

LUNAR ORIENTALE BASIN: ESTIMATION OF IMPACT CONDITIONS.

J. Echaurren¹. ¹Codelco Chile Chuquicamata Division. E-mail: jecha001@codelco.cl.

Introduction: The lunar Orientale basin, with a crater-rim diameter in excess of 900 km, is the best-preserved and youngest multi-ring basin on the Moon. It is located on the Moon's western limb and is centered at approximately 20°S, 265°E [1]. In this work are estimated the possible impact conditions, that could have given origin to this crater [2]. The models used here are based on: some equations postulated by *Holsapple* (crater depth) [4]; scaling; polynomial analysis; and an adaptation of quantum formalism for the mathematical representation of the energy pulse generated in the impact point, in where besides, is used one solution (soliton type) of the *Korteweg-De Vries's* equation [3].

Results obtained with the models: The development of this crater is realized in 4 stages [5], in which are specified the variables of impact more common [5], as follows: **a). Contact/Compression Stage:** In this stage the diameter of the impactor is estimated in ~ 111.62 km, the velocity of impact is ~ 5.53 km/s, the impact angle is $\sim 80.30^\circ$, the density of impactor is ~ 0.75 g/cm³, the crater depth is estimated in ~ 8.84 km [4], the melt volume is $\sim 179,191.6$ km³, the total energy of impact is estimated in $\sim 6.59 \times 10^{31}$ Erg ($\sim 1.57 \times 10^9$ megatons), pressure to 1 km of the impact point is $\sim 237,472.5$ Gpa, and the seismic shock-wave magnitude is >10.0 according the *Richter Scale*. **b). Modification/Excavation Stage:** In this stage the diameter of transient crater is ~ 623.56 km, the number of ejected fragments is $\sim 4.65 \times 10^{12}$, the average size of the fragments is ~ 12.53 m, the average density of fragments is ~ 1.74 g/cm³, the distance of ejection of the fragments is ~ 93.29 km, the velocity of ejection is ~ 672.33 m/s, the minimal angle of ejection is $\sim 9.76^\circ$, and the minimum height of ejection is ~ 4.01 km. **c). Collapse/Modification Stage:** In this stage the pressure toward the final crater rim decrease to ~ 1.1 Gpa. **d). Final Crater Stage:** The relation between the transient crater and the final crater is ~ 0.67 [5], the time of creation for the final crater can be estimated in ~ 4.46 minutes according to *Schmidt* and *Housen* [4], the hydrothermal zone could spread from ~ 311.78 km to ~ 440.8 km from the nucleus of impact, i.e., a hydrothermal band of ~ 129.02 km, the lifetimes estimated for this hydrothermal band are of ~ 16.23 Ma to ~ 25.34 Ma with uncertainties of $\sim (\pm) 0.74\%$ to $\sim (\pm) 2.08\%$, i.e., from $\sim (\pm) 0.12$ Ma to $\sim (\pm) 0.53$ Ma, hydrothermal temperatures from 0.25 years to 1,400 years, after of the impact, are estimated in ~ 135.41 °C to ~ 54.59 °C, the final temperature to the 25.34 Ma after of the impact is calculated in ~ 4.04 °C + environment temperature. In this stage is also possible to estimate (hypothetically) the diameter of one dust cloud of ~ 399.78 km.

References: [1] Nahm A. L. and Kring D. A. 2011. Abstract #1172. 42nd Lunar & Planetary Science Conference. [2] Echaurren J. C. and Ocampo A. C. 2003. *Geophysical Research Abstracts*, Vol. 5, 04450, EGS-AGU-EUG Joint Assembly. [3] Echaurren J. C. 2010. Abstract #5071. 73rd Annual Meteoritical Society Meeting. [4] Holsapple K. A. Theory and Equations for "Craters from Impacts and Explosions". [5] French B. M. 1998. *Traces of Catastrophe: A Handbook of Shock-Metamorphic Effects in Terrestrial Meteorite Impact Structures*. LPI Contribution No. 954, Lunar and Planetary Institute, Houston. 120 pp.