



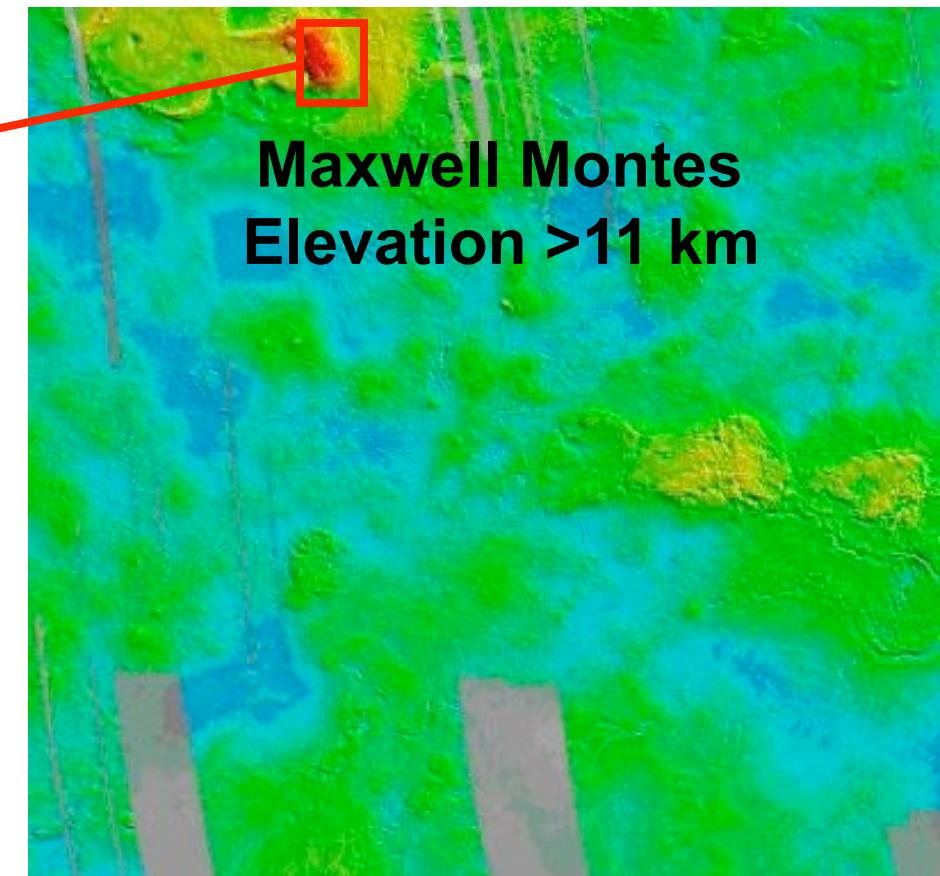
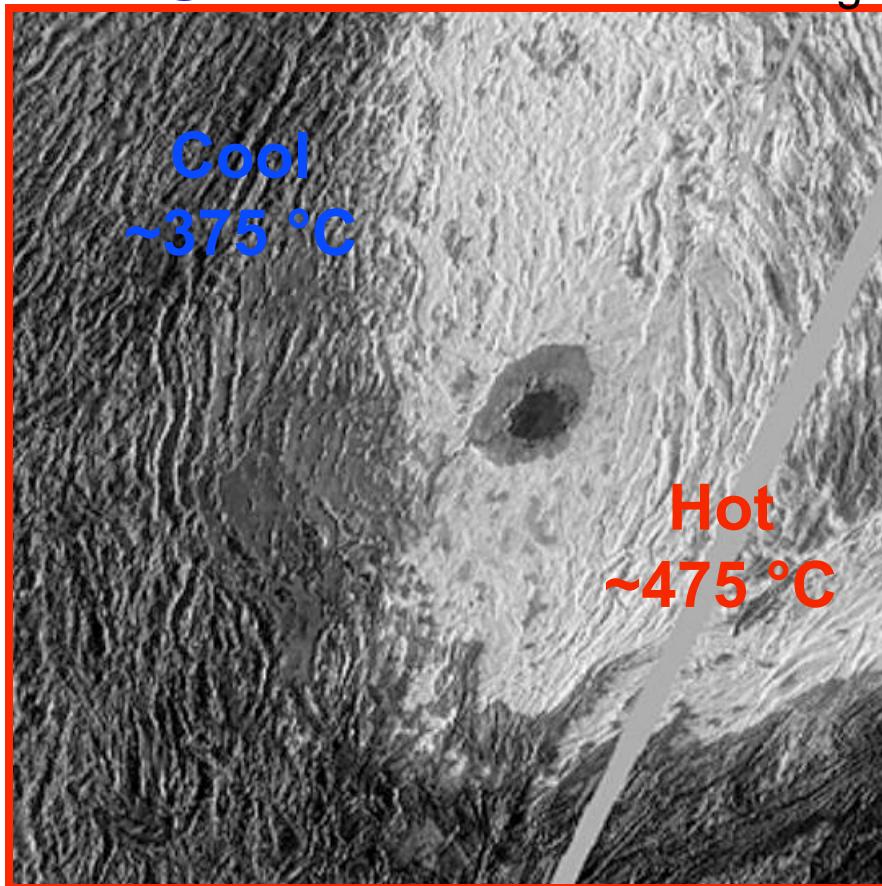
# Anomalous Radar Reflectivity of the Venus High Regions



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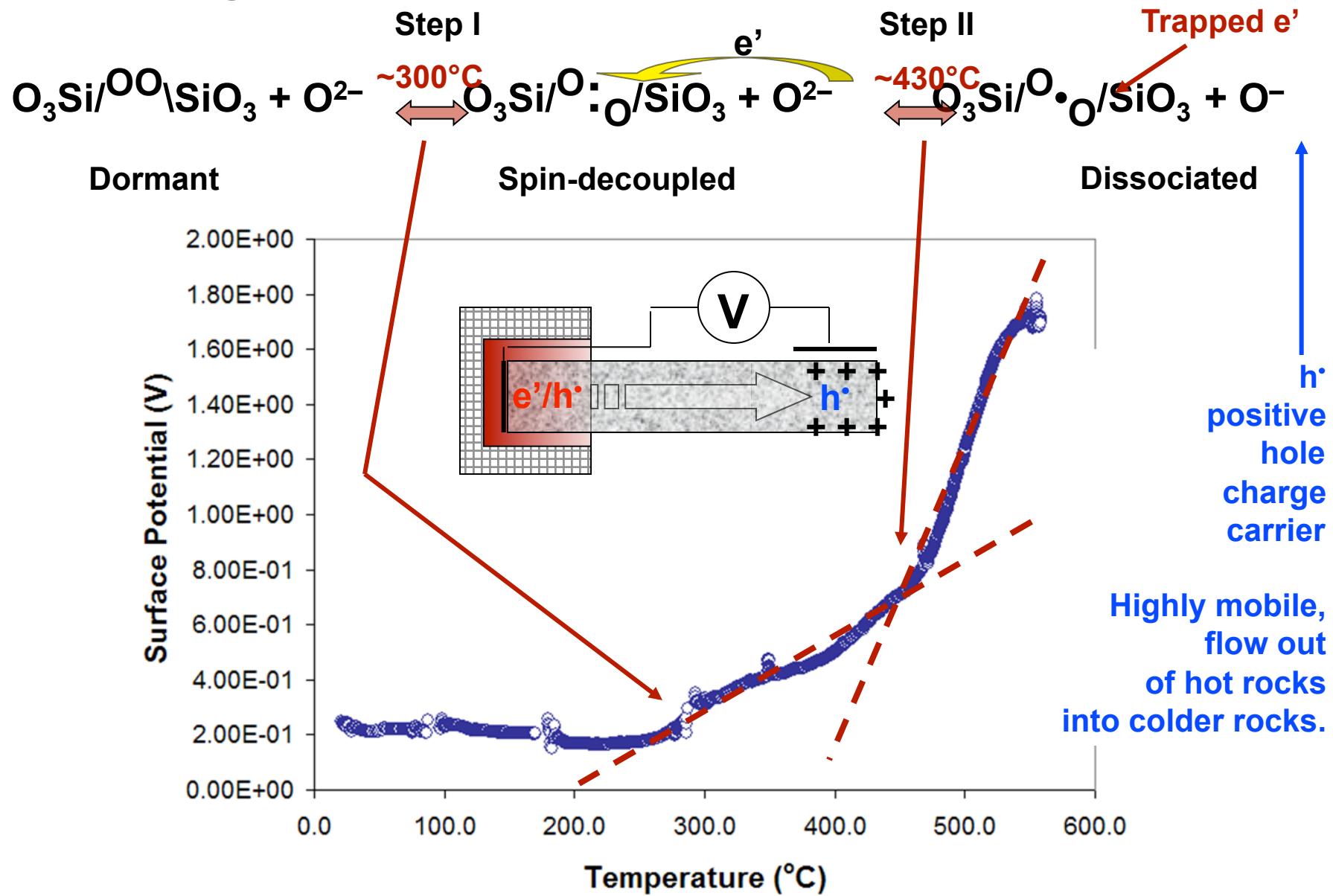
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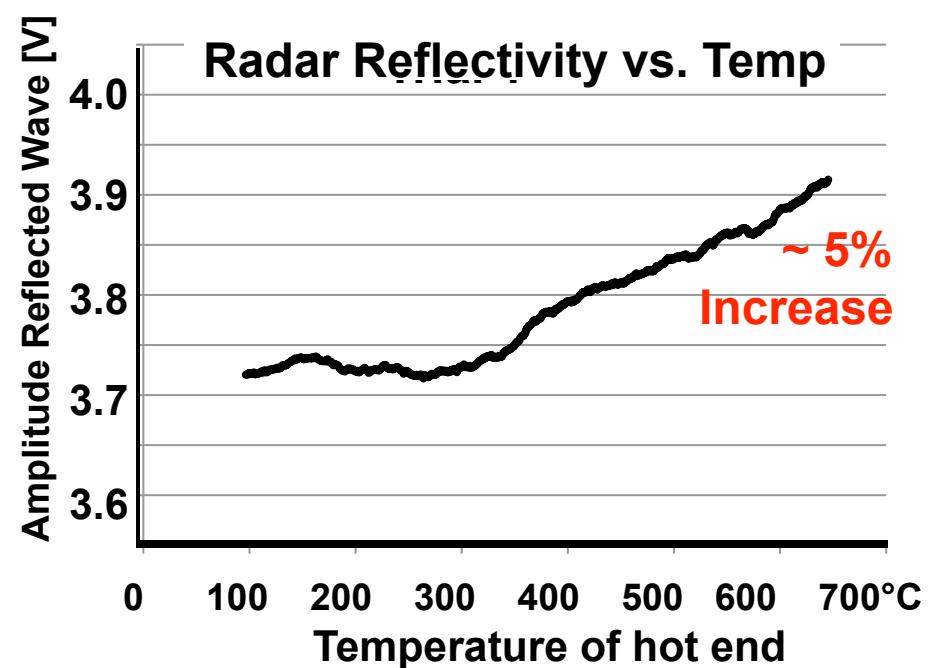
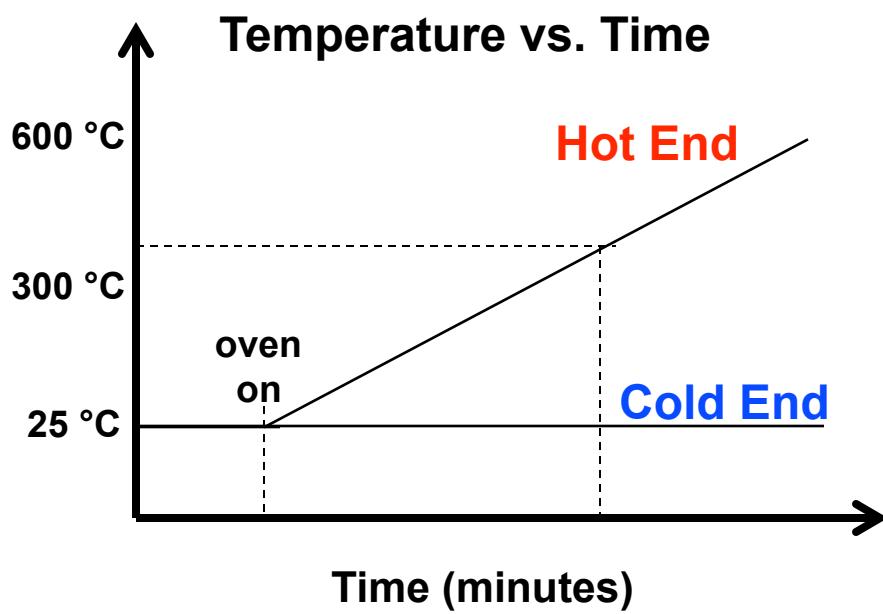
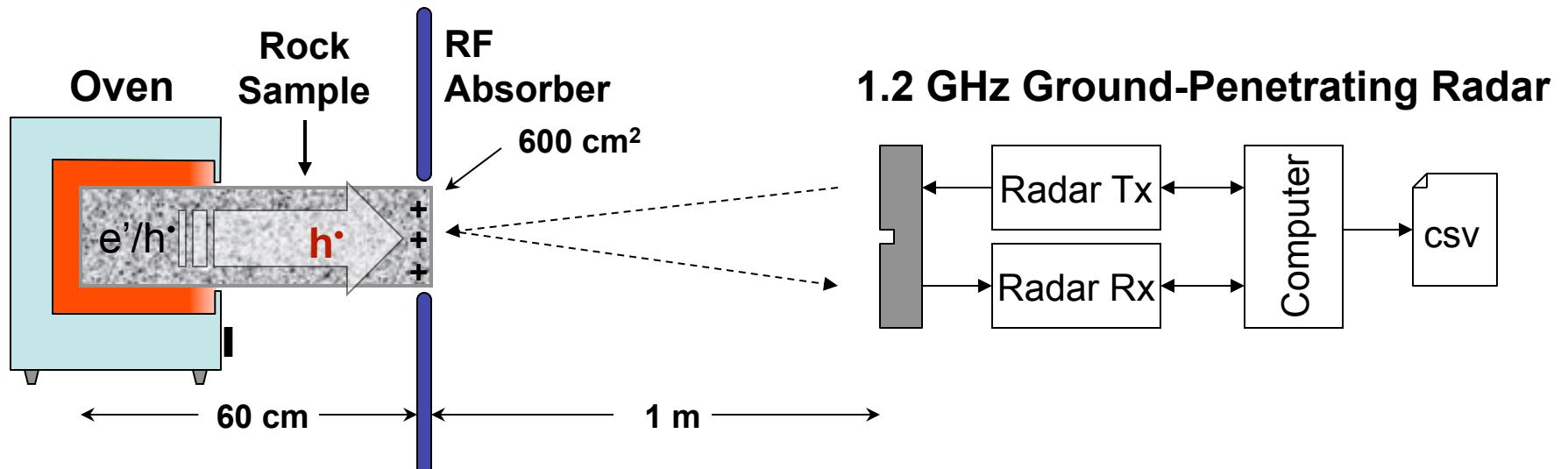


Images courtesy NASA Magellan

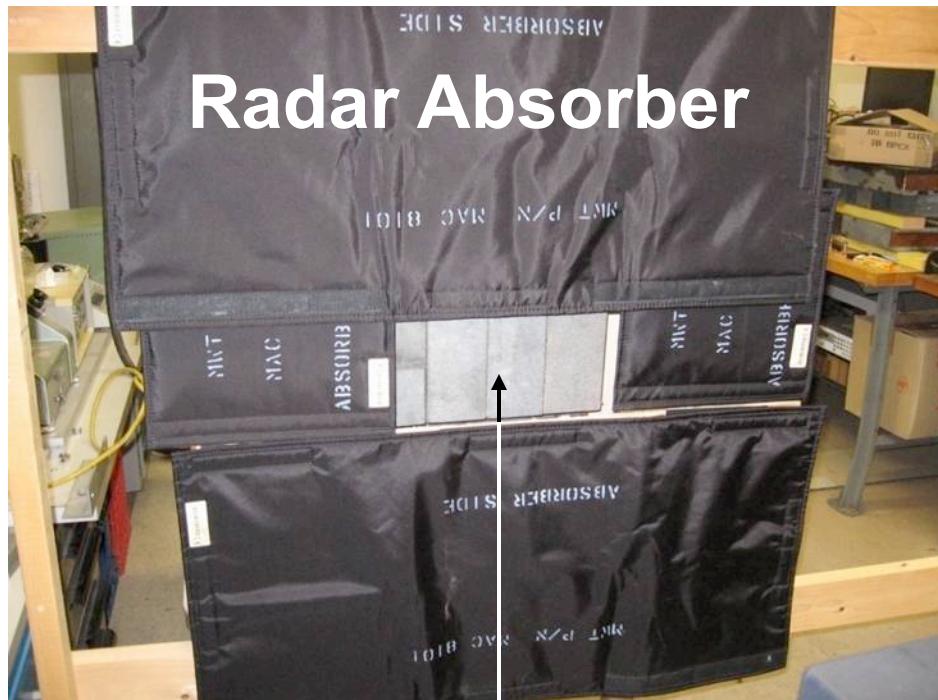
# Thermal 2-stage activation of positive hole charge carriers in rocks: Surface Potential



# Radar Reflectivity from the Cold End of Gabbro



# Experiment Photographs



Oven behind radar absorber  
not shown



Radar Controller

Transmitter  
& Receiver

# Summary and Hypothesis

- Igneous rocks contain dormant **peroxy** defects,  $\text{O}_3\text{Si}-\text{OO}-\text{SiO}_3$ .
- Temperature breaks the O<sup>-</sup>-O<sup>-</sup> bond, releasing **h<sup>•</sup>**, positive hole charge carriers in 2 steps: ~300°C and more profusely ~430°C.
- The **h<sup>•</sup>** are highly mobile, while electrons remain trapped.
- **h<sup>•</sup>** density ~ $10^{18}$  cm<sup>-3</sup>; phase velocity ~200 m/s; lifetime <10<sup>3</sup> s.
- When activated in the hot end, the **h<sup>•</sup>** flow out to the cool end.
- Due to mutual electrostatic repulsion, the **h<sup>•</sup>** repel each other.
- They diffuse to the surface forming a surface charge layer.
- On the surface the electronic states become more delocalized, leading to high electrical conductivity and high radar reflectivity.
- Anomalous high radar reflectivity of high mountains on Venus such as Maxwell Montes may be due to a surface charge layer.