

**73221** – 79.5 grams  
**73241** - 360 grams  
**73261** - 326 grams  
**73281** - 169 grams  
**Trench Soils**

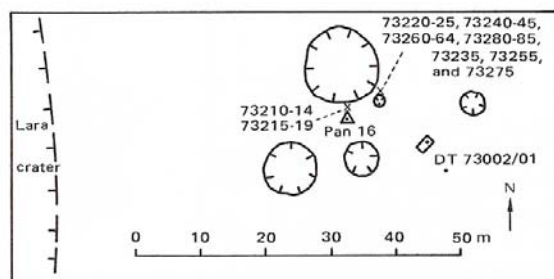


Figure 2: Map of station 3, Apollo 17, showing position of 10 meter crater with respect to ejecta blanket from Lara Crater.

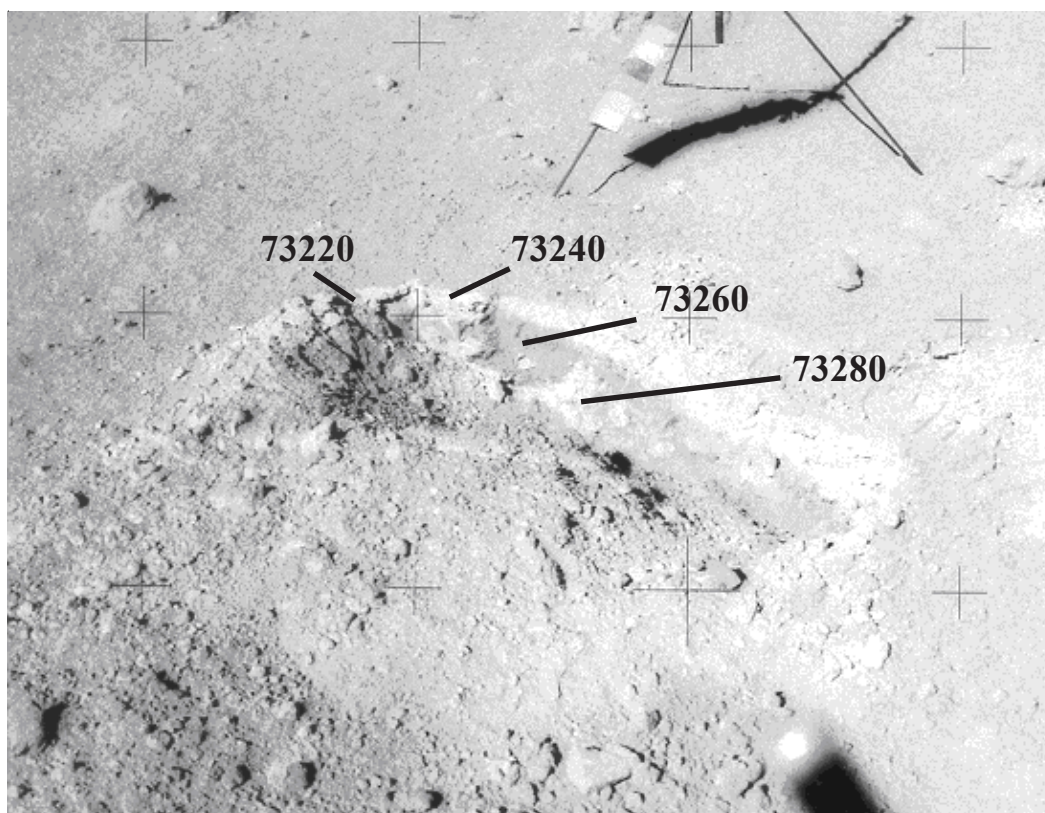


Figure 1: Trench dug at Apollo 17, station 3 (10-15 cm deep). AS17-138-21149. Gnomon legs are 50 cm apart (see also figure 3).

### **Introduction**

Station 3 was on the ejecta blanket of Lara Crater (Wolfe et al. 1981). Breccia samples 73235 and 73255 were located nearby as was a double drive tube (unopened).

The station 3 trench was dug in an area on the rim of a ten meter crater where there was noticeable color variation (figure 3). The nicely persevered color variation indicates that micro-gardening has not recently disturbed the regolith at this site. The color

variation is due to variation in agglutinate content and not due to a variation in chemical composition.

This trench showed a nice progression of cosmic-ray-induced activity – from high at the surface and low at depth - showing that the activity was due to low energy solar flares rather than high energy cosmic rays (figure 1). The only problem with trenching is that there is slumping from surface material, preventing perfect sampling.

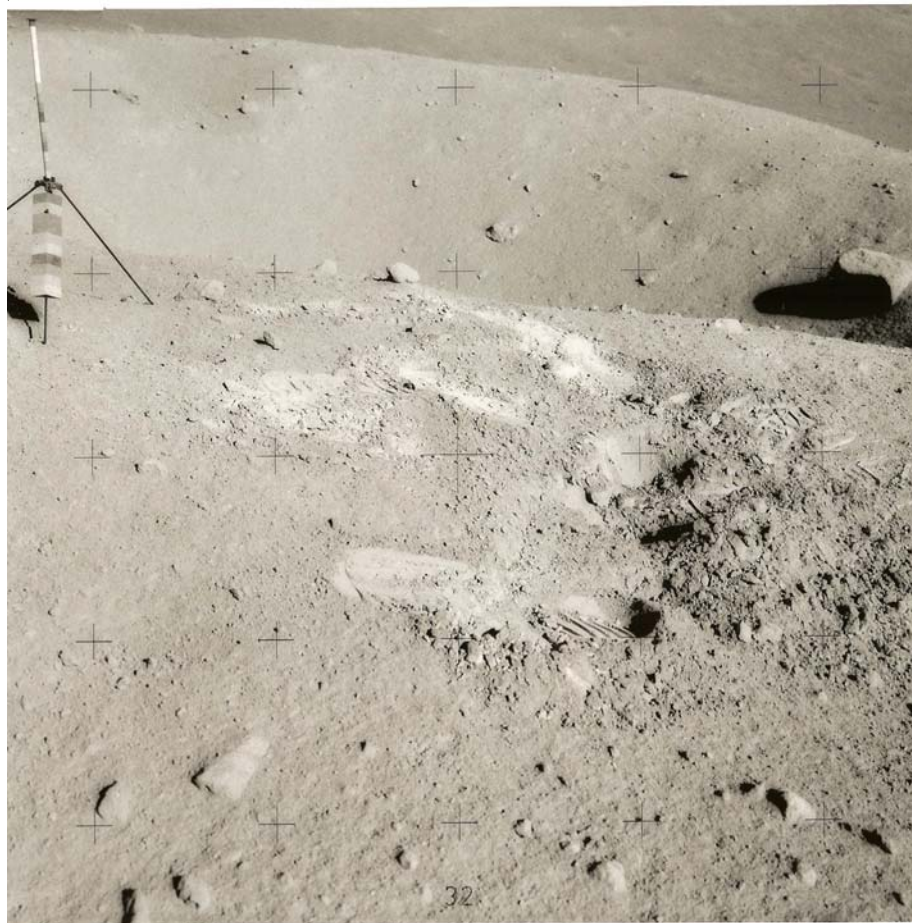


Figure 3: Location of trench on rim of crater showing streaks of light material (before trench was dug). AS17-138-21147

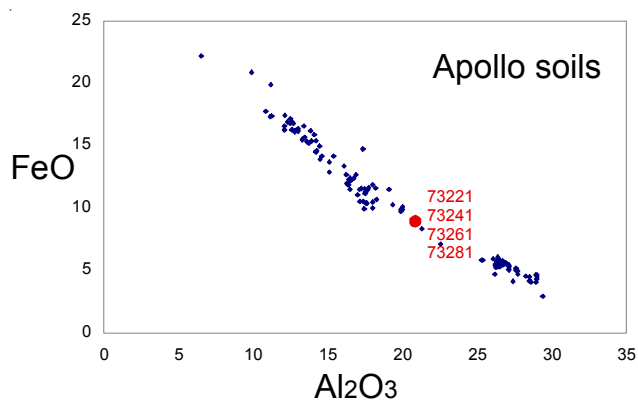


Figure 4: Composition of trench samples at station 3 compared with that of other Apollo soil samples.

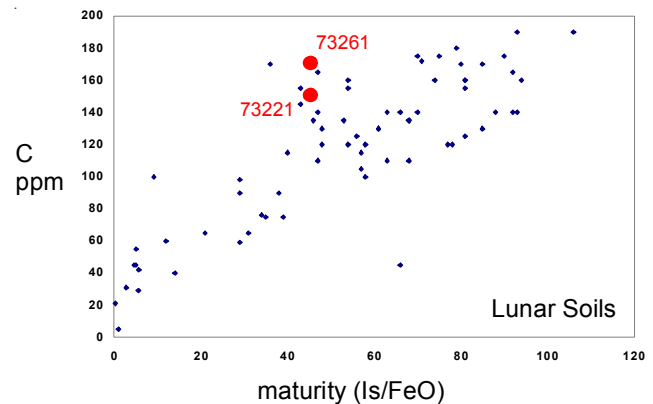


Figure 5: Carbon content and maturity index for two trench samples.

### **Petrography**

The maturity of 73221, 73241, 73261 and 73281 is  $I_s/FeO = 43, 18, 45$  and  $34$  and the average grain size is  $70, 81, 60$  and  $60$  microns, respectively (Morris 1978, Graf 1993).

Meyer (1973) found an abundance of feldspathic breccia in the  $4 - 10$  mm size fraction of 73244. Jolliff et al. (1996) and Bence et al. (1974) studied numerous coarse-fine particles from 73243 and 73263 respectively.

**Modal content of soil 73221 – 73281 (90-150 micron).**

*From Heiken and McKay 1974.*

	73221	73241	73261	73281
Agglutinates	26.3	8.4	34.3	24.6 %
Basalt	3	1	2	3.7
Breccia	46.6	61.5	39.1	46.6
Anorthosite	0.6	3	1.9	1.9
Norite				
Gabbro				
Plagioclase	11.3	11.4	9.7	9.3
Pyroxene	8	5.3	7	0.3
Olivine	1.3	0.7		0.3
Ilmenite	0.3		0.6	1.3
Orange glass		0.3	1.7	1.3
Glass other	3.5	6.7	2.5	2.9

**Chemistry**

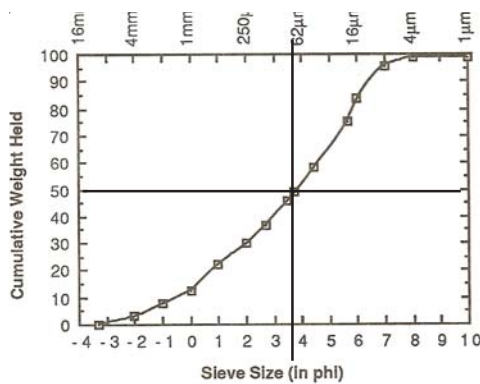
Figures 4 and 7 compare the composition of 73221 etc. with other Apollo soil samples. Rose et al. (1974) and Wanke et al. (1974) found that there was little variation in composition in spite of variation in color.

LSPET (1973) and Moore et al. (1974) reported 155 and 170 ppm carbon for 73221 and 73261 respectively (figure 5). Chang et al. (1974) reported 81 and 88 ppm

carbon for 73221 and 73261, while 73241 and 73281 had 42 and 81 ppm carbon. Chang et al. also reported 44 ppm nitrogen for 73221, 22 ppm nitrogen for 73241 and 54 and 40 ppm nitrogen for 73261 and 73281.

**Cosmogenic isotopes and exposure ages**

O’Kelley et al. (1974) determined the cosmic-ray-induced activity of <sup>22</sup>Na = 310 dpm/kg, <sup>26</sup>Al = 197 dpm/kg, <sup>46</sup>Sc = 33 dpm/kg, <sup>54</sup>Mn = 230 dpm/kg, and



average grain size = 70 microns

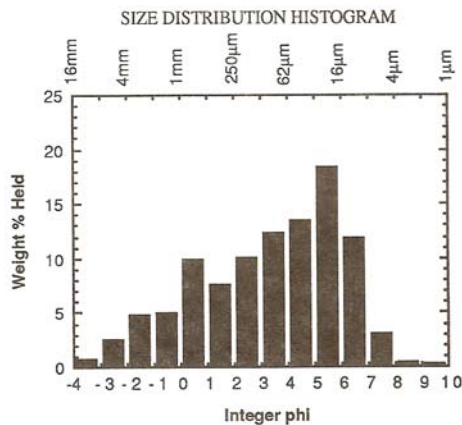
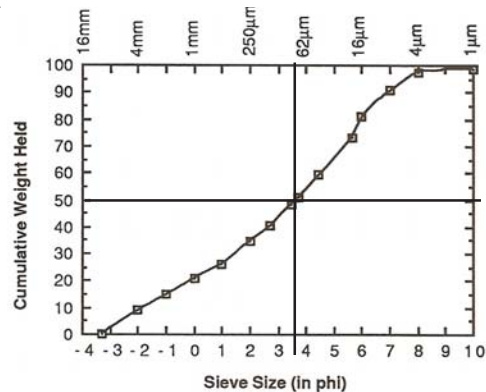


Figure 6a: Grain size distribution for 73220 (Graf 1993, data from McKay).



average grain size = 81 microns

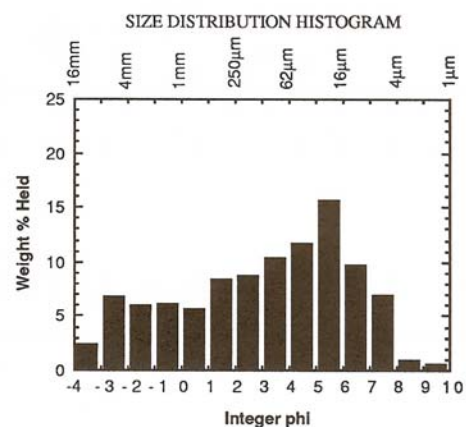
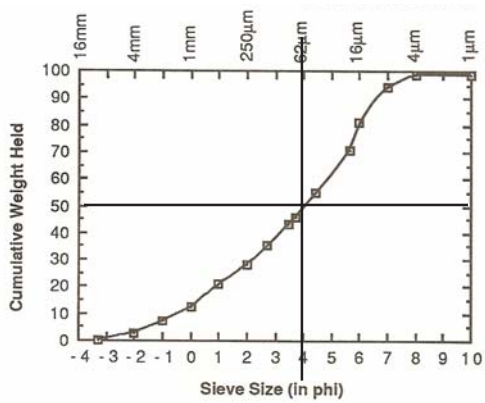
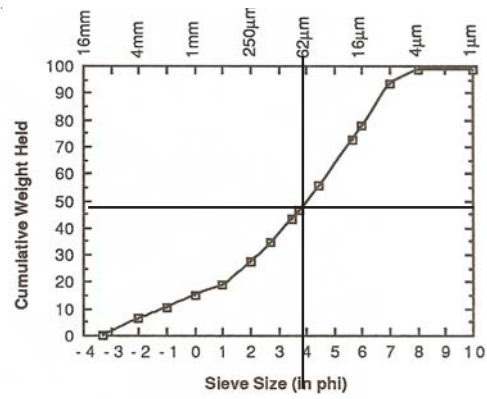


Figure 6b: Grain size distribution for 73240 (Graf 1993, data from McKay).



average grain size = 60 microns



average grain size = 60 microns

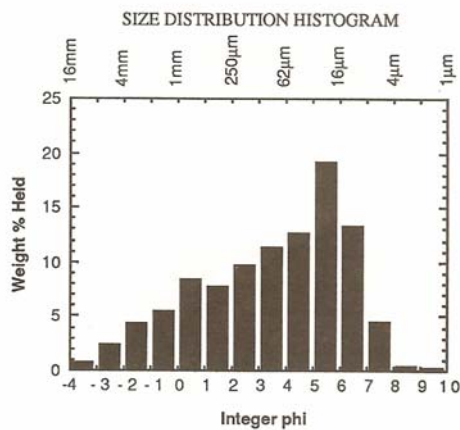


Figure 6c: Grain size distribution for 73260 (Graf 1993, data from McKay).

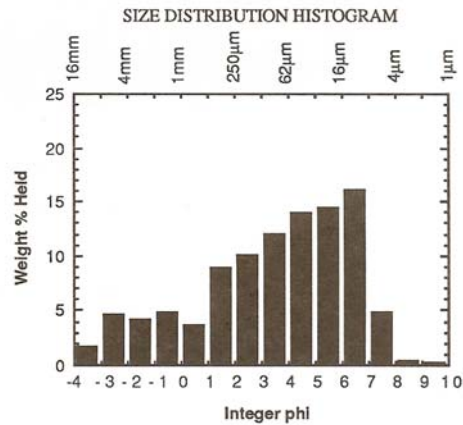


Figure 6d: Grain size distribution for 73280 (Graf 1993, data from McKay).

$^{56}\text{Co} = 810$  dpm/kg for the top soil 73221. For 73241 they determined  $^{22}\text{Na} = 110$  dpm/kg,  $^{26}\text{Al} = 92$  dpm/kg,  $^{46}\text{Sc} = 10$  dpm/kg,  $^{54}\text{Mn} = 80$  dpm/kg, and  $^{56}\text{Co} = 95$  dpm/kg. For 73261 they determined  $^{22}\text{Na} = 42$  dpm/kg,  $^{26}\text{Al} = 57$  dpm/kg,  $^{46}\text{Sc} = 8$  dpm/kg,  $^{54}\text{Mn} = 52$  dpm/kg, and  $^{56}\text{Co} = 5$  dpm/kg. For 73281 they determined  $^{22}\text{Na} = 42$  dpm/kg,  $^{26}\text{Al} = 46$  dpm/kg, and  $^{54}\text{Mn} = 50$  dpm/kg,

### Other Studies

Fireman et al. (1976, 1977) and Jull et al. (1995) studied short-lived radioactive  $^3\text{H}$  and  $^{14}\text{C}$  as function of depth in this trench.

Crozaz et al. (1974) studied the density of cosmic-ray induced nuclear tracks in feldspar crystals as function of depth in regolith.

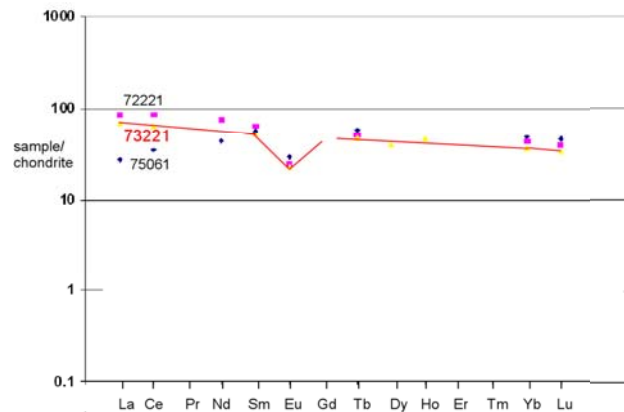


Figure 7: Normalized rare-earth-element diagram for 73221 compared with mare and highland soil samples from Apollo 17.

**Table 1. Chemical composition of 73221.**

reference weight	Rose74	Wanke74	Eldridge74
SiO2 %	45.2	(a) 45.15	(b)
TiO2	1.86	(a) 1.82	(b)
Al2O3	21.03	(a) 20.6	(b)
FeO	8.85	(a) 8.75	(b)
MnO	0.11	(a) 0.115	(b)
MgO	8.97	(a) 9.38	(b)
CaO	12.86	(a) 12.46	(b)
Na2O	0.41	(a) 0.46	(b)
K2O	0.16	(a) 0.14	(b) 0.142 (c)
P2O5			
S %			
sum		(a)	
Sc ppm	24	(a) 19.5	(b)
V	32	(a)	
Cr	1847	(a) 1500	(b)
Co	49	(a) 29.3	(b)
Ni	250	(a) 275	(b)
Cu	30	(a)	
Zn	21	(a)	
Ga	3.3	(a)	
Ge ppb			
As			
Se			
Rb	3.1	(a)	
Sr	167	(a) 160	(b)
Y	61	(a) 48	(b)
Zr	238	(a) 227	(b)
Nb	10	(a) 14	(b)
Mo			
Ru			
Rh			
Pd ppb			
Ag ppb			
Cd ppb			
In ppb			
Sn ppb			
Sb ppb			
Te ppb			
Cs ppm			
Ba	190	(a) 160	(b)
La	19	(a) 16	(b)
Ce		37.8	(b)
Pr			
Nd			
Sm		7.57	(b)
Eu		1.29	(b)
Gd			
Tb		1.7	(b)
Dy		9.6	(b)
Ho		2.6	(b)
Er			
Tm			
Yb	6	(a) 5.87	(b)
Lu		0.81	(b)
Hf		5.82	(b)
Ta		0.81	(b)
W ppb			
Re ppb			
Os ppb			
Ir ppb		9	(b)
Pt ppb			
Au ppb		13	(b)
Th ppm		2.7	(b) 2.13 (c)
U ppm			0.63 (c)

technique: (a) "microchemical", (b) multiple, (c) radiation count.

**Table 2. Chemical composition of 73241.**

reference weight	Rose74	Wanke74	Eldridge74
SiO2 %	44.55	(a) 45.37	(b)
TiO2	1.73	(a) 1.68	(b)
Al2O3	20.2	(a) 20.6	(b)
FeO	8.45	(a) 8.39	(b)
MnO	0.11	(a) 0.11	(b)
MgO	11.11	(a) 9.65	(b)
CaO	12.9	(a) 12.6	(b)
Na2O	0.46	(a) 0.45	(b)
K2O	0.16	(a) 0.145	(b) 0.146 (c)
P2O5	0.15	(a) 0.126	(b)
S %			
sum			
Sc ppm	15	(a) 17.7	(b)
V	40	(a)	
Cr	1710	(a) 1370	(b)
Co	37	(a) 27.7	(b)
Ni	320	(a) 170	(b)
Cu	9.8	(a) 6.25	(b)
Zn	18	(a) 20	(b)
Ga	4	(a) 4.7	(b)
Ge ppb		410	(b)
As		41	(b)
Se			
Rb	2.8	(a) 4.8	(b)
Sr	146	(a) 170	(b)
Y	56	(a)	
Zr	202	(a)	
Nb	12	(a)	
Mo			
Ru			
Rh			
Pd ppb		10	(b)
Ag ppb			
Cd ppb			
In ppb			
Sn ppb			
Sb ppb			
Te ppb			
Cs ppm		0.194	(b)
Ba	168	(a) 185	(b)
La		16.1	(b)
Ce		38	(b)
Pr		5.9	(b)
Nd			
Sm		7.8	(b)
Eu		1.25	(b)
Gd		9.1	(b)
Tb		1.6	(b)
Dy		10.5	(b)
Ho		2.3	(b)
Er		5.5	(b)
Tm			
Yb	4.6	(a) 5.84	(b)
Lu		0.81	(b)
Hf		5.72	(b)
Ta		0.76	(b)
W ppb		300	(b)
Re ppb		0.6	(b)
Os ppb			
Ir ppb			
Pt ppb			
Au ppb		4	(b)
Th ppm		2.3	(b) 2.25 (c)
U ppm		0.58	(b) 0.64 (c)

technique: (a) "microchemical", (b) multiple, (c) radiation count.

**Table 3. Chemical composition of 73261.**

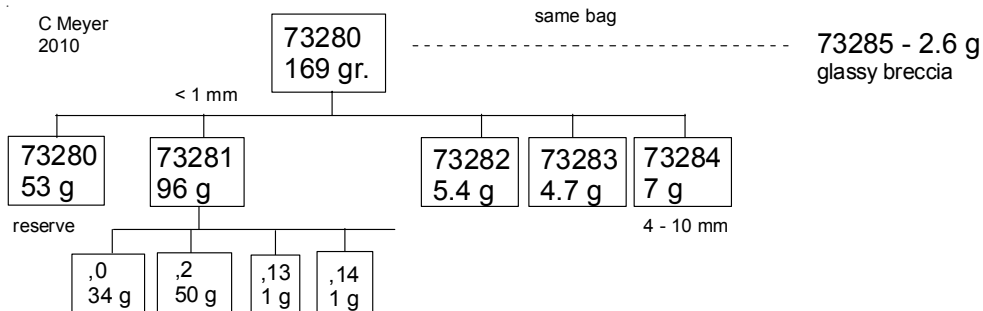
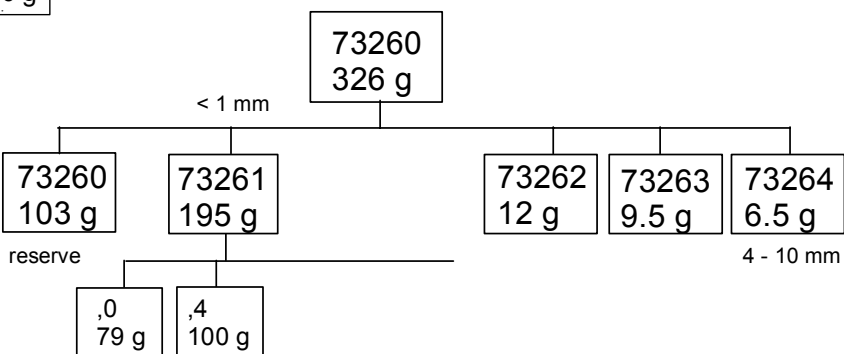
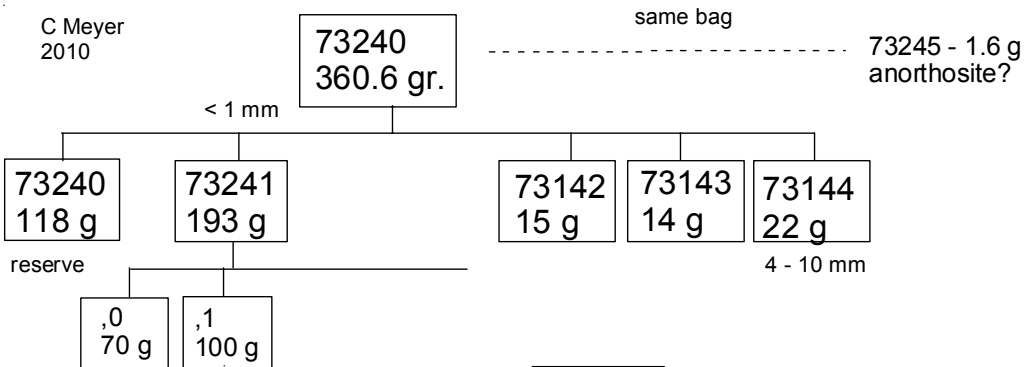
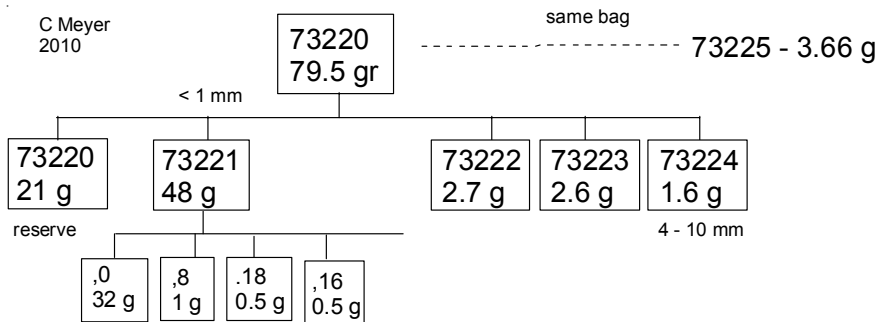
reference weight	Rose74	Wanke74	Eldridge74	
SiO2 %	44.71	(a) 45.15	(b)	
TiO2	1.9	(a) 1.82	(b)	
Al2O3	19.69	(a) 20.2	(b)	
FeO	8.86	(a) 8.84	(b)	
MnO	0.11	(a) 0.12	(b)	
MgO	10.95	(a) 9.65	(b)	
CaO	12.9	(a) 12.7	(b)	
Na2O	0.4	(a) 0.46	(b)	
K2O	0.16	(a) 0.15	(b)	0.13 (c)
P2O5	0.14	(a) 0.13	(b)	
S %				
sum				
Sc ppm	17	19.5	(b)	
V	46			
Cr	1642	1450	(b)	
Co	46	28	(b)	
Ni	450	240	(b)	
Cu	12			
Zn	18			
Ga	3.8			
Ge ppb				
As				
Se				
Rb	1.3	(a)		
Sr		130	(b)	
Y	68	(a)		
Zr	201	(a)		
Nb	12	(a)		
Mo				
Ru				
Rh				
Pd ppb				
Ag ppb				
Cd ppb				
In ppb				
Sn ppb				
Sb ppb				
Te ppb				
Cs ppm		0.18	(b)	
Ba	160	(a) 180	(b)	
La		16.5	(b)	
Ce		39.3	(b)	
Pr				
Nd				
Sm		7.85	(b)	
Eu		1.37	(b)	
Gd				
Tb		1.8	(b)	
Dy		9.8	(b)	
Ho		2.5	(b)	
Er				
Tm				
Yb	5.6	(a) 5.78	(b)	
Lu		0.82	(b)	
Hf		5.74	(b)	
Ta		0.78	(b)	
W ppb				
Re ppb				
Os ppb				
Ir ppb		15	(b)	
Pt ppb				
Au ppb				
Th ppm		2.8	(b)	2.4 (c)
U ppm				0.67 (c)

technique: (a) "microchemical", (b) multiple, (c) radiation count.

**Table 4. Chemical composition of 73281.**

reference weight	Rose74	Wanke74	Eldridge74	
SiO2 %	45.31	(a) 46	(b)	
TiO2	1.76	(a) 1.75	(b)	
Al2O3	20.23	(a) 20.8	(b)	
FeO	8.82	(a) 8.54	(b)	
MnO	0.11	(a) 0.11	(b)	
MgO	9.95	(a) 9.98	(b)	
CaO	12.91	(a) 11.76	(b)	
Na2O	0.41	(a) 0.44	(b)	
K2O	0.16	(a) 0.137	(b)	0.142 (c)
P2O5	0.14	(a)		
S %				
sum				
Sc ppm	15	(a) 17.5	(b)	
V	42	(a)		
Cr	1847	(a) 1410	(b)	
Co	46	(a) 26.7	(b)	
Ni	160	(a) 280	(b)	
Cu	7.8	(a)		
Zn	20	(a)		
Ga	3.6	(a)		
Ge ppb				
As				
Se				
Rb	2.8	(a)		
Sr	117	(a) 150	(b)	
Y	59	(a) 52	(b)	
Zr	207	(a) 245	(b)	
Nb	11	(a) 16	(b)	
Mo				
Ru				
Rh				
Pd ppb				
Ag ppb				
Cd ppb				
In ppb				
Sn ppb				
Sb ppb				
Te ppb				
Cs ppm				
Ba		160	(b)	
La		15	(b)	
Ce		35	(b)	
Pr				
Nd				
Sm		7.58	(b)	
Eu		1.22	(b)	
Gd				
Tb		1.7	(b)	
Dy		10.5	(b)	
Ho		2.4	(b)	
Er				
Tm				
Yb	4.7	(a) 5.52	(b)	
Lu		0.83	(b)	
Hf		5.78	(b)	
Ta		0.68	(b)	
W ppb				
Re ppb				
Os ppb				
Ir ppb		11	(b)	
Pt ppb				
Au ppb				
Th ppm		2.5	(b)	2.33 (c)
U ppm				0.58 (c)

technique: (a) "microchemical", (b) multiple, (c) radiation count.



## References for 73221, 73241, 73261 and 73281.

- Bence A.E., Delano J.W., Papike J.J. and Cameron K.L. (1974) Petrology of the highlands massifs at Taurus-Littrow: An analysis of the 2-4 mm soil fraction. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 785-827.
- Butler P. (1973) Lunar Sample Information Catalog Apollo 17. Lunar Receiving Laboratory. MSC 03211 Curator's Catalog. pp. 447.
- Butler J.C. and King E.A. (1974) Analysis of the grain size-frequency distributions of lunar fines. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 829-841.
- Chang S., Lennon K. and Gibson E.K. (1974) Abundances of C, N, H, He and S in Apollo 17 soils. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1785-1800.
- Crozaz G., Drozd R., Hohenberg C.M., Morgan C., Ralston C., Walker R.M. and Yuhas D. (1974) Lunar surface dynamics: Some general conclusions and new results from Apollo 16 and 17. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2475-2500.
- Eldridge J.S., O'Kelley G.D. and Northcutt K.J. (1974a) Primordial radioelement concentrations in rocks and soils from Taurus-Littrow. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1025-1033.
- Fireman E.L., DeFelice J. and D'Amico J. (1976) Solar wind 3H and 14C abundances and solar surface processes. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 525-531.
- Fireman E.L., DeFelice J. and D'Amico J. (1977) 14C in lunar soil: Temperature-release and grain-size dependence. *Proc. 8<sup>th</sup> Lunar Sci. Conf.* 3749-3754.
- Graf J.C. (1993) Lunar Soils Grain Size Catalog. NASA Reference Pub. 1265, March 1993
- Heiken G.H. (1974) A catalog of lunar soils. JSC Curator
- Heiken G.H. (1975) Petrology of lunar soils. *Rev. Geophys. Space Phys.* **13**, 567-587.
- Heiken G.H. and McKay D.S. (1974) Petrology of Apollo 17 soils. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 843-860.
- Jull A.J.T., Lal D. and Donahue D.J. (1995) Evidence for a non-cosmogenic implanted <sup>14</sup>C component in lunar samples. *Earth Planet. Sci. Lett.* **136**, 693-702.
- LSPET (1973a) Apollo 17 lunar samples : Chemical and petrographic description. *Science* **182**, 659-690.
- LSPET (1973c) Preliminary examination of lunar samples. Apollo 17 Preliminary Science Report. NASA SP-330, 7-1—7-46.
- McKay D.S., Fruland R.M. and Heiken G.H. (1974) Grain size and the evolution of lunar soils. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 887-906.
- Meyer C. (1973) Apollo 17 Coarse Fines (4-10 mm) Sample Location, Classification and Photo Index. Curator Report. pp. 182.
- Mitchell J.K., Carrier W.D., Costes N.C., Houston W.N., Scott R.F. and Hovland H.J. (1973) 8. Soil-Mechanics. In Apollo 17 Preliminary Science Rpt. NASA SP-330. pages 8-1-22.
- Moore C.B., Lewis C.F. and Cripe J.D. (1974a) Total carbon and sulfur contents of Apollo 17 lunar samples. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1897-1906.
- Moore C.B., Lewis C.F., Cripe J.D. and Volk M. (1974b) Total carbon and sulfur contents of Apollo 17 lunar samples (abs). *Lunar Sci.* **V**, 520-522. Lunar Planetary Institute, Houston.
- Morris R.V. (1976) Surface exposure indices of lunar soils: A comparative FMR study. *Proc. 7<sup>th</sup> Lunar Sci. Conf.* 315-335.
- Morris R.V., Score R., Dardano C. and Heiken G. (1983) Handbook of Lunar Soils. Two Parts. JSC 19069. Curator's Office, Houston
- Morris R.V. (1978) The surface exposure (maturity) of lunar soils: Some concepts and Is/FeO compilation. *Proc. 9<sup>th</sup> Lunar Sci. Conf.* 2287-2297.
- O'Kelley G.D., Eldridge J.S. and Northcutt K.J. (1974a) Cosmogenic radionuclides in samples from Taurus-Littrow: Effects of the solar flare of August 1972. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 2139-2147.
- Papike J.J., Simon S.B. and Laul J.C. (1982) The lunar regolith: Chemistry, Mineralogy and Petrology. *Rev. Geophys. Space Phys.* **20**, 761-826.
- Rose H.J., Cuttitta F., Berman S., Brown F.W., Carron M.K., Christian R.P., Dwornik E.J. and Greenland L.P. (1974a) Chemical composition of rocks and soils at Taurus-Littrow. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1119-1133.
- Schonfeld E. (1974) The contamination of lunar highland rocks by KREEP: Interpretations by mixing models. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1269-1286.
- Silver L.T. (1974) Patterns for U, Th, Pb distributions and isotopic relationships in Apollo 17 soils (abs). *Lunar Sci.* **V**, 706-708.



Wänke H., Palme H., Baddenhausen H., Dreibus G., Jagoutz E., Kruse H., Spettel B., Teschke F. and Thacker R. (1974) Chemistry of Apollo 16 and 17 samples: bulk composition, late-stage accumulation and early differentiation of the Moon. *Proc. 5<sup>th</sup> Lunar Sci. Conf.* 1307-1335.

Wolfe E.W., Bailey N.G., Lucchitta B.K., Muehlberger W.R., Scott D.H., Sutton R.L and Wilshire H.G. (1981) The geologic investigation of the Taurus-Littrow Valley: Apollo 17 Landing Site. US Geol. Survey Prof. Paper, 1080, pp. 280.