

# In Situ Exploration of Titan's Prebiotic Organic Chemistry and Habitability

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23 Feb 2017



### NASA Ocean Worlds Mission Themes Titan Science Objectives

- Understand the organic and methanogenic cycle on Titan, especially as it relates to prebiotic chemistry
- Investigate the subsurface ocean and/or liquid reservoirs, particularly their evolution and possible interaction with the surface

#### Titan is a high-priority target for prebiotic chemistry and habitability

500-1000 km

Diffusion through stratosphere

Clouds at 10-30km altitude H2 escapes to

C2H6, tholin etc drizzle down as

~1cm / 100,000 vrs

space

CH4 evaporation

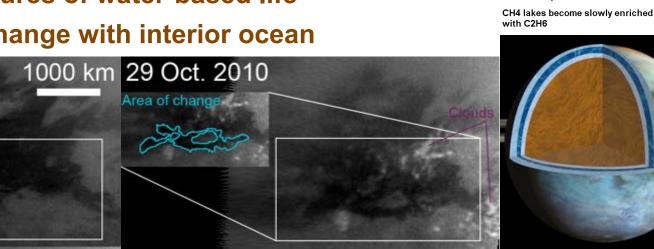
radiation converts CH4 into C2H6 etc

CH4 rainfall ~1 cm/yr

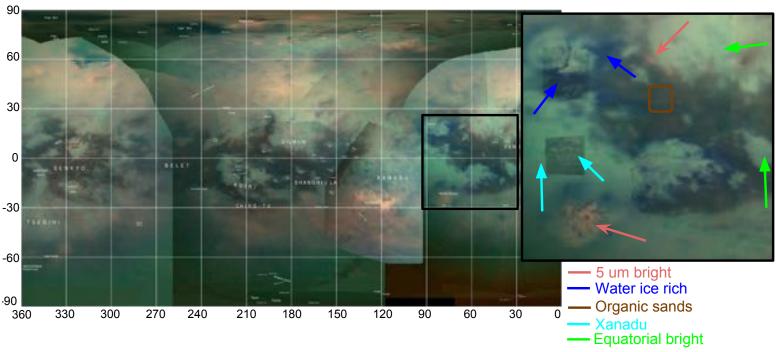
- Accessibility of complex organic material on the surface
- Earth-like system with methane cycle instead of water cycle
- Unique natural laboratory to investigate prebiotic chemistry and to search for biosignatures of hydrocarbon-based life
- Potential for organics to interact with liquid water near or at the surface → furthers prebiotic chemistry potential as well as search for signatures of water-based life
- Potential for exchange with interior ocean

27 Sept. 2010

Cloud



## Diversity of surface materials → scientific priority to sample diverse locations



- Cassini VIMS map showing spectral diversity of Titan's surface w/ higher-resolution inset from T114 (Nov 2015)
  - $\triangleright$  Red = 5  $\mu$ m, green = 2  $\mu$ m, blue = 1.3  $\mu$ m
  - Dark blue = higher water-ice content
  - > Dark brown = organic sands (Barnes et al. 2007; Soderblom et al. 2007)
  - Orange = 5-μm bright unit with characteristics consistent with evaporitic material (MacKenzie et al. 2014)



## Mobility is key to accessing material in different settings

- Compositions of solid materials on Titan's surface still essentially unknown
- Measuring the composition of materials in different geologic settings → how far has prebiotic chemistry progressed in environments that provide known key ingredients for life
- Multiple landers are an inefficient strategy, requiring multiple copies of instrumentation and sample acquisition equipment
- More efficient approach is to convey a single instrument suite to multiple locations
- Heavier-than-air mobility highly efficient @ Titan (Lorenz 2000; Langelaan et al. 2017)
  - ➤ Titan's atmosphere 4x denser than Earth's → reduces the wing/rotor area required to generate a given amount of lift → all forms of aviation are easier (lighter-than air as well as heavier-than-air)
  - ➤ Titan's gravity 1/7th Earth's → reduces the required magnitude of lift → powerful factor in favor of heavier-than-air vehicle

# Strategies considered for in situ Titan exploration in previous mission concepts

- Helicopter (Lorenz 2000)
- Titan airship (helium or hydrogen; Levine
   & Wright 2005; Hall et al. 2006)
- Montgolfière hot-air balloon (Reh et al. 2007)
- Airplane (Levine and Wright 2005; Barnes et al. 2012)
- Lander (TiME, Stofan et al. 2013)
   Flagship mission studies:
- NASA Titan Explorer Flagship Mission (Leary et al. 2007):
  - Lander + Montgolfière-type balloon
  - > Two landers
- NASA-ESA Titan Saturn System Mission (TSSM; Lunine, Lebreton et al. 2008):
  - Montgolfière + lander



#### Dragonfly Rotorcraft Lander



- Challenge is to get a capable mission suite to high-priority sites
- Most efficient approach is to convey a single instrument suite to multiple locations on a lander with aerial mobility
  - Modern control electronics make a multi-rotor vehicle (Langelaan et al. 2017) mechanically simpler than a helicopter, cf. proliferation of terrestrial quadcopter drones
  - > Improved flight control authority and surface sampling capability
  - > Redundant and failure tolerant
  - > Straightforward to test system on Earth
  - > Efficient to package in entry vehicle
- Atmosphere provides means to access different geologic settings 10s–100s km apart

#### Surface measurements

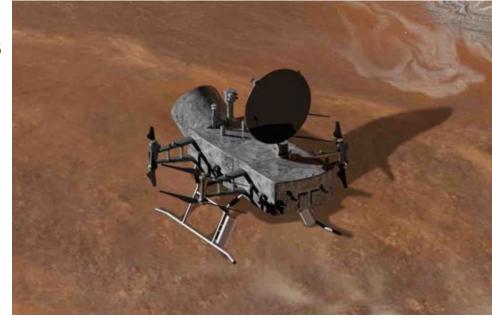


- Sample surface material into a mass spectrometer to identify chemical components available and processes at work to produce biologically relevant compounds.
- Measure bulk elemental surface composition with neutron-activated gamma-ray spectrometer.

 Monitor atmosphere and surface conditions with meteorology sensors and remote sensing instruments, including diurnal and

spatial variations

- Characterize geologic features with remote sensing instruments; also provide context for samples and scouting for scientific targets
- Perform seismic studies to detect subsurface activity and structure.







- Atmospheric profiles, including diurnal and spatial variations
- Aerial imagery for scouting landing sites and surface geology



- Concept of operations similar to rovers:
  - > Science activities while landed and some in flight
  - Use aerial scouting observations to identify sites of highest scientific potential for characterizing prebiotic chemistry, Titan's environment, and its habitability to inform prioritization of activities
- More relaxed pace with 16-day Titan-sols







- Participating Scientist Program
  - > Prockter et al. 2016, DPS
  - > Rathbun et al. 2017, Planetary Science Vision 2050 Workshop, Poster 8079: The Planetary Science Workforce: Goals through 2050
- E/PO
- Earth-based Titan weather campaign
- Extended mission
  - > Follow up on discoveries
  - > Access more distant targets



#### **Dragonfly**

A rotorcraft lander to investigate prebiotic organic chemistry and habitability using Titan's unique organic laboratory to understand how far chemistry can progress in environments that provide known key ingredients for life















