

61140, 61160 and 61180

Soils

229, 153 and 271 grams

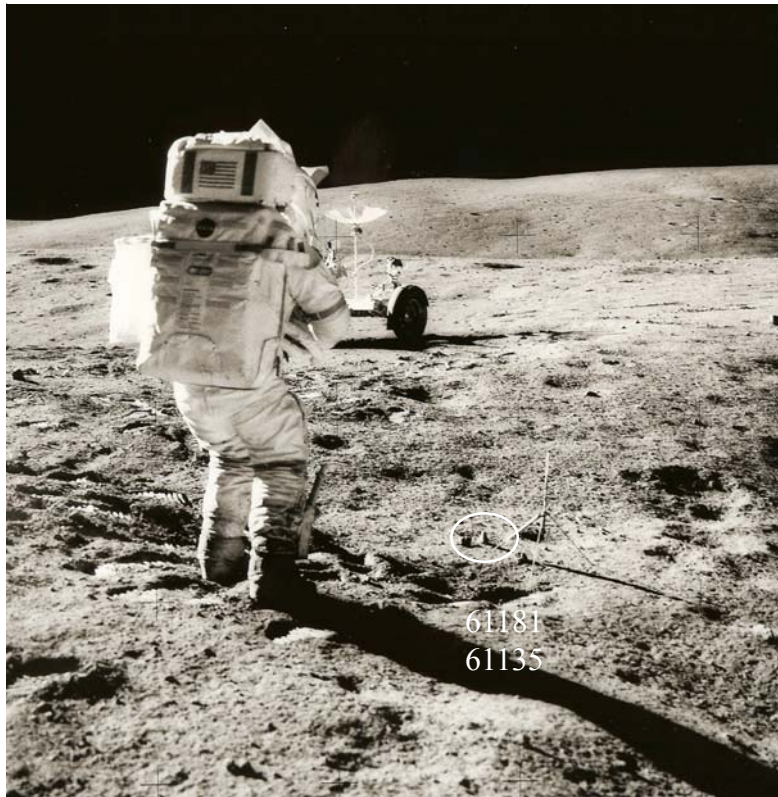


Figure 1: Apollo 16 commander taking picture of 61135 before collecting 61181. AS16-109-17800.

Introduction

61140, 61160 and 61180 were part of radial sample collected at Plum Crater, Apollo 16. 61180 was collected from beneath rock sample 61135 on the rim of Plum Crater (Sutton 1981). Plum Crater is on the rim of a large crater (Flag Crater, 200 m across, ~ 40 m deep) that, according to pre-mission planning, should

have penetrated the regolith and sampled the underlying Cayley Formation (Head 1974). 61140 was collected about 30 m from the crater rim. However, these samples do not appear to have been studied together as a suite; different samples were allocated to different investigators. Perhaps this is because there are several

Modal content of soils 61161 (90-150 micron).

	Heiken et al. 1973	Houck 1982
Agglutinates	37 %	35.4
Basalt	0.3	0.3
Breccia	33.3	27.8
Anorthosite	4.7	1.4
Norite	-	
Gabbro	0.9	
Plagioclase	14.7	21.7
Pyroxene	2.6	2.5
Olivine	-	
Ilmenite	-	
Glass other	4.1	7.2

Modal content of soils 61181 (90-150 micron).

	Heiken et al. 1973	Houck 1982
Agglutinates	59.6 %	56
Basalt	0.6	2.9
Breccia	25.9	27.8
Anorthosite	4.3	0.7
Norite	-	
Gabbro	0.9	
Plagioclase	6.9	10.5
Pyroxene	0.9	0.3
Olivine	-	
Ilmenite	-	
Glass other	0.6	4

craters in the area and a radial sample didn't quite make sense here.

Petrography

61161 and 61181 are surface soils with high maturity ($I_s/FeO = 82$) and an abundance of agglutinates. For some reason, 61141 apparently has lower maturity ($I_s/FeO = 56$) (Morris 1978). The mineralogical mode was determined by Houck (1982) and Heiken et al. (1973). Butler et al. (1973) reported the grain size distribution for 61141; McKay et al. (1973) reported 61161 and 61181. The average grain size is 110, 88 and 91 micron, respectively (figure 3 a,b,c).

Keller and McKay (1992) studied vapor deposits on grains in 61181 as part of their investigation of "space weathering". They found that the ultra-fine (<20 micron) fraction contained numerous tiny glass spheres with either refractory composition (HASP) or volatile-rich (VRAP). HASP stands for high alumina silica poor; VRAP stands for volatile-rich alumina-poor. They are apparently complementary material derived by high-temperature volatilization and re-condensation

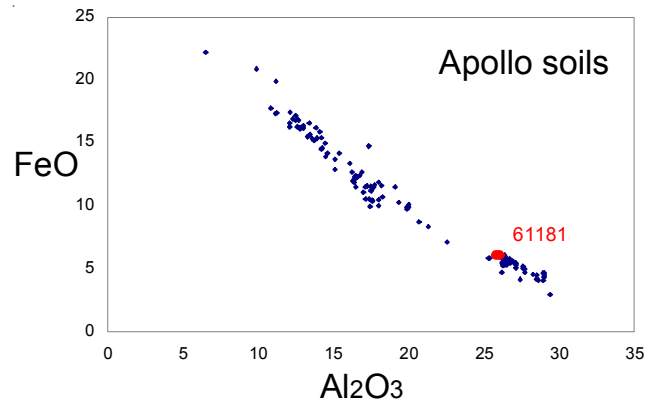
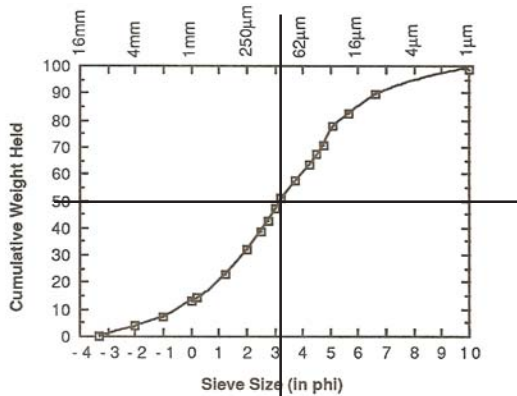


Figure 2: Chemical composition of lunar soils including 61181 (see table).

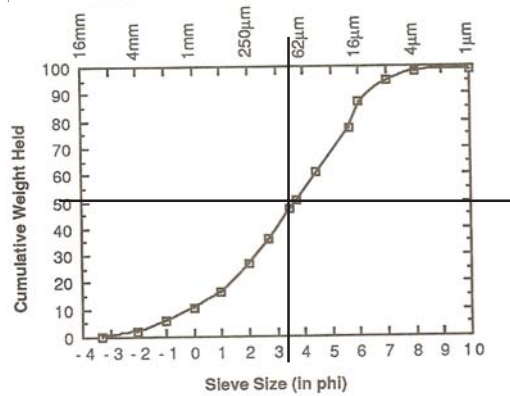
during impact into feldspathic highland rocks (figure 6).

Chemistry

Eldridge et al. (1973) and Wrigley (1973) reported the U, Th and K content of bulk samples, in general agreement with trace element analysis of sample splits (Rose et al. 1973, Wanke et al. 1973, Taylor et al. 1973 and Korotev 1982).



average grain size = 110 microns



average grain size = 88 microns

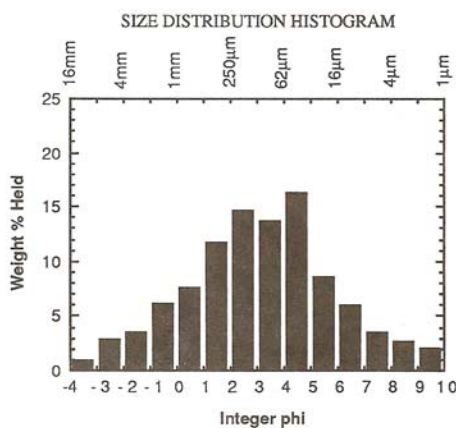


Figure 3a: Grain size distribution for 61141 (Graf, from data by Butler).

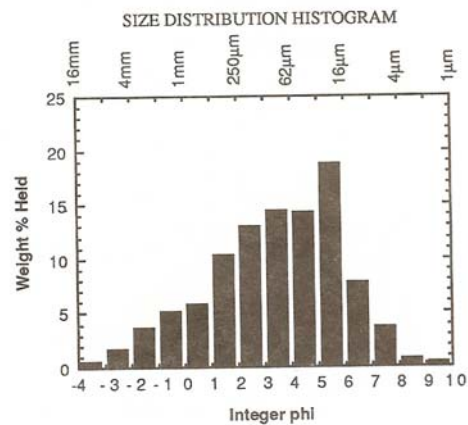


Figure 3b: Grain size distribution of 61161 (Graf 1991, from data by McKay et al.).

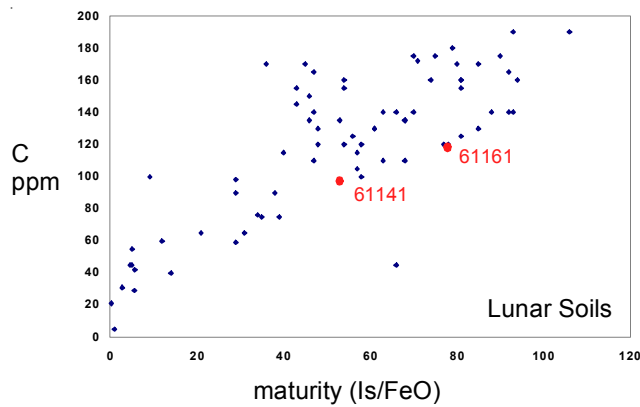
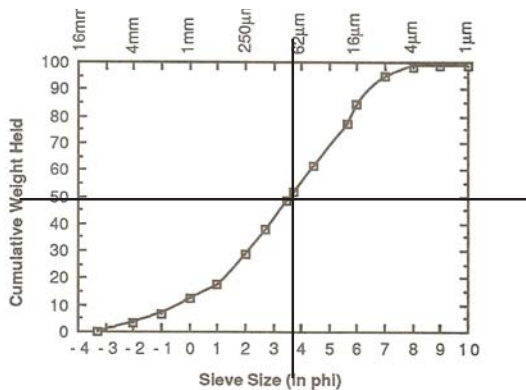


Figure 4: Carbon content and maturity index for 61141 and 61161 (Kerridge et al. 1978, Morris 1976).

Kerridge et al. (1978) and Becker (1980) reported 98, 111 and 116 ppm C and 59 - 73 ppm N for 61141 and 61161 (figure 4). C and N are closely correlated, indicating that they are both from solar wind implantation. Kerridge et al. (1975) reported 330 ppm S in 61141 and 470 ppm S in 61161. Cirlin and Housley (1981) reported excess, surface-correlated, Cd and Pb in 61161.



average grain size = 91 microns

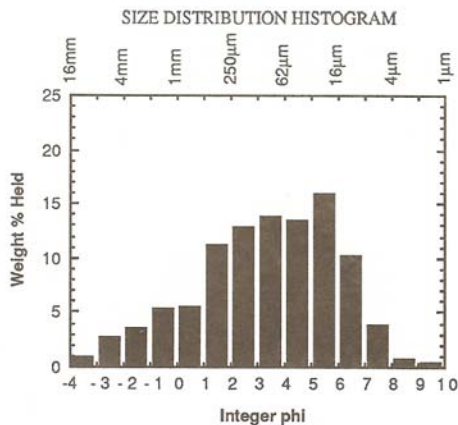


Figure 3c: Grain size distribution of 61181 (Graf 1991, data by McKay)

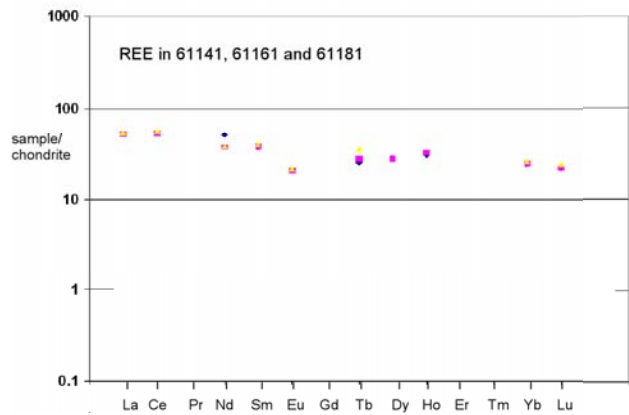


Figure 5: Normalized rare-earth-element diagram for 61141, 61161 and 61181 (data from tables).

Cosmogenic isotopes and exposure ages

The cosmic-ray-induced activity of 61181 is $^{22}\text{Na} = 40$ dpm/kg and $^{26}\text{Al} = 204$ dpm/kg (Eldridge et al. 1973). The activity of 61161 is $^{22}\text{Na} = 65$ dpm/kg and $^{26}\text{Al} = 202$ dpm/kg (Wrigley 1973).

Behrmann et al. (1973) and Bhattacharya et al. (1975) studied the cosmic-ray tracks in grains from 61181.

Other Studies

The rare gas content of 61181 was reported by Kirsten et al. (1973) and Walton et al. (1973). It has high Ne content.

Several investigators studied the magnetic properties (Tsay et al. 1973, Weeks 1973, Griscom et al. 1973).

Isotopic ratios for N and C were determined by Kerridge et al. (1978) and Becker (1980).

Processing

Marvin (1972) documented the coarse fine particles.

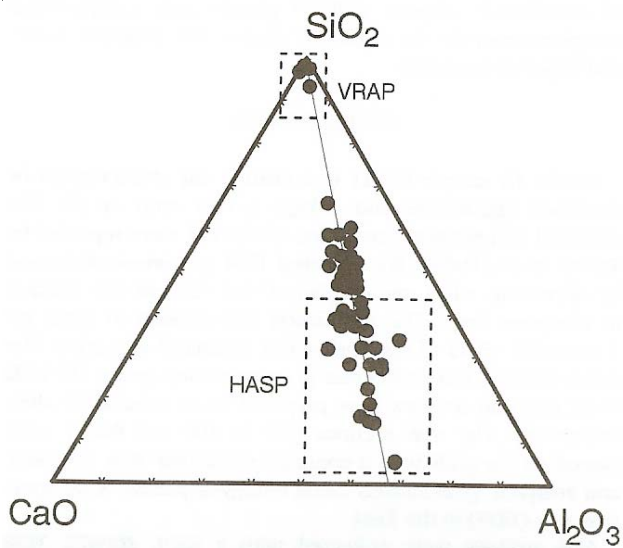


Figure 6: Chemical composition of glass condensates (VRAP) and refractory residual (HASP) in 61181 (Keller and McKay 1992).

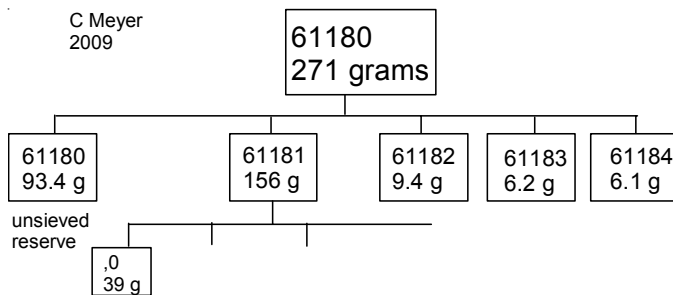
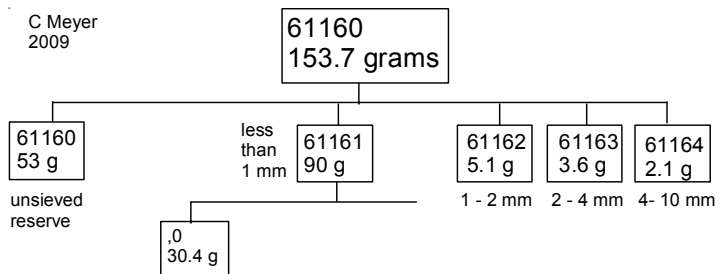
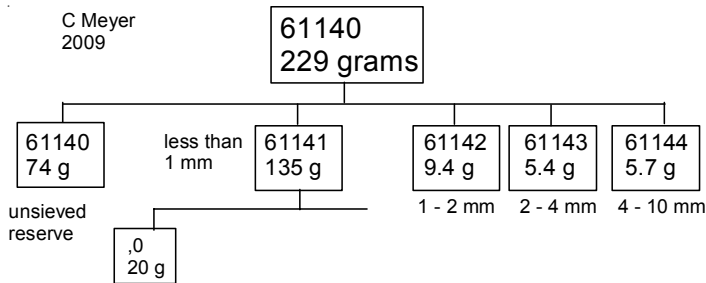


Table 1. Chemical composition of 61141.

reference weight	Rose73	Wanke73	Evenson73	
SiO2 %	45.2	(a) 44.9	(b)	
TiO2	0.58	(a) 0.53	(b)	
Al2O3	26.4	(a) 26.5	(b)	
FeO	5.29	(a) 5.33	(b)	
MnO	0.07	(a) 0.06	(b)	
MgO	6.1	(a) 6.2	(b)	
CaO	15.32	(a) 15.1	(b)	
Na2O	0.52	(a) 0.45	(b)	
K2O	0.14	(a) 0.11	(b) 0.11	(c)
P2O5	0.12	(a)		
S %				
sum				
Sc ppm	9.5	(a) 8.4	(b)	
V	20	(a)		
Cr	890	(a) 660	(b)	
Co	19	(a) 24.4	(b)	
Ni	310	(a) 400	(b)	
Cu	11	(a)		
Zn	28	(a)		
Ga	4.3	(a)		
Ge ppb				
As				
Se				
Rb	2.8	(a)	2.55	(c)
Sr	165	(a) 140	(b) 170	(c)
Y	39	(a) 38	(b)	
Zr	126	(a) 167	(b)	
Nb		11	(b)	
Mo				
Ru				
Rh				
Pd ppb				
Ag ppb				
Cd ppb				
In ppb				
Sn ppb				
Sb ppb				
Te ppb				
Cs ppm				
Ba		110	(b) 122	(c)
La		12.7	(b)	
Ce		33	(b)	
Pr				
Nd		23	(b)	
Sm		5.4	(b)	
Eu		1.17	(b)	
Gd				
Tb		0.92	(b)	
Dy		6.9	(b)	
Ho		1.7	(b)	
Er				
Tm				
Yb	2.6	(a) 3.9	(b)	
Lu		0.53	(b)	
Hf		4	(b)	
Ta		0.42	(b)	
W ppb				
Re ppb				
Os ppb				
Ir ppb		12	(b)	
Pt ppb				
Au ppb		10	(b)	
Th ppm		1.7	(b)	
U ppm				

technique: (a) "microchemical"; (b) INAA, (c) IDMS

Table 2. Chemical composition of 61161.

reference weight	Eldridge73	Wanke73	Wrigley73	Korotev81 C	F	
SiO2 %		44.7	(a)			
TiO2		0.58	(a)			
Al2O3		26.3	(a)			
FeO		5.65	(a)	5.23	5.44	(a)
MnO		0.065	(a)			
MgO		6.35	(a)			
CaO		15.9	(a)			
Na2O		0.47	(a)	0.499	0.515	(a)
K2O	0.1	(b) 0.1	(a) 0.11	(b)		
P2O5		0.12	(a)			
S %						
sum						
Sc ppm		8.7	(a)	9.4	9.4	(a)
V						
Cr		670	(a)	785	905	(a)
Co		29	(a)	22.2	23	(a)
Ni		400	(a)	305	390	(a)
Cu						
Zn						
Ga						
Ge ppb						
As						
Se						
Rb						
Sr		150	(a)			
Y		37	(a)			
Zr		184	(a)			
Nb		11.5	(a)			
Mo						
Ru						
Rh						
Pd ppb						
Ag ppb						
Cd ppb						
In ppb						
Sn ppb						
Sb ppb						
Te ppb						
Cs ppm						
Ba		120	(a)			
La		12.2	(a)	11.9	12.7	(a)
Ce		32	(a)	32	34	(a)
Pr						
Nd		17	(a)			
Sm		5.6	(a)	5.6	5.85	(a)
Eu		1.15	(a)	1.2	1.19	(a)
Gd						
Tb		1	(a)	1.27	1.27	(a)
Dy		6.7	(a)			
Ho		1.8	(a)			
Er						
Tm						
Yb		4	(a)	3.9	4.05	(a)
Lu		0.54	(a)	0.56	0.56	(a)
Hf		3.8	(a)	4.2	4.3	(a)
Ta		0.52	(a)	0.6	0.7	(a)
W ppb						
Re ppb						
Os ppb						
Ir ppb		19	(a)			
Pt ppb						
Au ppb		13	(a)			
Th ppm	1.97	(b) 1.5	(a) 1.97	(b) 1.8	2	(a)
U ppm	0.55	(b)	0.54	(b)		

technique: (a) INAA, (b) radiation counting

Table 3. Chemical composition of 61181.

reference	Eldridge73	Taylor73	Korotev82
<i>weight</i>			
SiO ₂ %		44.6 (b)	
TiO ₂		0.66 (b)	
Al ₂ O ₃		27.1 (b)	25.2 (c)
FeO		5.47 (b)	5.54 (c)
MnO			0.069 (c)
MgO		5.78 (b)	5.9 (c)
CaO		15.56 (b)	14.4 (c)
Na ₂ O		0.51 (b)	0.463 (c)
K ₂ O	0.11	(a) 0.25 (b)	
P ₂ O ₅		0.18 (b)	
S %			
<i>sum</i>			
Sc ppm		12 (b)	9.26 (c)
V		29 (b)	21 (c)
Cr		850 (b)	760 (c)
Co		27 (b)	31.5 (c)
Ni		340 (b)	435 (c)
Cu		8.5 (b)	
Zn			
Ga			
Ge ppb			
As			
Se			
Rb			
Sr			180 (c)
Y		44 (b)	
Zr		197 (b)	185 (c)
Nb		13.7 (b)	
Mo			
Ru			
Rh			
Pd ppb			
Ag ppb			
Cd ppb			
In ppb			
Sn ppb		0.11 (b)	
Sb ppb			
Te ppb			
Cs ppm			0.18 (c)
Ba			139 (c)
La		13.9 (b)	12.43 (c)
Ce		38.5 (b)	33.9 (c)
Pr		5.6 (b)	
Nd		22.1 (b)	
Sm		6.05 (b)	5.92 (c)
Eu		1.51 (b)	1.185 (c)
Gd		7.77 (b)	
Tb		1.15 (b)	1.28 (c)
Dy		7.63 (b)	
Ho		1.79 (b)	
Er		4 (b)	
Tm		0.79 (b)	
Yb		4.76 (b)	4.17 (c)
Lu		0.74 (b)	0.593 (c)
Hf		3.9 (b)	0.654 (c)
Ta			
W ppb			
Re ppb			
Os ppb			
Ir ppb			15.9
Pt ppb			
Au ppb			
Th ppm	2.02	(a) 2.12 (b)	2.09 (c)
U ppm	0.56	(a) 0.55 (b)	0.54 (c)

technique: (a) radiation counting, (b) spark source mass spec., (c) INAA

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