

14265
Breccia
65.79 grams



Figure 1: Top and bottom of 14265. Samples is 5 cm across. NASA S71-29151 and 29154.

Introduction

14265 is from the “comprehensive sample” taken near ALSEP station. A circle 14 meters in diameter was drawn and all large rocks were collected and placed in bag 1039 along with soil. It has a relatively high cosmic ray induced activity so it must have been at or near the surface.

Petrography

Phinney et al. (1975) found 14265 had a glass coating on one side and was pitted by micrometeorites on the other side as though the glass coating had eroded away (figure 1). It may have been a glass bomb. It is mostly matrix with only small clasts.

14265 has numerous fractures, some glass filled. The thin section shows vesicular glass bonding small fragments of a polymict breccia (Carlson and Walton 1977).

Chemistry

14265 appears to have the same composition as the soil at Apollo 14 (table 1, figure 2).

Cosmogenic isotopes and exposure ages

Eldridge et al. (1972) determined the cosmic-ray induced activity of $^{22}\text{Na} = 70$ dpm/kg and $^{26}\text{Al} = 102$ dpm/kg.

Processing

14265 is one of the comprehensive samples from weigh bag 1039 in ALRC 1007. There are two thin sections of 14265.

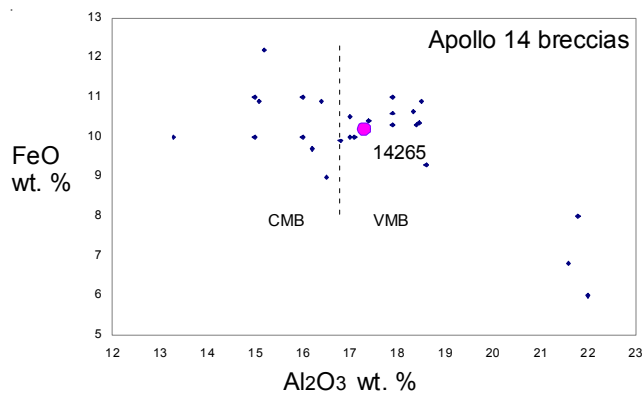
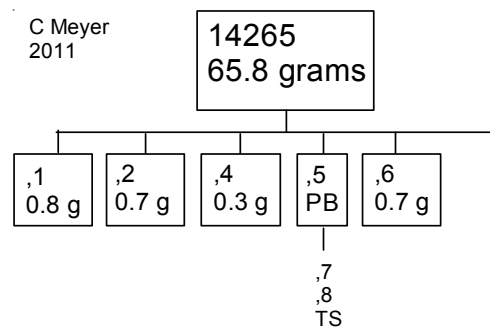


Figure 2: Composition of Apollo 14 breccias.



References for 14265

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Table 1. Chemical composition of 14265.

reference	Eldridge72	Simonds77		
weight		glass	matrix	
SiO2 %		47.53	(b) 47.36	(c)
TiO2		1.84	(b) 1.76	(c)
Al2O3		17.43	(b) 17.51	(c)
FeO		10.06	(b) 10.62	(c)
MnO			0.14	(c)
MgO		9.25	(b) 9.26	(c)
CaO		11.25	(b) 11.17	(c)
Na2O		0.6	(b) 0.68	(c)
K2O	0.49	(a) 0.46	(b) 0.47	(c)
P2O5			0.5	(c)
S %			0.09	(c)
sum				

Sc ppm

V

Cr 3210 (b)

Co

Ni

Cu

Zn

Ga

Ge ppb

As

Se

Rb

Sr

Y

Zr

Nb

Mo

Ru

Rh

Pd ppb

Ag ppb

Cd ppb

In ppb

Sn ppb

Sb ppb

Te ppb

Cs ppm

Ba

La

Ce

Pr

Nd

Sm

Eu

Gd

Tb

Dy

Ho

Er

Tm

Yb

Lu

Hf

Ta

W ppb

Re ppb

Os ppb

Ir ppb

Pt ppb

Au ppb

Th ppm 10.9 (a)

U ppm 3.3 (a)

technique: (a) radiation counting, (b) e. probe, (c) XRF