

## FIRST INVESTIGATION OF NOBLE GASES IN THE DENGLEI H3,8 CHONDRITE

M.A. Ivanova, S.S. Assonov, Yu.A. Shukolyukov. Vernadsky Inst. Geochem. Anal. Chem., Russian Acad. Sci.

The Dengle (H3,8) meteorite weighting 243,5 g is a find from the Karakum desert. It's a complex microbreccia contained unusual clasts which are more typical for regolithic breccias than for H-chondrites. Based on Xe and Kr contents and their isotopic composition, the Dengle doesn't differ significantly from other H-chondrites. Its exposure age is 7,6 Ma. That's near to common possible data of exposure age (6,2±0,2 Ma) of 350 H-chondrites. Dengle's K/Ar age (4,01 Ga) coincides with the same ages of many other H-chondrites. Thus the Dengle isn't regolithic breccia and it probably formed during accretion of its parent body.

Based on features of texture and chemical composition of olivine and low-Ca pyroxene he has been classified as H3,8 chondrite [1]. The meteorite has a chondritic texture, but irregular achondritic clasts and isometric SiO<sub>2</sub>-bearing inclusions distinguished on composition from the meteorite-host were discovered. Therefore this meteorite can consider a complex breccia, for example, a regolithic breccia.

The purpose of this investigation was a verification of this assumption using isotopic peculiarities of noble gases. The concentration of solar wind gases which are characterized by typical solar isotopic composition, should be much more in the regolithic breccia than in the ordinary chondrites.

The contents of Ar, Kr and Xe are very similar to other ordinary chondrites of petrological types 3-5 [2-5] (Tabl.1). Xe isotopic composition on any temperature step distinguishes from solar Xe (Fig.1) and resembles Xe of ordinary chondrites (there are average results for 16 H-chondrites [3-9]). Kr isotopic composition is within a range of other H-chondrites (Fig. 2) and it differs as strongly from solar Kr as Xe. Distinct features of cosmogenic

<sup>38</sup>Ar<sub>c</sub> (Tabl. 1) were founded in isotopic composition of Ar<sub>c</sub>. The majority of Ar<sub>c</sub> was isolated at 900°. The concentration of <sup>38</sup>Ar<sub>c</sub> is 4,18·10<sup>-9</sup> sm<sup>3</sup>/g, exposure age is 7,6 Ma (P<sub>38</sub>=0,055·10<sup>-8</sup> sm<sup>3</sup>/g) that is near to the maximum (6,3±0,2 Ma) of exposure age spread for 350 H-chondrites [10]. K/Ar age was estimated using the concentration of K=805±8 ppm and radiogenic

<sup>40</sup>Ar=(2,94±5)·10<sup>-5</sup> sm<sup>3</sup>/g. The age is 3,73±0,10 Ga which is near to the ages of some H-chondrites: Zaoyang (H5)-4,13±0,10 Ga, Lunan (H6)-4,15±0,10 Ga [3], Gueni (H4)-4,20 Ga, 474192 (H6)-4,10 Ga, 474418 (H6)-4,20 Ga [11], but the ages of other H-chondrites are older or younger.

Though the Dengle has some specifically petrological features different from most of the H-chondrites, it has no evidence of solar wind exposure as confirmed by our investigations of noble gases. That suggests that Dengle should not be interpreted as a regolithic breccia.

References: (1) Ivanova M.A. et al. (1992) Meteoritics v.27, N 4, c. 463-465, (2) Shultz L., Kruse H. (1983) Meteoritenforschung, p. 88, (3) Engster O. et al. (1987) EPSL, 84, p. 42-50, (4) Begemann F. et al. (1985) EPSL, 72, p. 247-262, (5) Takaoka N. et al. (1989), 44a, p.935-944, (6) Engster O. et al. (1968) (preprint), (7) Heymann D., Mazar E. (1968), GCA 32, p.1-19, (8) Ozima M., Podosek F. A. (1983) Cambridge University Press, (9) Freundel M. et al. (1986) GCA, 50, p.2663-2673, (10) Alekseev V. (1991) Astrologichesky vestnik (in Russian) 25, N 2, p.233-244, (11) Weber H.W. et al. (1982), 38a, p.262-272

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Tabl.1

Isotopic composition ( $^{130}\text{Xe}=1.000$ ,  $^{86}\text{Kr}=1.000$ ) and concentration of Xe, Kr, and Ar in the Dengli (H3.8), 365.1 mg.

T °C	$^{136}\text{Xe}$	$^{134}\text{Xe}$	$^{132}\text{Xe}$	$^{131}\text{Xe}$	$^{129}\text{Xe}$	$^{128}\text{Xe}$	$^{126}\text{Xe}$	$^{124}\text{Xe}$
800	2.24±4	2.63±7	6.47±13	5.18±10	7.11±14	0.472±20	0.0277±20	0.0222±20
900	2.20±8	2.40±10	6.63±25	5.43±22	7.25±30	0.504±40	0.0237±35	0.0247±35
1100	2.11±8	2.60±12	6.44±20	5.21±16	7.10±35	0.483±25	0.0232±30	0.0225±30
1300	2.01±3	2.38±3	6.24±7	5.08±5	8.34±8	0.507±10	0.0282±7	0.0272±7
1600	2.01±3	2.40±5	6.11±9	4.90±8	8.35±12	0.503±10	0.0293±12	0.0286±11
Total	2.04	2.41	6.24	5.06	8.17	0.502	0.0280	0.0269

T °C	$^{84}\text{Kr}$	$^{83}\text{Kr}$	$^{82}\text{Kr}$	$^{80}\text{Kr}$	$^{130}\text{Xe}$	$^{86}\text{Kr}$	$^{40}\text{Ar}$	$^{38}\text{Ar}/^{36}\text{Ar}$	$^{40}\text{Ar}/^{36}\text{Ar}$
					STP		STP		
					$\text{sm}^3/\text{g}\cdot 10^{-12}$		$\text{sm}^3/\text{g}\cdot 10^{-7}$		
600					1.7	11.5	0.5	0.225	1522±10
800	3.22±13	0.594	0.677±40	0.139±10	0.9	5.1	38.9	0.285	1755±28
900	3.17±16	0.653	0.550±44	0.136±14	1.7	6.0	54.5	0.451	5050±200
1100	3.12±15	0.662	0.659±33	0.148±14	19.2	35.6	81.6	0.410	5988±156
1300	3.23±7	0.707	0.664±20	0.164±6	7.3	22.5	153.0	0.340	1443±15
1600	3.27±7	0.695	0.694±20	0.145±6	30.8	80.5	52.9	0.271	369±4
Total	3.23	0.681	0.667	0.152			381.9	0.310	1287±12

