

**60095**  
Glass Sphere  
46.6 grams



Figure 1: Photo of 60095. Cube is 1 cm. NASA S72-39425.

### **Introduction**

Lunar sample 60095 was observed sitting on the lunar surface while placing the heat flow probe in the deep drill hole at Apollo 16 (Sutton 1981). It is a glass sphere that measures about 2.6 cm in diameter and is partly broken showing large vesicles inside (figure 1). Part of the surface is covered with micrometeorite craters.

### **Petrography**

Schaal et al. (1979abs), Mehta and Goldstein (1979) and Ryder and Norman (1980) describe 60095 as a broken sphere made of yellow-green to light brown glass with numerous internal vesicles. The outer surface is smooth, having formed by cooling in a vacuum, but with some cooling cracks and micrometeorite pits.

The internal glass is generally not devitrified except where there are residual inclusions of plagioclase that act as nucleation sites for devitrification. Small rounded bleb of metal with associated troilite and schreibersite are abundant, ranging in size from ~ 50 microns down to a few angstroms. Strings of metal particles are sometimes aligned in flow planes (schlieren).

### **Mineralogy**

**Glass:** The sphere is almost entirely homogeneous glass with only minor schlieren.

**Metallic Iron:** Two types of metal particles were found in 60095 glass (Mehta and Goldstein 1979). Finely divided metal/sulfide particles as small as 150 angstroms were observed in ion-thinned foils of 60095 glass. They range in Ni content 17 to 57%. Tiny cubes of Fe metal about <500 microns were found to be Ni free.

### **Chemistry**

See et al. (1986) and Morris et al. (1986) determined the composition of 60095 (table 1). Ganapathy et al. (1974) determined trace elements. It is similar in composition to other glass thought to be ejecta from South Ray Crater and similar to bulk soil composition (figure 2).

### **Cosmogenic isotopes and exposure ages**

Using a depth profiling technique Warhaut et al. (1979) found high  $^{22}\text{Ne}$  <50 angstroms below the surface, but little or no  $^4\text{He}$ .

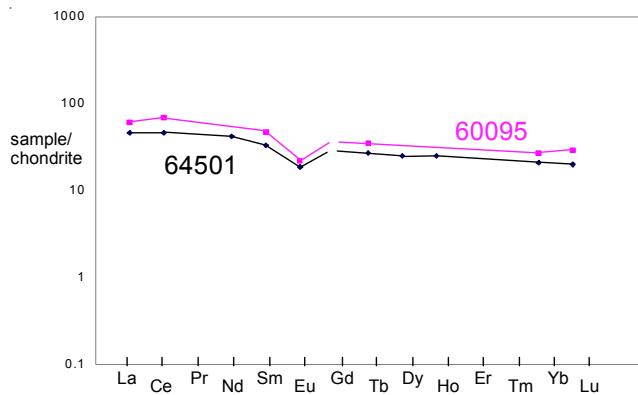


Figure 2: Normalized rare-earth-element pattern for 60095 glass compared with that of Apollo 16 soil (64501). Data from Morris et al. (1983) and Papike et al. (soil).

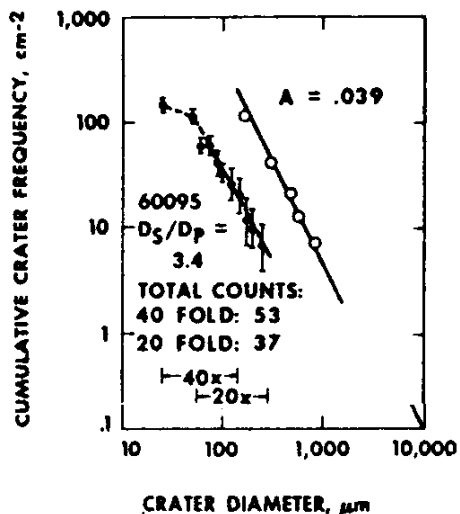


Figure 3: Crater count vs. size for 60095 (Neukum et al. 1973).

**Other Studies**

Neukum et al. (1973) and Brownlee et al. (1975) studied the micrometeorite craters that populate the top surface of 60095 (figures 3 and 4).

**Processing**

Sample 60095 was cut with a wire saw (figures 5 and 6). Various pieces broke off because of the pressure of the vise during sawing. These pieces were where the greatest density of micrometeorite pits were located and studied.

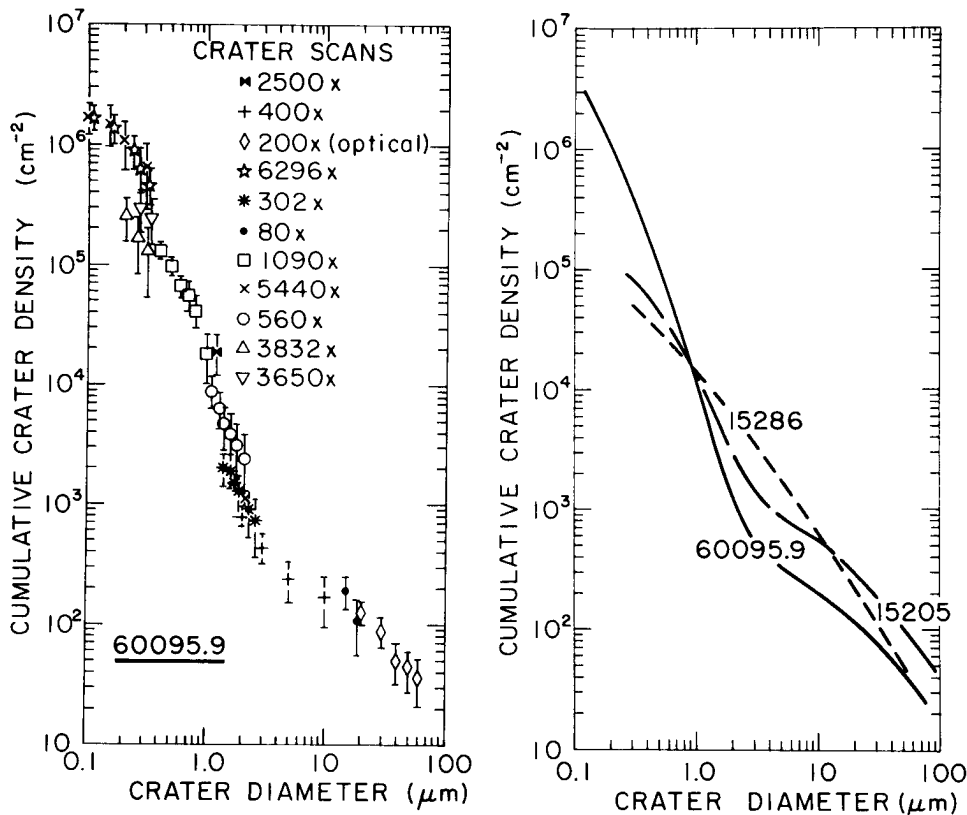


Figure 4: Crater frequency plot for 60095 compared with other lunar samples (Brownlee et al. 1975).

**Table 1. Chemical composition of 60095.**

reference	Ganapathy74	See 86	Morris 86	Uhlmann74
<i>weight</i>		glass	glass	
SiO <sub>2</sub> %		44.82	(a) 44.82	(a) 46.4 (d)
TiO <sub>2</sub>		0.48	(a) 0.48	(a)
Al <sub>2</sub> O <sub>3</sub>		25.94	(a) 25.94	(a) 23.5 (d)
FeO		5.72	(a) 5.53	(b) 6.9 (d)
MnO				
MgO		7.89	(a) 7.89	(a) 10.5 (d)
CaO		14.61	(a) 14.61	(a) 12.1 (d)
Na <sub>2</sub> O		0.32	(a) 0.32	
K <sub>2</sub> O		0.12	(a) 0.31	(b) 0.8 (d)
P <sub>2</sub> O <sub>5</sub>				
S %				
<i>sum</i>				
Sc ppm			7.08	(b)
V				
Cr			780	(b)
Co			40	(b)
Ni	560	(c)	412	(b)
Cu				
Zn	1.55	(c)		
Ga				
Ge ppb	306	(c)		
As				
Se				
Rb	1.67	(c)		
Sr				
Y				
Zr				
Nb				
Mo				
Ru				
Rh				
Pd ppb				
Ag ppb	1.2	(c)		
Cd ppb	1.8	(c)		
In ppb				
Sn ppb				
Sb ppb	2.62	(c)		
Te ppb	26	(c)		
Cs ppm	0.064	(c)		
Ba			200	(b)
La			14.4	(b)
Ce			41.7	(b)
Pr				
Nd				
Sm			6.98	(b)
Eu			1.22	(b)
Gd				
Tb			1.27	(b)
Dy				
Ho				
Er				
Tm				
Yb			4.27	(b)
Lu			0.71	(b)
Hf			4.9	(b)
Ta			0.68	(b)
W ppb				
Re ppb	2.17	(c)		
Os ppb				
Ir ppb	25.4	(c)		
Pt ppb				
Au ppb	7.11	(c)		
Th ppm			2.6	(b)
U ppm	0.67	(c)	0.99	(b)
<i>technique:</i>	(a) emp, (b) INAA, (c) RNAA, (d) ???			

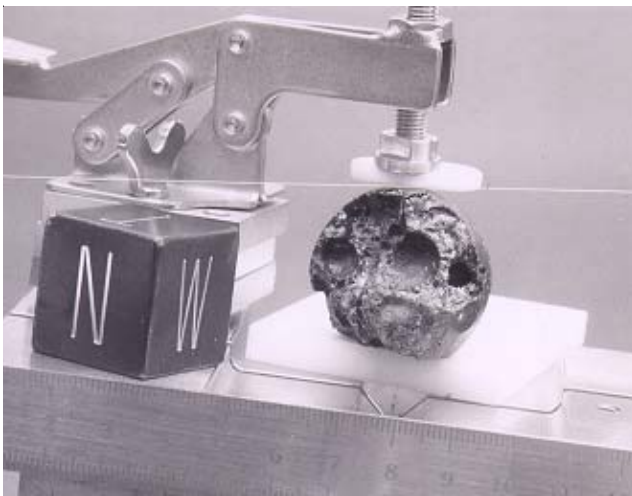
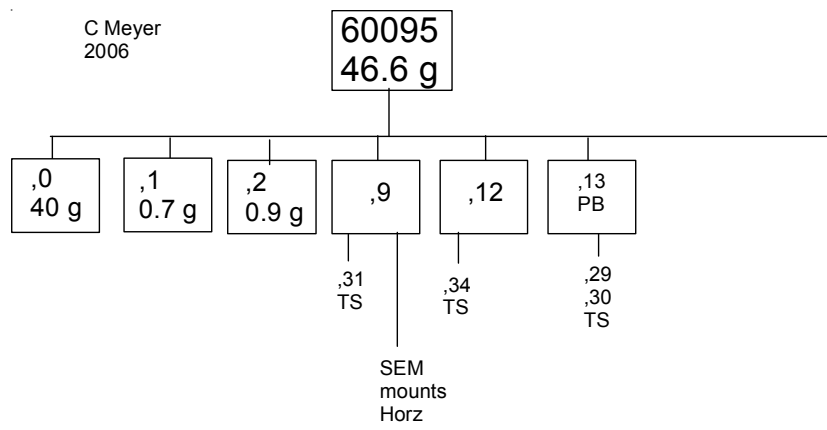


Figure 5: Wiresaw ready to cut 60095.

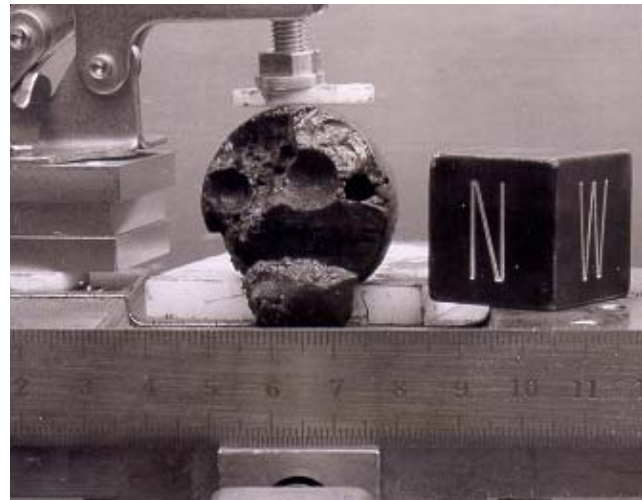


Figure 6: After wiresaw cut.



Figure 7: Numbering the splits of 60095. Splits ,1 and ,2 were from the large piece in figure 6. Other pieces broke off due to pressure of vise.

## References for 60095

- Blanford G.E., Fruland R.M., McKay D.S. and Morrison D.A. (1974a) Lunar surface phenomena: Solar flare track gradients, microcraters, and accretionary particles. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2501-2526.
- Brownlee D.E., Horz F., Hartung J.B. and Gault D.E. (1975) Density, chemistry and size distribution of interplanetary dust. *Proc. 6<sup>th</sup> Lunar Sci. Conf.* 3409-3416.
- Butler P. (1972a) Lunar Sample Information Catalog Apollo 16. Lunar Receiving Laboratory. MSC 03210 Curator's Catalog. pp. 370.
- Drozd R.J., Hohenberg C.M., Morgan C.J. and Ralston C.E. (1974) Cosmic-ray exposure history at the Apollo 16 and other lunar sites: lunar surface dynamics. *Geochim. Cosmochim. Acta* **38**, 1625-1642.
- Ganapathy R., Morgan J.W., Krahenbuhl U. and Anders E. (1973) Ancient meteoritic components in lunar highland rocks: Clues from trace elements in Apollo 15 and 16 samples. *Proc. 4<sup>th</sup> Lunar Sci. Conf.* 1239-1261.
- Ganapathy R., Morgan J.W., Higuchi H., Anders E. and Anderson A.T. (1974) Meteoritic and volatile elements in Apollo 16 rocks and in separated phases from 14306. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1659-1683.
- Hopper R.W., Onorato P. and Uhlmann D.R. (1974) Thermal histories and crystal distribution in partly devitrified lunar glasses cooled by radiation. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2257-2273
- Hunter R.H. and Taylor L.A. (1981) Rust and schreibersite in Apollo 16 highland rocks: Manifestations of volatile-element mobility. *Proc. 12<sup>th</sup> Lunar Planet. Sci. Conf.* 253-259.
- Klein L.C. and Uhlmann D.R. (1976) The kinetics of lunar glass formation, revisited. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 1113-1121.
- LSPET (1973b) The Apollo 16 lunar samples: Petrographic and chemical description. *Science* **179**, 23-34.
- LSPET (1972c) Preliminary examination of lunar samples. In Apollo 16 Preliminary Science Report. NASA SP-315, 7-1—7-58.
- Mehta S. and Goldstein J.I. (1979) Analytical electron microscopy study of submicroscopic metal particles in glassy constituents of lunar breccias 15015 and 60095. *Proc. 10<sup>th</sup> Lunar Planet. Sci. Conf.* 1507-1521.
- Morris R.V., See T.H. and Horz F. (1986) Composition of the Cayley Formation at Apollo 16 as inferred from impact melt splashes. *Proc. 17<sup>th</sup> Lunar Planet. Sci. Conf.* in J. Geophys. Res. **90**, E21-E42.
- Neukum G., Horz F., Morrison D.A. and Hartung J.B. (1973) Crater populations on lunar rocks. *Proc. 4<sup>th</sup> Lunar Sci. Conf.* 3255-3276.
- Ryder G. and Norman M.D. (1980) Catalog of Apollo 16 rocks (3 vol.). Curator's Office pub. #52, JSC #16904
- Schaal R.B., Thompson T.D., Hörz F. and Bauer J.F. (1979b) Experimentally shocked lunar basalt: Massive and particulate (abs). *Lunar Planet. Sci. X*, 1055-1057. Lunar Planetary Institute, Houston.
- See T.H., Horz F. and Morris R.V. (1986) Apollo 16 impact-melt splashes: Petrography and major-element composition. *Proc. 17<sup>th</sup> Lunar Planet. Sci. Conf.* in J. Geophys. Res. **91**, E3-E20.
- Sutton R.L. (1981) Documentation of Apollo 16 samples. In Geology of the Apollo 16 area, central lunar highlands. (Ulrich et al. ) U.S.G.S. Prof. Paper 1048.
- Uhlmann D.R., Klein L., Kritchevsky G. and Hopper R.W. (1974) The formation of lunar glasses. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2317-2331.