

LIQUID LINES OF DESCENT AND LIQUID IMMISCIBILITY IN HIGH Ti LUNAR BASALT; Malcolm J. Rutherford, Paul C. Hess, G.H. Daniel Dept. of Geological Sciences, Brown University, Providence, R.I.

Experiments have been performed on rock 70017,29 to derive a series of cogenetic salic liquids that represent potential end products of lunar basalt fractionated at low pressure. Rock 70017 (29 and 127) is a medium grained (< 3mm) porphyritic basalt composed of 22% ilmenite, 2% olivine (Fo₇₄₋₆₅), 49% Ca-pyroxene (zoned En₄₈Wo₃₅Fs₁₇ to En₃₁Wo₂₃Fs₄₆), 23% plagioclase (An₂₈Ab₁₁ to An₇₆Ab₂₁Or₃) and minor pigeonite (En₆₃Wo₇Fs₃₀), cristobalite and mesostasis. Glass containing exsolved Fe-rich spherules is present in the cristobalite. The composition of these immiscible melts (Table 1) indicates that they are closer to the critical state than those in Apollo 11 basalts (1) possibly due to higher temperature, or some compositional effect. In the experiments, chips of 70017,29 were wrapped in Mo foil, sealed in evacuated SiO₂ glass tubes and either partially melted or significantly melted and then recrystallized. Rocks with the composition of the resulting glasses were synthesized and used in subsequent experiments.

Representative compositions of the fractionated liquids (Table 1) show that as MgO decreases, TiO₂ decreases uniformly (Fig. 1) and SiO₂, FeO, MnO and K₂O increase. A liquid line of descent constructed with similar data for Apollo 12 basalts (2) contrasts with this, but where the paths intersect, the liquids have similar major element concentrations. If the highly differentiated Apollo 12 liquids were to be fractionated at this point, presumably the liquid would then follow a path similar to 70017.

The significant results of the experiments on 70017 are that (1) only minor SiO₂ enrichment was achieved in the highly fractionated liquids although 95% of the basalt was crystallized, and (2) the immiscible liquid field has been approached and probably reached from the iron rich side. This is demonstrated in Fig. 2, where the compositions of the two most fractionated liquids obtained (1-4 and 1-9) plot within the central field of liquid immiscibility in the K₂O-Al₂O₃-SiO₂-Fe-O System even though other components are included. More importantly, the tie lines defined by the two immiscible liquids in 70017 pass close by both these compositions. A simple calculation shows that in terms of all nine elements considered, the composition of the fractionated liquids in experiments 1-4 and 1-9 lie on

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the join between the 70017,127 immiscible melts at approximately 80% of the Fe-rich end member.

The geological significance of these results is that (1) the field of liquid immiscibility was apparently reached with a randomly chosen fractionation path and major SiO_2 enrichment is not necessary to achieve it* (2) The amount of melt present at the onset of liquid immiscibility can be several times the amount of SiO_2 rich mesostasis, the Fe-rich fraction being camouflaged as additions to Fe rich minerals. (3) The amount of total immiscible melt in a lunar soil could be as much as 4 times the amount of granitic component if the latter were the result of liquid immiscibility, a fact which has a bearing on the trace element content of the soil (3).

REFERENCES

- (1) ROEDDER E., and WEIBLEN P.W. (1971) Proc. of 2nd Lunar Sci. Conf., Vol. 1, p. 507-28.
- (2) BIGGAR G.M., O'HARA M.J., PECKETT A. and HUMPHRIES D. (1971) Proc. of 2nd Lunar Sc. Conf., Vol. 1, p. 617-43.
- (3) HESS P.C., and RUTHERFORD M.J., (1974, This Volume).
- (4) WEIBLEN P.W., and ROEDDER E., (1973) Proc. of 4th Lunar Sci. Conf., Vol. 2, p. 681-703.

* The presence of two immiscible liquids has not been confirmed, but submicroscopic exsolution may be present. Variations in SiO_2 counts on a similar glass produced in a somewhat longer experiment (3 days) indicate submicroscopic immiscibility.

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TABLE 1 Representative Liquid (Glass) Compositions

	70017		Rock ¹	70017,			Syn Expt. ²	
	Imm. melts			29-3	29-1	29,5	1-4	1-9
	H.Si	H.Fe						
SiO ₂	71.3	42.2	41.02	45.77	46.07	47.66	49.61	49.14
TiO ₂	1.9	4.1	11.04	9.35	7.08	4.85	3.51	2.71
Al ₂ O ₃	14.0	7.6	11.68	10.19	9.41	9.15	9.46	8.95
FeO	5.6	29.9	16.35	15.83	21.70	22.66	24.81	27.18
MnO	0.0	.46	.16	.25	.28	.35	.49	.57
MgO	.51	1.5	9.05	7.27	4.68	2.59	1.35	.68
CaO	2.6	11.6	10.02	11.65	9.49	10.48	10.03	9.15
Na ₂ O	.47	.02	.37	.19	.37	.29	.39	.30
K ₂ O	3.2	.55	.13	.17	.24	.28	.66	1.15
Total	99.58	97.93	99.82	100.67	99.31	98.31	100.31	99.83
Temperature °C				1152	1100	1070	1030	1006
Percent Melt				65.8	15.4	7.4	38.5	22.4

1. Calculated from 29-3, mineral and glass compositions and mode.
2. Experiments on synthetic rock of 29-1 composition.

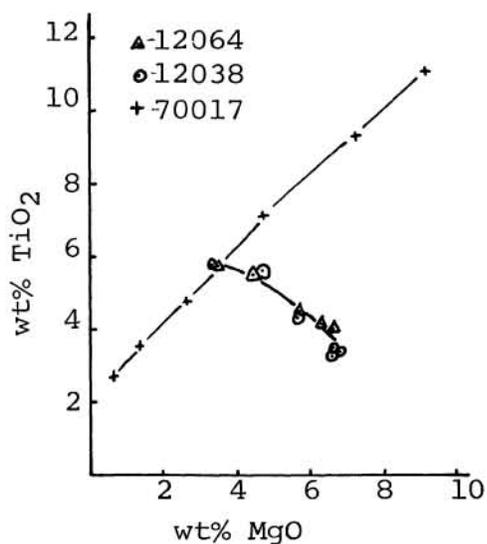


Fig. 1

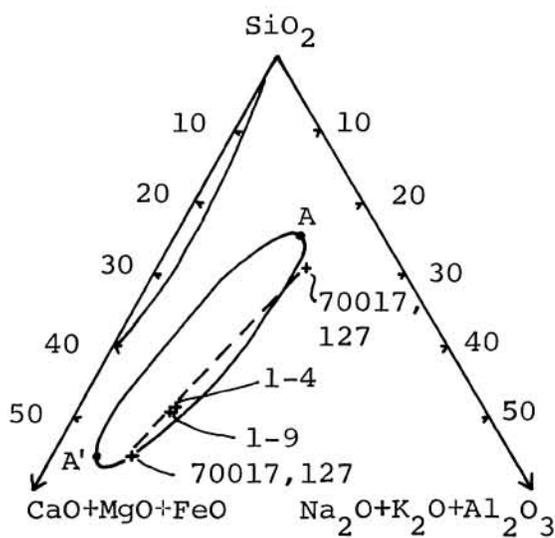


Fig. 2 (after (4))