

MAss Spectrometer for Planetary EXploration

# MASPEX/Europa

OPAG @ JHU/APL

J. Hunter Waite and the MASPEX Europa Team

August 24, 2015

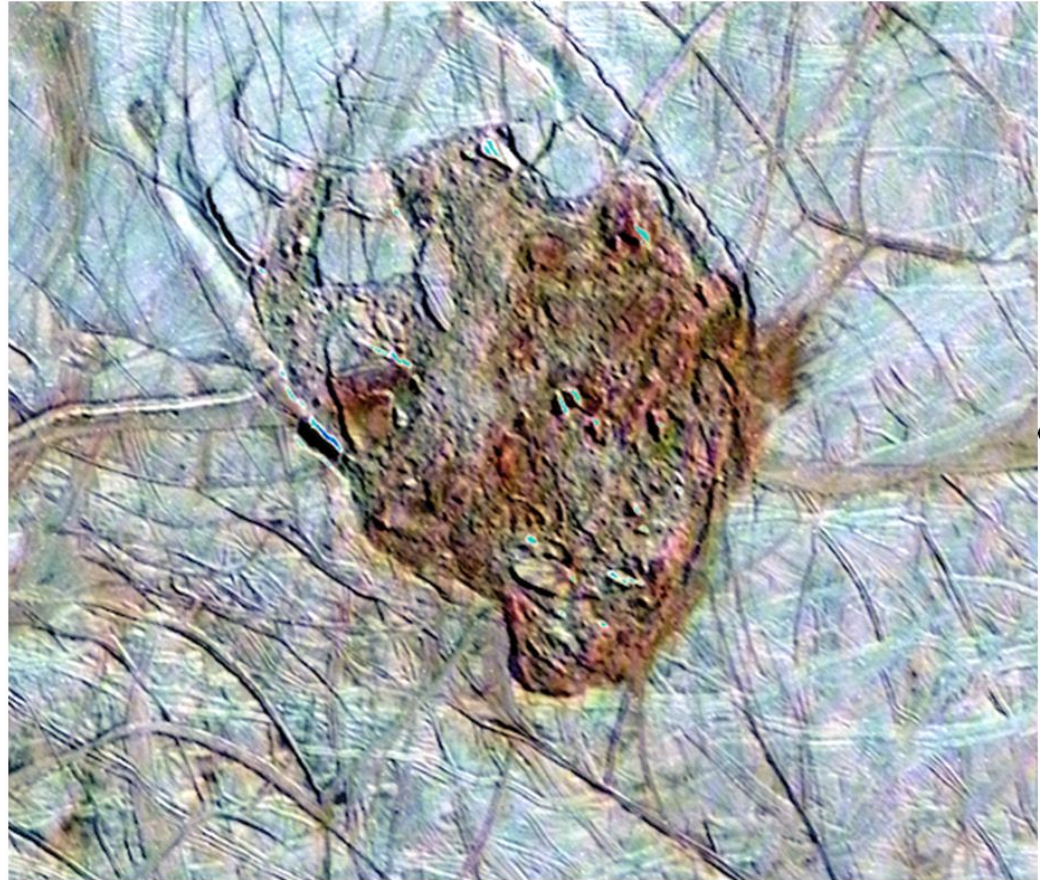
# Science Traceability Matrix

Goal	Science Objectives	Observables	Scientific Measurements
			Physical Parameters
To constrain conditions for habitability on Europa through high-resolution, high-sensitivity measurements of the volatile composition of the exosphere and possible plumes.	Determine the distribution of major volatiles and key organic compounds in Europa's exosphere/plumes and their association with geological features	1A) Density variations of major volatiles ( $H_2O$ , $O_2$ , $H_2$ , $CH_4$ , $C_2H_4$ , $C_2H_6$ , HCN) on a global scale to provide exospheric structure, plume identification (both during and after active plume release), and correlation with geological features.	Determine density variations of $H_2O$ , $O_2$ , $H_2$ , $CH_4$ , $C_2H_4$ , $C_2H_6$ , and HCN, where mixing ratios exceed 1 ppm, to 25% precision with a spatial resolution of 100 km S/C track (23 s @ 4.43 km/s).
		1B) $H_2O/^{40}Ar$ ratio to determine surface-interior overturn.	Detect and measure $^{40}Ar$ at mixing ratios >1 ppb.
	Determine the relative abundances of key compounds to constrain the chemical conditions of Europa's ocean	2A) Densities of $H_2O$ , $H_2$ , $CH_4$ and their D/H ratios from plume or outgassing sources to establish the level of communication with an internal aqueous system.	Determine D/H in $CH_4$ , $H_2$ , and $H_2O$ to 5% precision.
		2B) The molecular and isotopic composition of the volatiles $NH_3$ , $N_2$ , CO, $CH_4$ , $CH_3OH$ , $C_2H_5OH$ , $CH_3CO_2H$ , $C_2H_5NH_2$ , and $CO_2$ , from plume or outgassing sources that serve as indicators of redox state and degree of redox disequilibrium in the internal ocean.	Measure the volatiles $NH_3$ , HCN, $N_2$ , CO, $CH_4$ , $CH_3OH$ , $C_2H_5OH$ , $CH_3CO_2H$ , $C_2H_5NH_2$ , and $CO_2$ , for mixing ratios >10 ppb, to 25% precision.
		2C) The radiolysis products ( $NH_3$ , HCN, $N_2$ , CO, $CH_4$ , $CH_3OH$ , $C_2H_5OH$ , $CH_3CO_2H$ , and $CO_2$ ) of surface compounds and plume deposits to divine likely synthesis pathways and surface residence times.	Where mixing ratios are $>1 \times 10^{-3}$ measure isotopic ratios to precisions of 0.5% [C]; 30% [H]; 0.5% [O]; 25% [N]
		2D) Abundances for the $C_1$ - $C_8$ hydrocarbons: $CH_4$ , $C_2H_2$ , $C_2H_4$ , $C_2H_6$ , $C_3H_4$ , $C_3H_6$ , $C_3H_8$ , $C_4H_8$ , $C_4H_{10}$ , $C_5H_{10}$ , $C_5H_{12}$ , and $C_8H_{18}$ from plume or outgassing sources to determine the extent of synthesis, the hydrothermal temperature, and degree of equilibrium among species.	Measure abundances to 25% precision where mixing ratios >10 ppb.

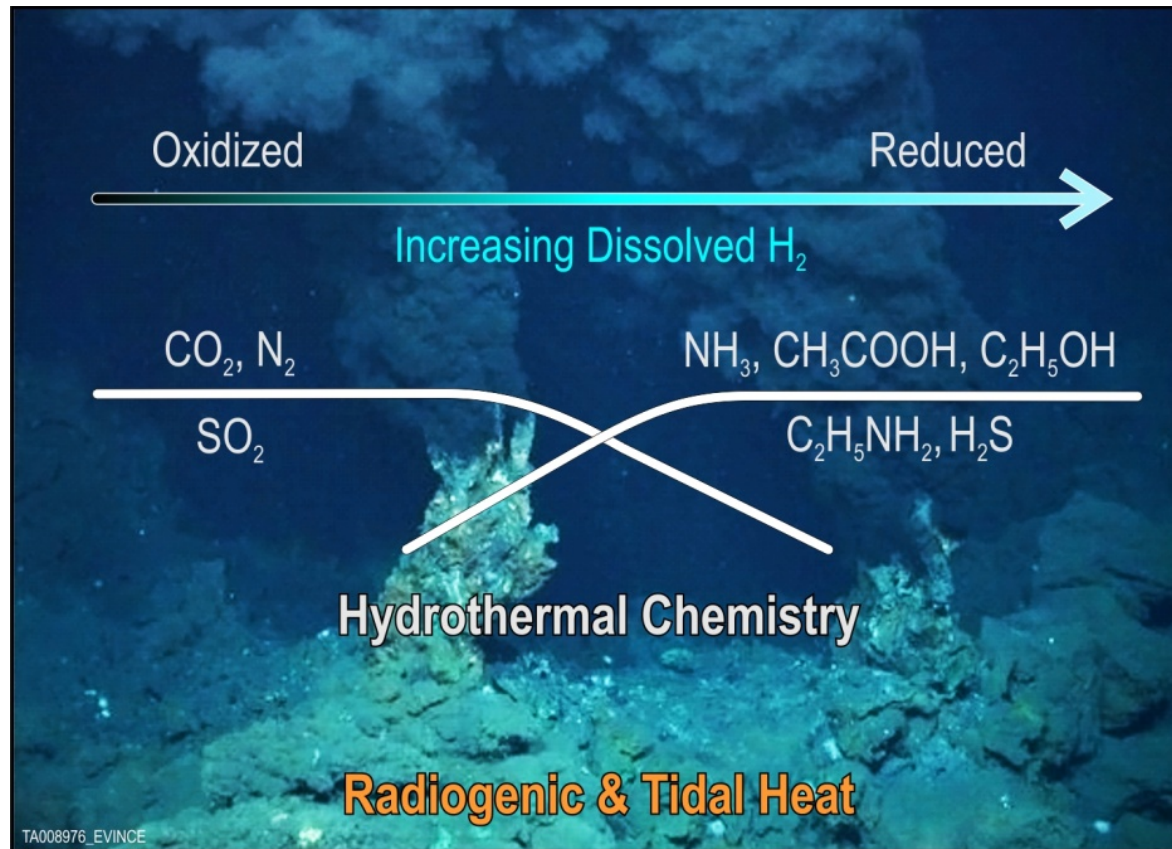


# Surface Composition

Thera Macula, an example of disrupted, “chaotic” terrain that may be the site of exchange between the Europa’s surface and a subsurface liquid layer. MASPEX measurements of exospheric composition will be correlated with such geological features and contrasted with those from older, less disturbed terrain to provide insight into possible interior- surface exchange.



# Ocean Chemistry

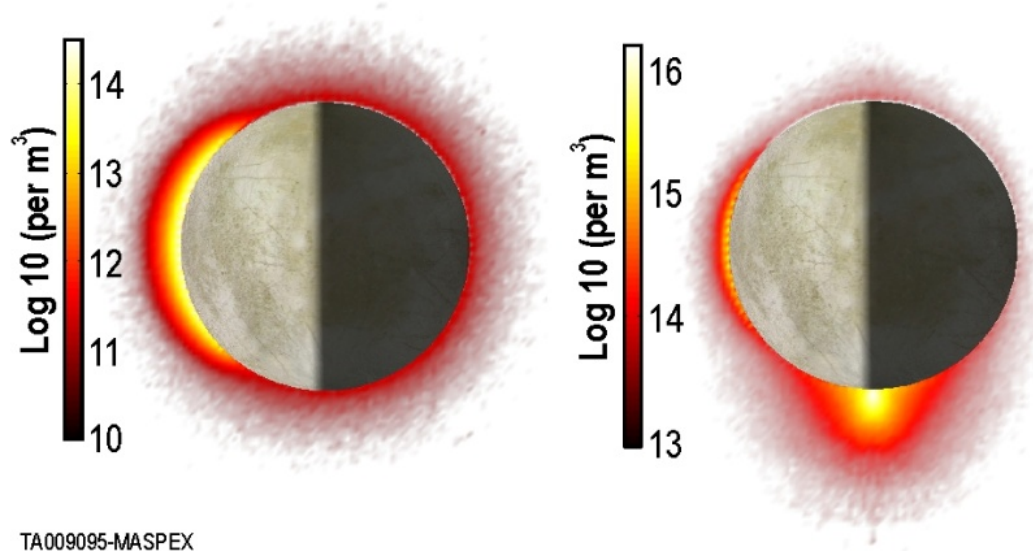


- The oxidation state and temperature of the water-rock system determines the products formed from dissolved aqueous C, N, S, etc.
- Typical terrestrial hydrothermal systems operating at moderate temperatures favor reducing conditions and the synthesis of organic compounds



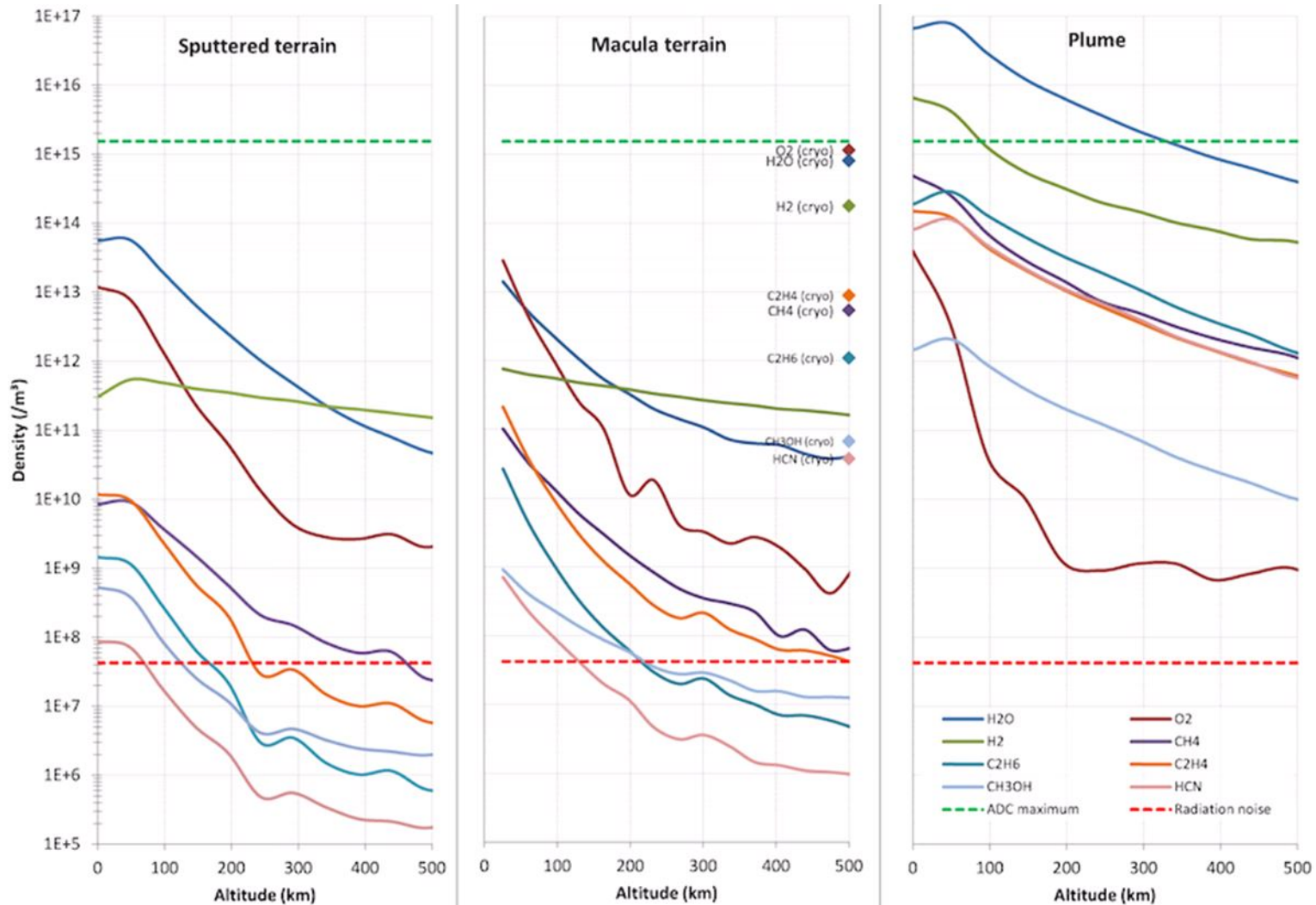
# Global Exosphere Modeling

H<sub>2</sub>O density



- Energetic particle sputtering combined with surface sublimation
- Creates a significant density enhancement over the sunlit region
- Sputter-produced exosphere combined with a south polar plume
- Plume material is transported equatorward and deposited at lower latitudes, where it is sputtered together with older material.

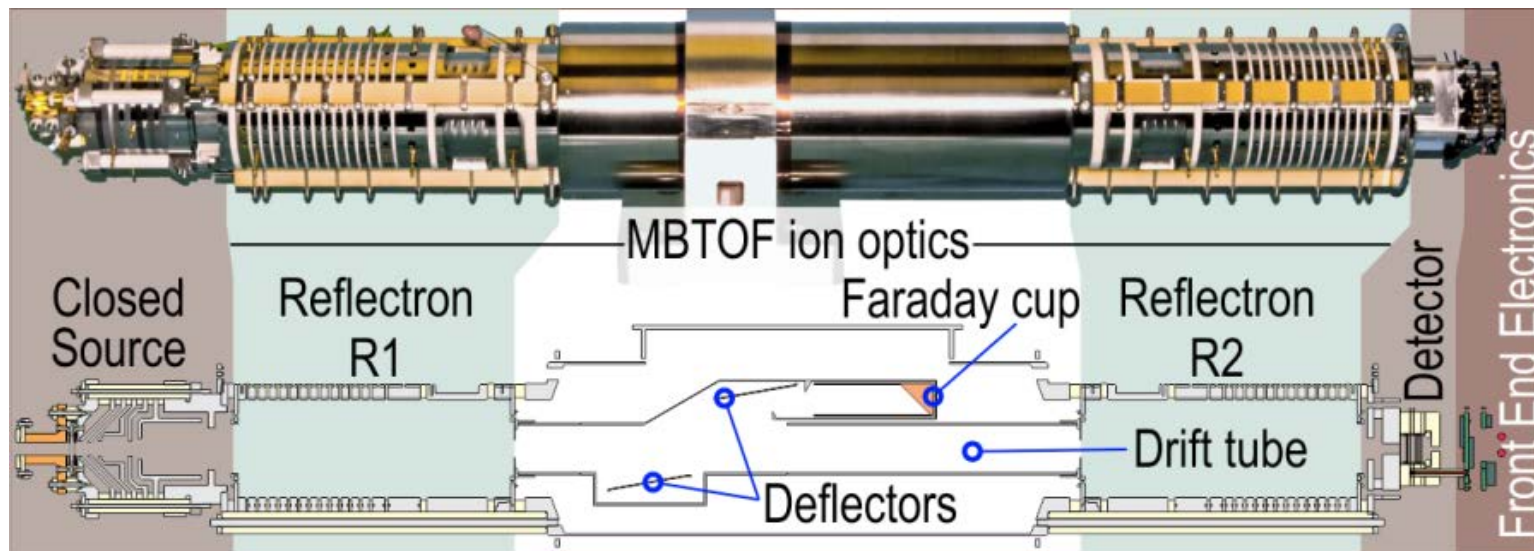
# Modeled Densities



Density profiles for representative species in three different European exospheric environments

# Instrument Characteristics

System Characteristics		
Subsystem	Type	Specification
GIS - thermalizer	Spherical antechamber, single use cover	FOV: $\pm 90^\circ$ in Y, $\pm 25^\circ$ in X, + Z boresighted
GIS - cryotrap	Cryotrapping/direct inlet, neutral gas	Sensitivity boost: up to $200 \text{ counts s}^{-1} \text{ per mol cm}^{-3}$
MS - closed source	Storage, electron ionization, redundant emitters	Ambient sensitivity: $0.02 \text{ counts s}^{-1} \text{ per mol cm}^{-3}$
MS - ion optics	Multibounce time-of-flight	Resolution: $24,540 \text{ m}/\Delta m$ 10% valley definition
MS - detector	Electron multiplier	MCP, 3 plate z-stack, $2 \mu\text{m}$ pore
Resources - Current Best Estimate (CBE)		
Mass & Volume	Power (W)	Average Data Rate
11.71 kg, instrument, 4.5 kg harness, 4.0 kg radiation shielding, $0.0203 \text{ m}^3$	45.4 (Peak), 35.4 (Average), 8.0 (Standby), 4.5 (Survival)	57 kbps (peak), 14 kbps (average)

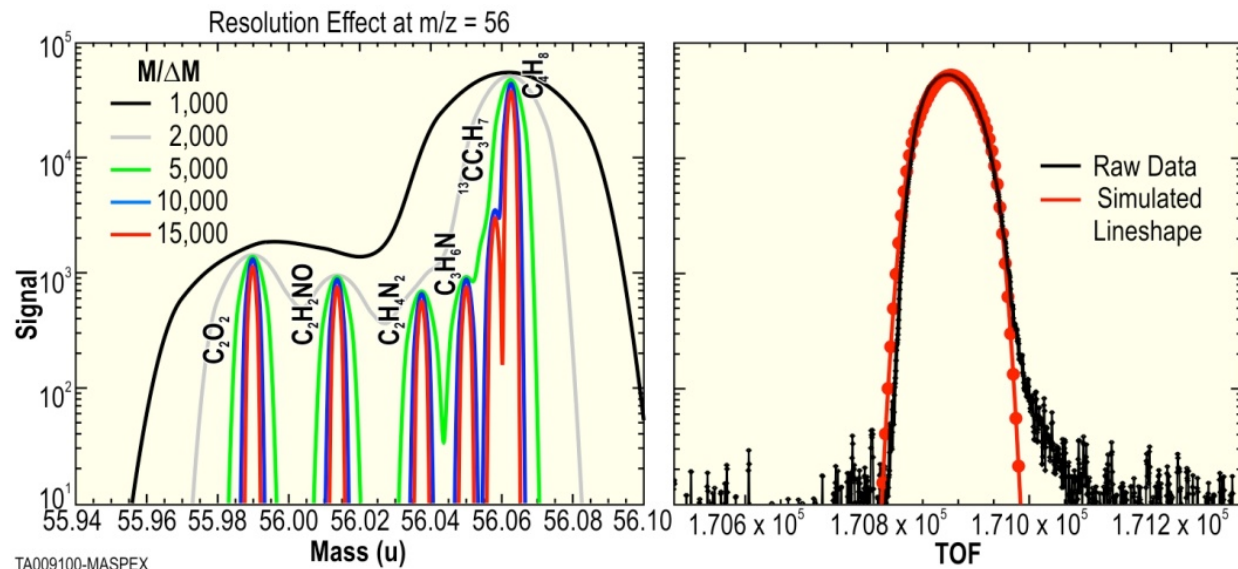


MBTOF cross section and photo of the third-generation instrument identifying the arrangement of the major components



# The Importance of Mass Resolution

Studies of the Enceladus plume hint at the presence of a large number of possible organic compounds whose identification is hindered by the low resolution of Cassini INMS. The simulation (left) demonstrates how increasing resolution enables the individual fragment ions (created by electron impact ionization within the source) of parent organic molecules to be fully separated and quantified. The spectrum (right) shows how the peak shape used in the simulation derived from MASPEX laboratory data.



A comprehensive simulation of the expected Europa environment was performed in the same way to identify the optimum analyte fragments and the resolution necessary for full separation. This information is used to define MASPEX's science operation modes.



# Science Operations Modes

Obj.	Mode	Description	Characteristics	Data (kbits)	Driving Requirements
1	SOM-A	Ambient mode. Specific chemical composition analysis with spatial resolution	23 s of varying length co-added spectra in each of 4 ROIs, mass range 2-32 u, m/Δm 6,948.	3,723 /flyby	Spatial variations – selected species every 23 s. Mass range – 2-32 u.
2	SOM-T	Cryotrapping mode. Prior to mode C, parallel to mode A.	Traps sample on adsorber at ~65K; Sample density determines trap time	n/a	Sensitivity: 400× compression factor boosts <sup>40</sup> Ar to 1×10 <sup>6</sup> for SOM-C.
2	SOM-C	Cryotrapped sample survey. Enables method updates to suit environment. Comprehensive mass survey, ppm mixing fraction species.	0.5 s co-added spectra in each of 34 ROIs, mass range 2-200 u, m/Δm ~12,000, total time 17 s.	74,834 / sample	Sampling time – 1 per flyby. Mass range – 16-114 u. Resolution – C <sub>2</sub> HD needs m/Δm 23,822.
		Cryotrapped sample analysis. Targeted ROI method minimizing analysis time and data volume.	2028 s of varying length co-added spectra in each of 13 ROIs, mass range 16-114 u, m/Δm up to ~23,822.	20,607 / sample	<sup>40</sup> Ar analysis drives dynamic range to 1×10 <sup>9</sup> , and detection limit to 1×10 <sup>6</sup> /m <sup>3</sup> .
Note: Obj. = Science Objective. Data are telemetry volumes, compressed data in transmission packets; CBE. m/Δm uses 10% valley definition.					

On every flyby MASPEX executes a series of ambient measurements (SOM-A) and cryotrap an exospheric sample (SOM-T). The SOM-T sample is held and analyzed at a point in the orbit where the radiation-induced noise on the detector is at low levels.

# Method Characteristics for SOM-C

Anticipated measured compounds for both the ambient (SOM-A) and cryotrapped (SOM-C) modes.

ROI	Integration time (s)	Analyte	Mass (u)	Resolution (m/ $\Delta$ m)
1	16.71	HCN	27.01	3,690
		C <sub>2</sub> H <sub>4</sub>	28.03	6,948
		C <sub>2</sub> H <sub>6</sub>	30.05	6,948
2	0.55	O <sub>2</sub>	31.99	3,042
3	1.36	CH <sub>4</sub>	16.03	1,464
		H <sub>2</sub> O	18.01	91
4	4.37	H <sub>2</sub>	2.02	31
		HD	3.02	37

Note: m/ $\Delta$ m uses 10% valley definition.

ROI	Integration time (s)	Analyte	Mass (u)	Resolution (m/ $\Delta$ m)
1	26.20	CO( <sup>13</sup> C)	29.00	35,800*
		N <sub>2</sub> ( <sup>15</sup> N)	29.00	73,805*
		C <sub>2</sub> H <sub>6</sub> ( <sup>2</sup> H)	31.05	25,158*
2	4.67	C <sub>3</sub> H <sub>8</sub> ( <sup>13</sup> C)	45.07	11,699
		C <sub>3</sub> H <sub>8</sub> ( <sup>2</sup> H)	45.07	37,120*
		C <sub>2</sub> H <sub>6</sub> O	46.04	10,539
3	27.26	CO( <sup>18</sup> O)	30.00	33,113*
4	8.13	C <sub>2</sub> H <sub>2</sub> ( <sup>2</sup> H)	27.02	23,822
5	57.40	NH <sub>3</sub>	17.03	1,154
		CH <sub>4</sub> ( <sup>13</sup> C)	17.03	4,351
		CH <sub>4</sub> ( <sup>2</sup> H)	17.04	11,811
6	5.35	H <sub>2</sub> O	18.01	91
		NH <sub>3</sub> ( <sup>15</sup> N)	18.02	2,283
		NH <sub>3</sub> ( <sup>2</sup> H)	18.03	12,882
7	27.42	H <sub>2</sub> O( <sup>17</sup> O)	19.01	9,851
		H <sub>2</sub> O( <sup>2</sup> H)	19.02	13,228
8	28.99	C <sub>4</sub> H <sub>8</sub>	56.06	13,048
		C <sub>3</sub> H <sub>4</sub> O <sub>2</sub>	60.02	2,167
		C <sub>5</sub> H <sub>10</sub>	70.08	4,630
		C <sub>5</sub> H <sub>12</sub>	72.09	16,988
9	22.90	C <sub>3</sub> H <sub>6</sub>	42.05	10,061
		CO <sub>2</sub>	43.99	143
		C <sub>3</sub> H <sub>8</sub>	44.06	10,996
		C <sub>2</sub> H <sub>7</sub> N	45.06	10,424
		C <sub>4</sub> H <sub>10</sub>	58.08	10,305
		C <sub>8</sub> H <sub>18</sub>	114.14	690
10	1756.90	Ar( <sup>40</sup> Ar)	39.96	2,312
		C <sub>3</sub> H <sub>4</sub>	40.03	9,392
11	7.95	C <sub>2</sub> H <sub>2</sub>	26.02	110
		CO	27.99	2,846
		N <sub>2</sub>	28.01	2,846
		C <sub>2</sub> H <sub>4</sub>	28.03	6,948
		C <sub>2</sub> H <sub>6</sub>	30.05	6,948
12	1.56	CH <sub>2</sub> O	32.03	7,427
13	53.01	CH <sub>4</sub>	16.03	1,464
		H <sub>2</sub> O( <sup>18</sup> O)	20.01	96

\*Determined using peak deconvolution  
Note: m/ $\Delta$ m uses 10% valley definition.



# Data Products and Science Closure

Observable	Data Products	Science Closure
1A	H <sub>2</sub> O, O <sub>2</sub> , H <sub>2</sub> , CH <sub>4</sub> , C <sub>2</sub> H <sub>2</sub> , C <sub>2</sub> H <sub>4</sub> , C <sub>2</sub> H <sub>6</sub> , HCN density @ 25, 50, 100 km altitude spatially localized maps for flyby mission (37 flybys); near-global coverage for the orbital mission	<b>Determine</b> exospheric composition, structure, and locations of plume/outgassing [Johnson et al., 2009] <b>Determine</b> compositional variation associated with geological features [Johnson et al., 2009]
1B	Density of <sup>40</sup> Ar every flyby	<b>Determine</b> surface to interior exchange rates [Waite et al., 2005]
2A	Densities of CH <sub>4</sub> , H <sub>2</sub> , and H <sub>2</sub> O and their D/H for every flyby ≤100 km or 156 times per orbit	<b>Identify</b> active plume connected to an internal ocean from H <sub>2</sub> and CH <sub>4</sub> isotopes relative to H <sub>2</sub> O. <b>Determine</b> hydrothermal temperature [Proskurowski et al., 2006]
2B	Densities of: <ul style="list-style-type: none"> <li>• CO<sub>2</sub>, N<sub>2</sub>, NH<sub>3</sub>, CO, CH<sub>4</sub></li> <li>• <sup>15</sup>N/<sup>14</sup>N in N<sub>2</sub> &amp; NH<sub>3</sub>; D/H in NH<sub>3</sub></li> <li>• <sup>13</sup>C/<sup>12</sup>C and <sup>18</sup>O/<sup>16</sup>O in CO</li> <li>• <sup>13</sup>C/<sup>12</sup>C and D/H in CH<sub>4</sub></li> <li>• CH<sub>3</sub>OH, C<sub>2</sub>H<sub>5</sub>OH, HCOOH, CH<sub>3</sub>COOH, C<sub>2</sub>H<sub>5</sub>NH<sub>2</sub> for all flybys ≤50 km or once per orbit</li> </ul>	<b>Determine:</b> <ul style="list-style-type: none"> <li>• oxidation state of the ocean from plume outgassing [Shock and McKinnon, 1993]</li> <li>• role of minerals in compound transformations (alkane/alkene, alcohol/ketone) [Shock et al., 2013]</li> <li>• temperature of oxidation/reduction reactions (alkane ratios) [Shock et al., 2013]</li> <li>• pH acidic: low ratio (alkenes, alkynes, esters, ethers, amides, imides and nitriles) / (alcohols, carboxylic acids, ammonia), pH basic: hemi-acetal bonds of polymers are hydrolyzed [Shipp et al., 2013]</li> <li>• free energy from disequilibria of volatile compounds [Shock et al., 2013]</li> <li>• origin of N- and C-bearing material [Mandt et al., 2014; Mumma and Charnley, 2011]</li> </ul>
2C	Densities and isotopes of major C- and N-bearing compounds at the ppm level	<b>Determine</b> origin and radiolytic age of organic compounds on the surface and test for habitability [Sephton et al., 2014; Mumma and Charnley, 2011].
2D	Density of hydrocarbon compounds (plume gas) with mixing ratios greater than 10 ppb when ≤50 km from cryotrapped samples or once per orbit	<b>Identify</b> the formation of organics through hydrothermal organic transformations [Shock et al., 2013; McCollom, 2013]

# Instrument Block Diagram

A block diagram of the instrument showing the major subsystems:

- Gas inlet system (GIS)
- Mass spectrometer (MS).

