

MAGNETITE-BEARING LAYERS IN ALLAN HILLS 84001 CARBONATE GLOBULES: BULK AND MINERAL COMPOSITIONS. A. H. Treiman¹ and L. P. Keller², ¹Lunar and Planetary Institute, 3600 Bay Area Boulevard, Houston TX 77058, USA (treiman@lpi.usra.edu), ²MVA Inc., 5500/200 Oakbrook Parkway, Norcross GA 30093, USA (lkeller@mva-inc.com).

Introduction: Controversy surrounds the origin of submicron magnetite grains in the carbonate globules of ALH 84001. The magnetites may have been emplaced *chemically* by decarbonation/oxidation of original Fe-rich carbonates [1], or *physically* by concentration of grains formed elsewhere [2]. The validity of these mechanisms can be tested through chemical analyses of bulk magnetite-rich (mt-rich) layers and of minerals within the layers.

Analyses: “Bulk” compositions were analyzed by electron microprobe (SX-100 at JSC) for Ca, Mg, Fe, and Mn (3 nA focussed beam, 15 kV, 30–40 s on peak). Mt-rich layers are fine-grained compared to the path lengths of electrons and X-rays, so that ϕ - ρ -Z corrections can be applied. TEM analyses of carbonate grains (10 nm probe diam.) and “bulks” (0.5 μ m probe diam.) were from a ~70-nm thick ultramicrotomed section (JEOL 2010 at MVA; 200 kV; Noran thin window EDS). Counting statistical errors were <3%; Cliff-Lorimer k-factors were derived from well-characterized standards.

Results: Microprobe analyses (TS ,145), as in earlier reports [3], give low-Ca carbonates in a linear array from ~Fe₀Mg₉₅Ca₅ to ~Fe₅₀Mg₃₉Ca₁₁ with nearly constant Ca/Fe. TEM analyses (split ,255) of carbonates outside mt-rich layers are consistent with EMP data (Fig. 1).

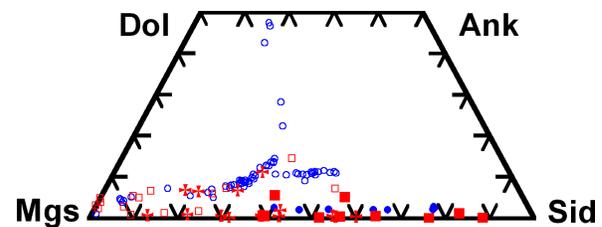
TEM analyses of individual carbonate grains within the mt-rich layers fall in two groups: most are nearly Ca-free (and Mn-free) from Mg₈₆Fe₁₄ to Mg₄₀Fe₆₀; a few are like carbonates outside mt-bearing layers ~Ca₇Mg₆₅Fe₂₈ (Fig. 1). Other phases in the mt-rich layers include Ca-bearing brucite (outer mt-rich layer), a Fe-S mineral, Mg-Fe phyllosilicate (smectite?), and Fe-rich amorphous material containing Si and S (latter two in inner mt-rich layer). Ca-rich phases were sought and not found.

TEM “bulk” compositions of mt-rich layers are qualitatively consistent with mixtures of these phases. Re-examination of EMP data showed several analyses (earlier rejected for high S and/or Si contents) which qualitatively match the TEM “bulk” analyses (Fig. 1). The concordance of TEM and EMP analyses suggests that they approximate the true bulk compositions of the mt-rich layers.

Inferences: The mt-rich layers are not merely mixtures of magnetite and magnesian carbonate, and so cannot be explained by simple physical or chemical emplacement of the magnetites. The presence of brucite is most easily explained as the hydration product

of Ca-bearing periclase. Perhaps this hydration episode also involved deposition of the clay(?) and amorphous Fe-rich material. But the putative periclase could have formed during decarbonation/oxidation of ferroan carbonate (chemical emplacement mechanism) or during decarbonation of calcian magnesite adjacent to nano-phase magnetites (physical emplacement mechanism).

Figure 1. Chemical compositions in Ca-Mg-Fe. Circles are by EMP; angular symbols are by TEM. Open symbols are 'massive' carbonate; filled symbols are 'bulk' of magnetite-rich layers; crosses are carbonates within magnetite-rich layers.



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References: [1] Brearley et al. (1998) *LPS XXIX*, Abstract #1451. Bradley et al. (1998) *MAPS*, 33, 765; Golden et al. (2000) *LPS XXXI*, Abstract #1799. [2] McKay D. et al. (1996) *Science*, 273, 924; Friedman et al. (1998) *Workshop Martian Meteorites*, p. 14, LPI Contrib. No. 956; Kirschvink et al. (1999) *LPS XXX*, Abstract #1681; Thomas-Keprta et al. (2000) *LPS XXXI*, Abstract #1683. [3] Harvey and McSween (1996) *Nature*, 382, 49; McKay G. and Lofgren (1977) *LPS XXVIII*, 921. Scott et al. (1998) *MAPS*, 33, 709; Golden et al. (2000) *MAPS*, 35, 457.