

**NOBLE GASES IN A GIBEON IRON METEORITE
FRAGMENT HEAVILY SHIELDED TO COSMIC-RAYS**

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Introduction: Gibeon iron meteorite classified as IVA has been discovered as a large number of fragments in a strewn field in Namibia, Africa. Total weight of the meteorite becomes about 21 tons [1], and its preatmospheric body was estimated to be larger than 3 meters in radius (applying the model calculation from [2] for our data). The scarcity of cosmogenic noble gases may provide us with a chance to investigate primordial component in iron meteorite parent body. We will report noble gas data recently obtained for a plate-shaped sample (about $240 \times 170 \times 8$ mm in size), which shows extraordinary low concentration of cosmogenic noble gas, e.g., $^3\text{He} < 10^{-12}$ cm³STP/g [3]. This belongs to a group with relatively short cosmic-ray exposure age ($\sim 10^7$ y) [4]. Our purpose of this study is to investigate the primitive noble gas component in iron meteorite parent body.

Noble gas analysis: Meteorite pieces weighing 90–130 mg each, drilled out from 8 points in the sample plate, were analyzed for noble gases on a modified-VG5400 (MS-III). We used a Mo crucible of which inner wall was coated with zirconium oxide to avoid melting of the crucible by forming alloy with the iron meteorite.

Results and Discussion: Measured $^3\text{He}/^4\text{He}$ ratios were as low as $(3-5) \times 10^{-4}$, which resemble isotopic ratio of “primordial helium”. Several cases can be considered for the low $^3\text{He}/^4\text{He}$: 1) mixing of cosmogenic helium and radiogenic ^4He , 2) true primordial helium, and 3) air contamination. Combination of these cases can also be considered. If ^3He is totally cosmogenic, cosmogenic ^3He was generated as low as $(2.2 \pm 0.4) \times 10^{-13}$ cm³STP/g in a time spans of 10 My [4]. Concentrations of ^4He were in the order of 10^{-9} cm³STP/g, homogeneously distribute in the sample plate. Given that the ^4He was produced from U, Th (Th/U = 3) for 4 Ga, concentration of uranium can be estimated as low as 0.5 ppt. Contamination of terrestrial air is improbable, because $^4\text{He}/^{20}\text{Ne}$ ratios in our sample are >1.3 much larger than the atmospheric value of 0.32 and contamination generally occurs selectively in heavier noble gas.

References: [1] Buchwald V. 1975. Handbook of Iron meteorites Vol. 2, 584-593. [2] Leya I. et al. 2002. Meteorit. Planet. Sci. 37, 1015-1025. [3] Nagao K. and Honda M. 2005. 29th Symp. Antarct. Meteorites. 49-50 (NIPR). [4] Honda M. et al. 2007. Meteorit. Planet. Sci. 42: A68.