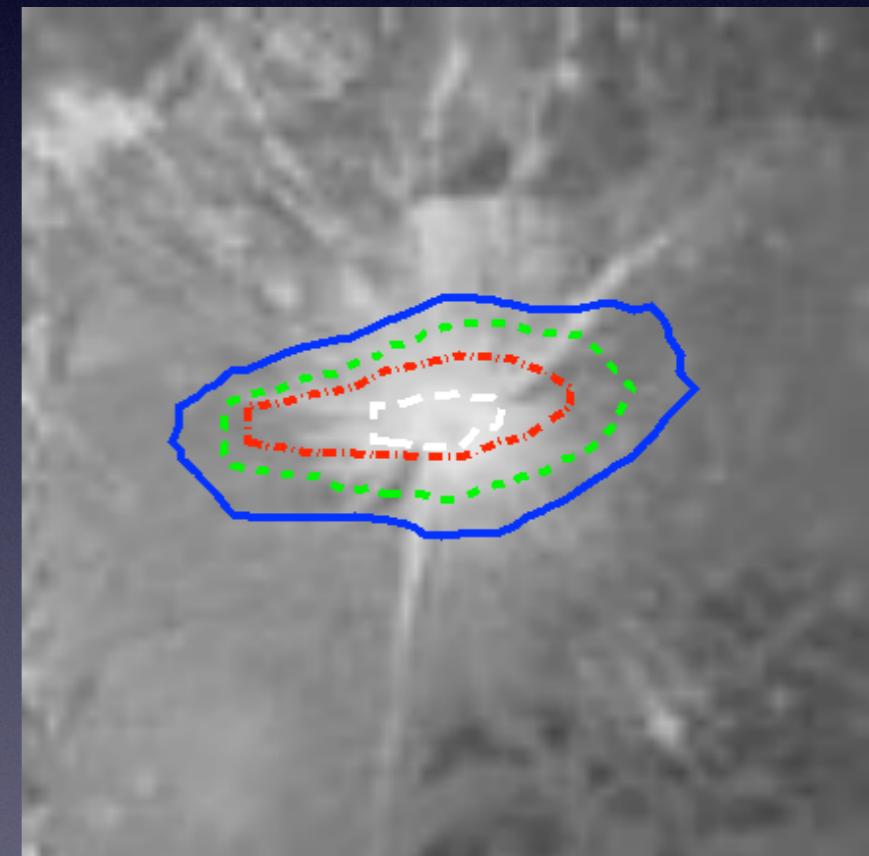
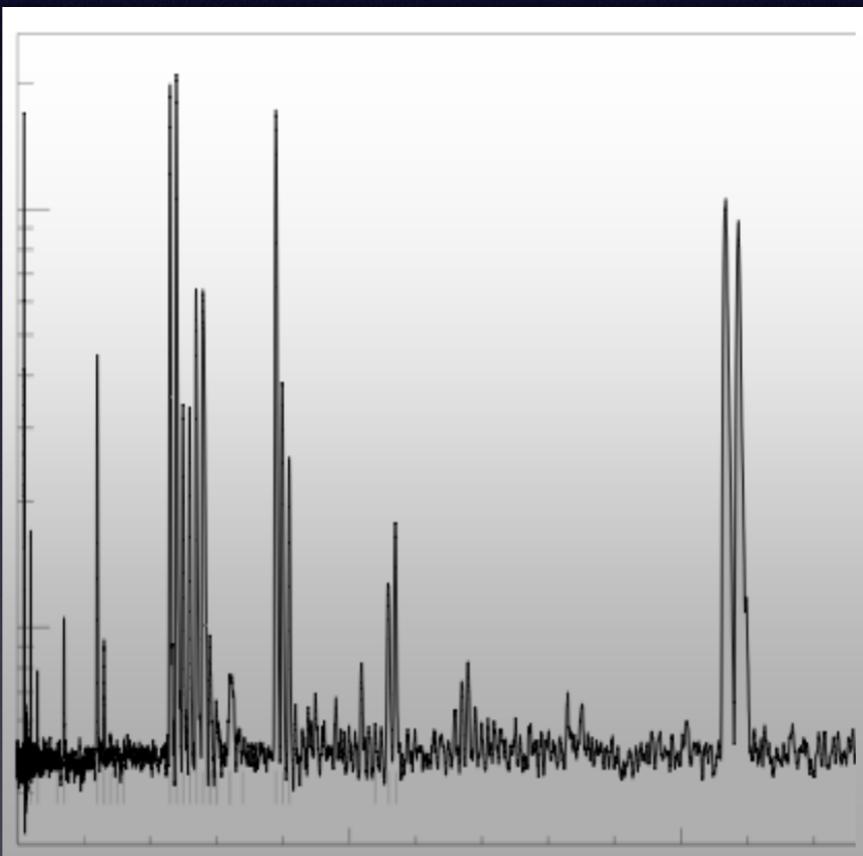
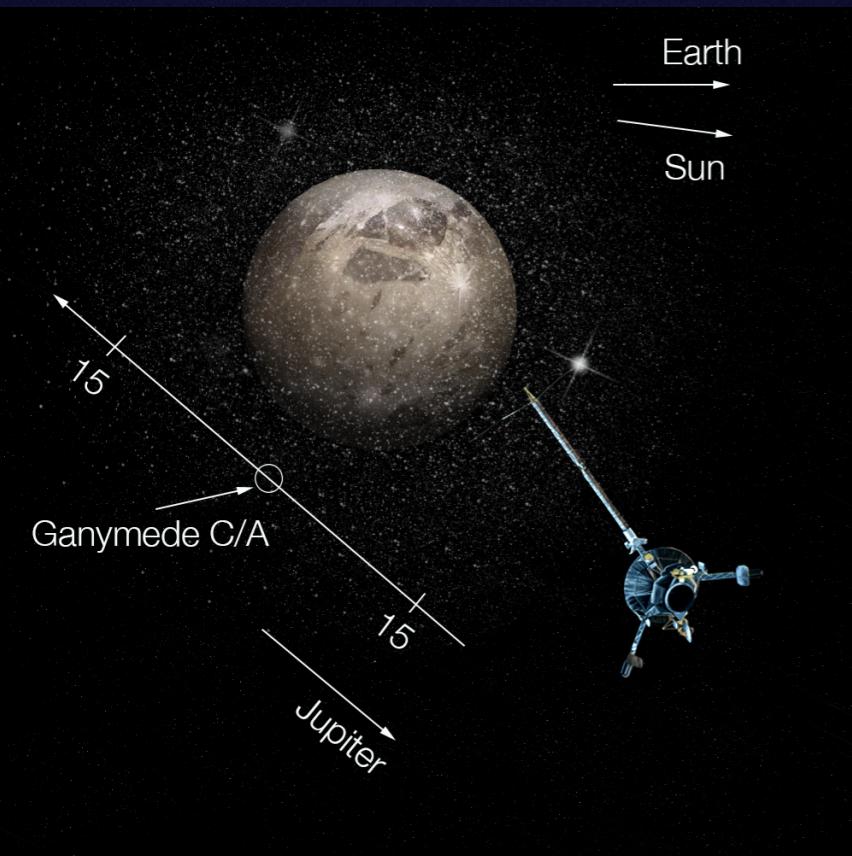




Compositional Mapping

Determine Surface Composition from Orbit „in-situ“

Ejecta Clouds Ejecta Composition Ejecta Backtracking



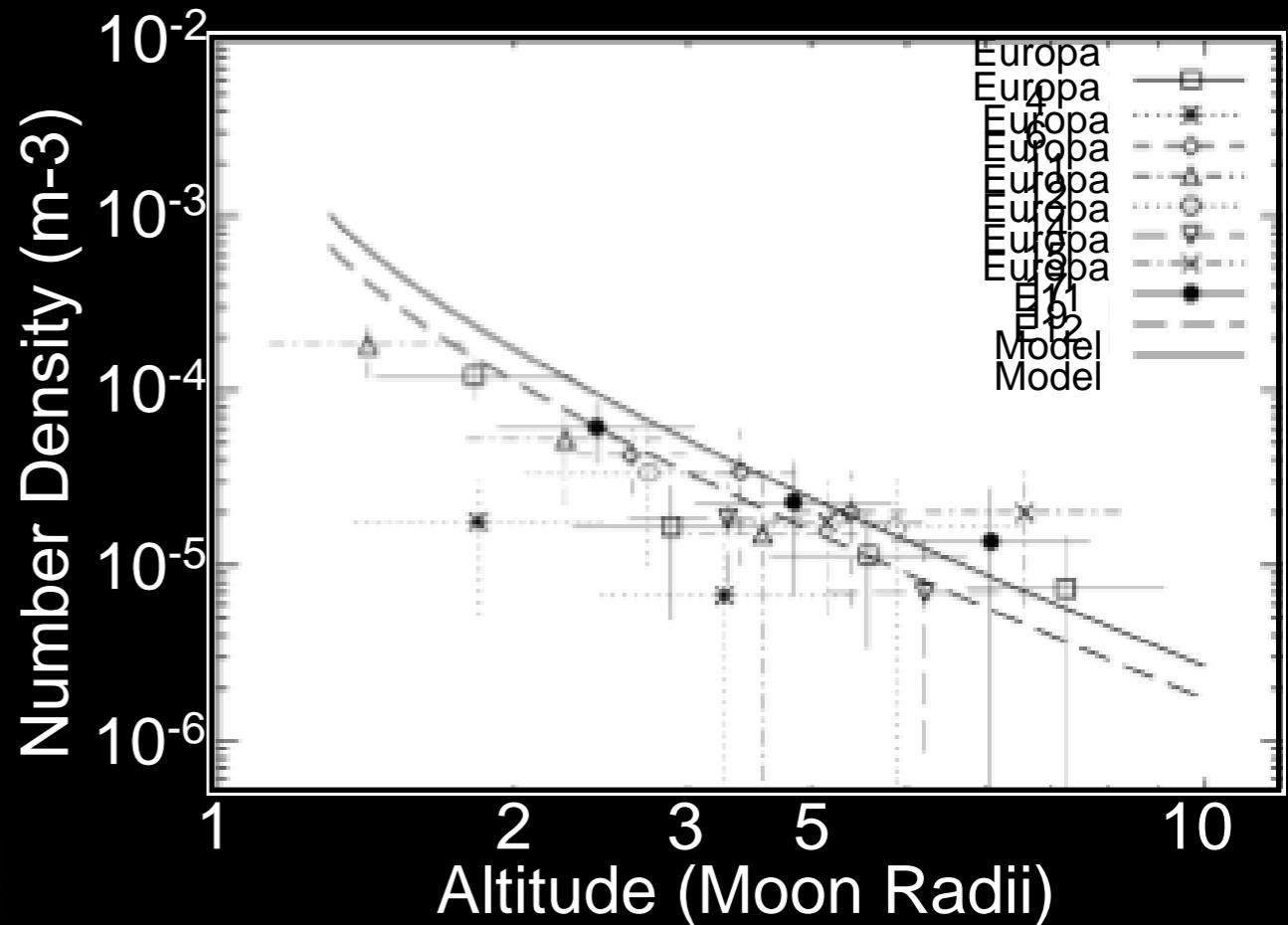
The Next Best Thing to a
Lander

1: Ejecta Clouds



Galileo Dust Detector:
Galilean Satellites
Wrapped in Dust
Clouds

(Krüger et al., Nature, 1999)



Almost Isotropic Clouds
Composed of Surface
Ejecta

Ejecta Production

Meteoroid Impacts
Produce Surface Ejecta

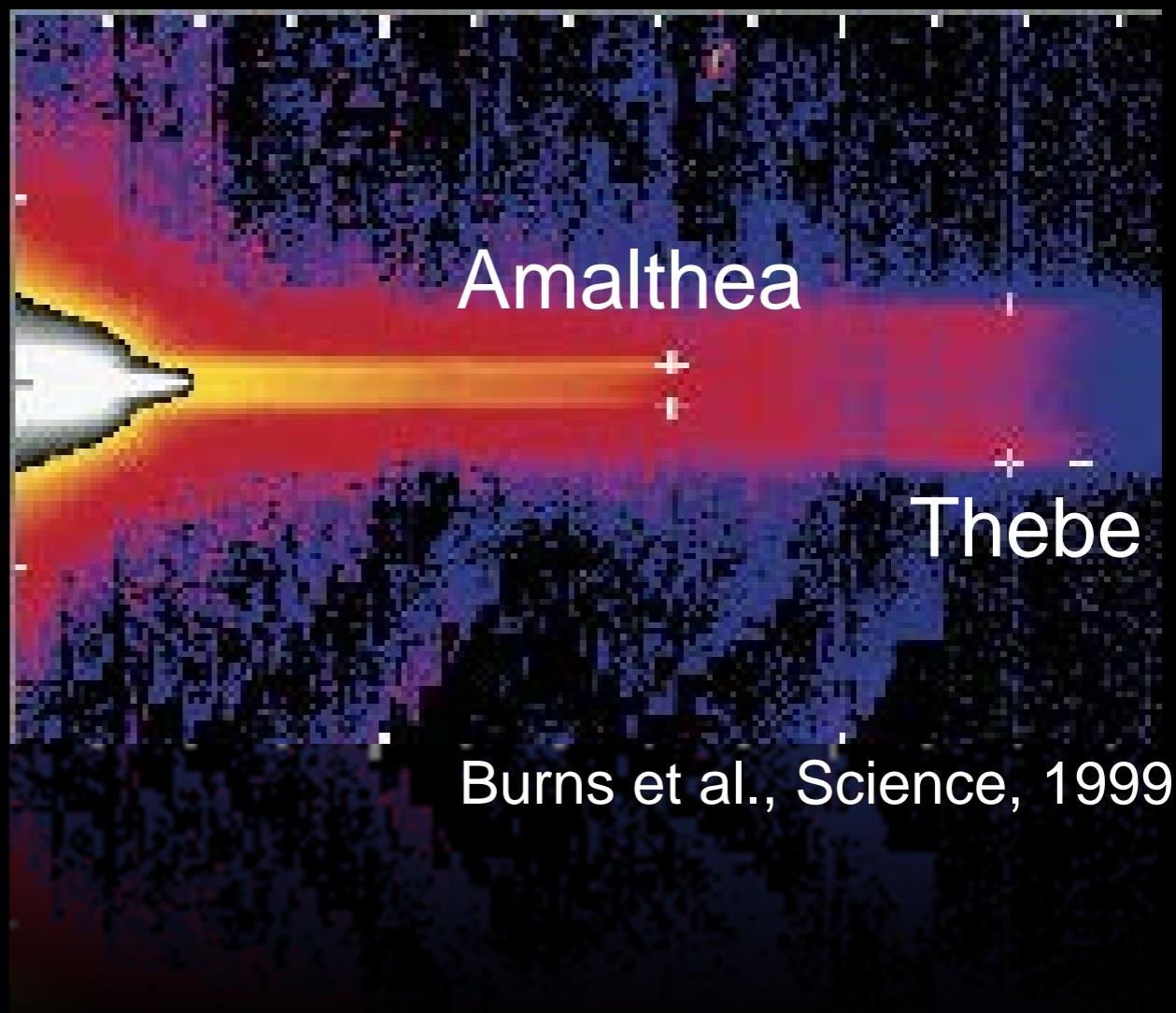
100 μm micrometeoroid impacts generate
 $\sim 500 \text{ kg ejecta/second}$



Mass Yield ~ 4000

Koschny & Grün, Icarus, 2001; Krivov et al.,
Icarus, 2003

Ejecta Escaping from
Moon's Gravity feed
Rings



Dust Composition

Cassini Dust Detector
CDA

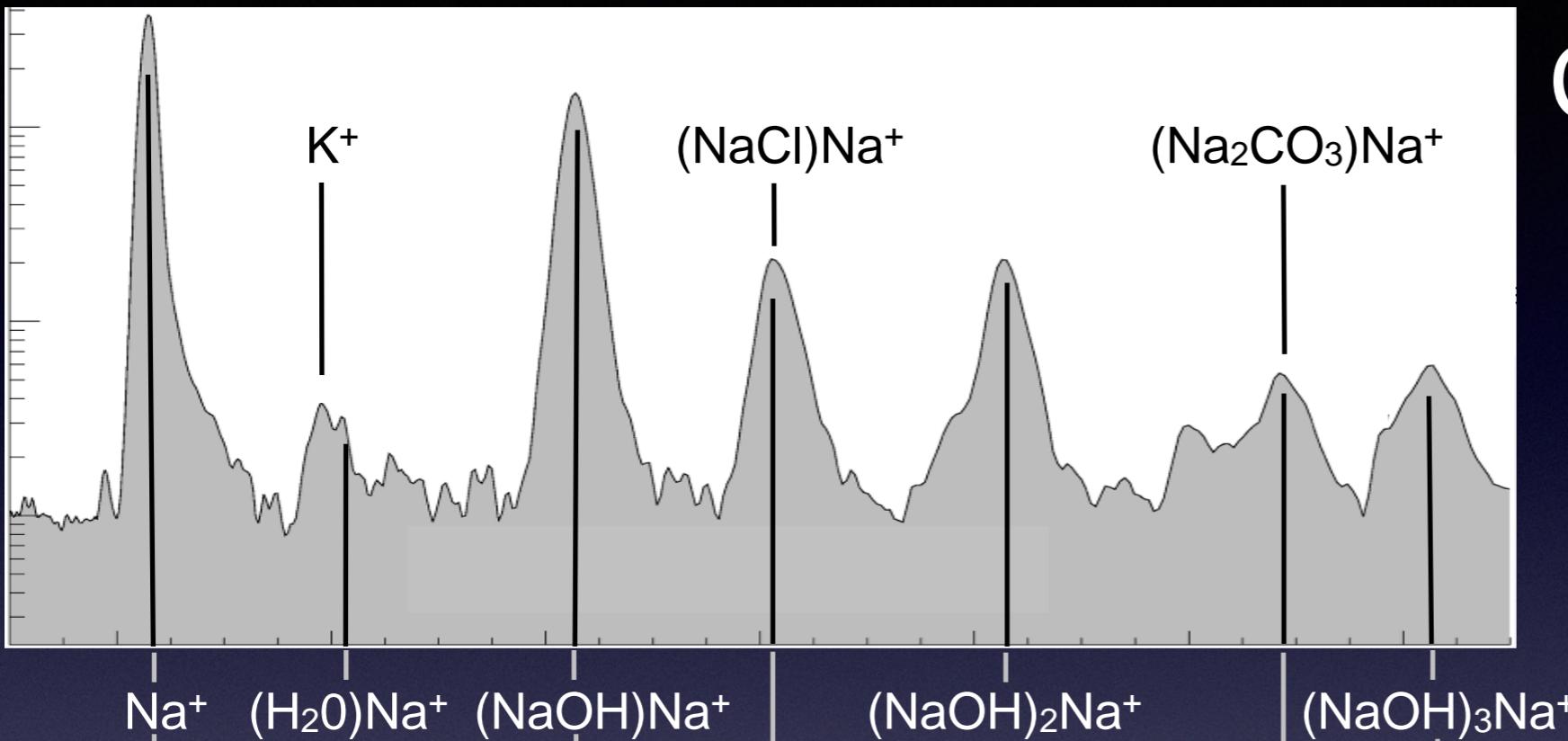


Composition of Enceladus
Plume Particles

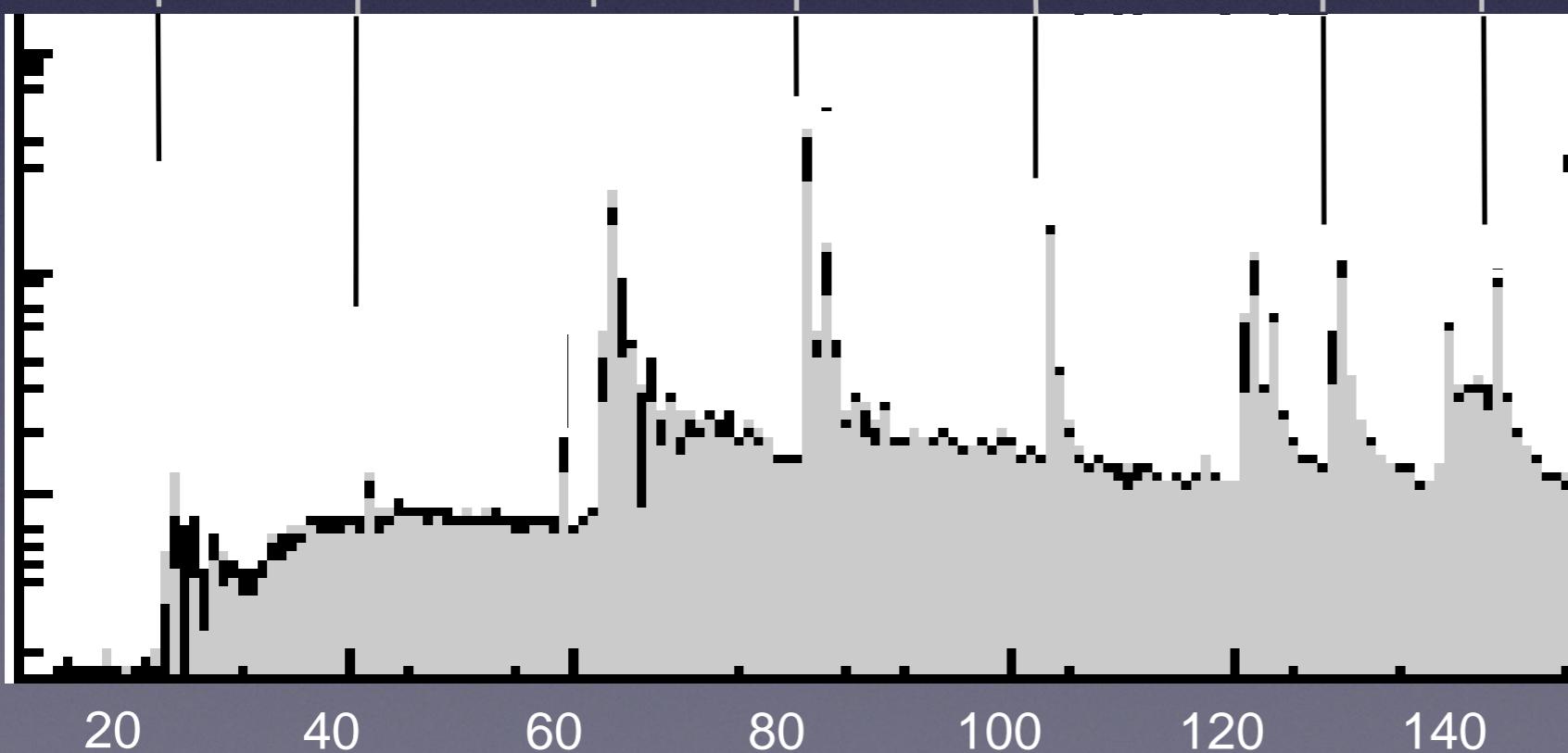


NASA/JPL/Space Science Inst.

Enceladus Dust Composition



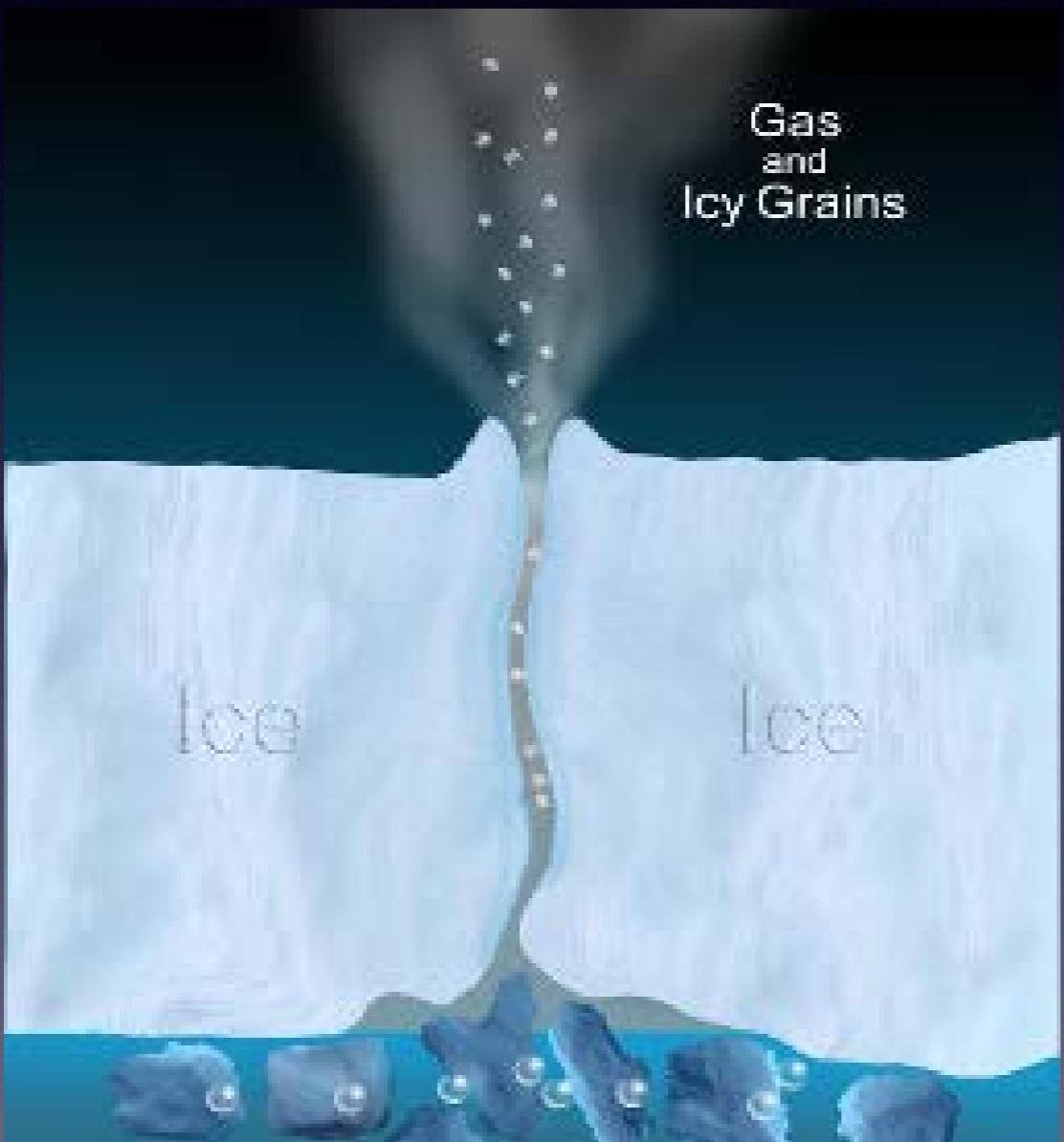
Co-Added CDA
Spectrum:
Salt-rich Geyser
Ice Grains
(6%)



Lab Spectrum:
Laser Dispersion
of Salt Water

Postberg et al., Nature,
2009

The Enceladus Ocean



„Soda“ Ocean

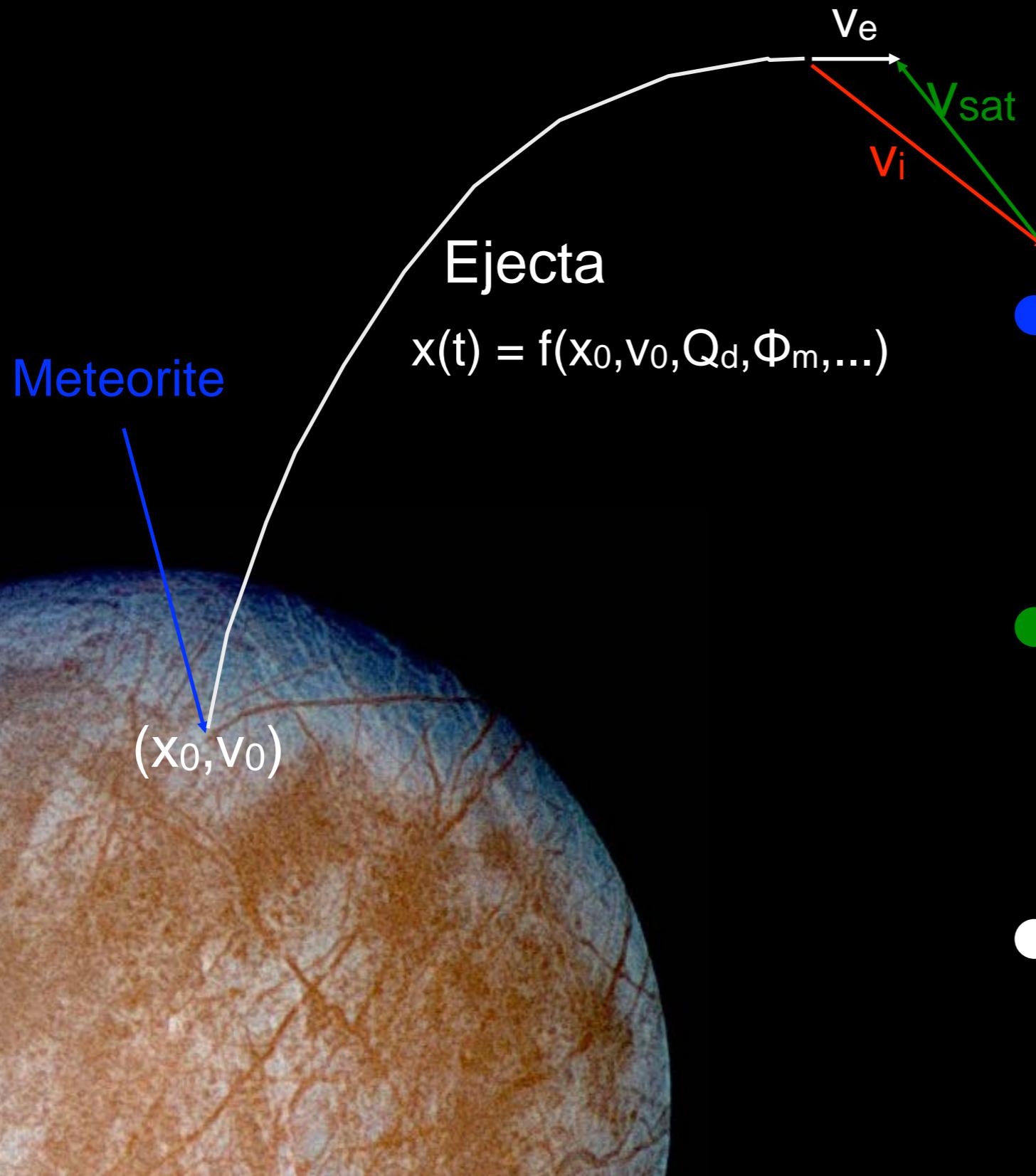
Rich in Carbonates

pH ~ 9

Salinity ~1% (Earth
1...4%)

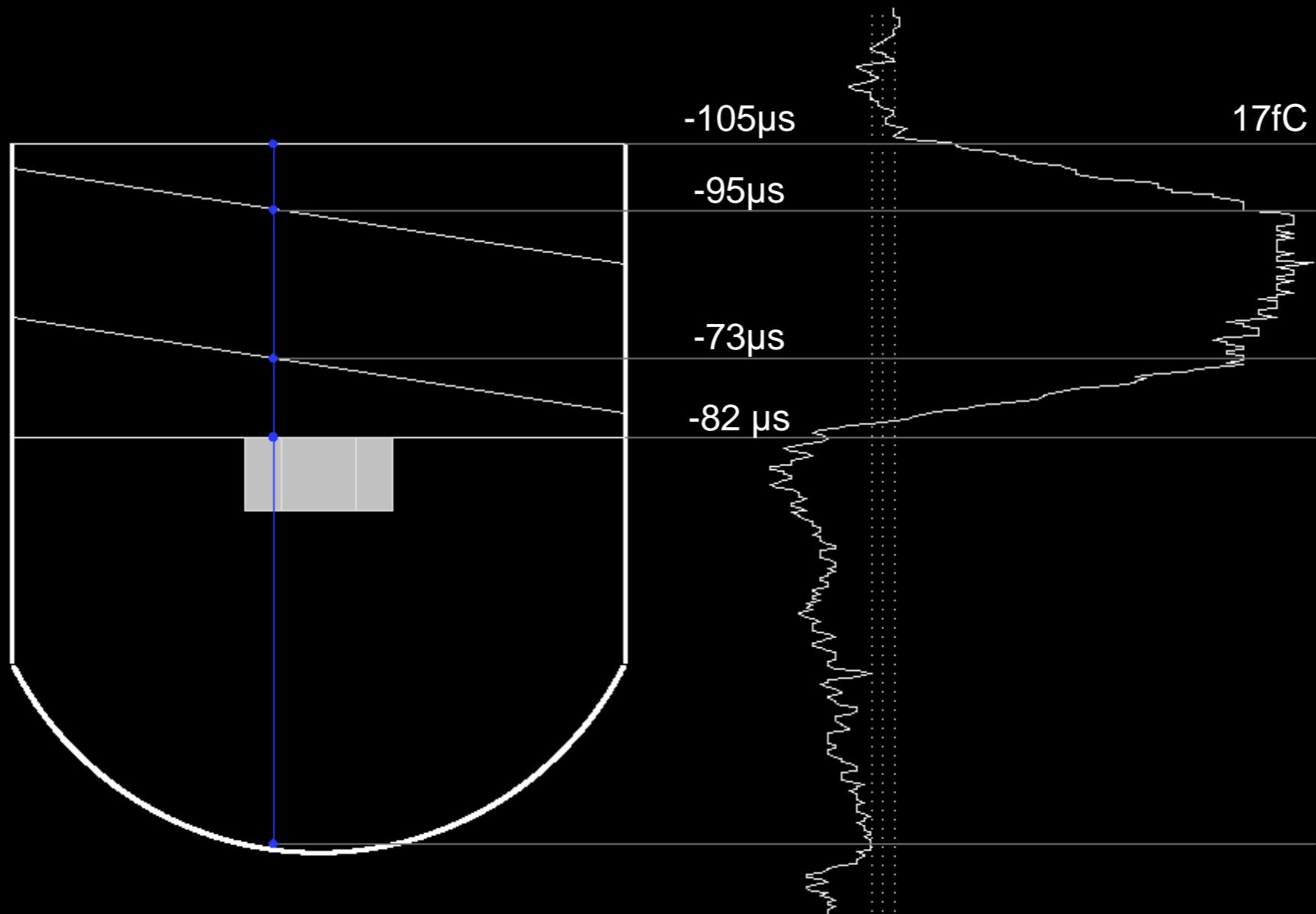
Postberg et al., Nature,
2009

Ejecta Backtracking



- Meteorite impact splashes up multiple ejecta
- Satellite moves relative to ejecta:
 $v_i = v_e - v_{sat}$ (\approx Apex)
- Know Starting Position:
 $x_0 = f(x_i, v_i, Q_d(t_i), t_i, \dots)$

CDA Measures Velocity of Charged Dust



$$v_d = 4.3 \text{ km/s} \quad R_d = 0.6 \mu\text{m}$$

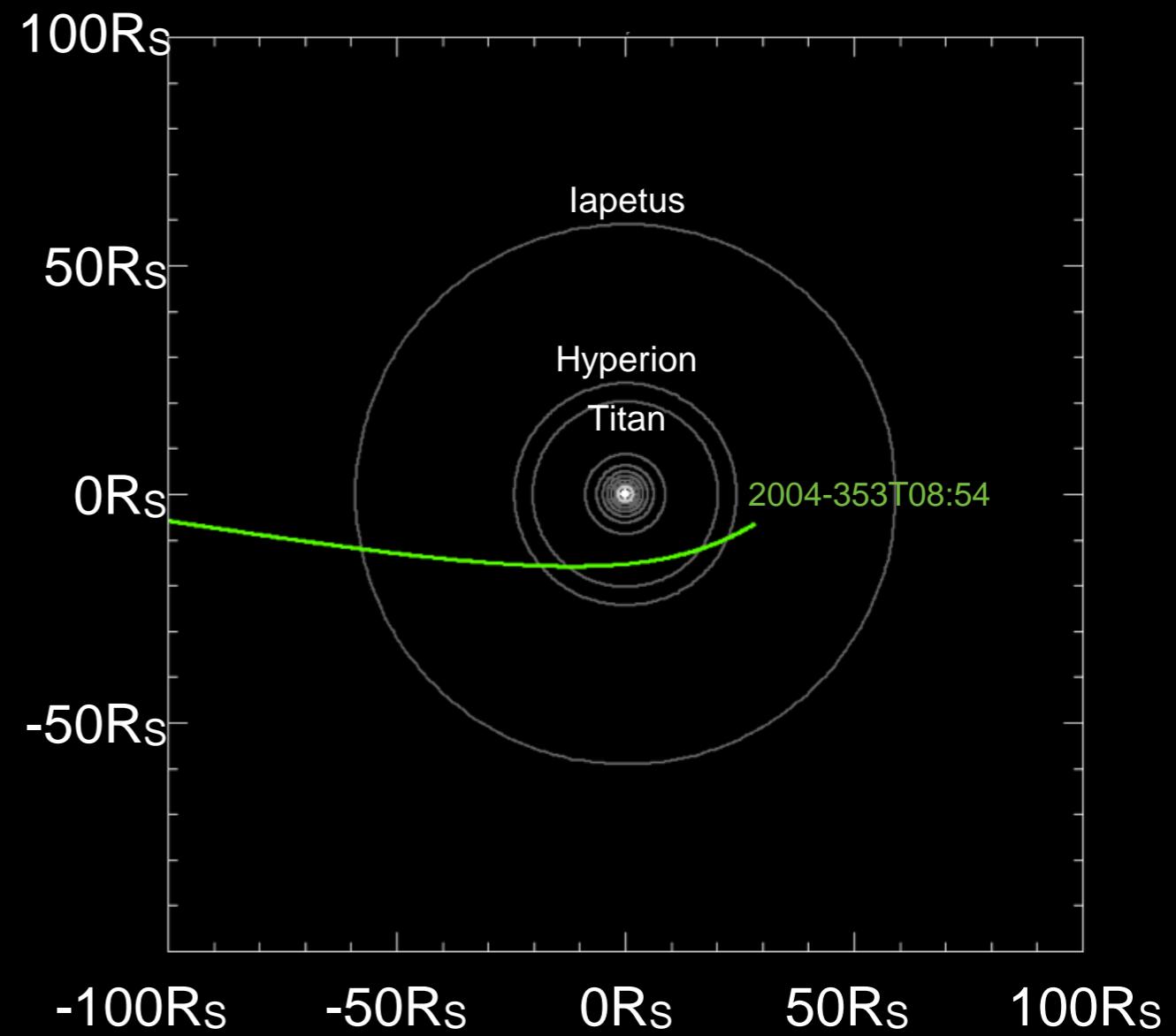


Dust Orbit Reconstruction

2004-353T08:54



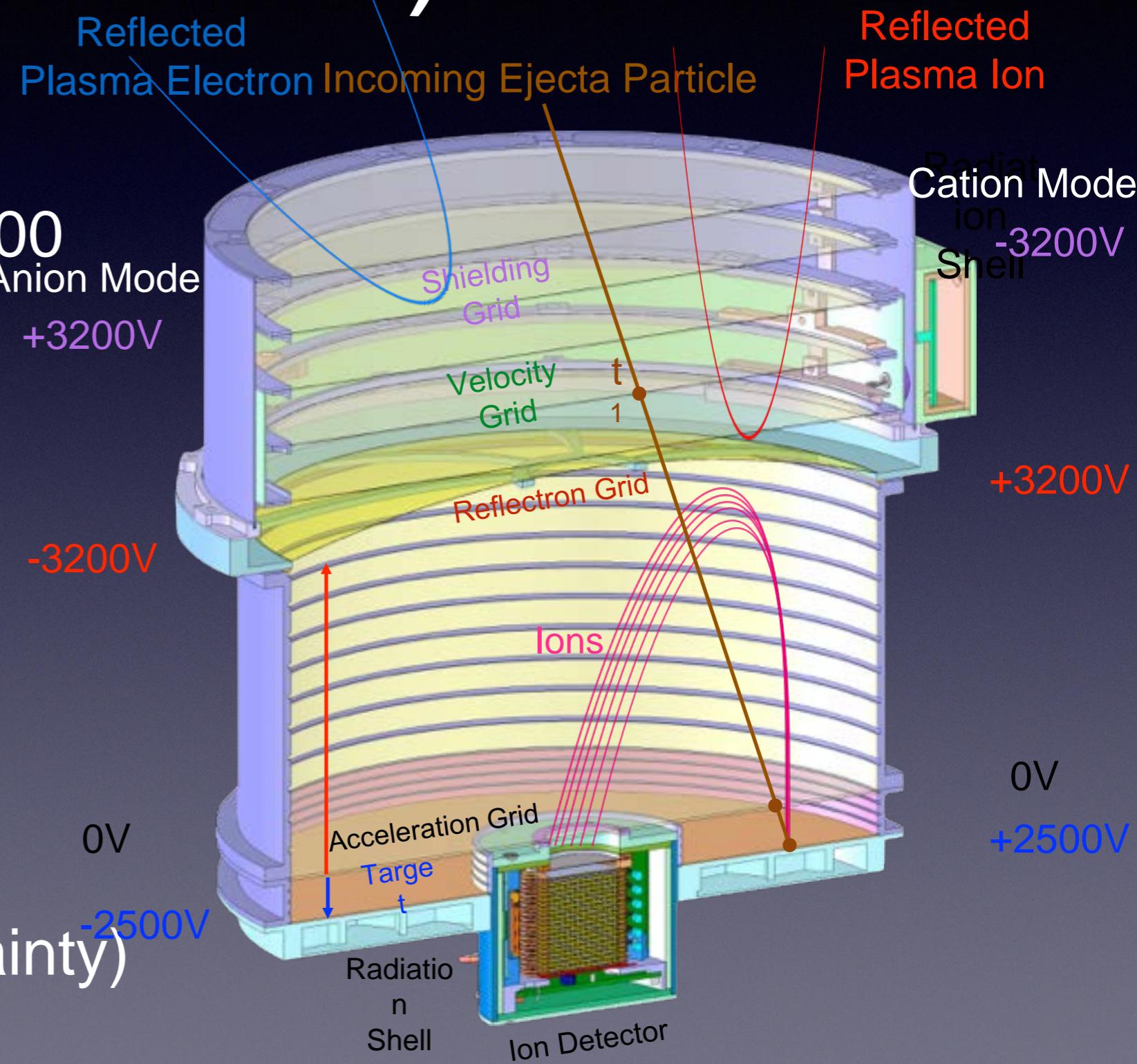
$R=18.2\mu\text{m}$ $Q=5.7\text{fC}$ $\Phi=+2.8\text{V}$



Kuiper Belt Particle

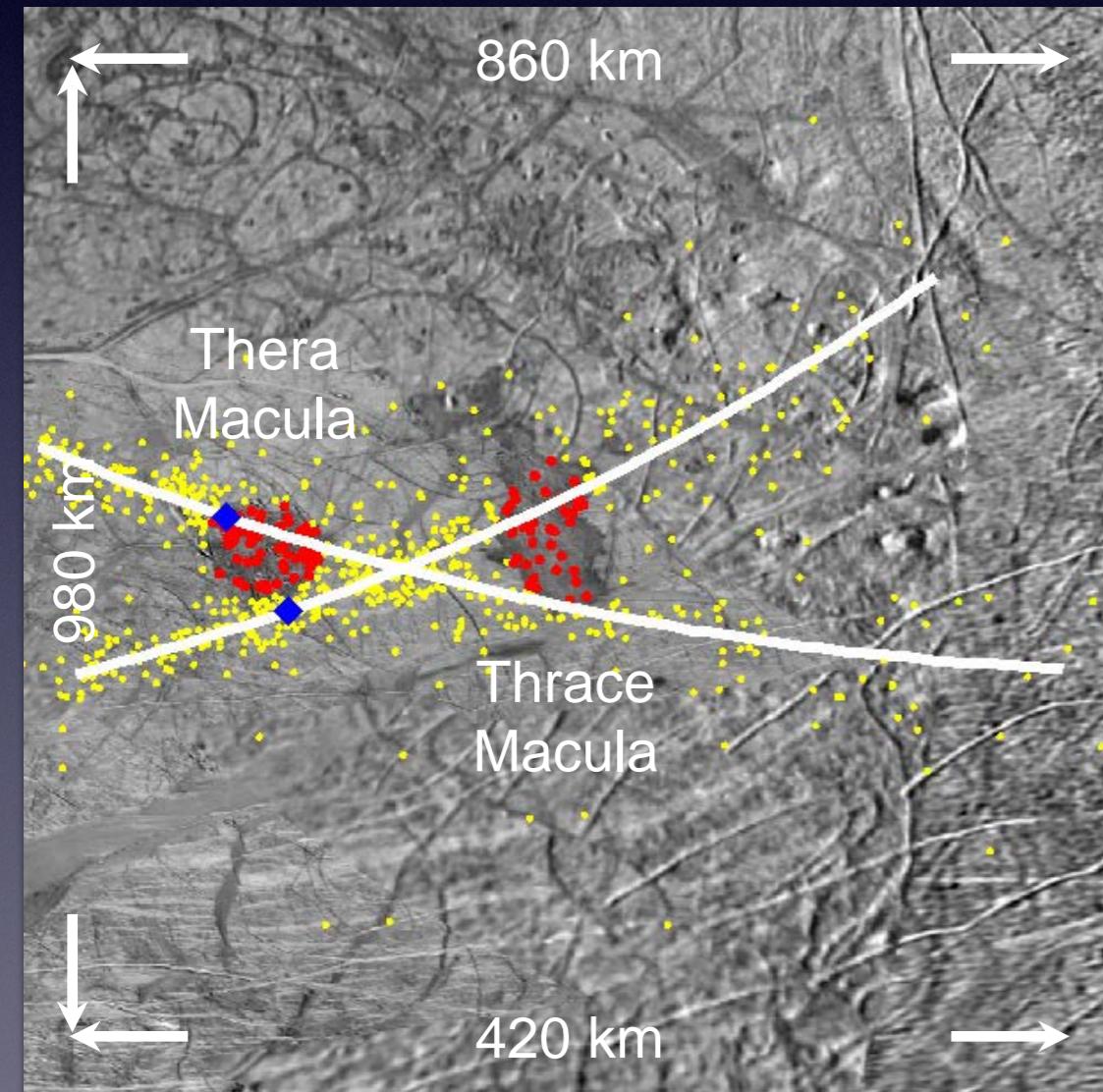
SURface Dust Analyzer (SUDA)

- Mass Spectrometer:
- Mass Resolution ~ 200
- Electrostatic Mirror:
- Parabolic Grid
- Ring Electrodes
- \pm Polarity
- Trajectory Sensor:
- Velocity (1% Uncertainty)

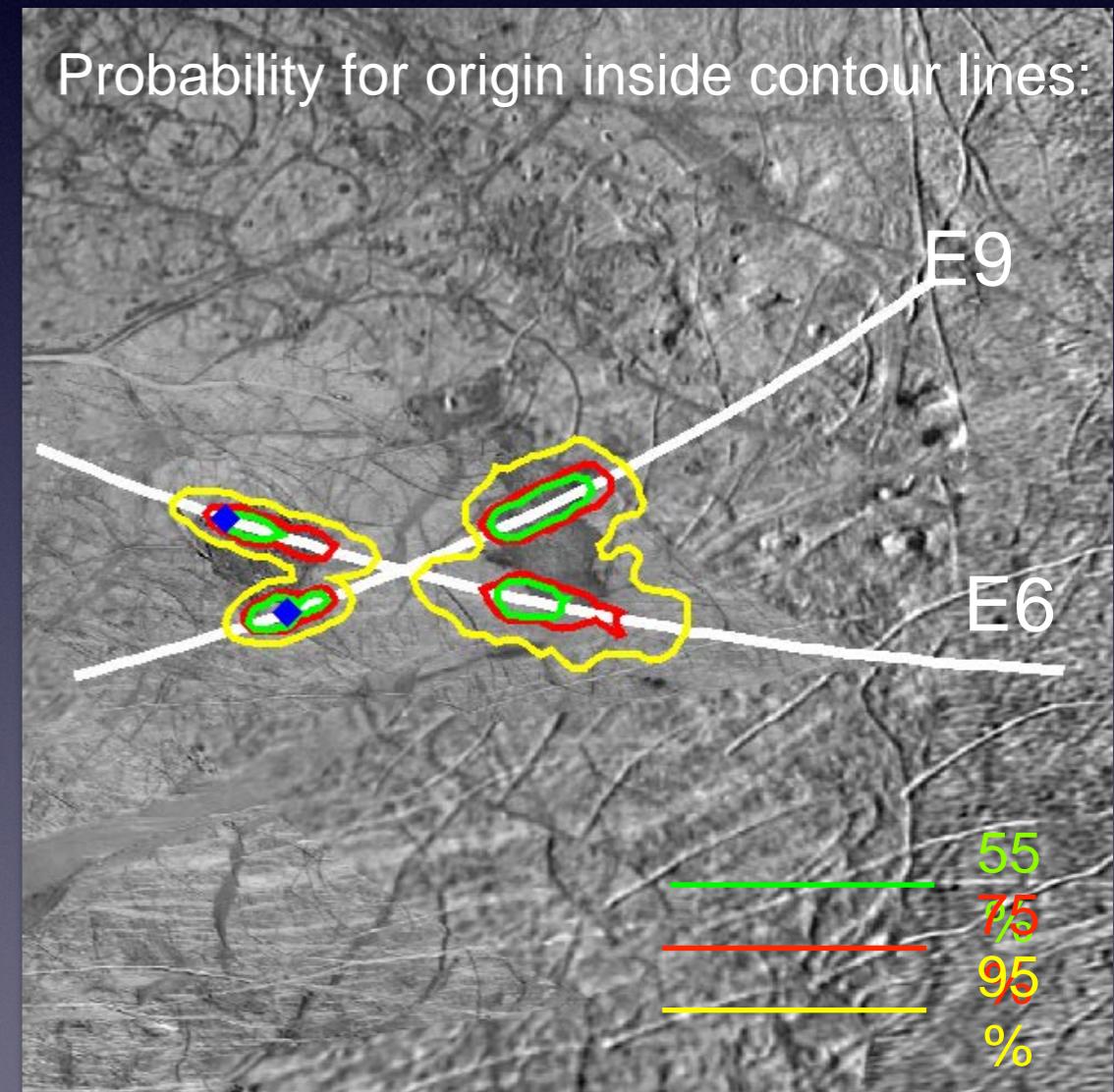


SUDA Composition Map

Dark Lobated Features Thrace Macula and Thera Macula on Europa



MC Simulation for SUDA Compositional Mapping



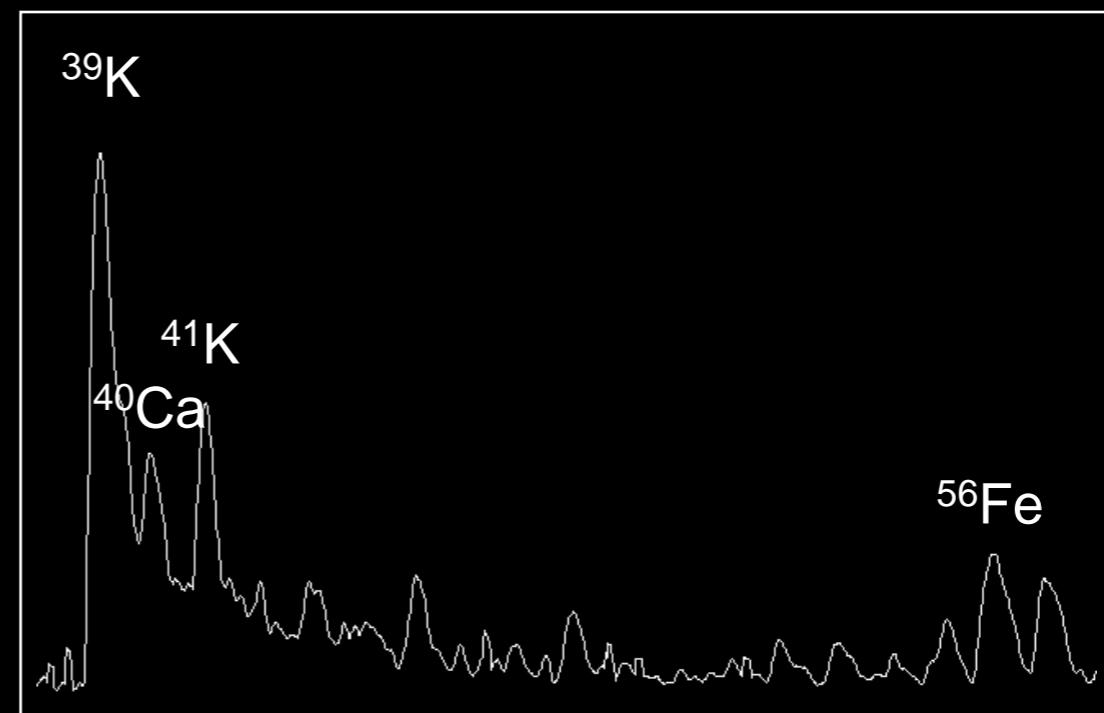
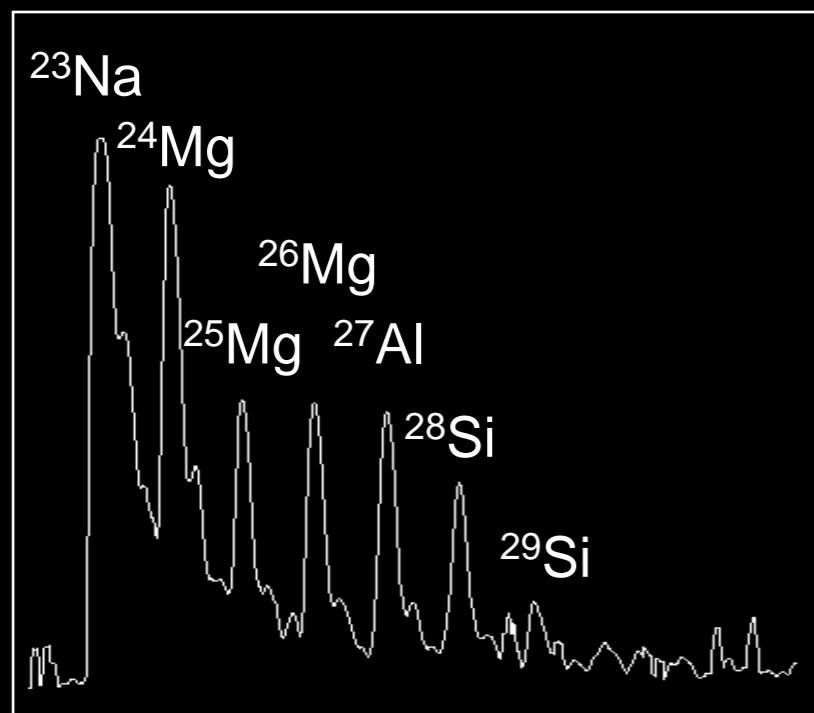
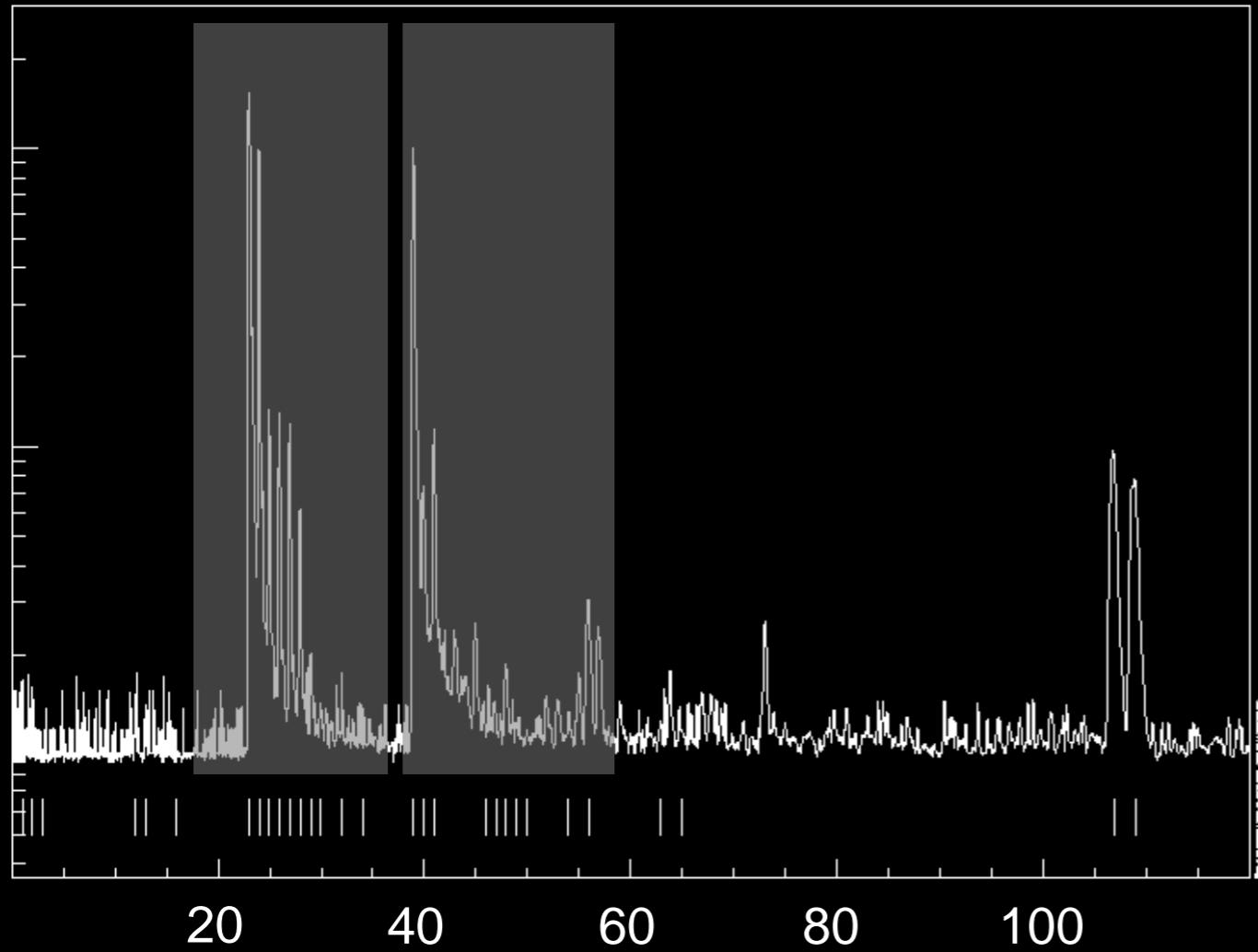
SUDA Will Collect

Europa Clipper Flybys:

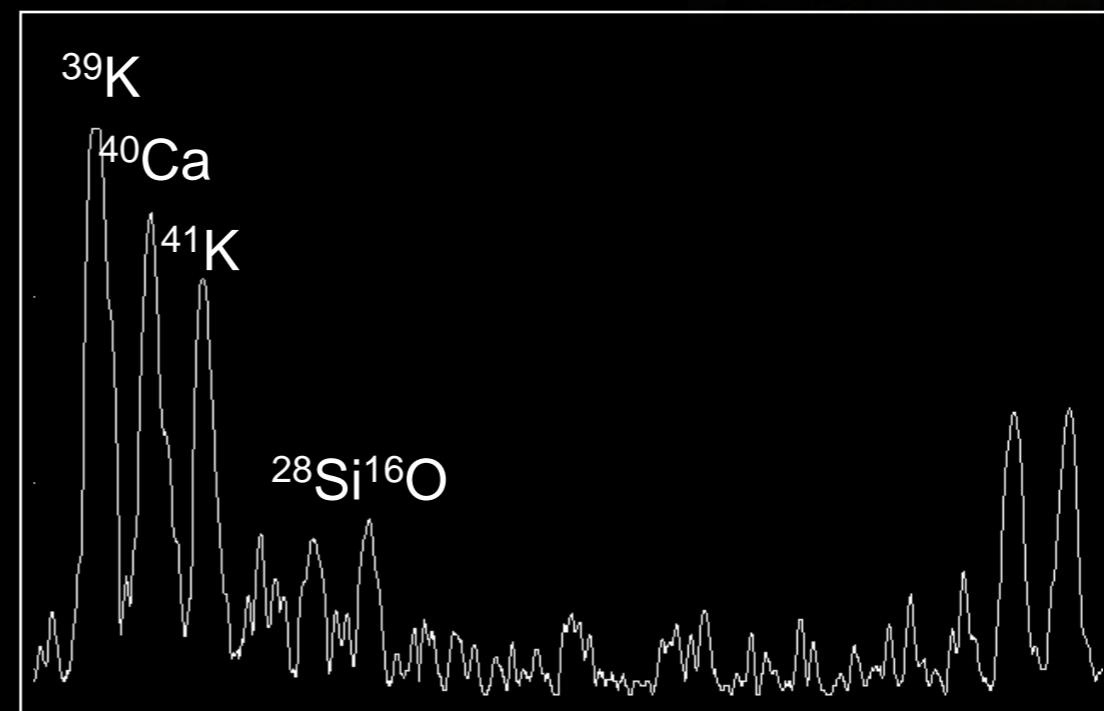
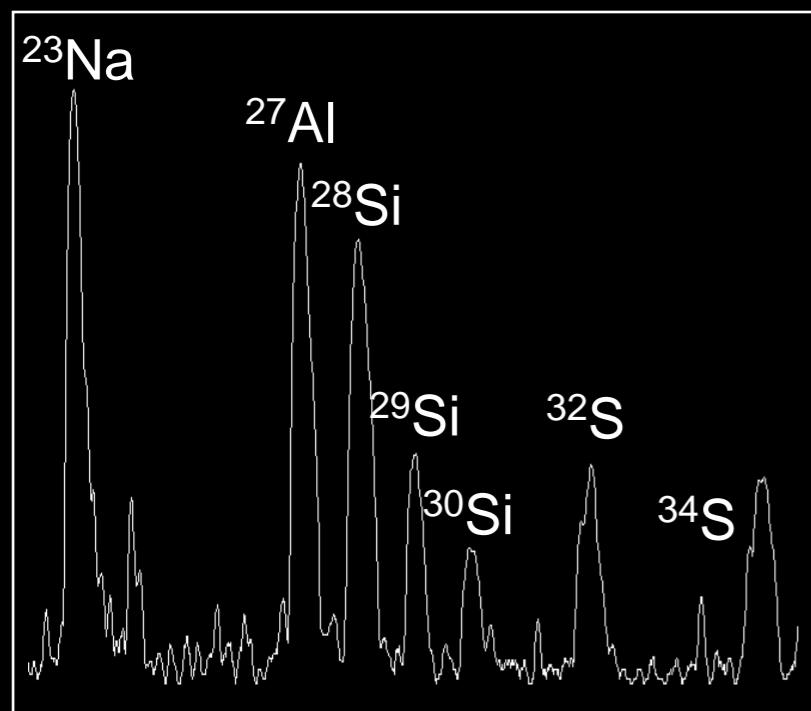
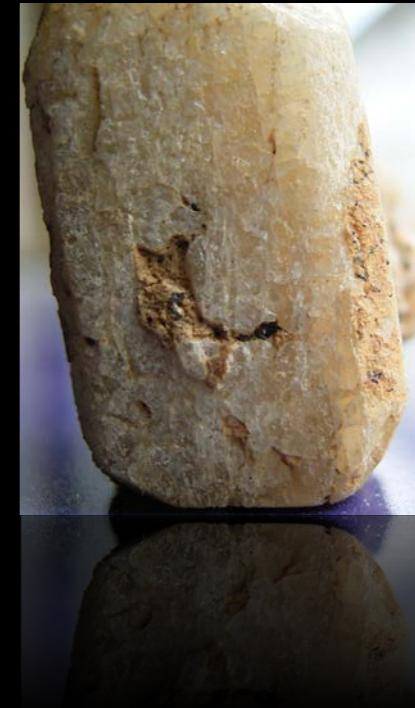
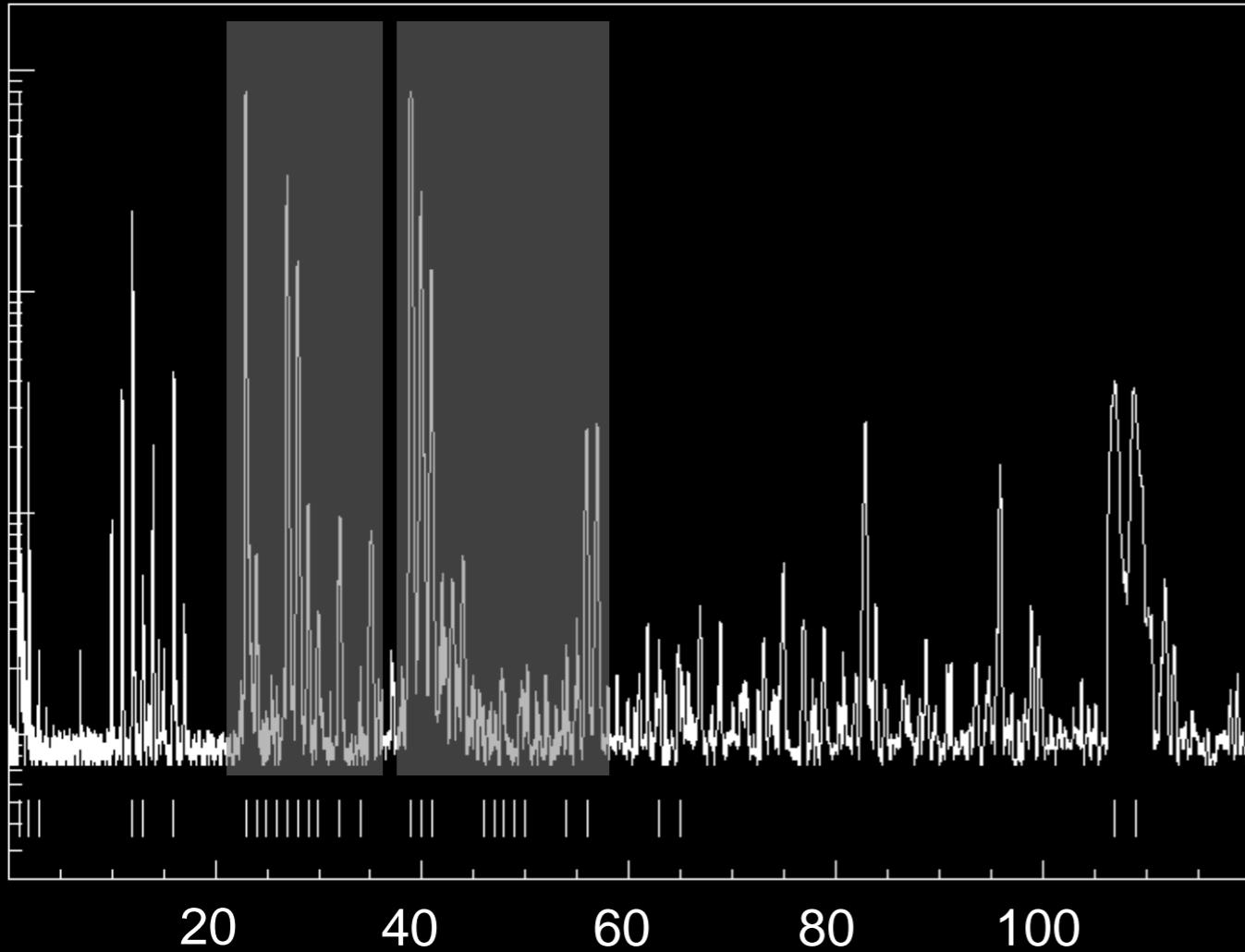
Flyby	Impact Rate	Total Sample #
Europa 25 km	40 per second	5300
Europa 50 km	14 per second	2700
Europa 100 km	5 per second	1350

In total, SUDA will collect about 120 000 samples from Europa's surface

Pyroxene



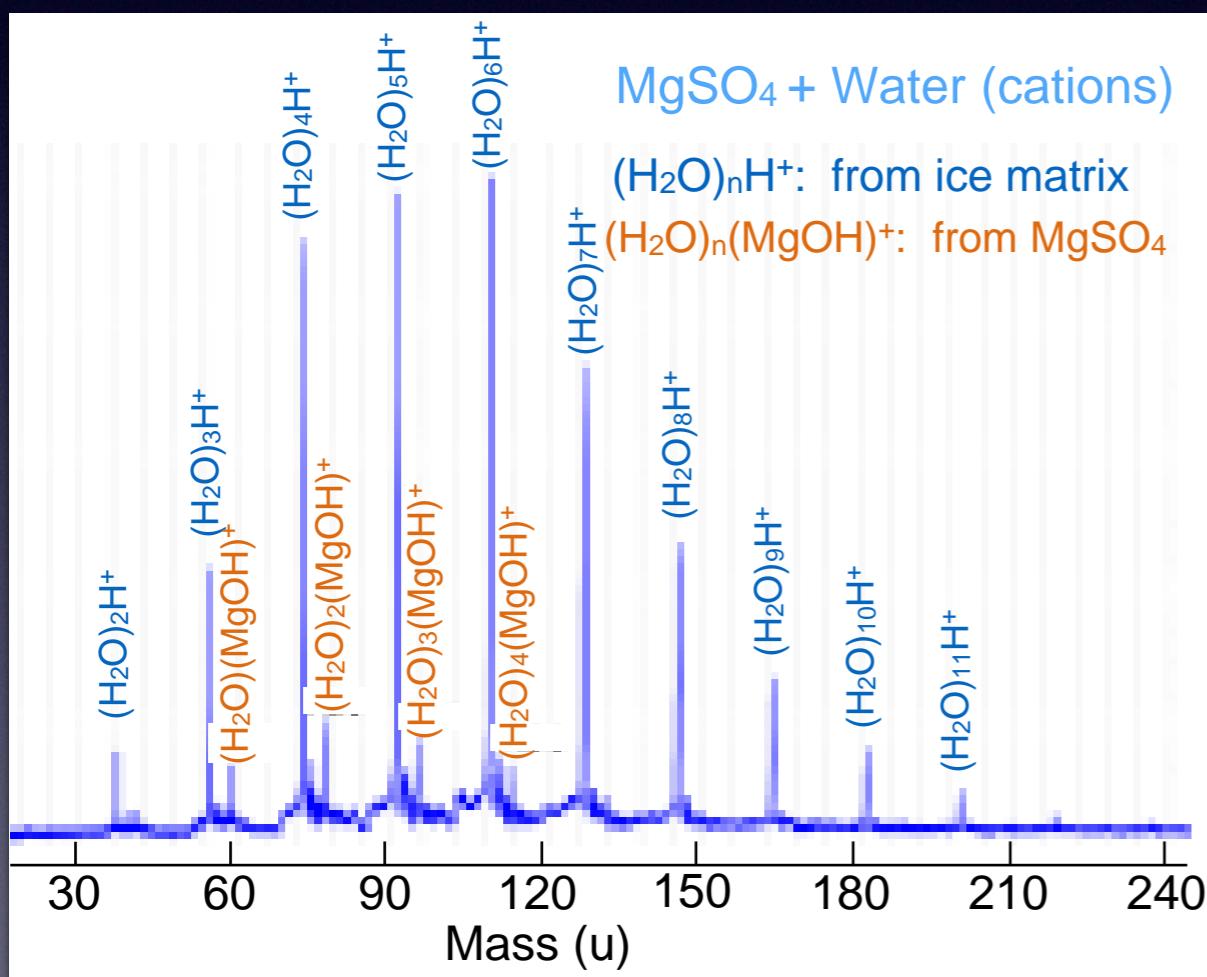
Anorthite

$$\text{CaAl}_2\text{Si}_2\text{O}_8$$


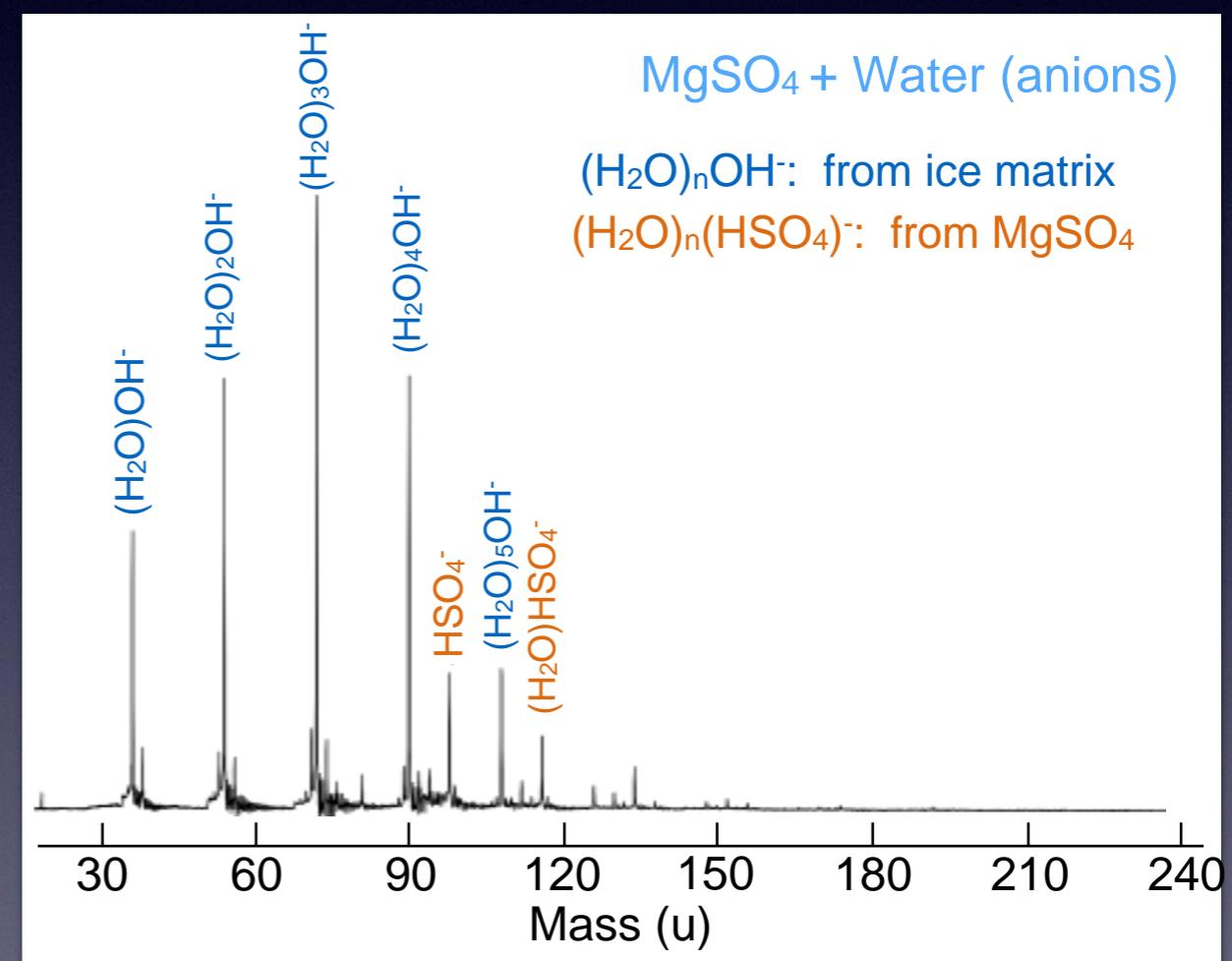
SUDA @ Europa

Water + MgSO₄

Cations:



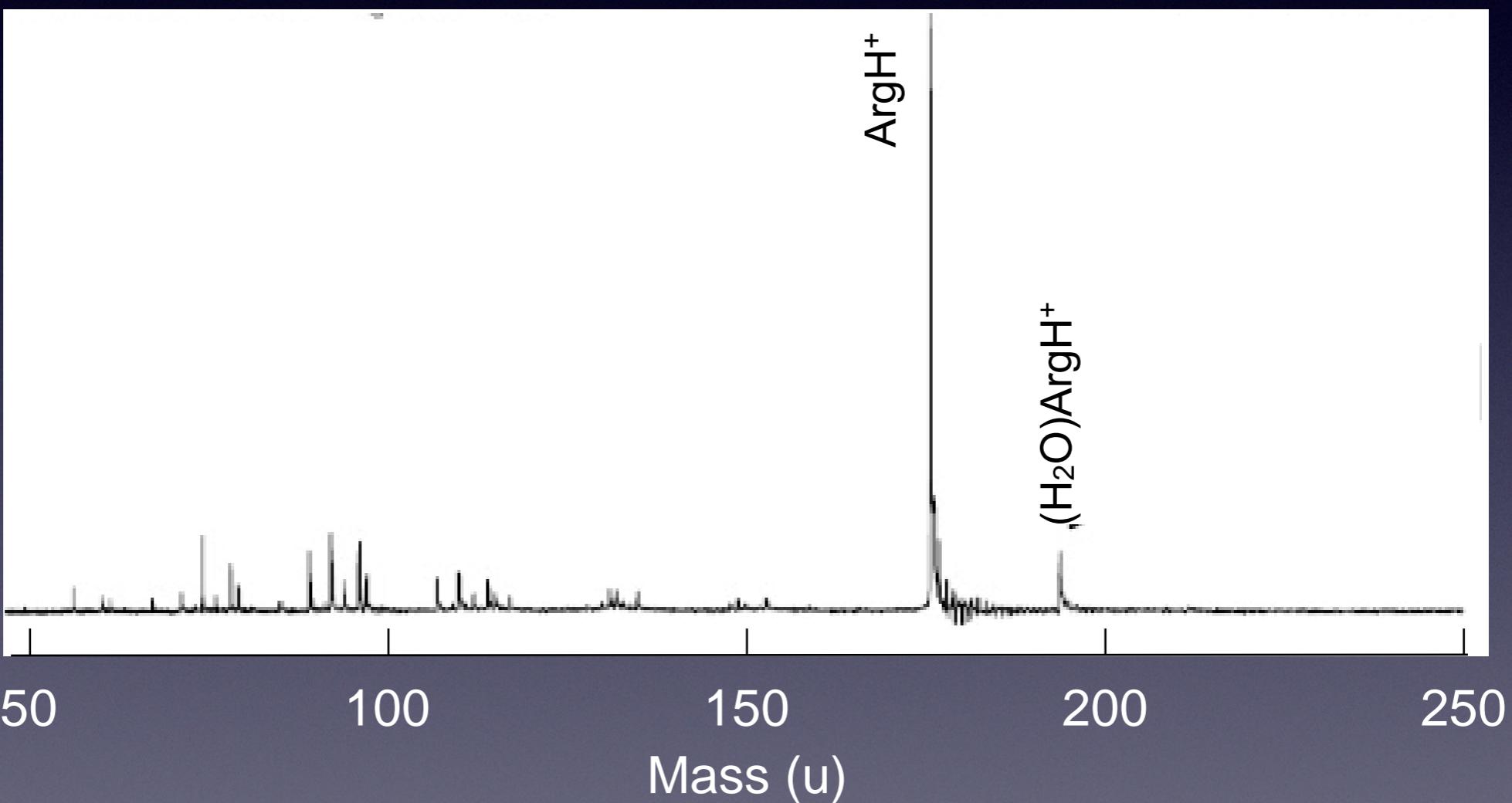
Anions:



Laser-assisted dispersion spectra of MgSO₄ at a concentration of 0.1 ppm in water

SUDA @ Europa

Arginine + Water (Cations)



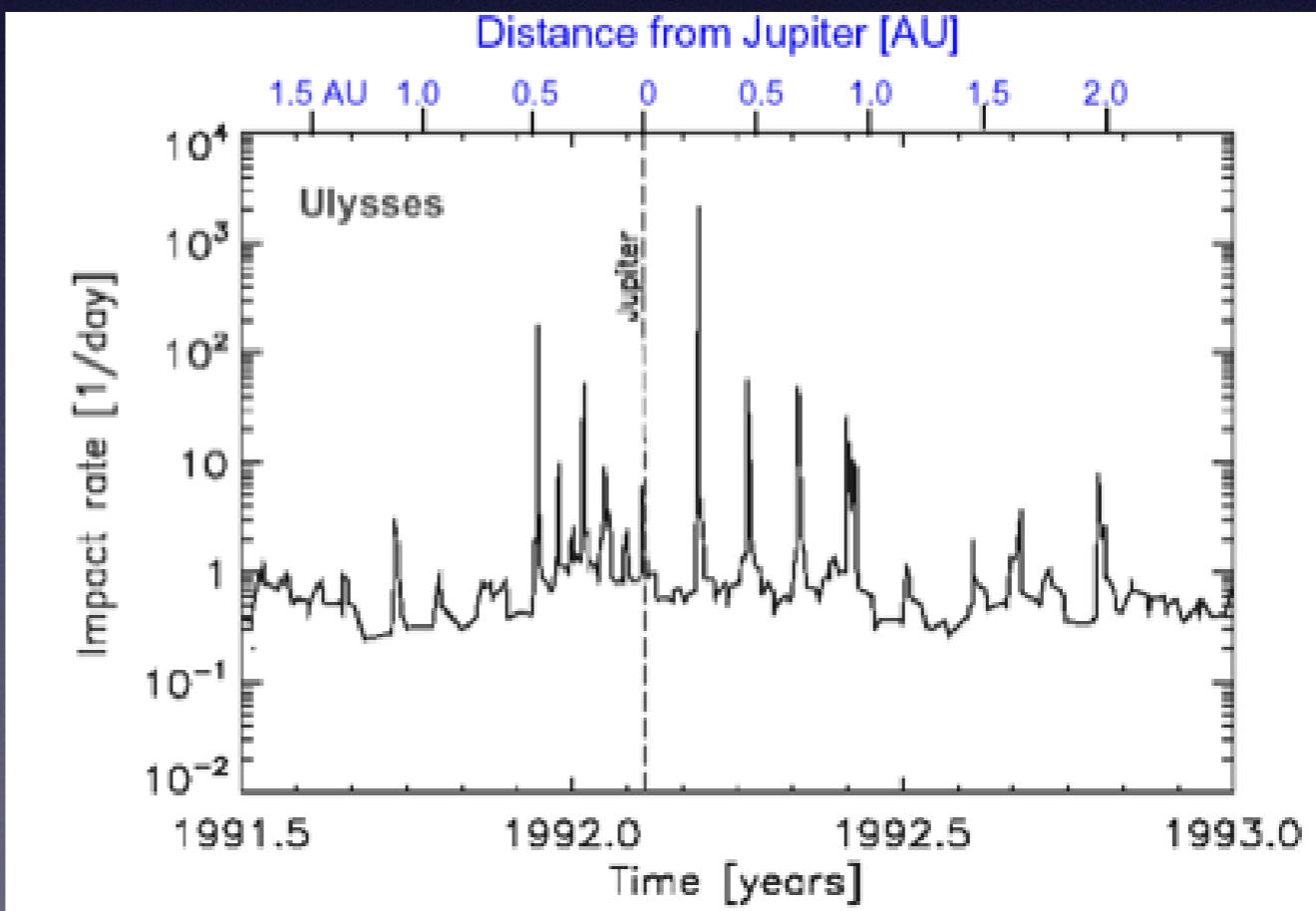
Laser-assisted dispersion cation spectrum of the amino acid arginine ($\text{C}_6\text{H}_{14}\text{N}_4\text{O}_2$) dissolved in water at a concentration of 10^{-4} mol/l.



SUDA Europa Science Starts at the Asteroid

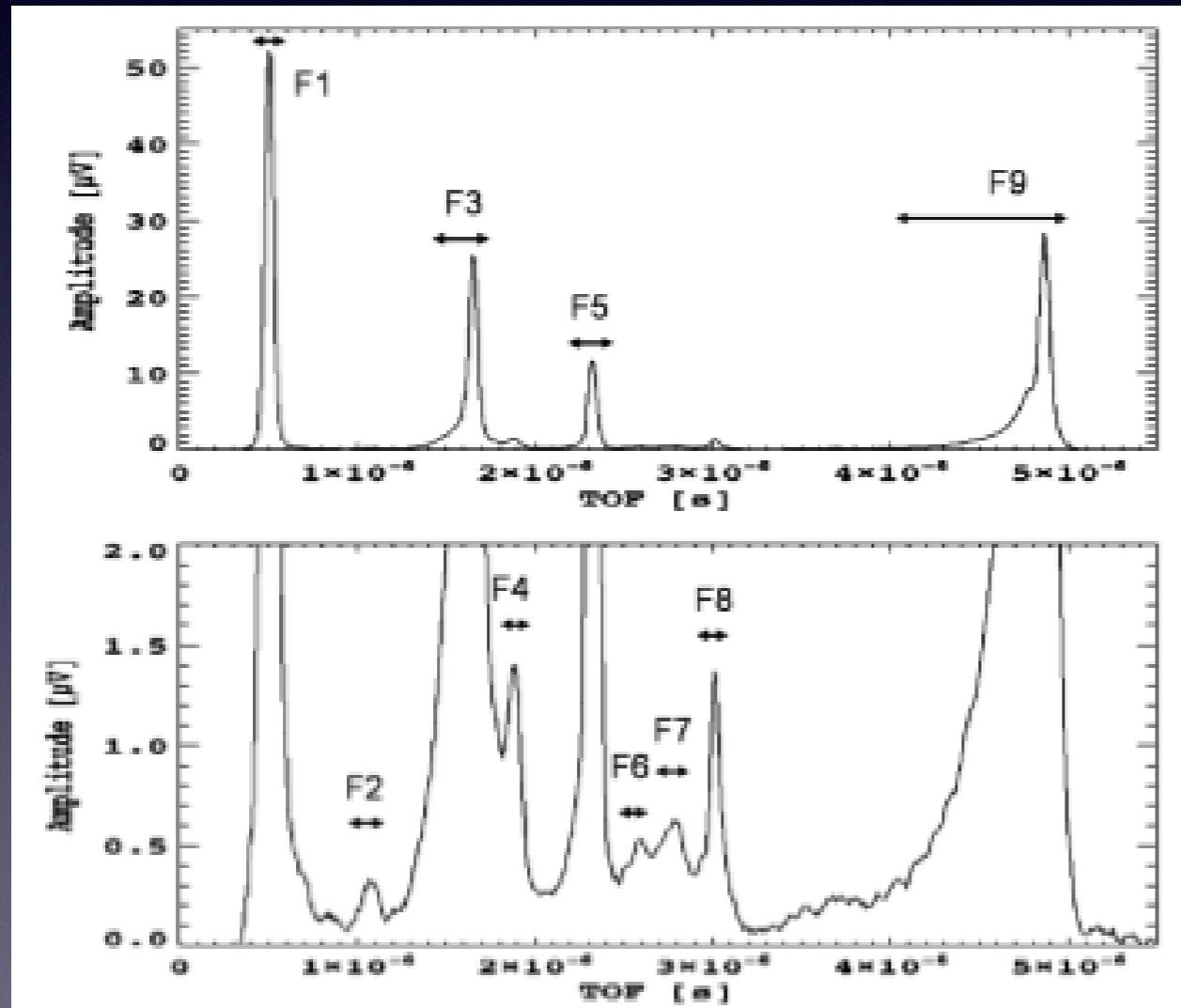
Belt

Ulysses discovered "periodic streams" of nano-dust emerging from the Jovian system



Grün et al., 1993, Nature, 362

CASSINI Determined the Nano-Dust Composition



- Primary source: Io's volcanoes
- Composition: NaCl