

**A STUDY OF THE FRAGMENT DISPERSAL AND TRAJECTORY OF THE SAYH AL UHAYMIR 001 METEORITE SHOWER.** A. V. Korochantsev<sup>1</sup>, D. A. Sadilenko<sup>1</sup>, M. A. Ivanova<sup>1</sup>, C. A. Lorentz<sup>1</sup>, and E.V. Zabalueva<sup>1</sup>, <sup>1</sup>Vernadsky Institute of Geochemistry and Analytical Chemistry, Moscow 119991, Russia (korochantsev@geokhi.ru).

*Sayh al Uhaymir 001 (SAUH 001) is one of Oman's largest known meteorite showers. We obtained hundreds of GPS find locations, analyzed the fragment distribution, and developed a model of the meteorite's trajectory.*

SAUH 001 is a stony meteorite shower (L4/5) found March 16, 2000. More than 2670 samples weighing more than 450 kg have been collected. The collected samples may be a large portion of the total mass of the meteorite body, because the surface in this region is hard and it is likely that only a small portion of the fragments penetrated the surface. GPS locations were obtained for 748 found samples to within 0,001 minutes (~2 meters) (Fig. below). We assume that the fragments were not moved after falling.

The ground track of the meteorite was modeled as follows: (1) Determined center of mass (CM) of the fragment distribution. (2) Fitted a line, weighted for fragment mass, through the CM along the long axis of the distribution in both directions. Increasing fragment mass along track indicated a flight direction of 233°.

The fragments' mass distribution was used to model the meteorite trajectory. In the model, the distribution depends on the speed of entry into the atmosphere ( $v$ ), angle to the surface ( $\alpha$ ), and breakup height ( $H$ ). We developed a computer program to calculate trajectories based on the distribution. For seven mass classes, we calculated the average mass and position along the ground track. We then varied  $v$ ,  $\alpha$  and  $H$  to find solutions closest to the observed locations of the mass classes. To limit the solutions we used the relationship between  $v$  and  $H$  at the destruction air flux  $A=(\rho v^2)_{\max} \cdot H = h \cdot \ln(A/v \cdot \rho_0)$ , where  $\rho_0$  is the sea-level air density and  $h$  is the scale height.  $A$  was taken as  $6.5 \times 10^6$ , the average of several measurements for L chondrites [1]. The model yielded trajectory angle of 70-75° and breakup heights of 40 km with a meteorite velocity of 27 km/s, 30 km altitude with  $v=15$  km/s, and 25 km with  $v=11$  km/s. The meteorite may have broken up in several stages. For example, Pribram broke up gradually from 44 km down to 23 km altitude [2]. SAUH 001 samples show evidence of multi-stage fusion crust formation, suggesting multi-stage breakup.

**Conclusion:** Our model indicated the SAUH001 bolide travelled northeast to southwest, bearing 233°. Trajectory angle was ~70°, with a probable velocity of 27 km/s. Our model gave a range of breakup altitudes similar to observations of Pribram. Breakup appears to have occurred in several stages.

**References:** [1] Tseplekha Z. (1961) *Meteoritika*, 30, 178-184 (Russ.); [2] Tsvetkov V. I. (1985) *Thesis*, 120-122 (Russ.).

