

Present Status of Japanese Venus Climate Orbiter

Current status of VCO development:

Mechanical and Thermal models were evaluated in 2007.

Critical Design Review has been carried out through Dec. 2007 to March 2008 and the mission has successfully been phased up.

The flight model is being manufactured with minor modifications from the mechanical and thermal models.

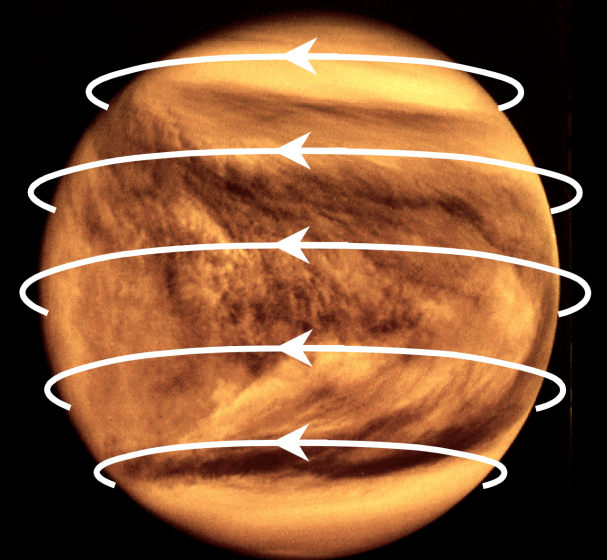
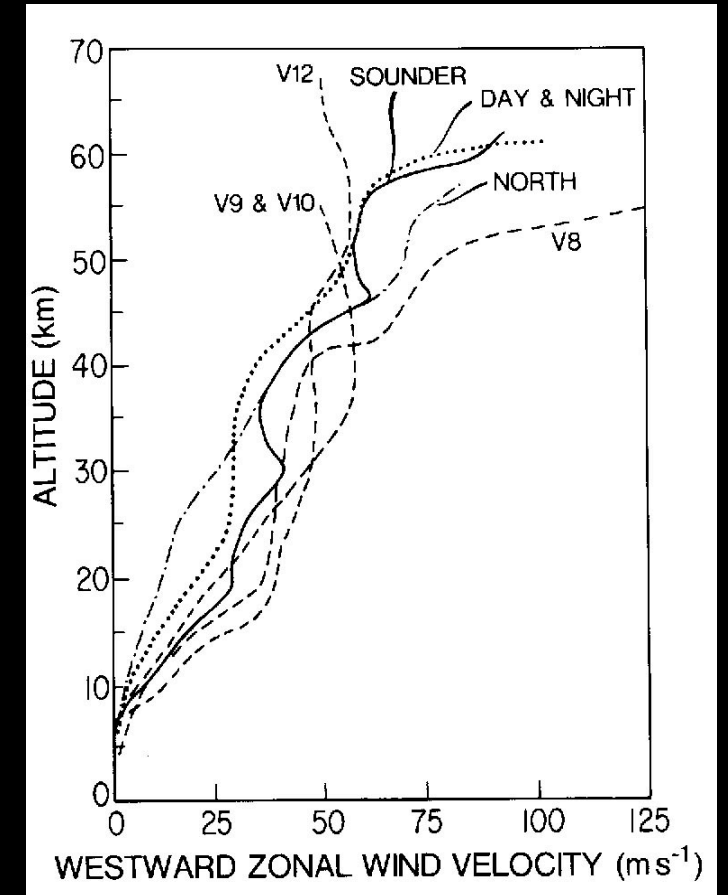


- *VCO will be the first Japanese Venus mission*
- *It is dedicated to the meteorological survey*
- *It will be launched in May 2010 and arrive at Venus in December 2010*
- *Nominal mission life is 2 years*
- *It has 5 cameras (UV, Visible, NIRx2, LIR) to investigate the atmospheric*

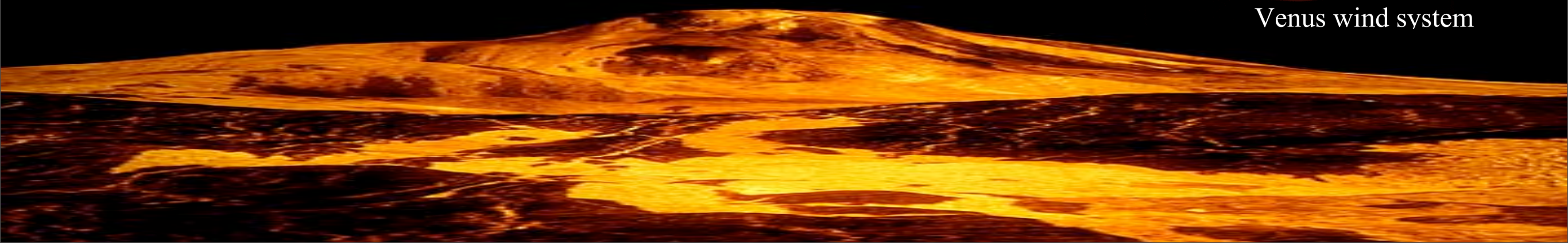
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Scientific Targets

- Atmospheric dynamics
 - Super rotation
 - Circulations
 - Meso-scale phenomena
- Lightning
- Cloud physics
- Active volcano
- Geological survey

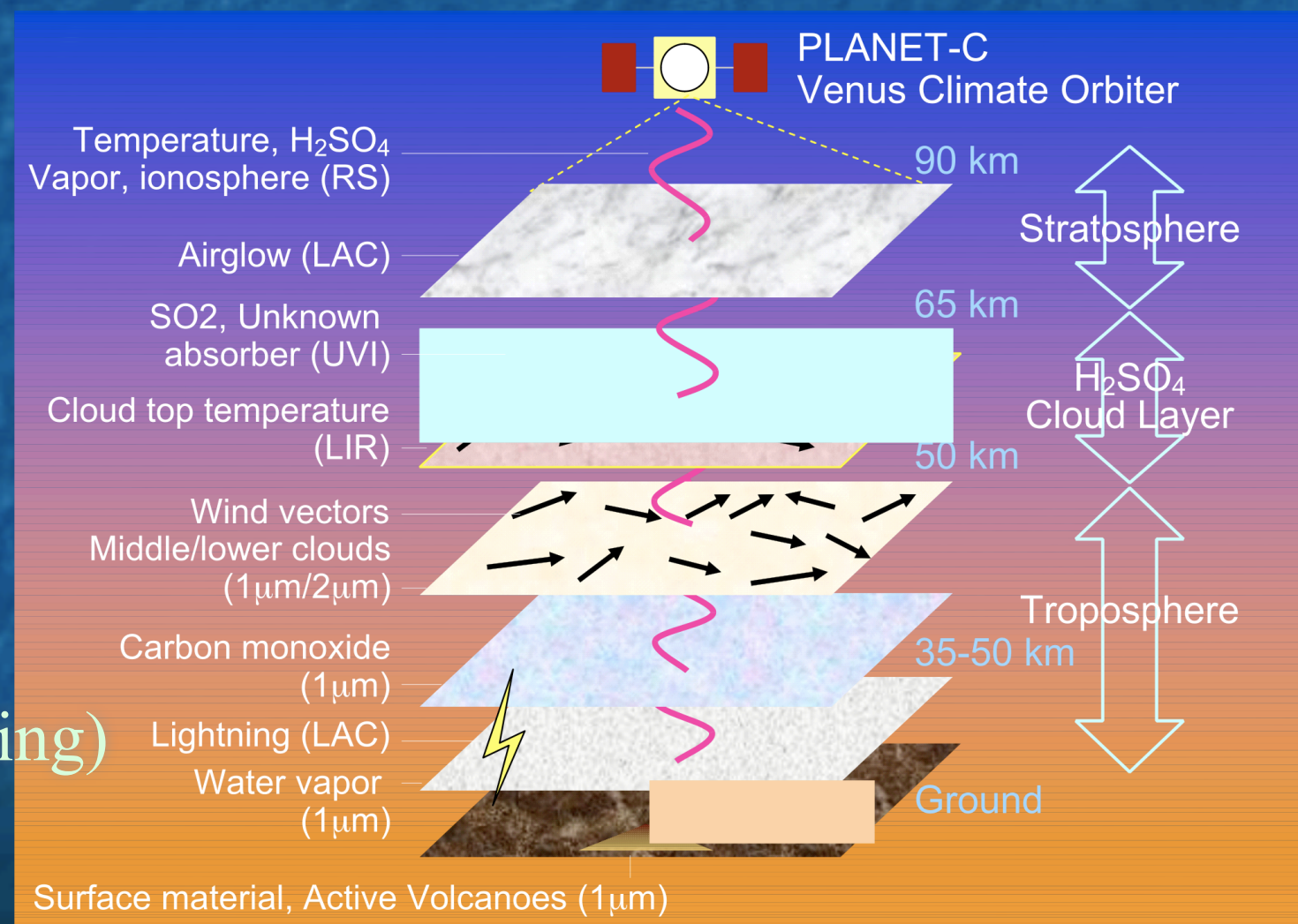


Venus wind system

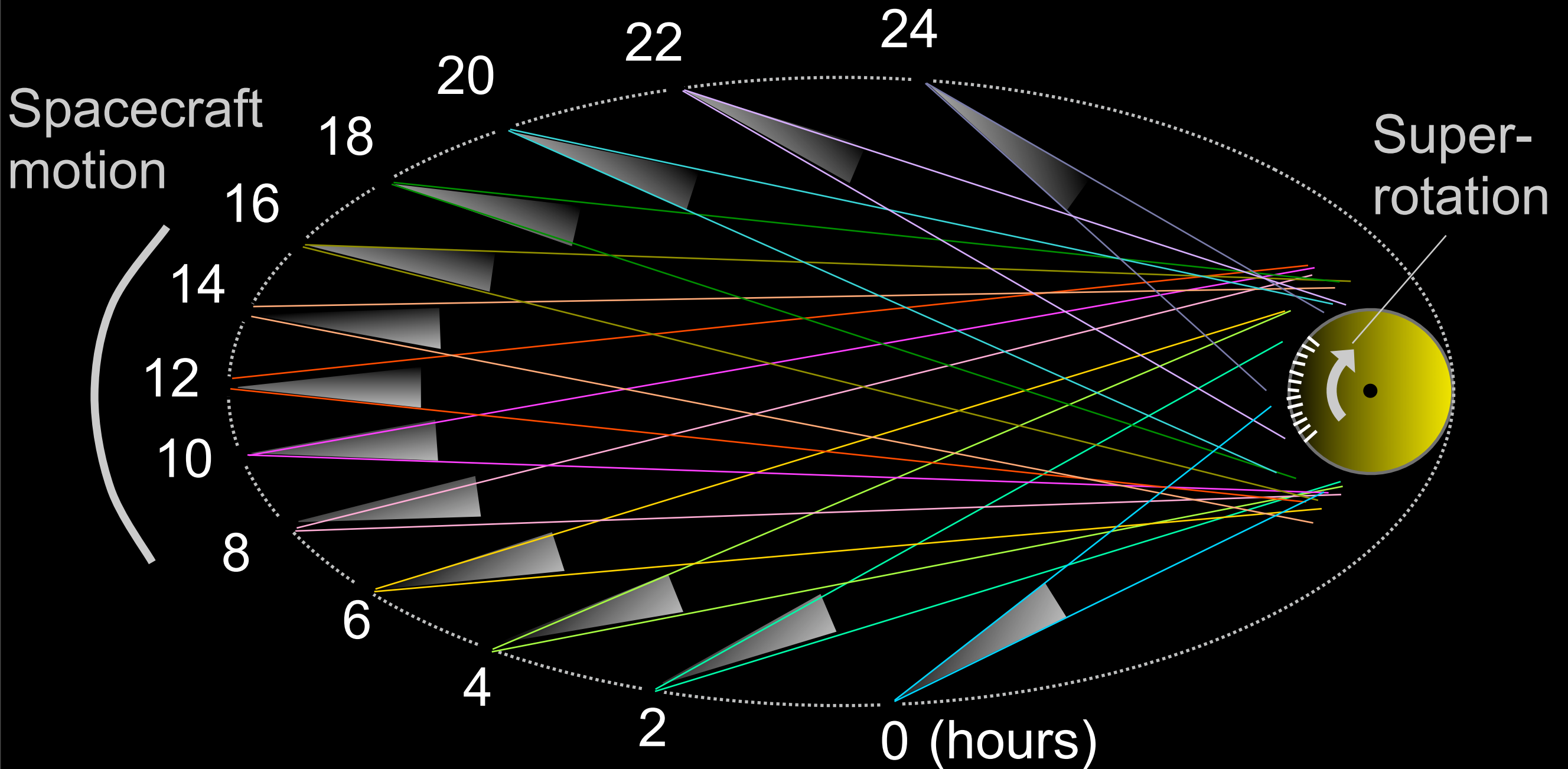


5 Cameras onboard PLANET-C

- **IR1** $\lambda = 0.9, 1.01 \mu\text{m}$
- 1024x1024 SiCCD
- **IR2** $\lambda = 1.73, 2.26, 2.32 \mu\text{m}$ (Venus atmosphere.),
- $2.02 \mu\text{m}$ (CO_2), $1.65 \mu\text{m}$ (IPD)
- 1024x1024 PtSi
- **LIR** $\lambda = 8-12 \mu\text{m}$
- 240x320 micro bolometer array
- **UVI** $\lambda = 280, 365 \text{ nm}$
- 1024x1024 UV intensified SiCCD
- **LAC** $\lambda = 777, 551, 558 \text{ nm}$
- 8x8 Av. photo-diode (50kHz sampling)
- **DE & USO**



Viewed from the north pole



Optimization for detecting small deviations of atmospheric motions from the background flow

Evaluation of the H-IIA mechanical environment

- Acoustic & Vibration test
- Marman clamp shock test
- SAP extension shock test

■ *Flight load factors*

Lift off

*Main engine
cut-off*

	<i>Axial</i>	<i>Lateral</i>
<i>compression</i>	$-3.2G$	$\pm 1.8G$
<i>tension</i>	$-0.1G$	$\pm 1.8G$
<i>Before MECO</i>	$-4.0G$	$\pm 0.5G$
<i>During MECO</i>	$+1.0G$	$\pm 1.0G$

■ *Sinusoidal vibration levels*

Axial

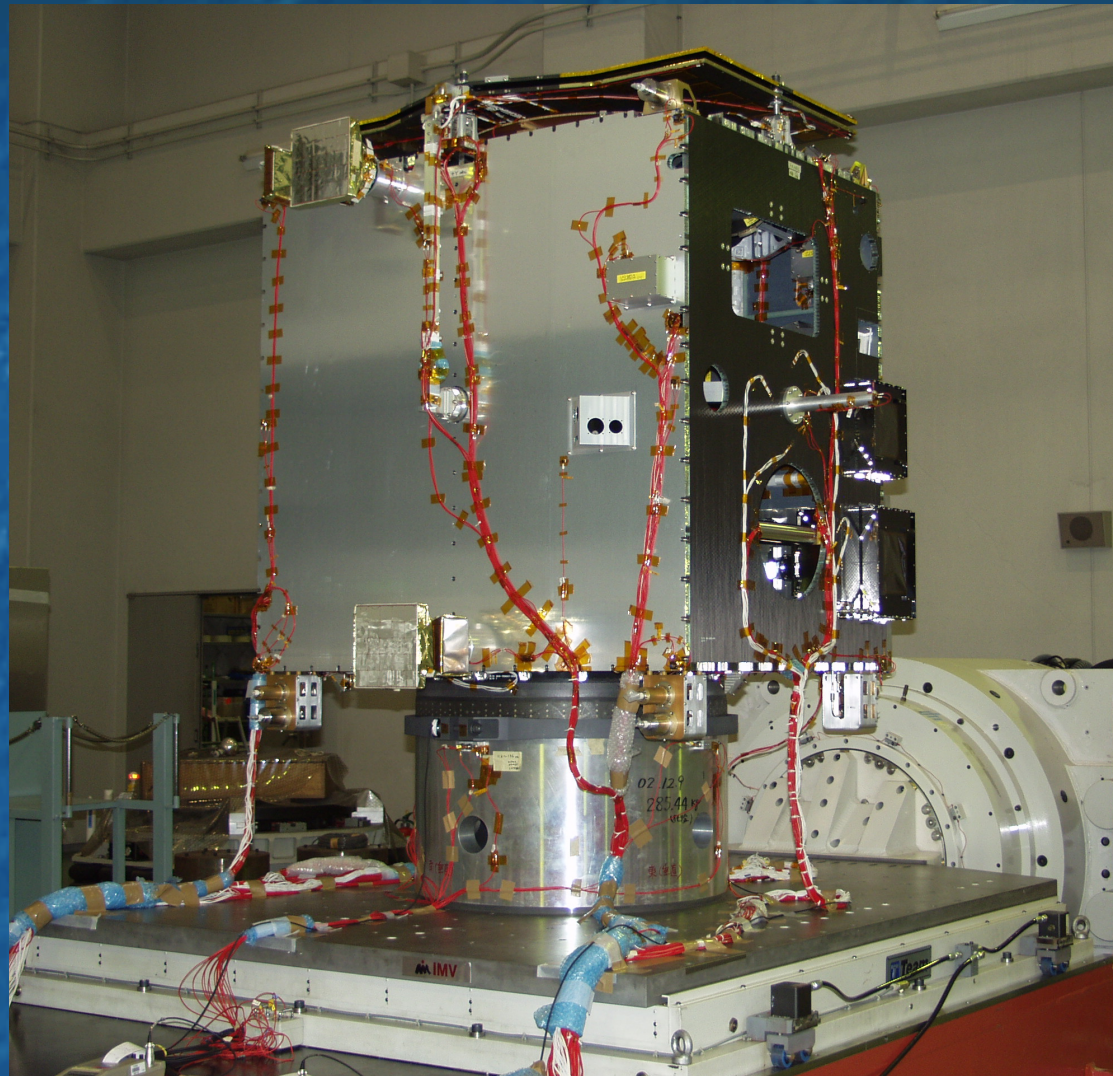
Lateral

<i>5~30Hz</i>	<i>1.0 Go-p</i>
<i>30~100Hz</i>	<i>0.8 Go-p</i>
<i>5~18 Hz</i>	<i>0.7 Go-p</i>
<i>18~100 Hz</i>	<i>0.6 Go-p</i>

Sound pressure levels

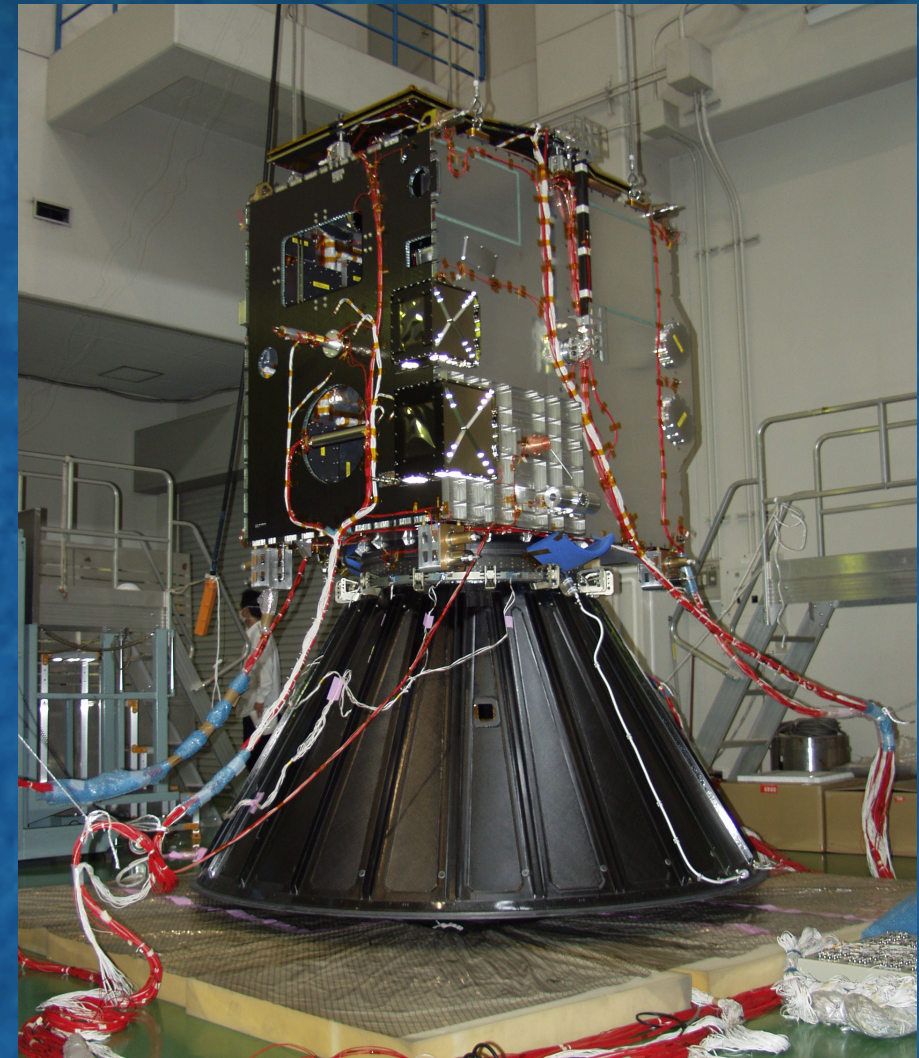
<i>Center Frequency(Hz)</i>	<i>SPL(dB)</i>
<i>31.5</i>	<i>125.0</i>
<i>63</i>	<i>126.5</i>
<i>125</i>	<i>131.0</i>
<i>250</i>	<i>133.0</i>
<i>500</i>	<i>128.5</i>
<i>1000</i>	<i>125.0</i>
<i>2000</i>	<i>120.0</i>
<i>4000</i>	<i>115.0</i>
<i>8000</i>	<i>113.0</i>
<i>O.A.</i>	<i>137.5</i>

PLANET-C Mechanical Test Model



Vibration test

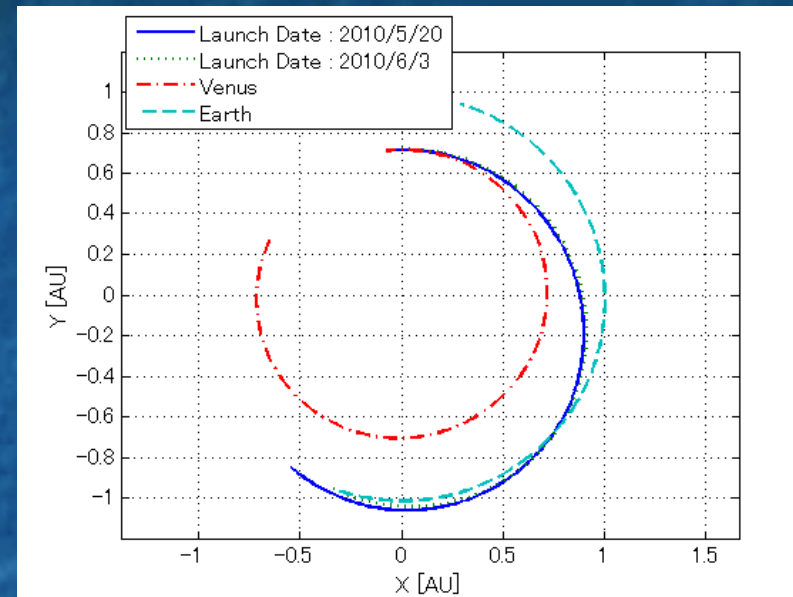
Eigen frequency



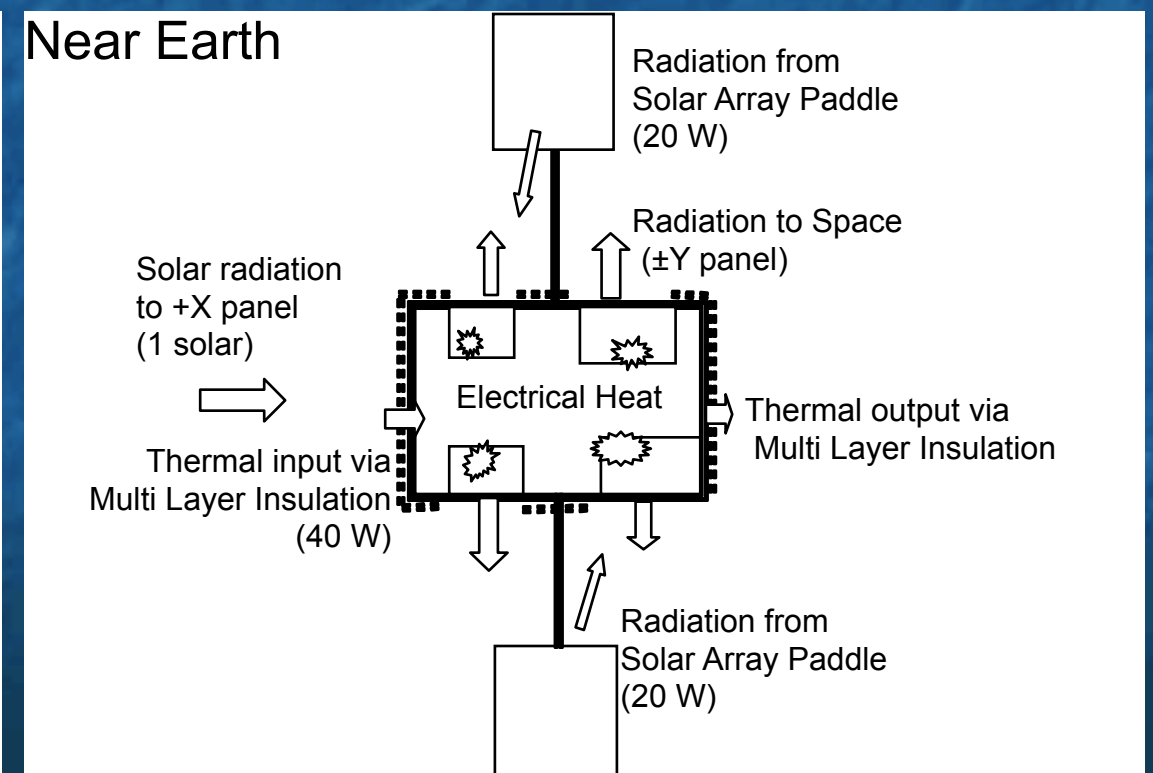
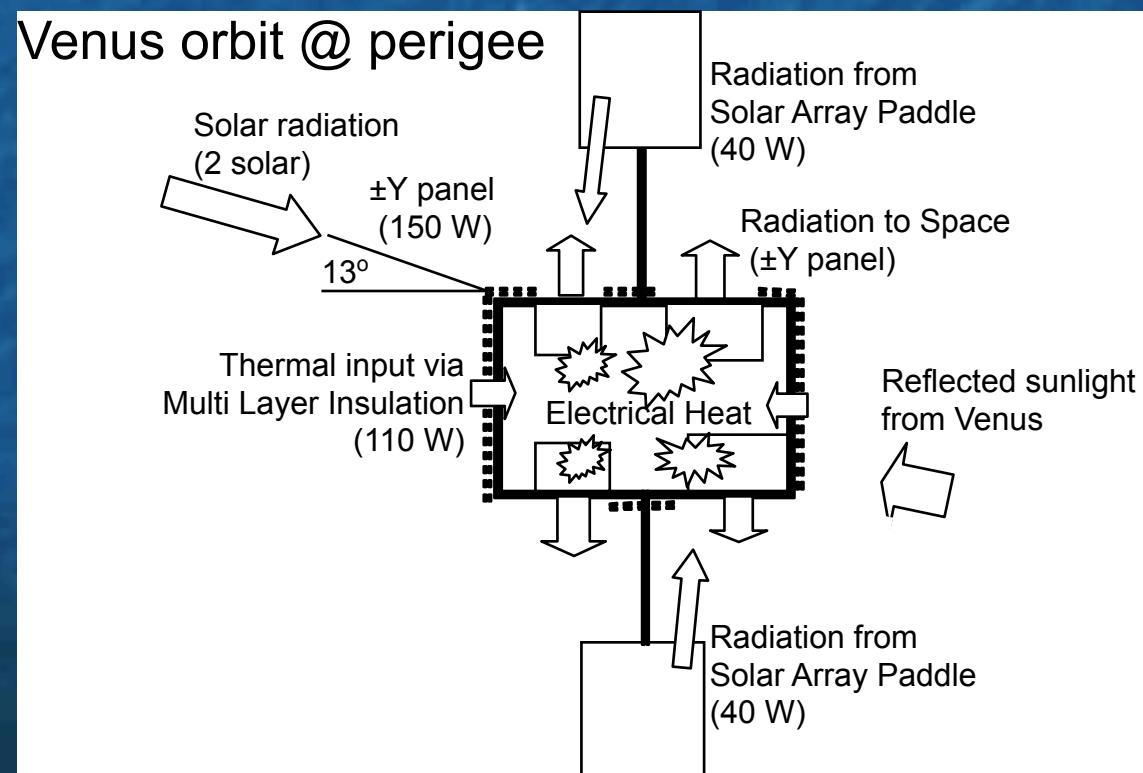
Marman clamp shock
test

	<i>Requirement</i>	<i>Analysis result</i>	<i>Test result</i>
<i>Axial</i>	>30Hz	63.7Hz (Z)	64.0Hz
<i>Lateral</i>	>10Hz	22.9Hz (X)	22.8Hz
		26.7Hz (Y)	26.2Hz

Transfer orbit from Earth to Venus



Heat input at a Venus orbit is twice larger than at an earth orbit

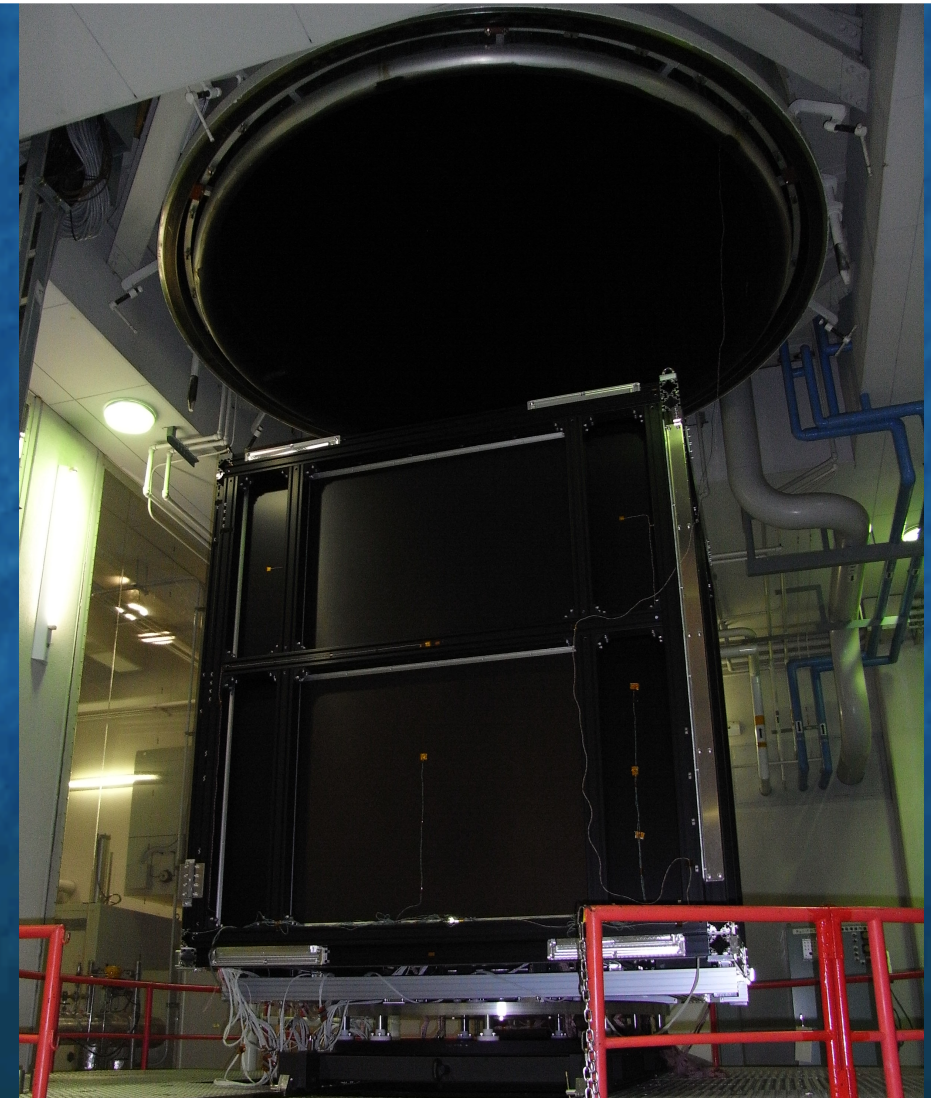


Evaluation of the thermal environment

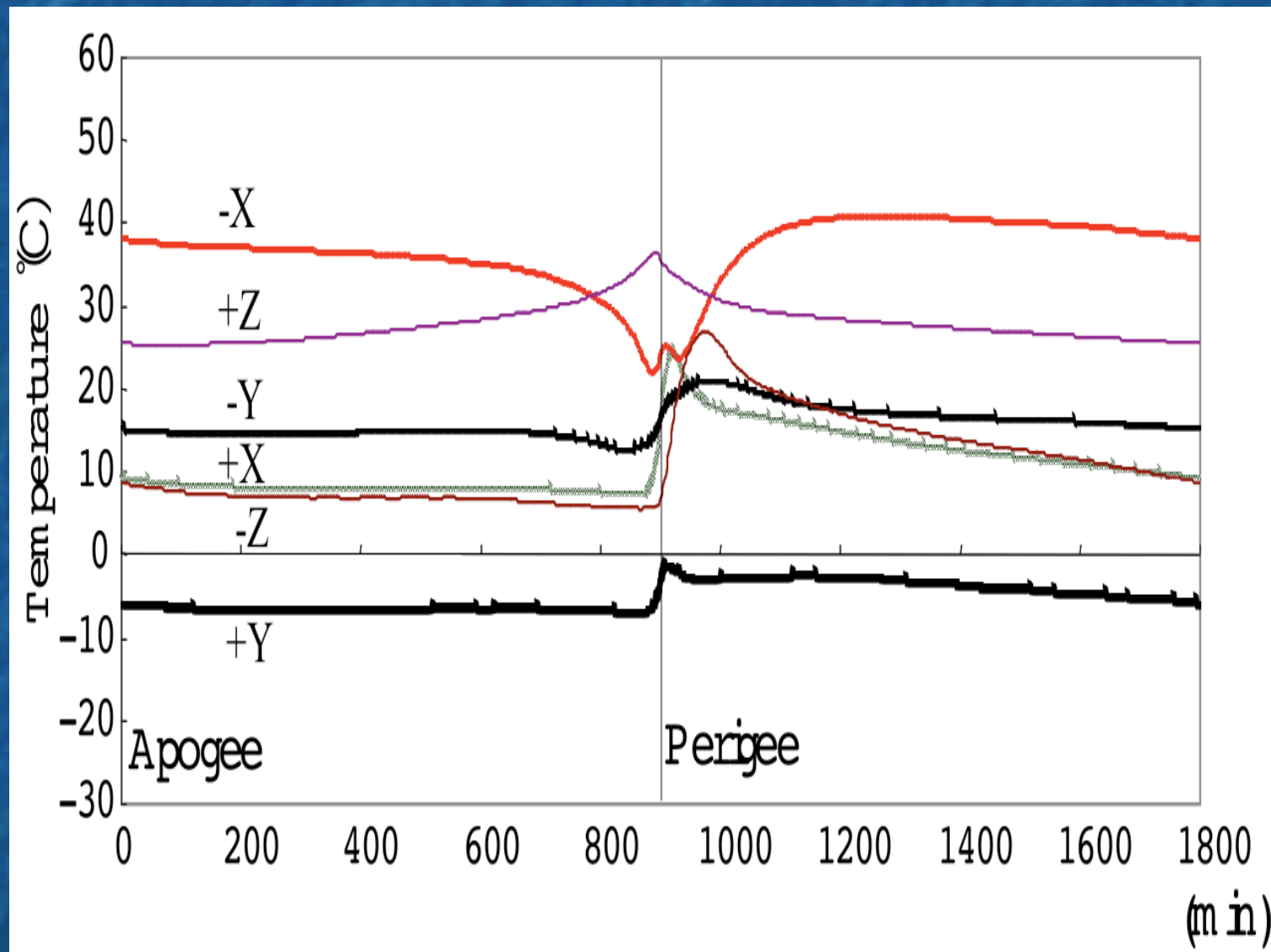
<i>Orbit</i>	<i>MODE</i>	<i>Cameras</i>	<i>Telecom</i>	<i>Attitude Control</i>	<i>DHU</i>
<i>Venus</i>	<i>Data Aquisition (HOT1)</i>	<i>ON</i>	<i>SSPA</i>	<i>RW</i>	<i>ON</i>
	<i>Telecom (HOT2)</i>	<i>OFF</i>	<i>XTWT</i>	<i>RW</i>	<i>ON</i>
	<i>Safe hold (HOT3)</i>	<i>OFF</i>	<i>SSPA</i>	<i>RCS</i>	<i>ON</i>
<i>Transfer</i>	<i>Safe hold (COLD1)</i>	<i>OFF</i>	<i>XTWT</i>	<i>RCS</i>	<i>ON</i>



PLANET-C Thermal Test Model



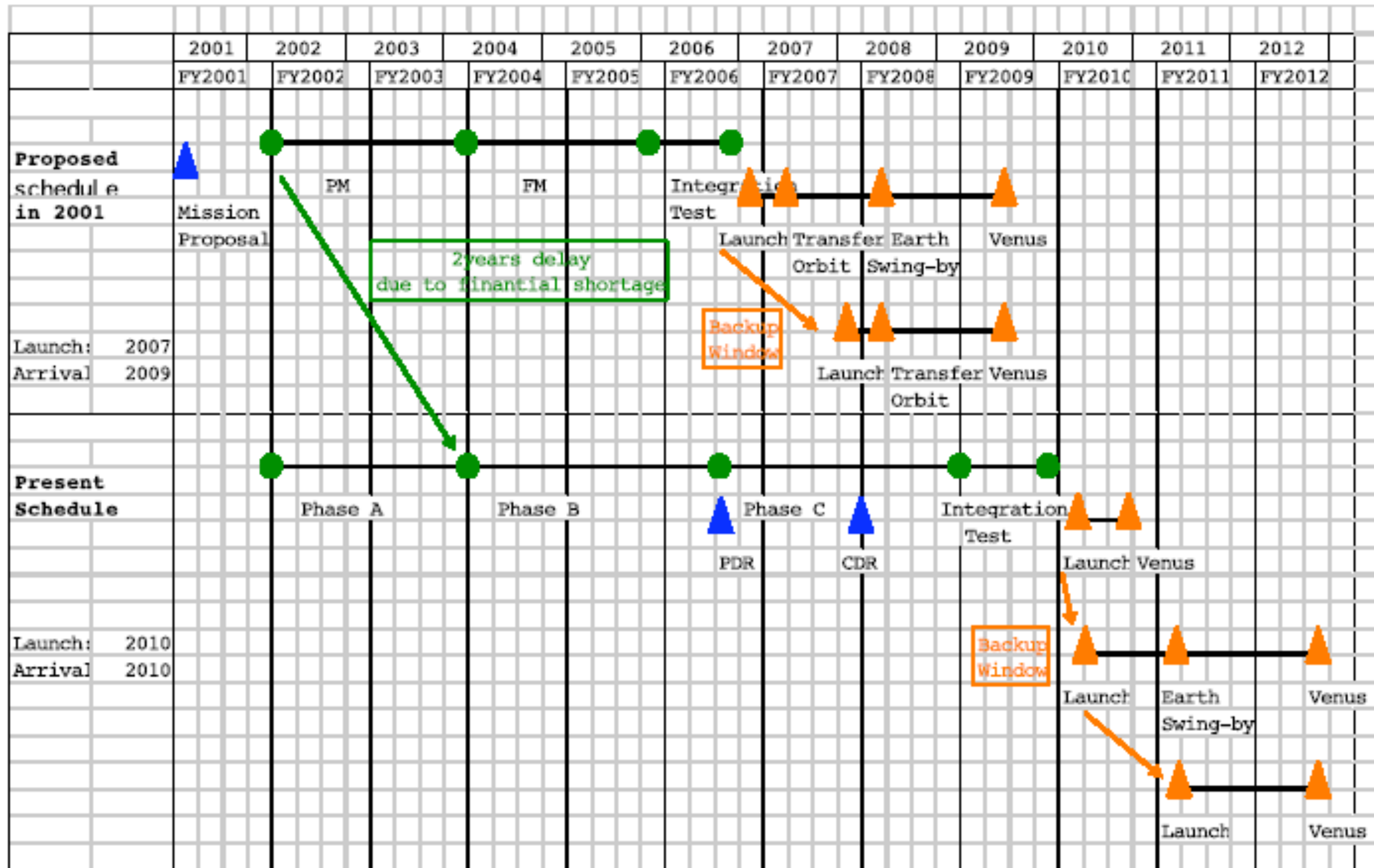
Typical temperature profile of each panel at Venus orbit



	Temperature range (°C)	
	(HOT3) ←	→ (HOT1)
+X	-11.7 -11.1	26.1 25.3
-X	-18.9 -16.2	26.9 26.9
+Y	-3.4 -3.4	28.2 28.7
-Y	-19.3 -19.1	22.5 23.3
+Z	-13.9 -12.4	27.4 29.3
-Z	-32.6 -33.7	12.3 12.1

Test result ← →
Analysis result

Schedule



VCO is looking forward
to meeting VEX in 2010
at Venus!

