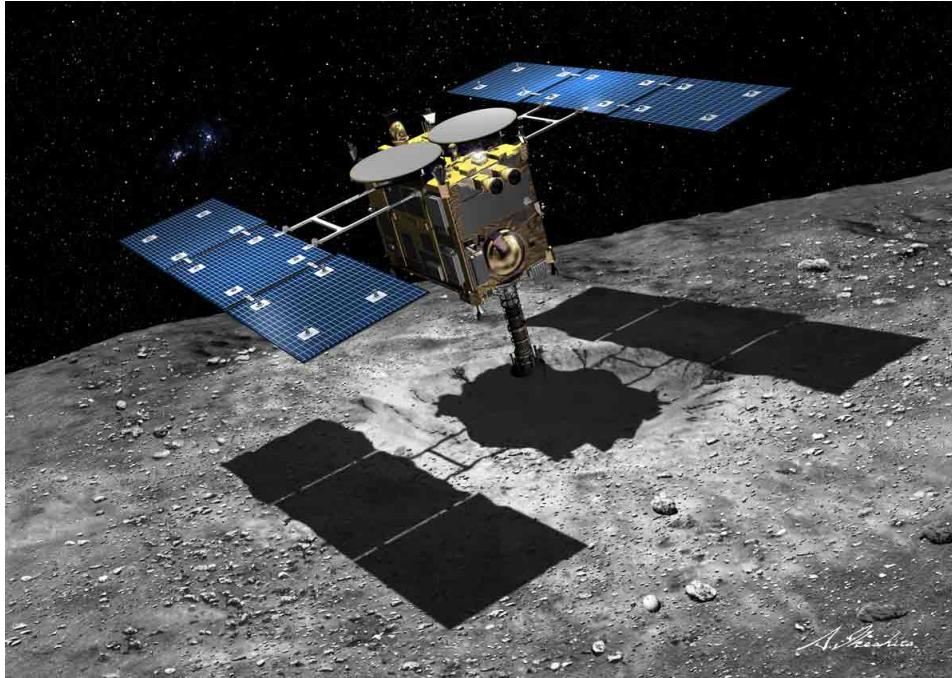


# Hayabusa2



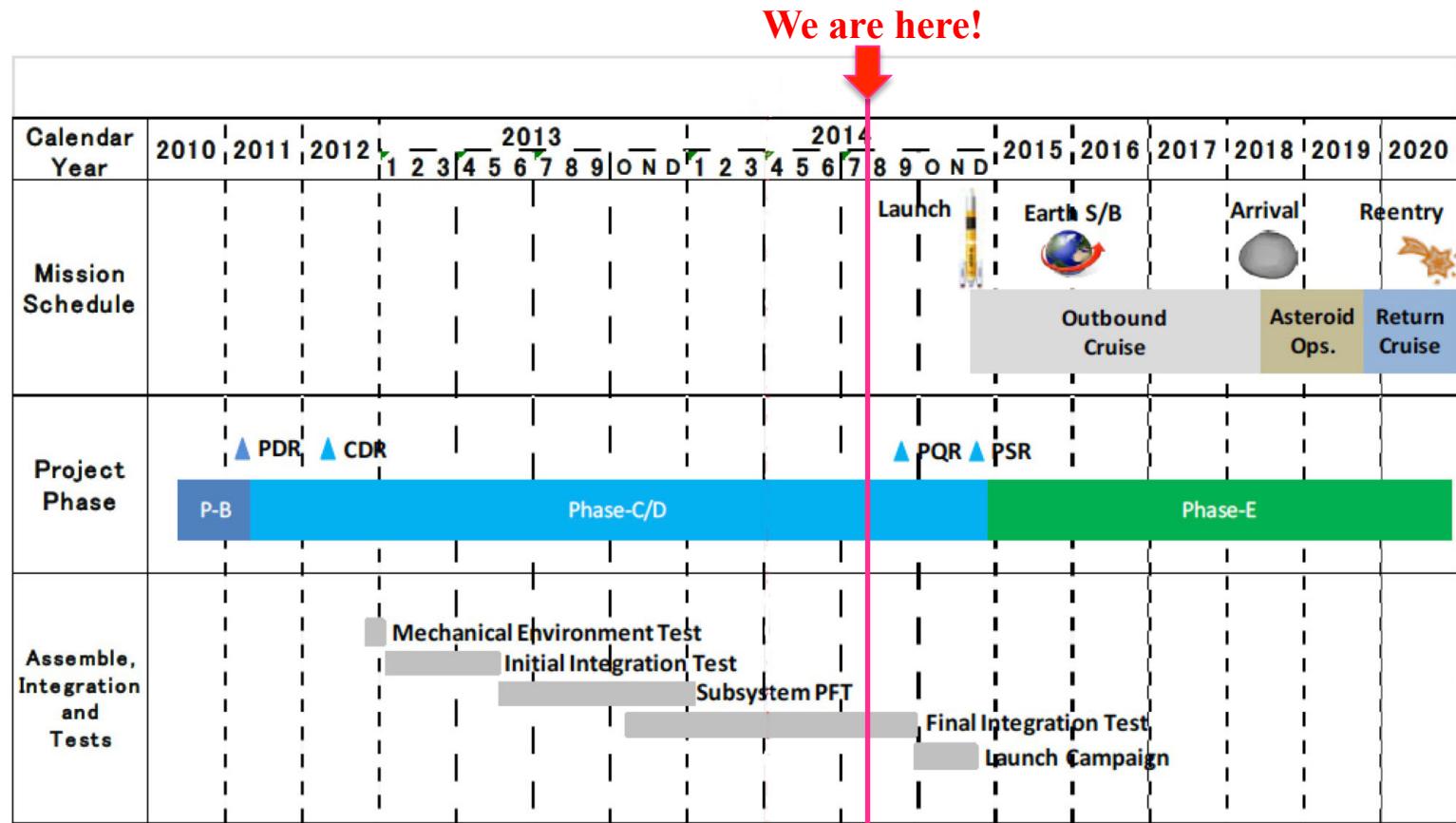
**SBAG, Washington DC, USA  
29 July 2014**

**Presented by Paul Abell**

**H. Kuninaka, N. Inaba, Y. Tsuda, M. Yoshikawa (JAXA)**

# Current Status

- We are now doing final integration test, which will be continued to September.
- Then the spacecraft is shipped to Tanegashima Space Center.
- The launch is scheduled in the winter of 2014.



# Objectives : Hayabusa vs Hayabusa2

## Hayabusa

### Technological demonstrator

- Round-trip to asteroid
- Sample return

### Engineering

- Ion engine
- Autonomous navigation
- Sample collection
- Reentry capsule

### Science :

Origin and evolution of the solar system

- Remote sensing observation
- Sample analysis

## Hayabusa2

### 1. Science

- Origin and evolution of the solar system
- Organic matter, H<sub>2</sub>O

### 2. Engineering

- Technology : more reliable and robust
- New challenge : ex) impactor

### 3. Exploration

- Extend the area that human can reach
- Spaceguard, Resources, Research for manned mission, etc.

C-type Asteroid

S-type Asteroid

# Hayabusa2 Mission Outline

Launch

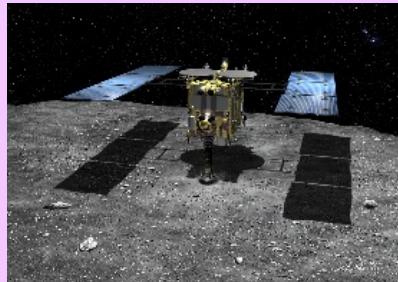
2014



July 2018 : Arrival at 1999 JU3

Sample analysis

- *In-situ remote science*
- *Rovers deployment*
- *Surface Sampling*
- *Impact*



Summer 2018 – Winter



Earth Return

Dec. 2020



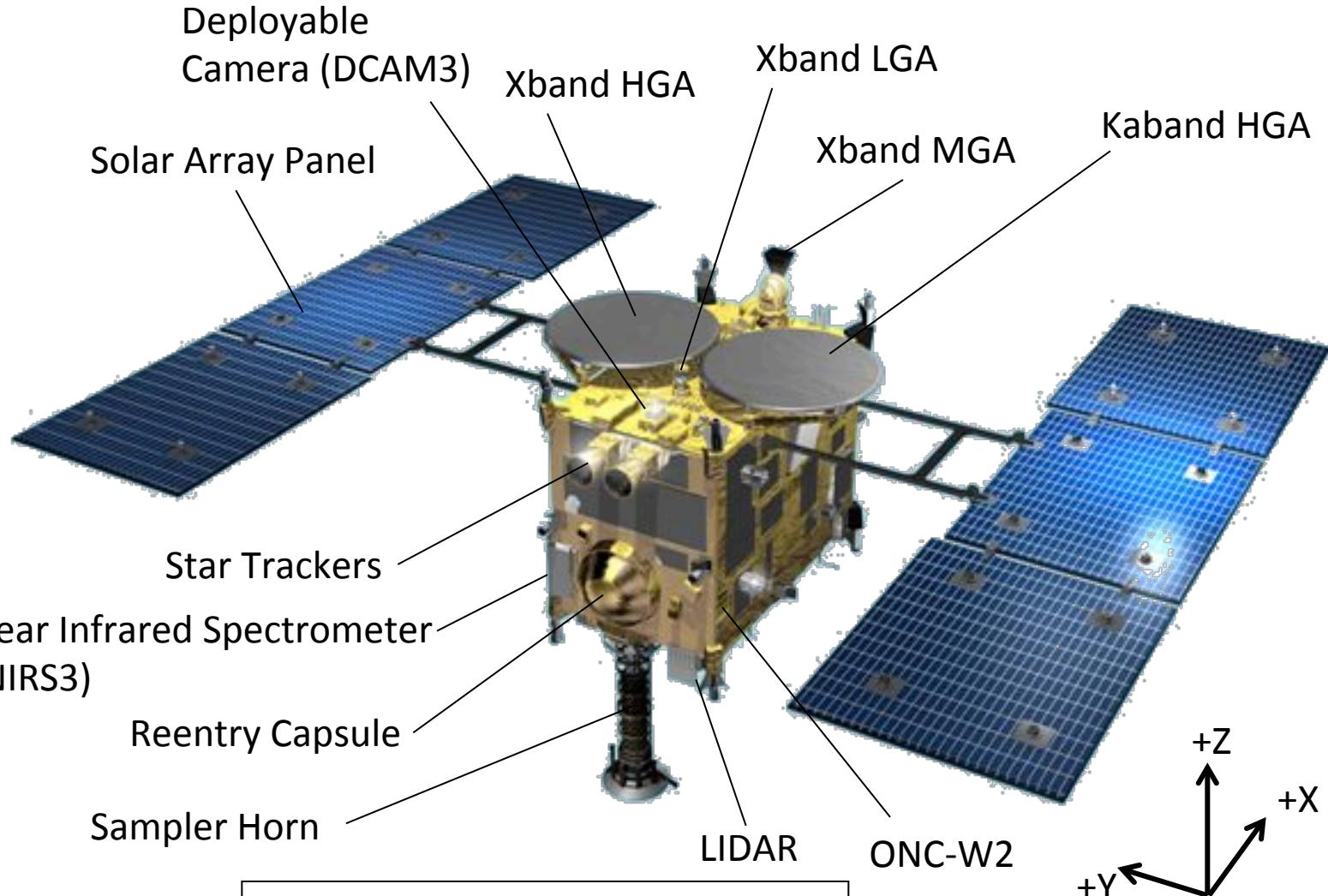
Dec. 2019 : Departure



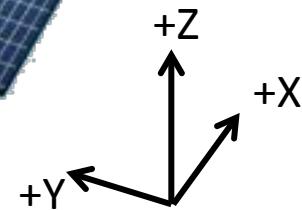
# Hayabusa2 Mission CG



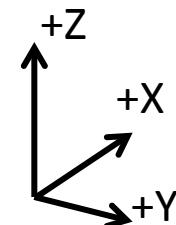
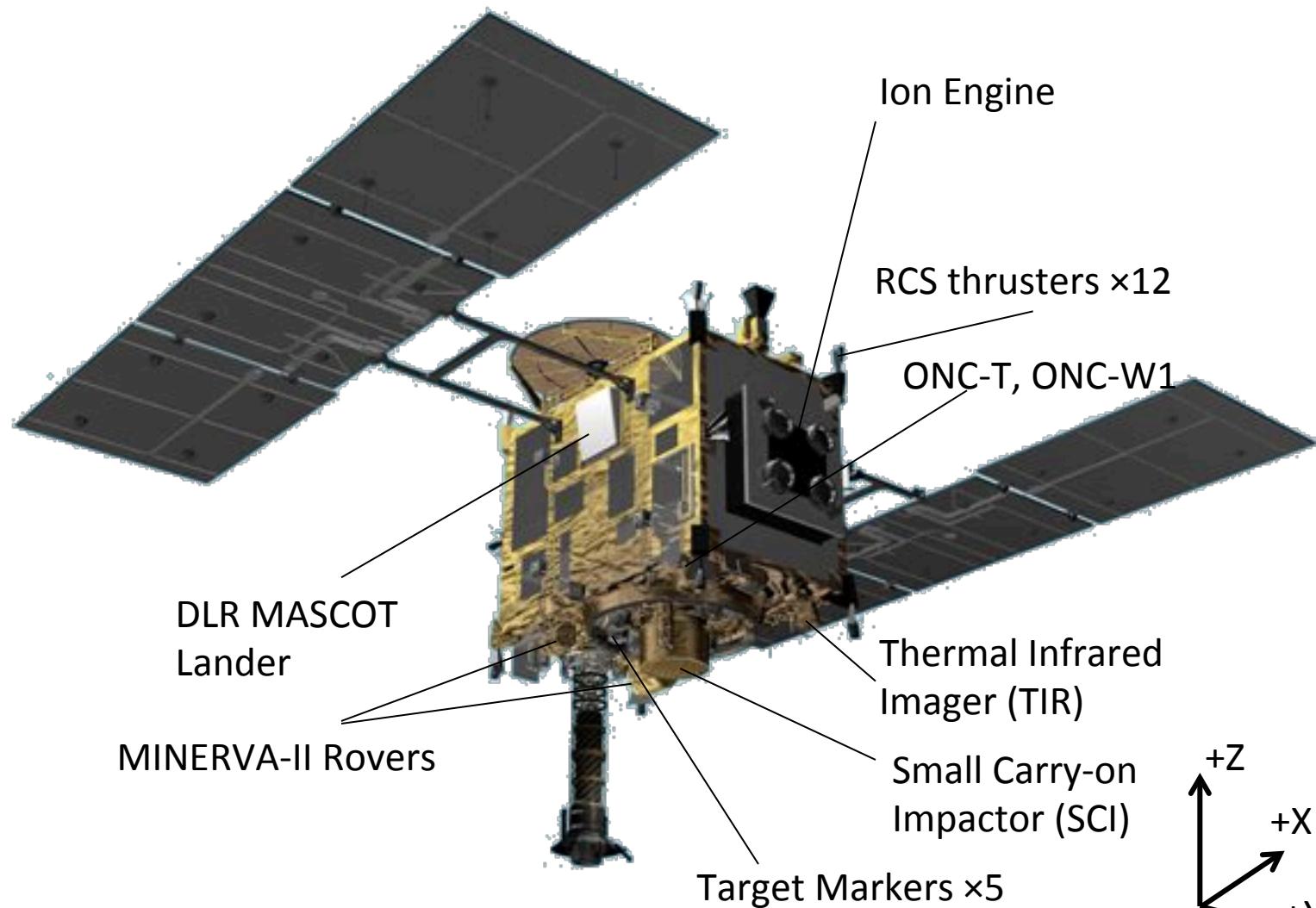
# External View of Spacecraft (1/2)



Size : 1m×1.6m×1.25m (body)  
Mass: 600kg (Wet)



# External View of Spacecraft (2/2)



# Spacecraft System Specifications

|  |   |                             |  |
|--|---|-----------------------------|--|
| <b>Structure</b>                         | -1.6m × 1.0m × 1.4m(Height) box structure with two fixed SAPs<br>-Mass 600kg(wet), 500kg(dry)   | <b>Communication System</b> | -Xband TT&C (Coherent Xup/Xdown), 8bps-32Kbps, Double redundant.<br>-Kaband Telemetry (Coherent Xup/Ka-down), 8bps-32Kbps<br>-Normal/Regenerative Ranging System<br>-DDOR support<br>-1 X-HGA, 1 Ka-HGA, 1 two-axis gimballed X-MGA, 3 X-LGA |
| <b>Data Handling System</b>              | -COSMO16 based DHU-PIM bus system<br>-Onboard autonomous command generation feature<br>-1Gbyte Data Recorder  |                             |  |
| <b>Attitude and Orbit Control System</b> | -HR5000S based processor, Double redundant.<br>-4 Reaction Wheels, 2 IRUs, 2 Star Trackers (STT), 4 Coarse Sun Aspect Sensors (CSAS), 4 accelerometers (ACS).<br>-Sensors for Proximity Operation LIDAR, LRF, 5 Target Makers (TM), Flash lamp (FLASH)<br>-Optical Navigation Cameras (ONC)<br>Wide: ONC-W1, ONC-W2 (FOV 54deg × 54deg, 1Mpix)<br>Telescopic: ONC-T (FOV 5.4deg × 5.4deg, 1Mpix, 5 band filter) | <b>Power System</b>         | SAP<br>-1.4kW@1.4AU, 2.6kW@1AU.<br>BAT<br>-Li-ion Battery 13.2AH.<br>Power Bus<br>-Series Switching Regulator(SSR) System, 50V Bus.  |
| <b>Propulsion System</b>                 | RCS<br>-Bi-propellant hydrazine system<br>- 20N thruster × 12.<br>IES<br>-Xe microwave discharge ion engine system<br>-Maximum thrust 28mN, Isp=2800sec.<br>-4 thruster heads on gimballed stage<br>-3 operative at once (4/3 redundant)  | <b>Mission Payload</b>      | -Sampler Horn (SMP)<br>-Small Carry-on Impactor(SCI)<br>-Near Infrared Spectrometer(NIRS3)<br>-Thermal Infrared Imager(TIR)<br>-3 Rovers(MINERVA-II-1A/1B/2)<br>-Lander(MASCOT)<br>-Deployable Camera(DCAM3)<br>-Reentry Capsule(CPSL)       |

# Science Instruments

- **Sampler Horn** ● Newly developed for HYB2
  - Increased sample container rooms (2→3 rooms)
  - Metal seal adopted for gas-sealing performance
  - Several improvements in mechanical design to increase yield performance
- **NIRS3(Near Infrared Spectrometer)** ●
  - Newly developed for HYB2
  - Observation wavelength 1.7-3.4μm. (Water)
  - Passive cooling
- **TIR(Thermal Infrared Imager)**
  - Planet-C (AKATSUKI)-derived component
  - Observation wavelength 7-14μm. (Thermal distribution, Hydrated minerals distribution)
- **Lander(MASCOT)** ●
  - Developed and provided by DLR
  - 10kg surface exploration lander
  - Detail observation of local surface (MARA, MicrOmega, MAG, CAM)
- **Rover(MINERVA-II 1A / 1B / 2)** ●
  - Developed by JAXA and a Japanese university consortium
  - Three 1.5kg class surface exploration rovers
- **Optical Navigation Cameras(W1/W2/T), LIDAR**
  - Multipurpose components both for bus and scientific operation
  - ONC:Global mapping/asteroid relative navigation, LIDAR:altimeter/gravity determination
- **Small Carry-on Impactor(SCI) ● , Deployable Camera(DCAM) ●**
  - 18kg kinetic impact system (detonator+Cu liner+Electronics), 2kg impact mass at 2km/s impact speed.
  - Impact observation by IKAROS-derived DCAM(remote camera)

# Trajectory Design

**Earth Departure:** 2014

**Earth Swing-by:** 2015/12

**1999JU3 Arrival:** 2018/7

**1999JU3 Dep.:** 2019/12

**Earth Reentry:** 2020/12

**Departure C<sub>3</sub>=21km<sup>2</sup>/s<sup>2</sup>**

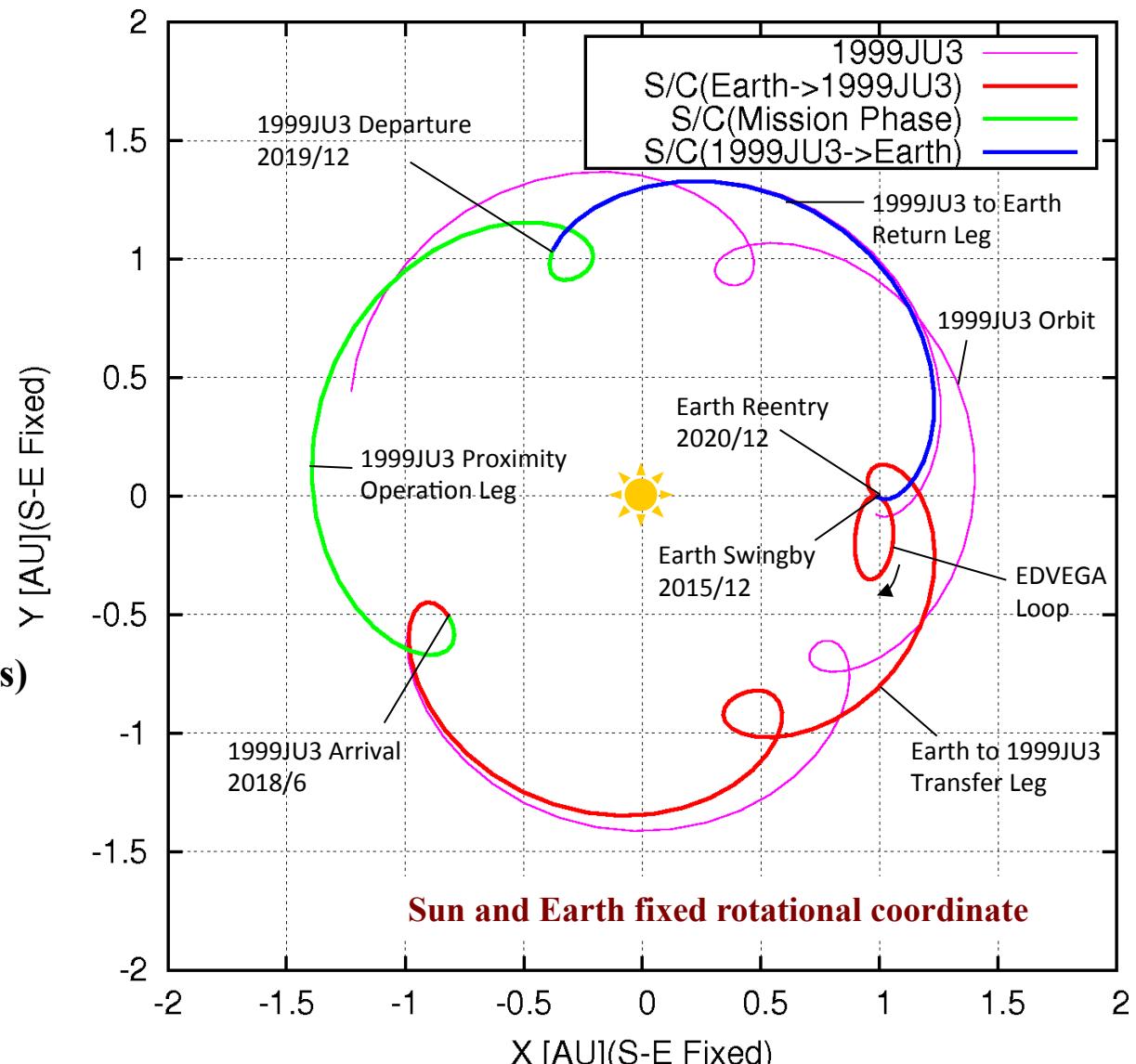
**IES Total Impulse=2km/s**

**Reentry Speed=11.6km/s**

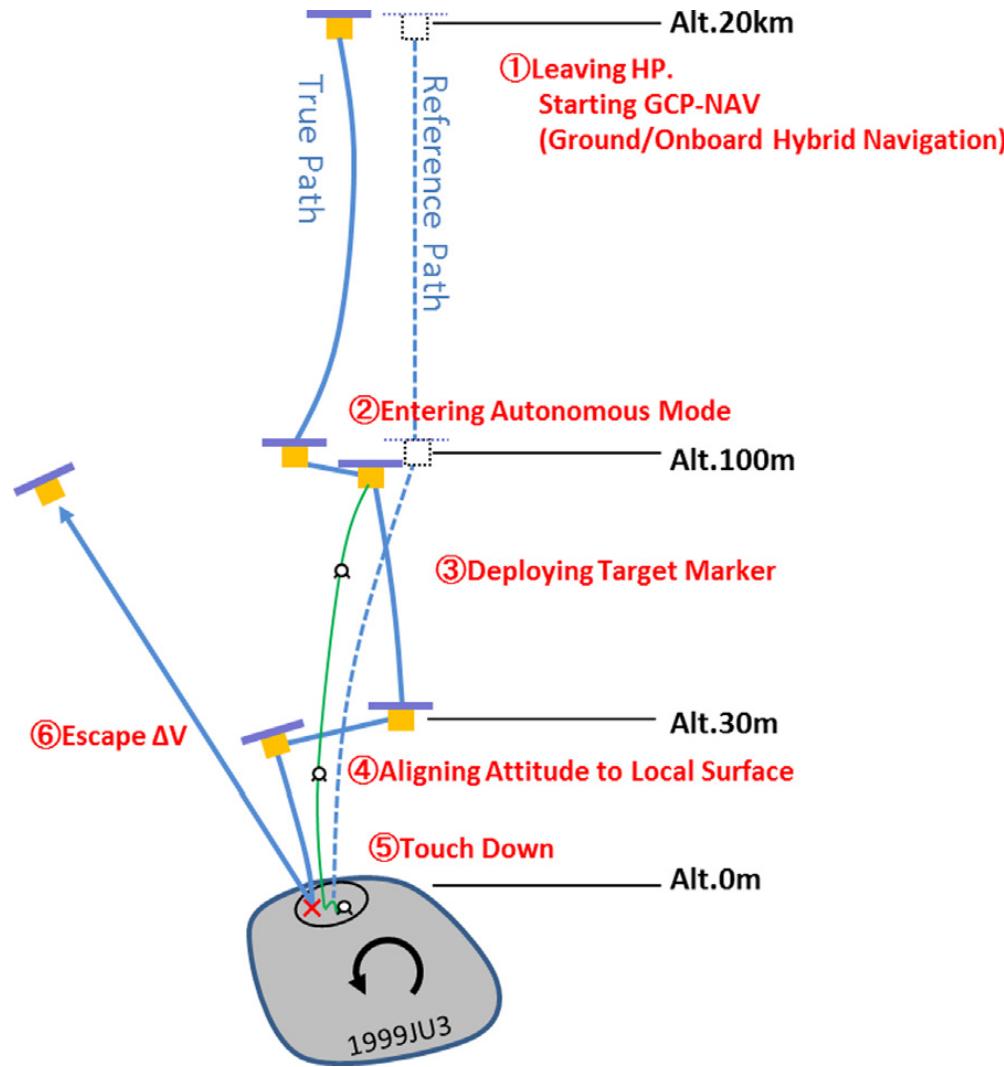
**Total Flight Time=6yr**

**(Cruising 4.5yrs)**

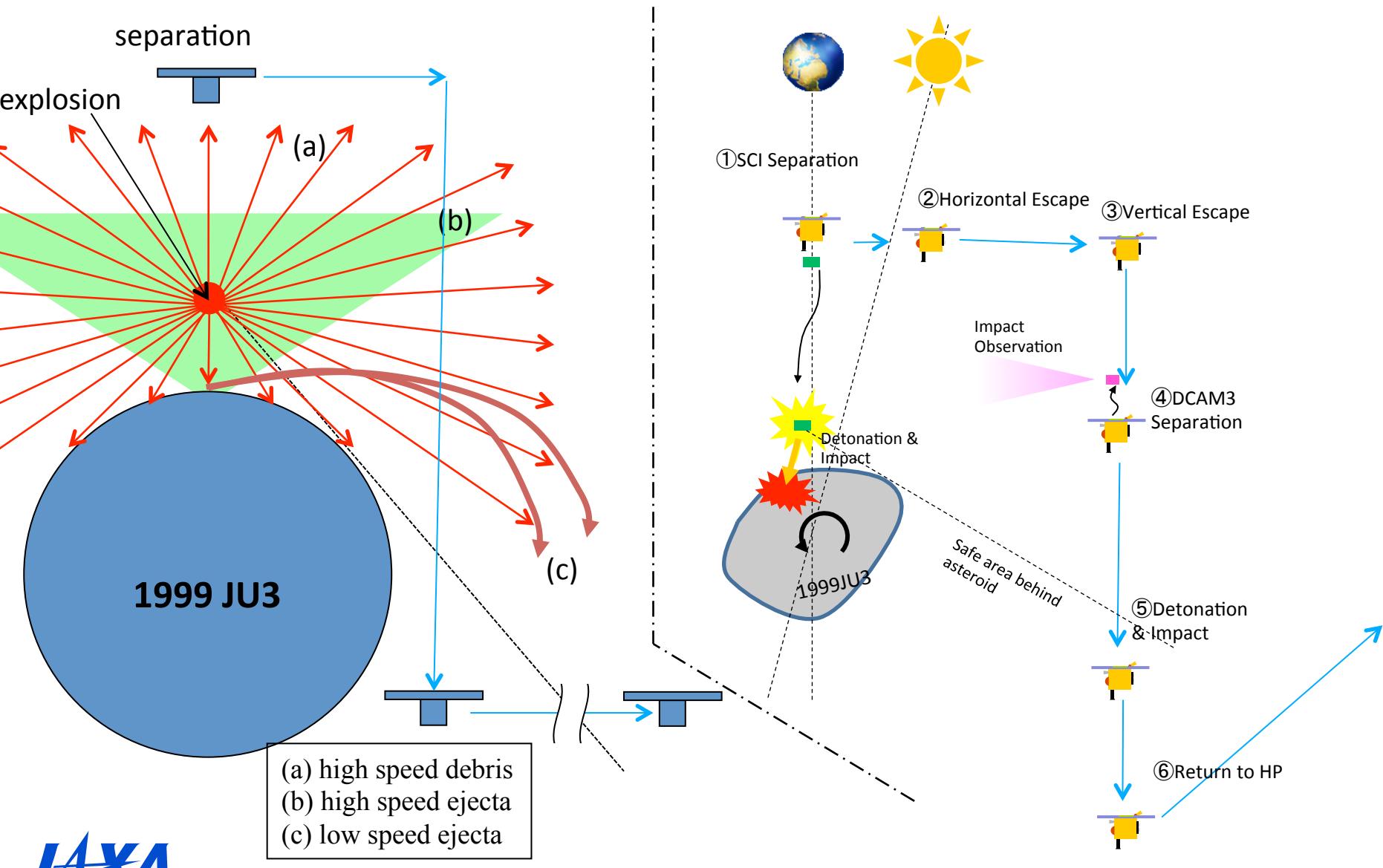
**Total Powered Flight Time  
=1.5yr**



# Sampling Operation Sequence

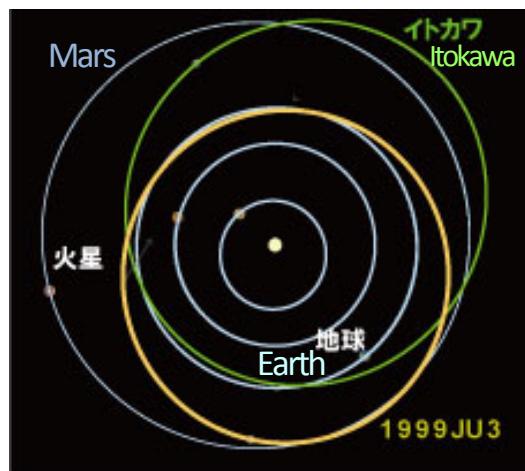


# Artificial Crater Generation Operation



# Target : (162173) 1999 JU<sub>3</sub>

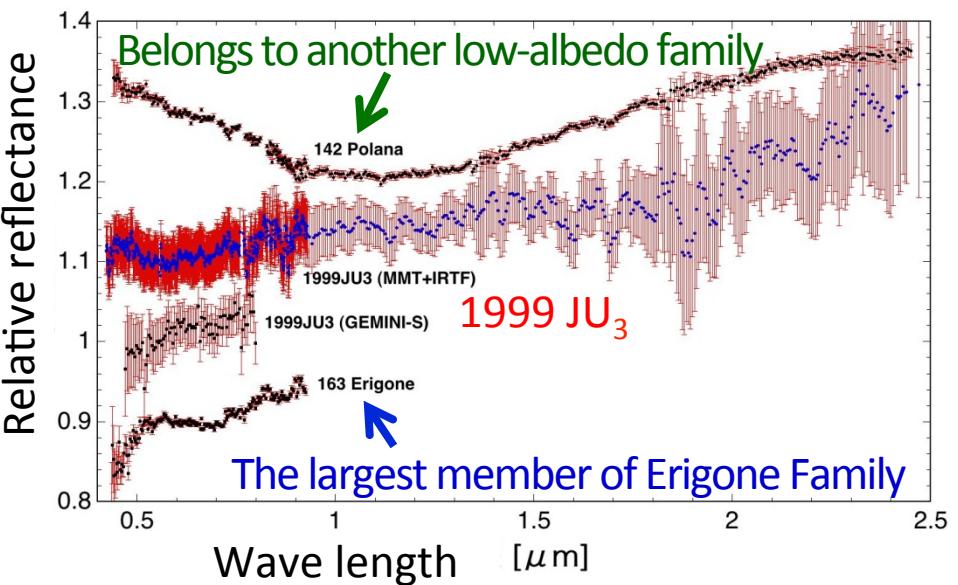
- C-type near-Earth asteroid
- Diameter: 0.9 km (nearly spherical?)
- Hydrated minerals, organic matter ?
- Origin: inner MB,  $v_6$  secular resonance  
Erigone Family ( $\sim 300$ Ma) ?



$$\begin{aligned}a &= 1.19 \text{ AU}, T = 1.30 \text{ yr} \\q &= 0.96 \text{ AU}, Q = 1.42 \text{ AU} \\e &= 0.19, i = 5.88^\circ\end{aligned}$$

P. Michel and M. Delbo 2010  
H. Campins, et al. 2012

origin of 1999 JU<sub>3</sub>  
92%  $v_6$   
8% 3:1



Data from Vilas 2008, Sugita+ 2012, Abe+ 2008

# Physical Properties of 1999 JU3

Observation results : by M. Ishiguro & D. Kuroda

- Photometric Properties

- Rotational period  $7.6312 \pm 0.0010$  hr (\*1)
- $H_0$  in V-band:  $19.25 \pm 0.03$  (\*2)
- G parameter ( $\alpha < 30^\circ$ ):  $0.13 \pm 0.02$  (\*2)
- Geometric albedo:  $0.047 \pm 0.003$  (\*1-2)

(As of 2014 July)

- Spectroscopic Properties

- Spectral type: C-type (\*3-6)
- Absorption around  $0.7 \mu\text{m}$  (\*4)
- Rotational variability: Not obvious (\*5-6)

- Thermal Properties (\*1)

- Effective diameter  $875 \pm 15$  m
- Thermal inertia  $250 \pm 50 \text{ J m}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$
- Pole orientation Probably retrograde (ongoing)

[1] Mueller et al. ACM 2014 [1] Ishiguro et al submitted [3] Binzel et al. 2001, Icarus 151, 139

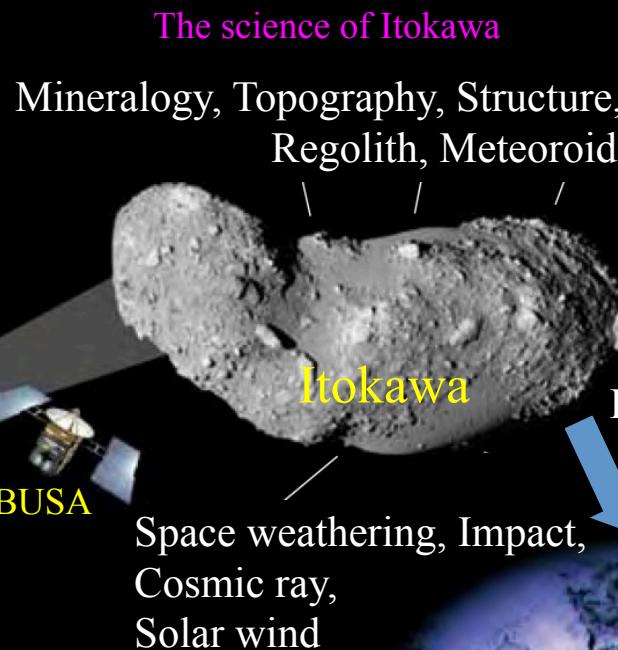
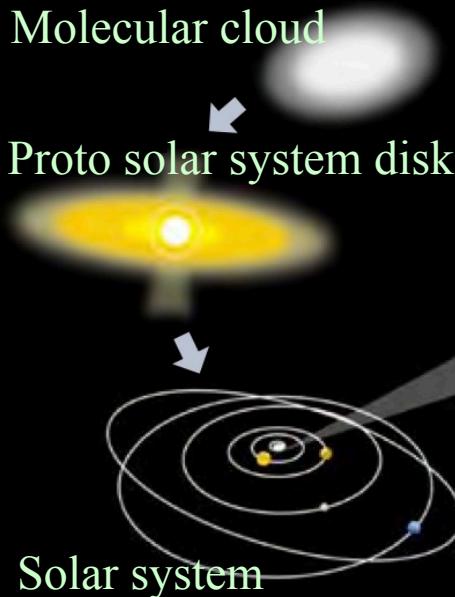
[4] Vilas 2008, AJ 135, 1191 [5] Lazzaro et al. 2013, A&A 549, L2 [6] Moskovitz et al. 2013, Icarus 224, 23,

# Science of Hayabusa and Hayabusa2

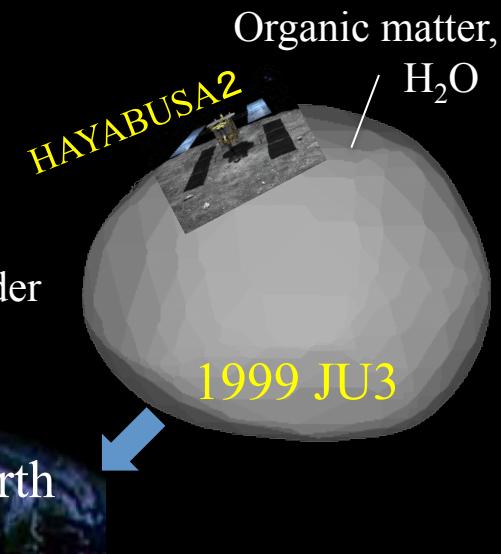
## Study of the origin and evolution of the solar system

- Planetesimal formation : Destruction and accumulation
- Thermal evolution from planetesimal to near earth asteroid
- Diversification of organics through interactions with minerals and water on planetesimal
- Material circulation in the early solar system

**4.6 billion years ago...**

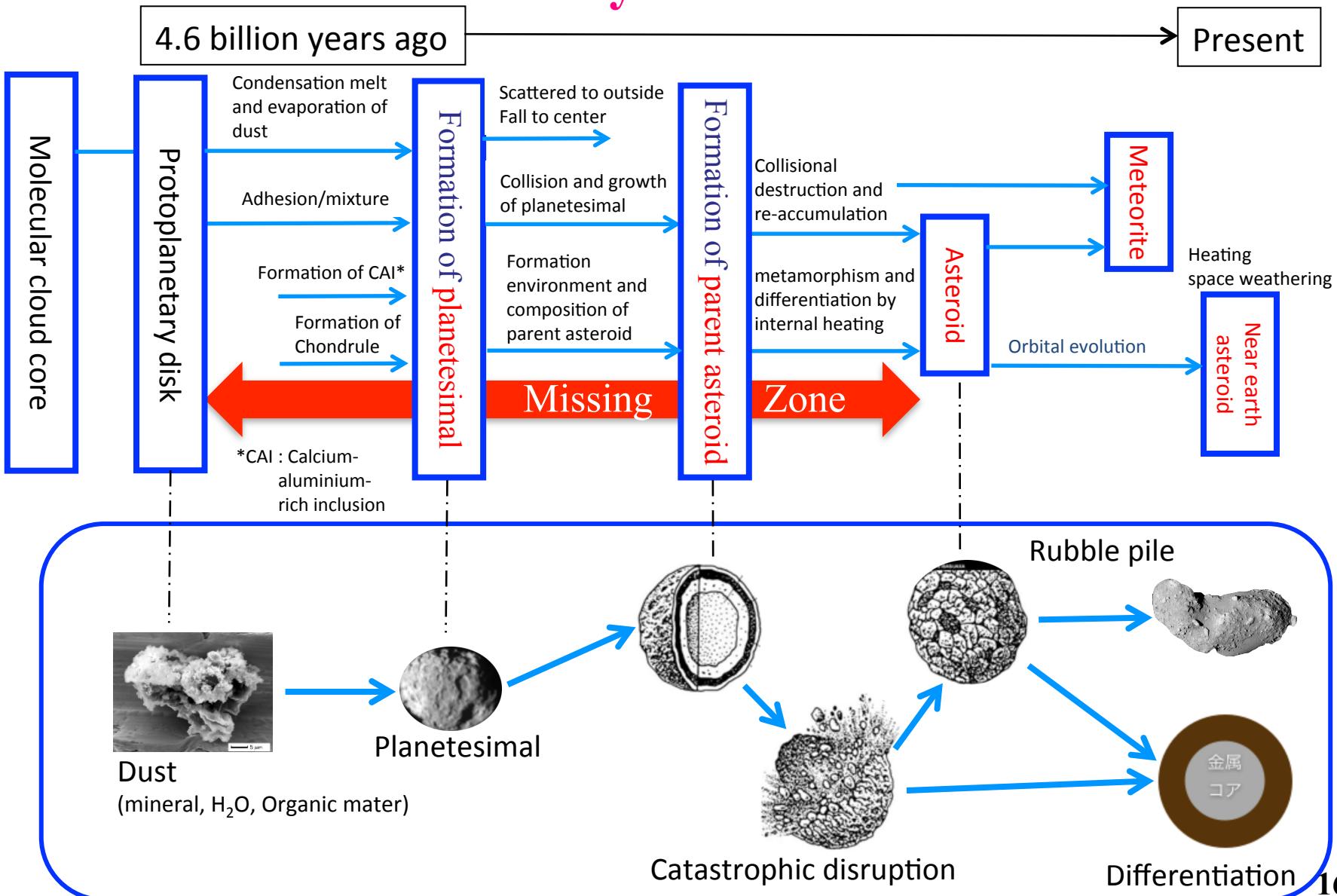


In addition to the  
science of Itokawa...

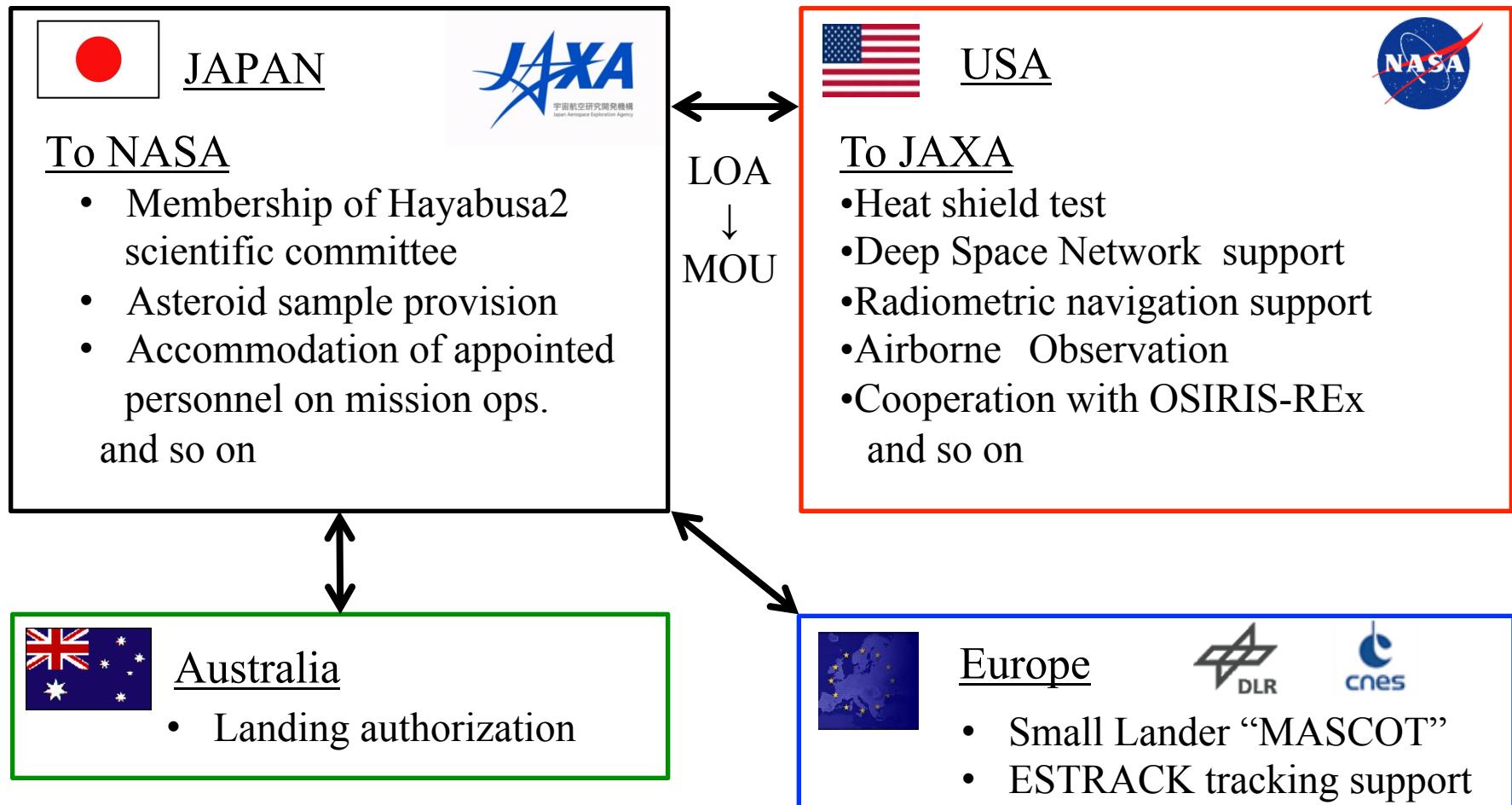


# Hayabusa2 will solve the "Missing Zone"

## History of Asteroid



# International Collaboration Structure



# Hayabusa2 Joint Science Team (HJST) Meeting



1<sup>st</sup> : November 26-27, 2012



2<sup>nd</sup> : September 19-20, 2013

3<sup>rd</sup> : March 18, 2014

# Collaboration with OSIRIS-REx Team

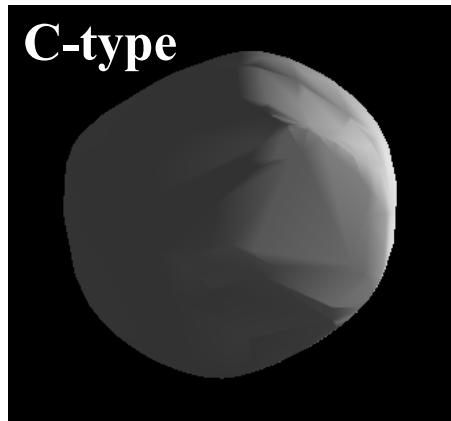
**Hayabusa2**



**OSIRIS-REx**



**C-type**



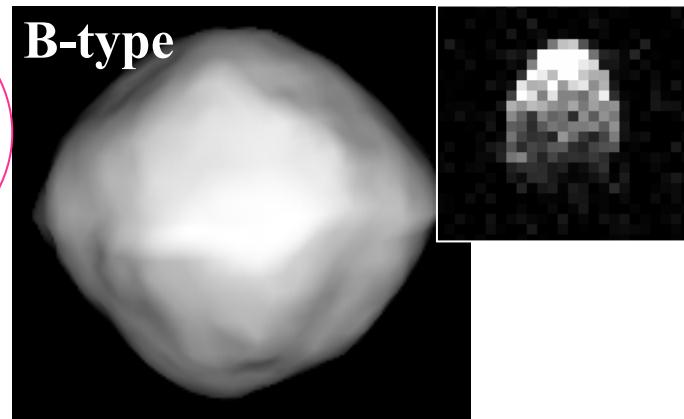
**(162173) 1999 JU3**

**Collaborations**

**Science  
Education  
Outreach**

...

**B-type**



**(101955) Bennu**

# **Summary**

- Hayabusa2 is the modified spacecraft based on the lessons learned of Hayabusa, and it will aim for much more advance science to understand the origin and evolution of the solar system.
- We are now in the final stage in the integration of the spacecraft, and it will be launched in the winter of 2014.

**Thank you!**