104862583R M102500639L M102500639R

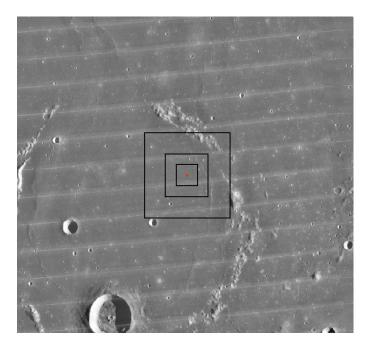
~10 km

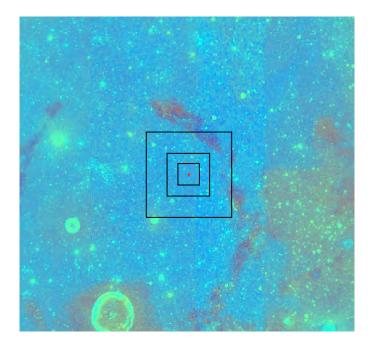
Cx ROI: Aristarchus 2 NAC Images: M111945148R LRO Altitude: 47 km Resolution: 0.54 m/pixel



Constellation High-Fe or High-Ti Mare Regions of Interest







Flamsteed Crater

Location (longitude, latitude): -43.22, -2.45

Scientific Rationale:

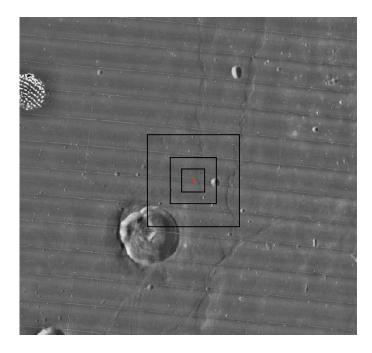
Surveyor 1 site Young basaltic lavas (Eratosthenian–Copernican?) Thin regolith Nearby Flamsteed P crater ring

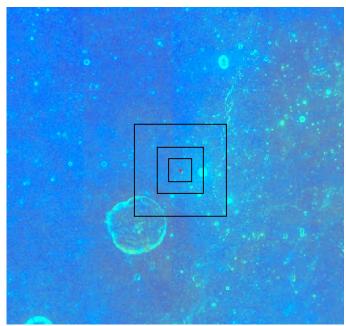
Resource Potential: High-Ti basalts

Operational Perspective: Mare terrain Near side location

NASA References:

Exploration Systems Architecture Study (2005) Geoscience and a Lunar Base (1990)





Mare Tranquillitatis

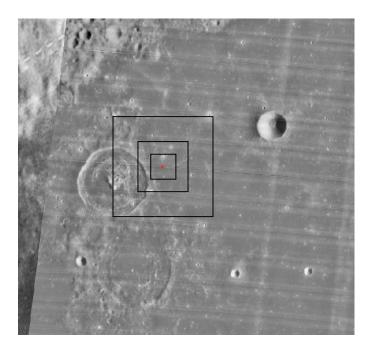
Location (longitude, latitude): 22.06, 6.93

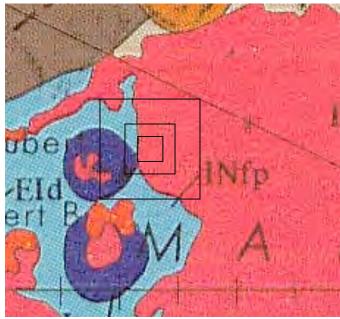
Scientific Rationale: High-Ti basaltic lavas Volcanic processes Wrinkle ridges, low basaltic shields

Resource Potential: High-Ti mare regolith

Operational Perspective: Mare terrain Near side location

NASA References: Exploration Systems Architecture Study (2005)





(Clementine uvvis color ratio image not available)

Mare Smythii

Location (longitude, latitude): 85.33, 2.15

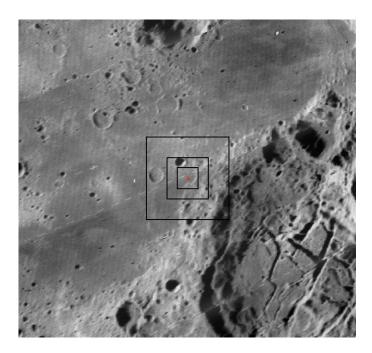
Scientific Rationale: Young basaltic lavas Nearby floor-fractured crater

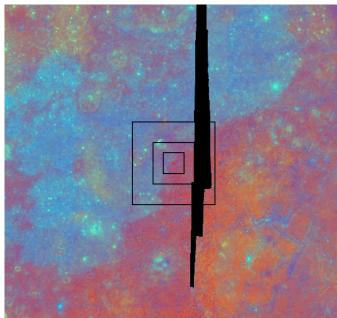
Resource Potential: High-Fe mare regoilth

Operational Perspective: Mare terrain Limb location

NASA References:

Exploration Systems Architecture Study (2005) A Site Selection Strategy for a Lunar Outpost (1990)





Mare Moscoviense

Location (longitude, latitude): 150.47, 26.19

Scientific Rationale:

Mare age and composition (e.g., far side mare) Basin geology (e.g., inner ring)

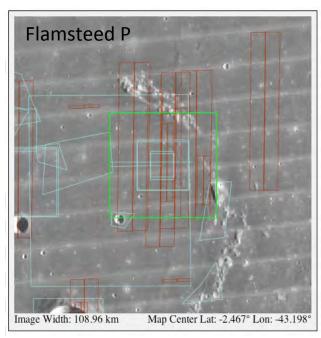
Resource Potential: High-Ti mare regolith

Operational Perspective:

Mare terrain Highlands terrain Far side location

NASA References: Geoscience and a Lunar Base (1990)

LROC NAC Coverage of Constellation High-Fe or High-Ti Mare Regions of Interest

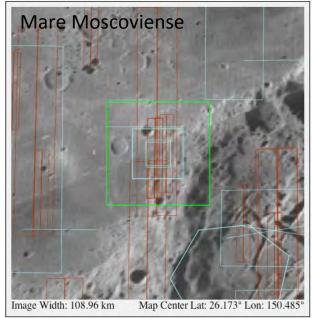


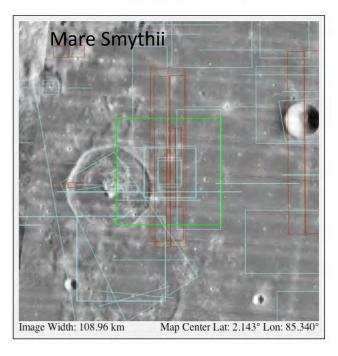
 Mare Tranquillitatis

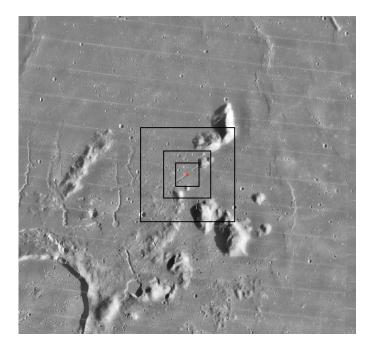
 Image Width: 108.96 km

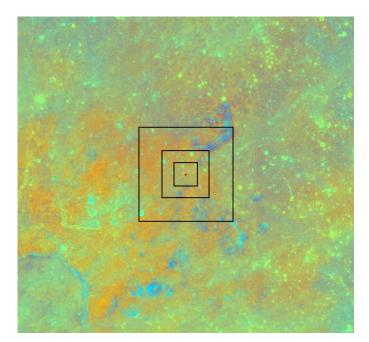
Map Center Lat: 6.912° Lon: 22.074°

(as of November 6, 2009)









Rimae Prinz

Location (longitude, latitude): -41.72, 27.41

Scientific Rationale: Rille Possible lava tube Nearby highlands massifs (Imbrium basin related)

Resource Potential: Mare regoilth

Operational Perspective:

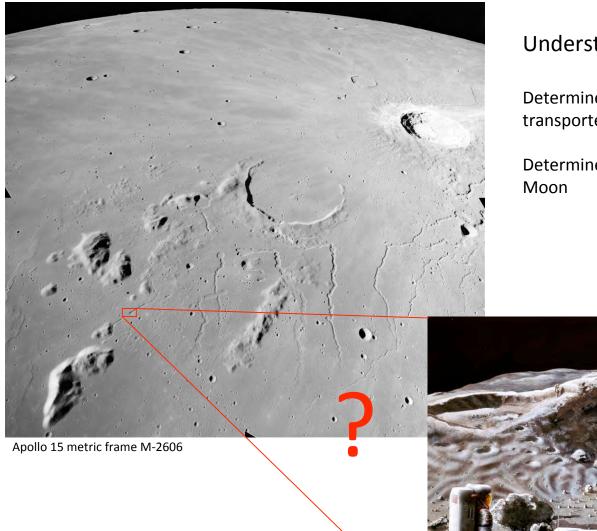
Mare terrain Sinuous rille (e.g., similar to Apollo 15 Hadley rille) Near side location

NASA References:

Other References:

Hörz, Lunar Bases and Space Activities of the 21st Century (1985)

Rimae Prinz



Understanding Volcanic Processes

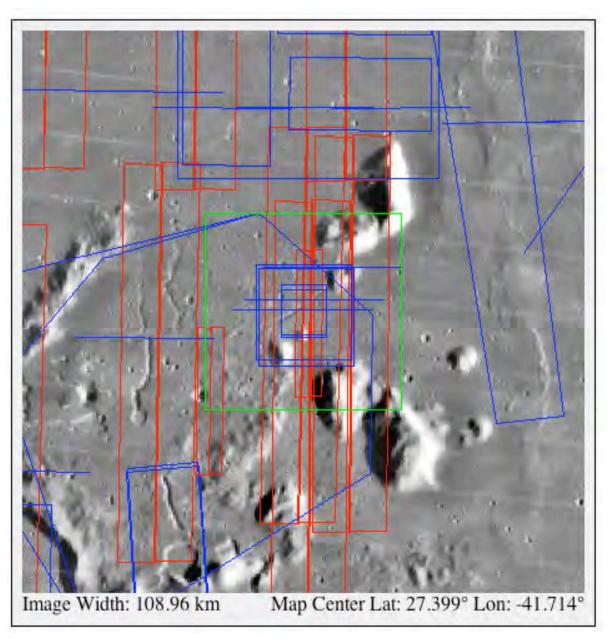
Determine how magma is generated and transported to the surface

Determine how lava flows are emplaced on the Moon

Artist concept of the discovery of a lunar lava tube



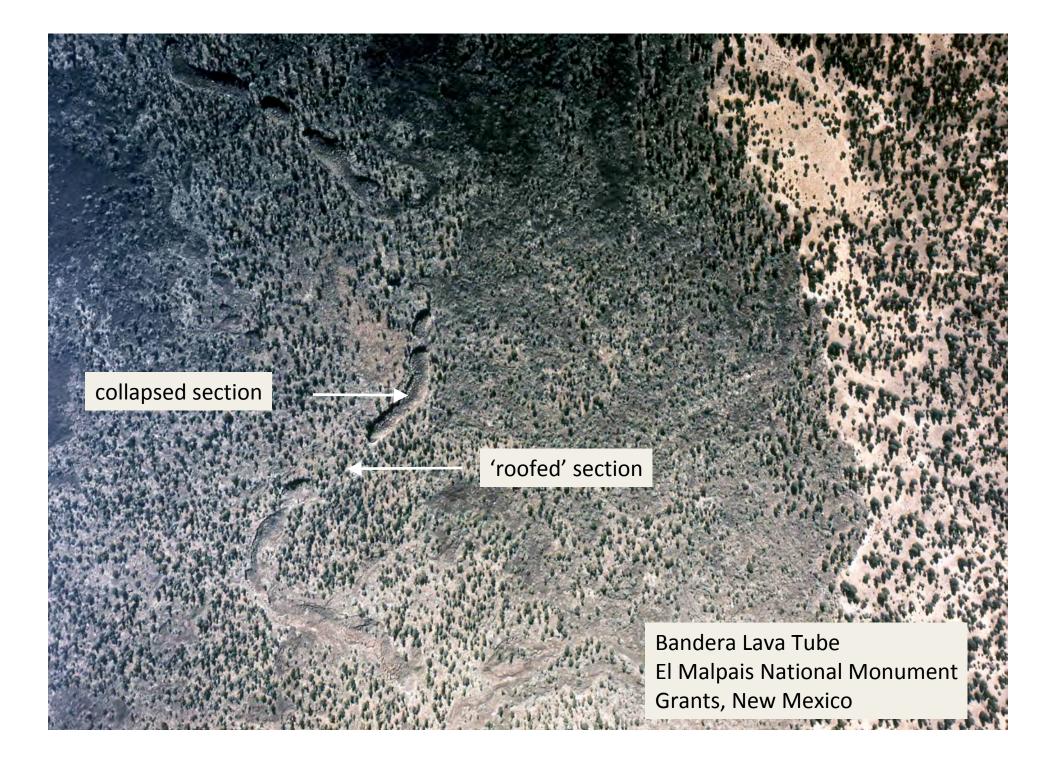
LROC NAC images acquired of Rimae Prinz region of interest (as of 10-20-09)





Simple Mosaic using NAC images: (left to right) M102436231L M102436231R M102429075L M102429075R

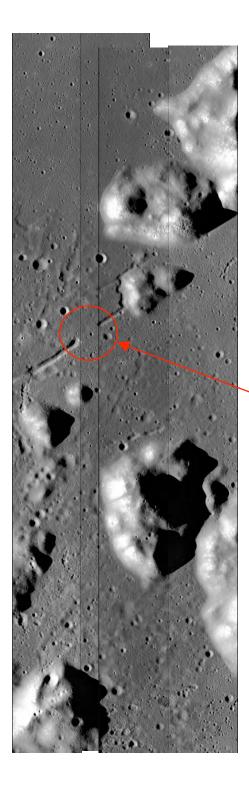
~10 km





opening into tube at end of collapse section (courtesy of BLM)

opening into tube through a 'skylight' (courtesy of NPS)

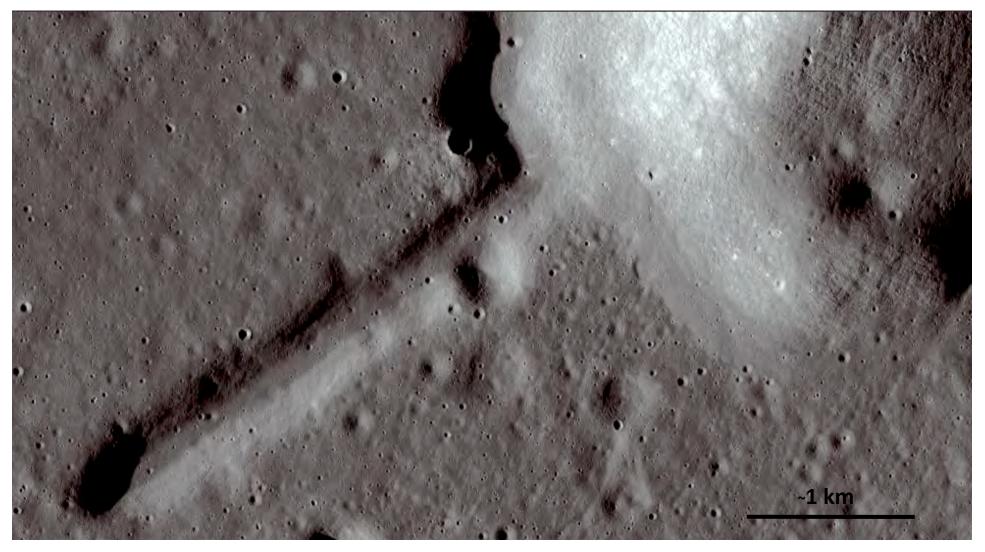


Simple Mosaic using NAC images: (left to right) M102436231L M102436231R M102429075L M102429075R

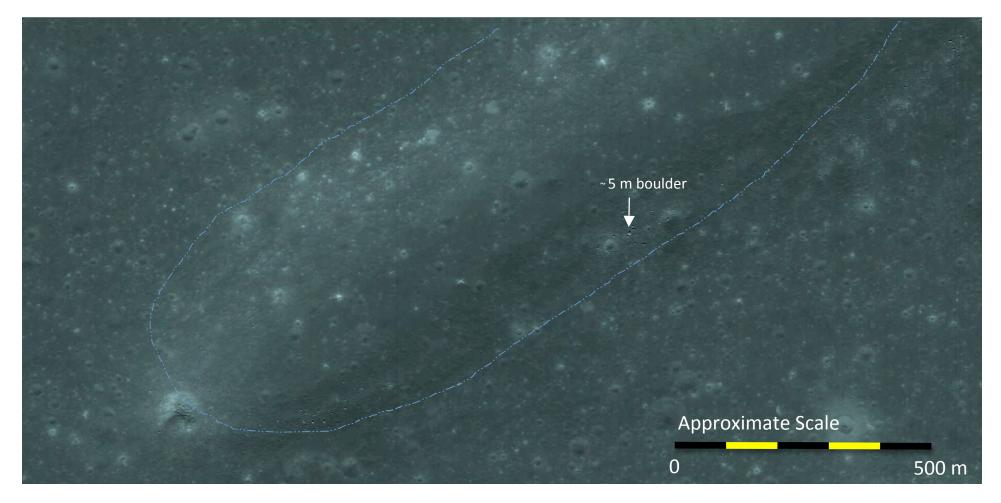
The next five slides will focus on this area

~10 km

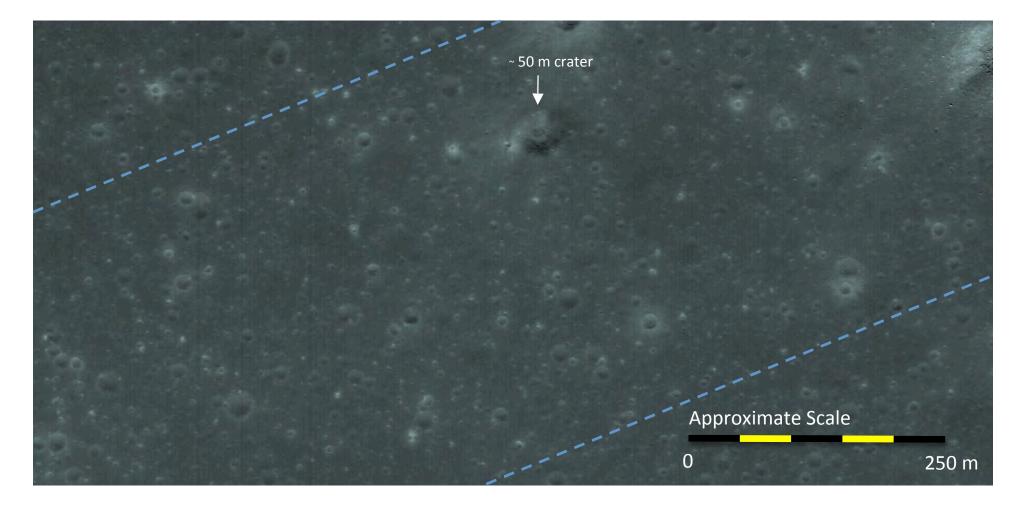
NAC Image: M102436231R LRO Altitude: 146.18 km Resolution: 1.47 m/pixel



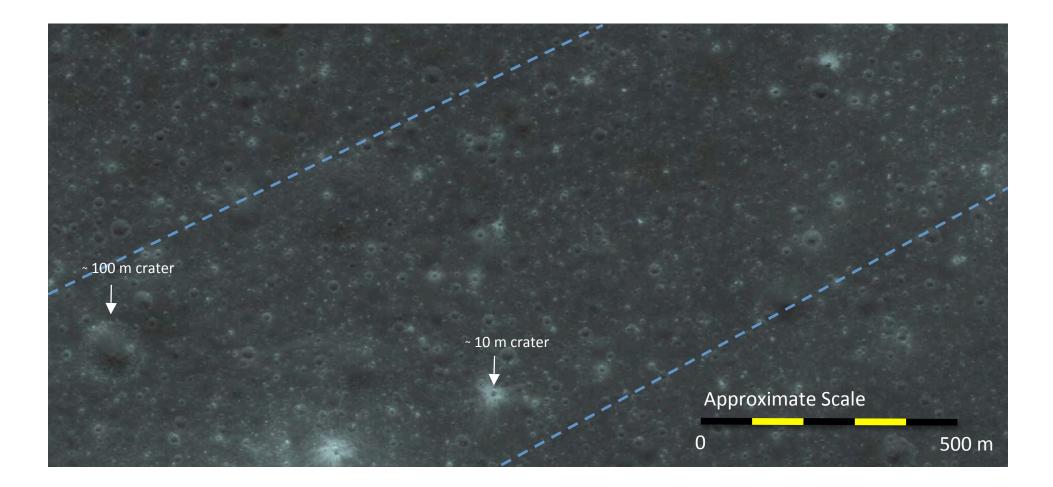
NAC Image: M109507800L LRO Altitude: 50.52 km Resolution: 0.53 m/pixel



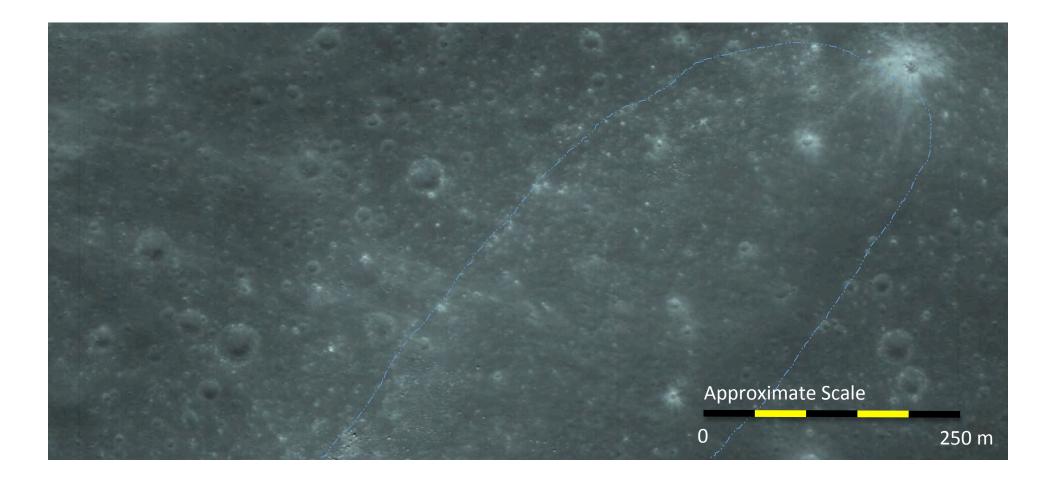
NAC Image: M109507800L LRO Altitude: 50.52 km Resolution: 0.53 m/pixel

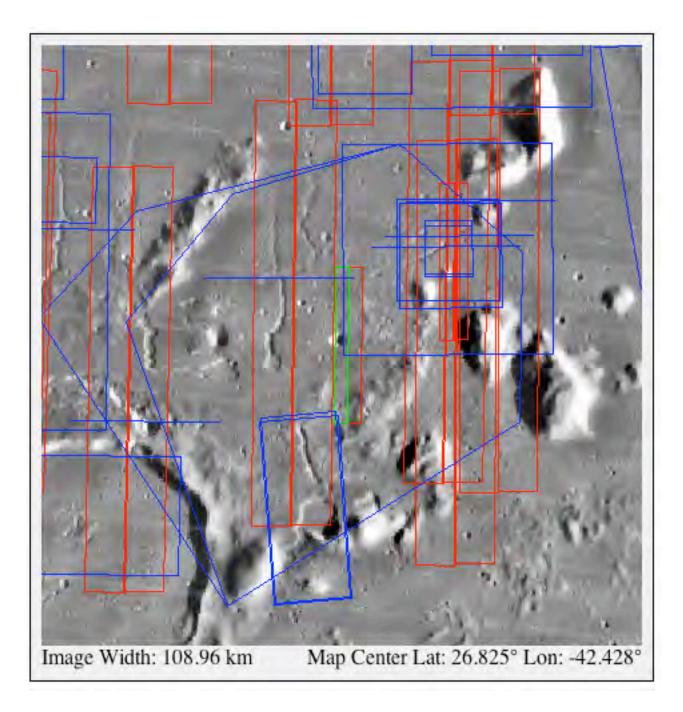


NAC Image: M109507800R LRO Altitude: 50.52 km Resolution: 0.53 m/pixel



NAC Image: M109507800R LRO Altitude: 50.52 km Resolution: 0.53 m/pixel



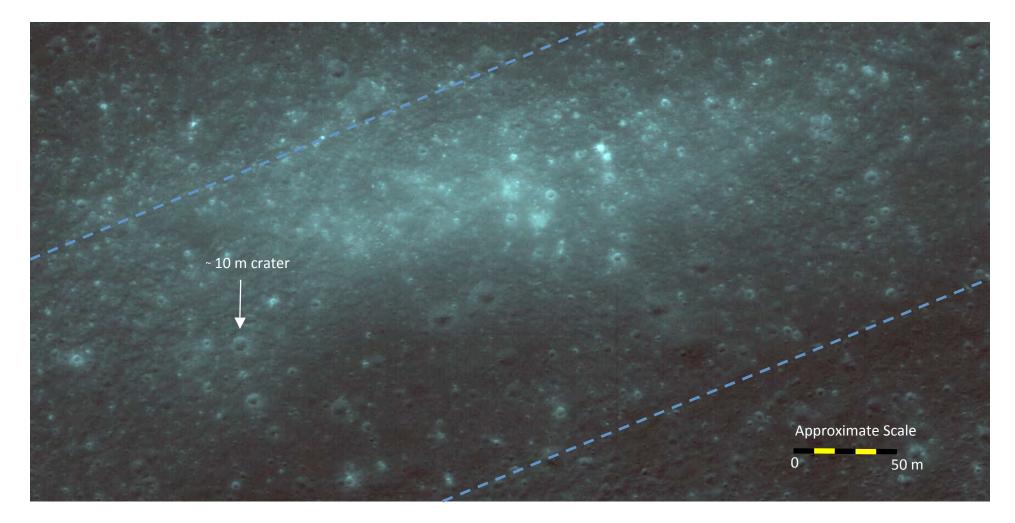


Approximate Scale

500 m

Simple Mosaic using NAC images: (left to right) M109514603R M109514603L

LRO altitude: 50.34 km Resolution: 0.52 m/pixel NAC Image: M109514603R LRO Altitude: 50.34 km Resolution: 0.52 m/pixel



Summary

 In coordination with the LROC team and the lunar science community, NASA's Constellation program (CxP) has designated 50 regions of interest for LROC NAC imaging

- As a set, they Illustrate the diversity of the lunar surface, and form a representative basis for scientific exploration, resource development, and mission operations
- The 50 CxP regions of interest are a subset of over 16,000 LROC targets
- The CxP regions of interest are getting excellent quality images and consistent imaging from the LROC NAC
- The CxP regions of interest <u>DO NOT</u> represent the initial step in a site selection process for future human missions to the Moon
- Many of the CxP regions of interest support the sustainability theme of the LEAG-coordinated Lunar Exploration Roadmap
- Resources
- Long-term habitation
- Initial analysis has begun on the 50 CxP regions of interest
- Much of the Aristarchus 2 ROI appears to be relatively boulder free at 1-meter scale
- There are no apparent openings into one of the lava tube candidates at the Rimae Prinz ROI