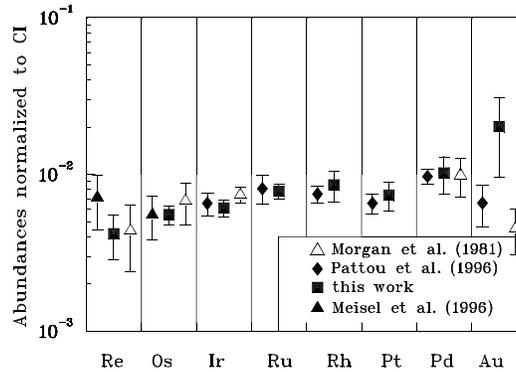


## COMPOSITION OF THE LATE INFLUX OF THE EARTH.

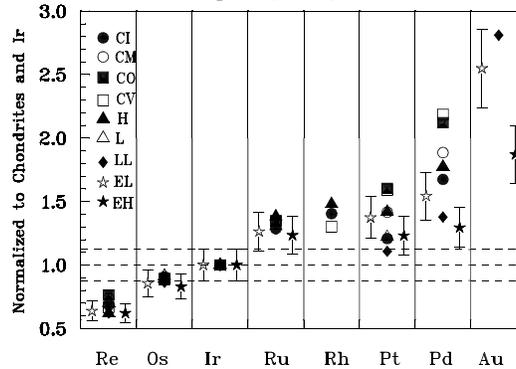
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**Introduction:** Many authors explain the relative “high” abundances of HSE and their broadly chondritic proportions in the Earth’s mantle (PUM) by the addition of a late chondritic veneer after core formation. Many recent studies have documented significant regional variations in absolute HSE abundances and inter-element ratios in mantle samples. Such variations may have been caused by complex geochemical processes such as partial melting, melt percolation and aqueous metasomatism in the subcontinental lithosphere [1]. For a better characterisation of the late veneer component(s) of the Earth I review here selected neutron activation data from our own studies for orogenic spinel lherzolites [2,3] that have suffered only slight melt depletion ( $Ca/Si > 0.086$ ) and compare this data with selected data from the literature [4-6].

**Results:** Abundances of HSE in PUM sampled by lherzolites from our studies [2,3] are as follows;  $0.16 \pm 0.05$  ng/g Re;  $2.69 \pm 0.38$  ng/g Os,  $2.80 \pm 0.34$  ng/g Ir,  $5.60 \pm 0.61$  ng/g Ru,  $1.20 \pm 0.27$  ng/g Rh,  $7.33 \pm 1.55$  ng/g Pt,  $5.68 \pm 1.52$  ng/g Pd and  $3.06 \pm 1.60$  ng/g Au [7].



In a large number of „fertile“ samples [2-6] the abundance distribution of the HSE is remarkably uniform with slightly increasing abundances with decreasing refractory character of the elements from Re to Pd, except Pt (FIG. 1).



In FIG. 2 mantle HSE abundances (this work) are plotted normalized to Ir and chondrites. The HSE systematics show that the late veneer closely resembles E-chondrites or LL-chondrites [7,8].

**References:** [1] Schmidt G. et al. (2003) *Chem. Geol.*, 196, 77-105. [2] Lorand J.P. et al. (2000) *Lithos*, 53, 149-164. [3] Snow J.E. et al. (2000) *EPSL*, 175, 119-132. [4] Morgan J.W. et al. (1981) *Tectonophysics*, 75, 47-67. [5] Meisel T. et al. (1996) *Nature*, 383, 517-520. [6] Pattou L. et al. (1996) *Nature*, 379, 712-715. [7] Schmidt G. (2003) *Annual Report 2002, Institut für Kernchemie, IKMz 2003 - ISSN 0932-7622*. [8] Morgan J.W. et al. (2001) *Meteorit. Planet. Sci.*, 36, 1257-1275.