

**NANOSIMS ELEMENT MAPPING OF ACFER 094**

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**Introduction:** Acfer 094 is a type 3.0 ungrouped carbonaceous chondrite with affinities to CO and CM type chondrites. It has a fine-grained matrix composed of Fe,Ni sulphides, low-Ca-pyroxenes, forsteritic olivines and amorphous material [1]. Matrix plots with elements in order of increasing volatility suggest Acfer094 is compositionally closer to CM than CO [2]. Although a primitive specimen, Acfer 094 shows evidence of terrestrial weathering [3] Given the extreme fine grain-size of matrix components, in the current work we have used NanoSIMS for element mapping (data acquired at The Centre for Microscopy, Characterisation and Analysis (CMCA) at the UWA).

**Method:** A thin section of Acfer 094 was gold coated (6nm) and nine 30 $\mu$ m x 30 $\mu$ m areas were sputtered for element mapping in a Cameca NanoSIMS 50 ion microprobe. A Cesium (Cs<sup>+</sup>) ion source with a resolution of 50nm and an Oxygen (O<sup>-</sup>) ion source with a resolution of 200nm were used to analyse both electro-negative (H-, C-, O-, N-, F-, Cl-, P-, Te-, Ni- & S-) and electro-positive (Mg<sup>+</sup>, Al<sup>+</sup>, Si<sup>+</sup>, Ca<sup>+</sup>, Fe<sup>+</sup>, Co<sup>+</sup> & Ni<sup>+</sup>) elements.

**Results:** Over fifty individual element maps were arithmetically combined using ImageJ software highlighting a series of matrix features on a micron and sub-micron scale: Ni-veins, Ni-rings, S-inclusions,  $\mu$ CAIs, Cr, Ni & Ti hotspots, as well as surface contaminants, rust veins & other weathering features.

**Discussion:** Among many unusual features observed during the course of this analysis, one interesting result was the identification of two isolated areas showing sub- $\mu$ m sulphide inclusions (Figure 1), embedded in a silicate background material, which incorporates minor FeO and Ni. Similar features - described as metal/sulphide-silicate shock-melt in CBs QUE94411 and HAH237 - were found interstitial between silicate and metal [4]. However the inclusions in Acfer are significantly smaller and sulphide is more angular. Their origin, and additional sub- $\mu$ m matrix features revealed by this technique, will be discussed at the conference.

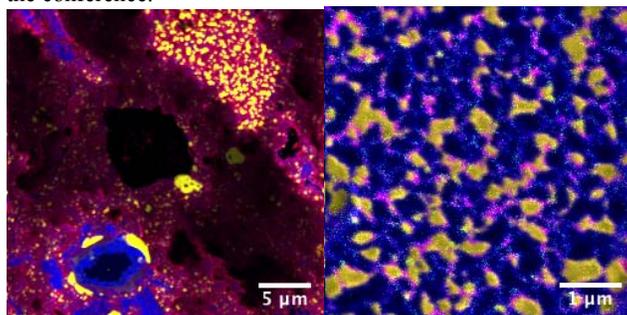


Figure 1: *Left:* O=blue, S=yellow, Si=red. *Right:* Close-up of S inclusion; O=blue, S=yellow, Ni=green & FeO=red

**References:** [1] A. Greshake. 1997. *Geochemica et Cosmochimica Acta* 61:437-452 [2] P. A. Bland. 2005. *Proceedings of the National Academy of Science* 102:13755-13760 [3] J. T. Wasson & A.E. Rubin. 2010. *Meteoritics & Planetary Science*. 45:73-90 [4] A.Meibom et al. 2000. *31<sup>st</sup> Lunar and Planetary Science Conference*. Abstract #1420