

14264
Vitric-matrix Breccia
 117.8 grams



Figure 1: Photo of 14264. Sample is 5 cm. NASA S71-29217.

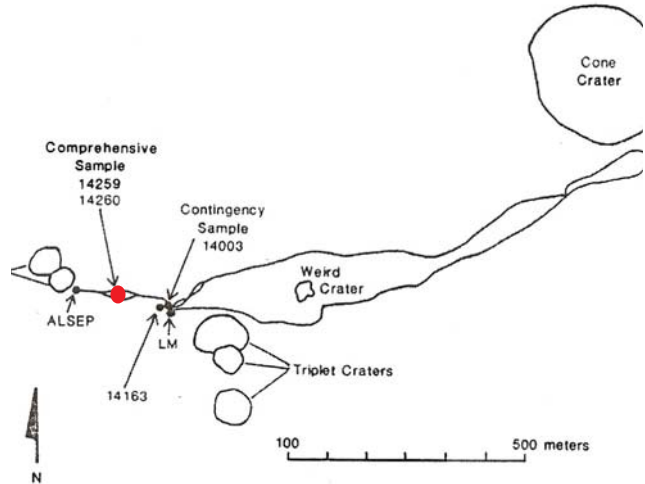


Figure 2: Location of comprehensive sample on Apollo 14 traverse map.

Introduction

14264 (figure 1) was collected as part of the “comprehensive sample” taken near the ALSEP station (figure 2). It is dark grey regolith breccia with large rounded lithic clasts. It has micrometeorite pits on all surfaces.

Petrography

Simonds et al. (1977) and Phinney et al. (1975) termed 14264 a vitric matrix breccia and reported a few percent agglutinate.

Warren and Wasson (1980) studied the large white clast – finding it non-pristine. They determined the composition of the pyroxenes, plagioclases and olivine (figure 3).

Figure 6 is a photomicrograph of one of the lithic clasts.

Processing

Sample 14264 is the largest sample from the comprehensive soil. It was returned in weigh bag 1039 along with the soil. 1039 may have been in ALSRC 1007. There are 7 thin sections.

Mineralogical Mode for 14264

	Simonds et al 1977
Matrix	83 %
Clasts	
Plagioclase	3
Mafic	
Breccia	5
Glass	4
Agglutinate	2
Granulite	3
Mare basalt	
Felds basalt	

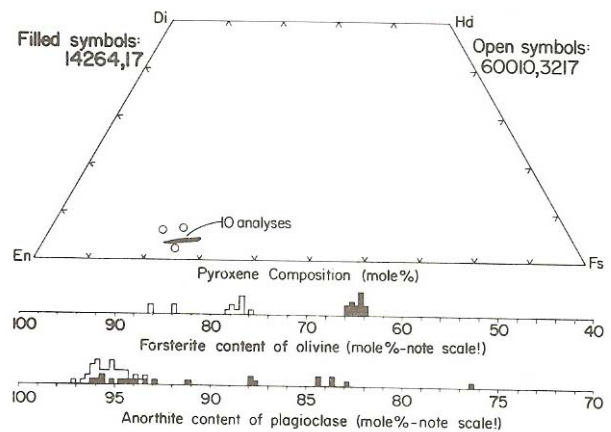


Figure 3: Pyroxene composition of 14264 and 60010 (Warren and Wasson 1980).

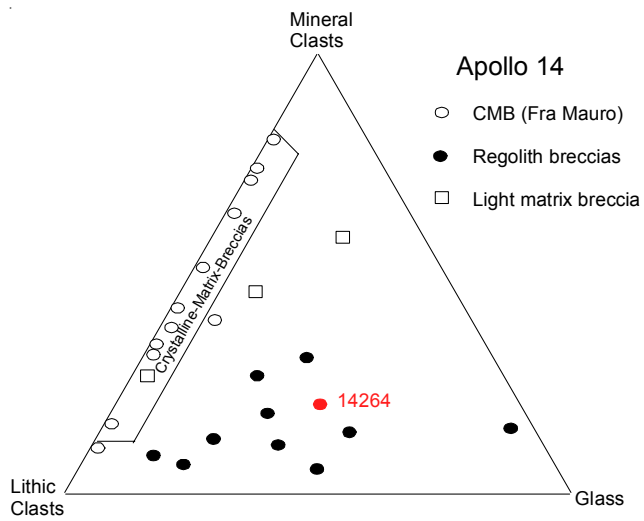


Figure 4: Simonds diagram for Apollo 14 breccias.

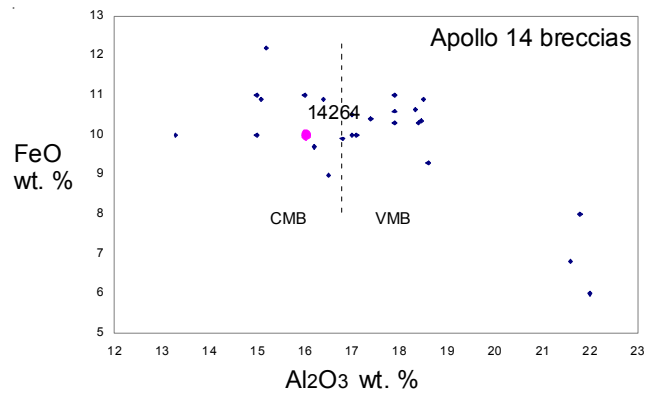
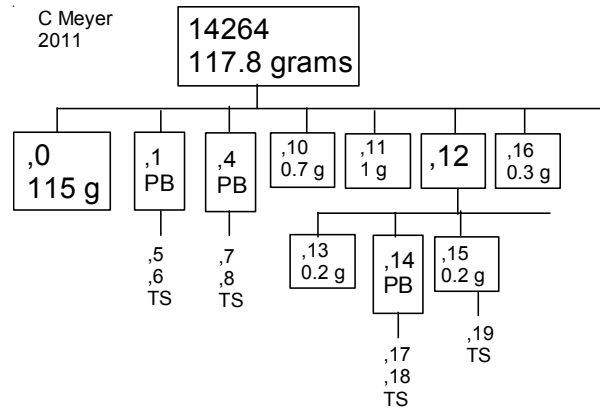


Figure 5: Composition of Apollo 14 breccias.



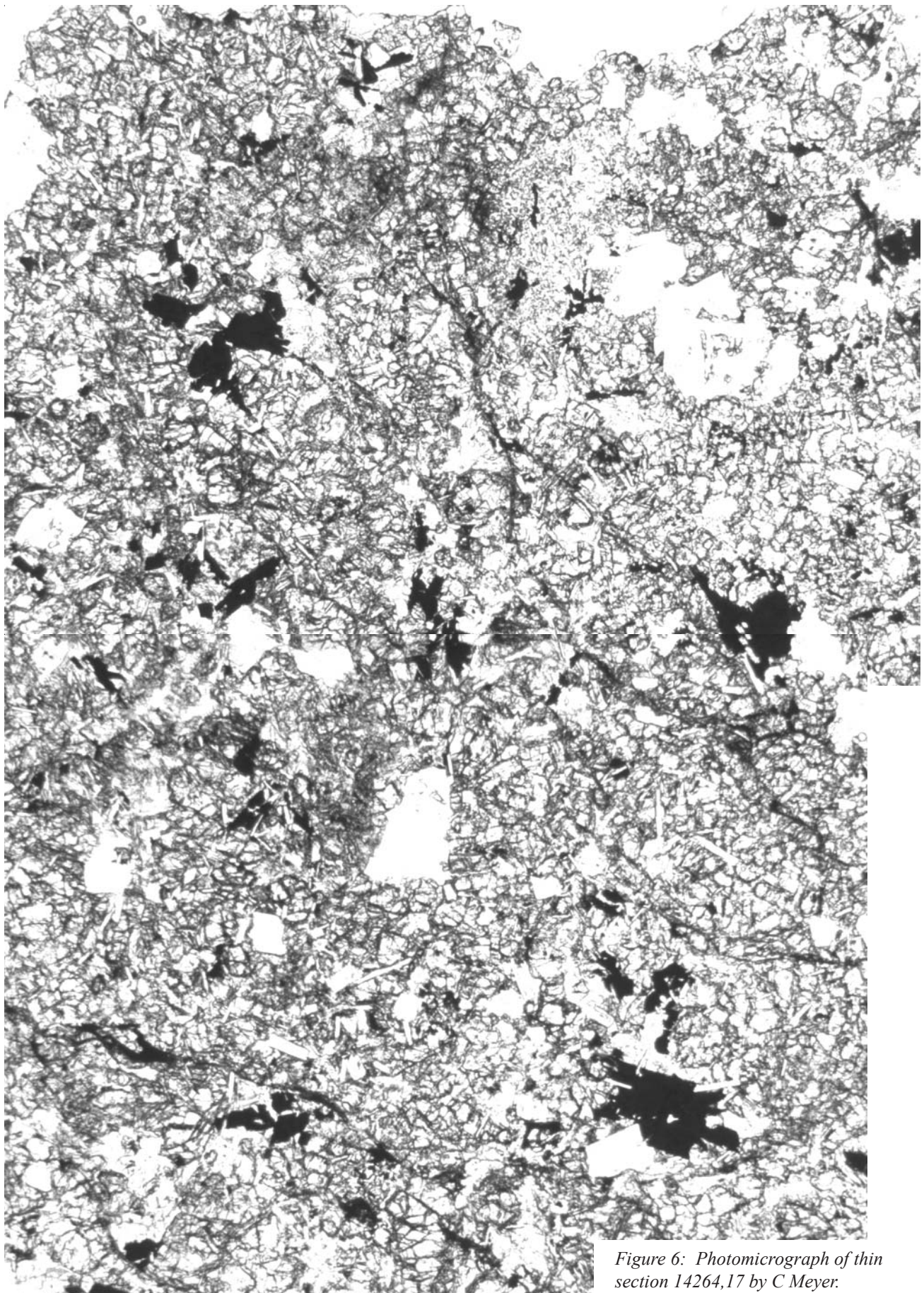


Figure 6: Photomicrograph of thin section 14264,17 by C Meyer.



Figure 7: Processing photo of 14264 showing clasts. NASA S75-24428. Large clast is 1 cm.

References for 14264

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Table 1. Chemical composition of 14264.

reference weight	Warren80		Simonds77		
	clast		matrix	clast	
SiO2 %	48.8	(a)	49.28	(c)	46.19 (b)
TiO2	2	(a)	2.05	(c)	1.8 (b)
Al2O3	15.1	(a)	16.29	(c)	11.81 (b)
FeO	10.2	(a)	10.23	(c)	16.43 (b)
MnO	0.13	(a)	0.13	(c)	
MgO	12.4	(a)	9.08	(c)	11.36 (b)
CaO	9.2	(a)	10.12	(c)	10.26 (b)
Na2O	0.9	(a)	0.93	(c)	0.37 (b)
K2O	0.95	(a)	0.82	(c)	0.61 (b)
P2O5			0.66	(c)	
S %			0.09	(c)	
sum					
Sc ppm	22.4	(a)			
V					
Cr	1300	(a)		7900	(b)
Co	24.3	(a)			
Ni	220	(a)			
Cu					
Zn	4.1	(a)			
Ga	5	(a)			
Ge ppb	290	(a)			
As					
Se					
Rb					
Sr					
Y					
Zr	1480	(a)			
Nb					
Mo					
Ru					
Rh					
Pd ppb					
Ag ppb					
Cd ppb	15	(a)			
In ppb	18	(a)			
Sn ppb					
Sb ppb					
Te ppb					
Cs ppm					
Ba	1090	(a)			
La	99	(a)			
Ce	229	(a)			
Pr					
Nd	135	(a)			
Sm	39.8	(a)			
Eu	2.2	(a)			
Gd					
Tb	8.6	(a)			
Dy					
Ho					
Er					
Tm					
Yb	30.4	(a)			
Lu	4.26	(a)			
Hf	31	(a)			
Ta	3.56	(a)			
W ppb					
Re ppb	0.44	(a)			
Os ppb					
Ir ppb	6	(a)			
Pt ppb		(a)			
Au ppb	5.3	(a)			
Th ppm	18.6	(a)			
U ppm	5.4	(a)			

technique: (a) INAA, (b) e. probe, (c) XRF

revisited, or breccias aren't so bad after all. *Proc. 8th Lunar Sci. Conf.* 1869-1893.

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