

K-RICH PHASE IN NIHO-3(MIYANO) H3 CHONDRITE;

Y. MIURA and G. SHIBUYA

Department of Mineralogical Sciences and Geology, Faculty of Science, Yamaguchi University, Yoshida, Yamaguchi, 753, Japan.

The Niho-3(Miyano) meteorite of 19.18 grams fell in Miyano, Yamaguchi-shi, Yamaguchi-ken Japan ($34^{\circ}12'N$, $131^{\circ}34'E$) on August 8, 1897. The meteorite is considered to be the same fragment with the Niho-1 and -2 (H3) chondrites discovered at a distance 6.2 km north-east away. The ellipsoidal dark-gray meteorite was originally 3x2x1.5cm (density 3.59 gr/cm³). Electron microprobe analyses of 165,600 analyzed points by the JCMA-733 yield the bulk composition of the meteorite as follows in weight percent: SiO₂ 35.57, Al₂O₃ 2.54, Cr₂O₃ 0.59, FeO 22.04, MnO 0.39, MgO 21.23, CaO 2.31, Na₂O 0.95, K₂O 0.08, NiO 0.23 and BaO 0.04. Modal mineral contents (in volume percent) are orthopyroxene 32, clinopyroxene 24, olivine 24, plagioclase(-like) 8, metal phases 9.0 and K-rich phase 1.0. The meteorite contains numerous glassy and crystalline chondrule (up to 2.2x1.2 mm in size). Various sizes of orthopyroxenes of inhomogeneous composition (En₈₃₋₈₇ Fs₁₀₋₁₆ Wo₀₋₅) show microporphyric chondrules. The large pyroxene chondrule (2.2x1.2mm) consists of orthopyroxene (En₈₅Fs₁₄) at the rim and clinopyroxene (Wo₄₆ En₄₆) in the core. Olivine occurs in the microporphyric chondrules and has inhomogeneous composition (Fo₈₂₋₉₉Fa₁₋₁₈). The metal phases consist of kamacite(86 vol.%), plessite(11 vol.%) and taenite(3 vol.%). The fine-grained plagioclase(-like) grains consist of glassy (maskelynite, An₁₇₋₁₉Or₂₋₃) and clear crystal of An₁₁Or₁₃ (40x30μm in size). The latter plagioclase(-like) crystal grains are enriched with Mg, K and Fe, but deficient in Al, Na and Ca; that is, plausible chemical formula is (Na_{0.6}K_{0.1}Ca_{0.1}Mg_{0.1}Fe_{0.1})(Al_{0.8}Si_{2.9}Mg_{0.3})O₈. K-rich phase is irregularly distributed in the matrix, and shows composition of Or₃₃An₄₀, but enriched with Mg and Fe (Table 1).

Frequency distribution of iron contents of olivine and orthopyroxene, and textural characteristics suggest that the Niho-3 meteorite belongs to the H-group and petrologic type 3 of Dodd(1981) classification [1]. It is confirmed that the Niho-3 meteorite is the same type and group with the Niho-2 (H3) reported by Shima et al. (1984)[2].

The Niho meteorites might be differentiated fragments within the IIE parent body at the primordial stage by the following reasons: (1) The K-rich phase of alkali-feldspar has been discovered in type IIE iron meteorite and Naklites (but not in chondrites) so far. (2) Oxygen-isotope data for the group IIE silicate inclusions suggest that they are related to H-group chondrites[3]. (3) The Niho-2 chondrite shows gas-rich of ⁴He and trapped Ne, and age of 4.51 b.y.(by ⁴⁰Ar dating)[2]. (4) The values of Mg/(Fe+Mg) in the Niho-3 are almost the same in orthopyroxenes(0.85), olivine (0.85), plagioclase(-like) grains (0.81) and K-rich phase(0.80) (cf. Table 1).

If there is the different parts of the Niho parent body, it is exceedingly valuable to explain the anomalous compositions of

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K-rich phase and plagioclase(-like) grain in the Niho-3 chondrite. The Kuga iron meteorite might be the different fragments from the Niho parent body.

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Table 1. Chemical compositions of representative K-rich phase and plagioclase(-like) grain in the Niho-3(Miyano) H3 chondrite.

K-rich phase		plagioclase(-like) grain	
(wt.%)	(Cations per 80)	(wt.%)	(Cations per 80)
SiO ₂	64.26	Si 2.93	65.24
TiO ₂	0.37	Ti 0.01	0.65
Al ₂ O ₃	11.91	Al 0.64	15.03
Cr ₂ O ₃	0.48	Cr 0.02	0.20
FeO*	3.09	Fe 0.12	2.56
MnO ₂	0.55	Mn 0.02	0.10
MgO	7.00	Mg 0.48	6.52
BaO	0.00	Ba 0.00	0.01
CaO	6.14	Ca 0.30	1.77
Na ₂ O	2.25	Na 0.20	6.67
K ₂ O	4.28	K 0.25	1.78
Total	100.33	4.97	100.53
Or(mol %)	33.3		13.3
Ab(mol %)	26.6		75.7
An(mol %)	40.1		11.0
Mg/(Fe+Mg)	0.80		0.81

* Total iron oxide as FeO.