

76315
Impact Melt Breccia
671 grams

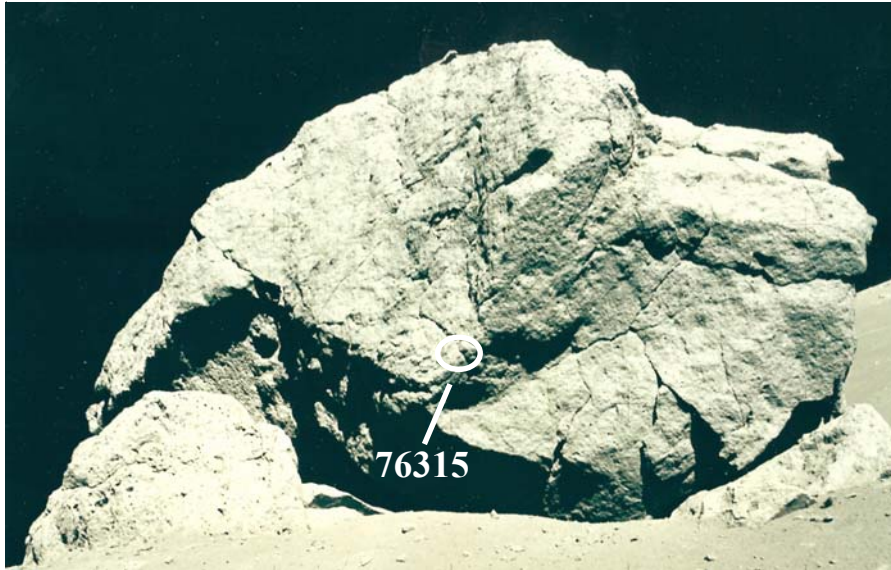


Figure 1: Location of 76315 on station 6 boulder (block 2). AS17-140-21436.



Figure 2: Photo of front, exposed, surface of 76315. The lack of zap pits on the upper left corner, indicate that this corner was facing down on the boulder. Cube is 1 cm. S73-17108.

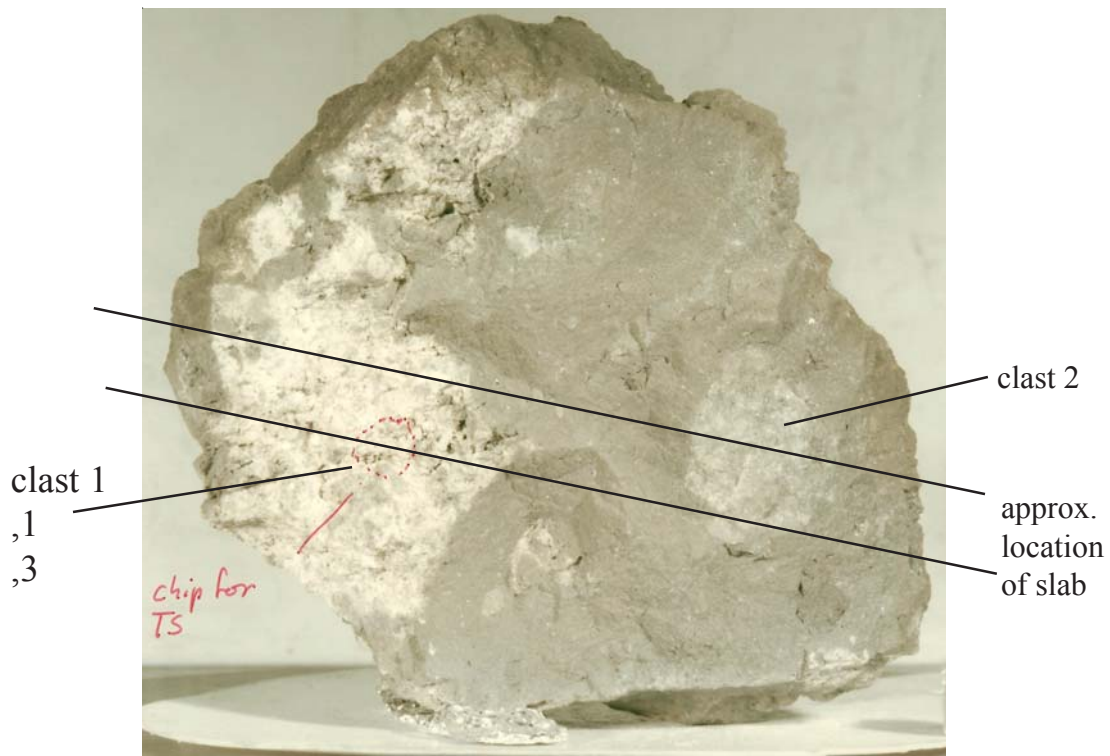


Figure 3: Fresh broken side of 76315, showing thin remnants of large white clast. About 10 cm across. S73-17109.



Figure 3: Photo of top of 76315, showing portion of thin white clast. Cube is 1 cm. S73-17104.

CDR This is the easiest part of the rock in the world to work. Here's a big white clast. There's one on top about a foot and a half across, and here's one – must be 2 feet across – 3 feet, and that's in the blue-grey.

LMP Well, Bob, I think I've done the best I can. I would – I'd say that there're pretty clearly inclusions of the blue-grey in the anorthositic gabbro here near the contact.

CC OK. And Gene, your bag is hanging by one hook there. Be careful, if you can – or LMP --

LMP OK, Bob, by accident – I didn't think I could do it but I got a sample of the inclusion. And it's in bag 539 (76315).

CDR Hey, Jack, that's your bag that's hanging by one hook. Let me get it.

LMP Oh, they're talking to me, huh?

CDR I didn't think they could see me. I'm way up on top.

LMP And it's blue-grey with light colored inclusions in it. But the whole thing seems to be pretty well altered, or metamorphosed – compared to the major rock we sampled – to the other grey rock.

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LMP OK, Bob, I think that inclusion will give you an example of what this thing – what the anorthositic gabbro did to the blue-grey breccia.

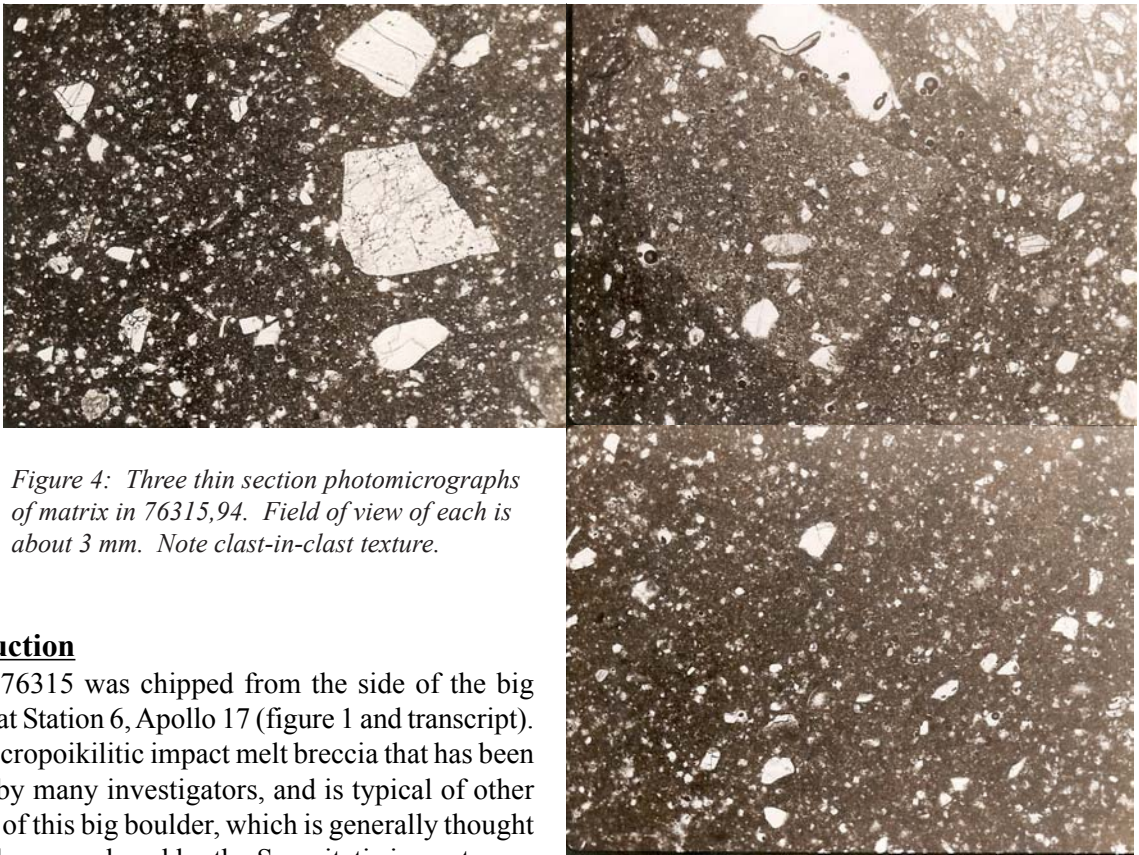


Figure 4: Three thin section photomicrographs of matrix in 76315,94. Field of view of each is about 3 mm. Note clast-in-clast texture.

Introduction

Sample 76315 was chipped from the side of the big boulder at Station 6, Apollo 17 (figure 1 and transcript). It is a micropoikilitic impact melt breccia that has been studied by many investigators, and is typical of other samples of this big boulder, which is generally thought to have been produced by the Serenitatis impact.

The crystallization age of 76315 is 3.9 b.y. with an exposure age of 22 m.y.

Petrography

Heiken et al. (1973) mapped the visible features of the Station 6 boulder and identified this blue-gray breccia sample to be from the “transitional zone” (lithology AB). The surface of 76315 was covered with patina (figure 2) such that the underlying lithology could not be discerned except on the freshly broken B1 face (figure 3). The broken surface was composed of dark grey breccia with a large irregular patch of “pink grey” material (clast 1) and a 1 x 2 cm light grey clast (2).

McGee et al. (1977), Phinney (1981) and Meyer (1994) previously made summaries of the studies conducted on 76315. The breccia matrix has finely-divided flow banding which can be seen best in the sawn surface (figure 12a). The distinct foliation is due to variations in matrix color and trains of minute vesicles. The large clasts seen in figures 3 and 4 were found to be very thin and disappointingly small in volume.

The modal mineralogy of the matrix of 76315 is about 50% plagioclase and 40% low-Ca pyroxene with minor amounts of augite, olivine, ilmenite, armalcolite and

metallic iron. The grain size of the matrix feldspar is ~10 microns while pyroxene is 25-35 microns (figure 4).

Simonds et al. (1974) studied numerous small lithic clasts in 20 thin sections of 76315, including two poikilitic 70-80% plagioclase fragments, three granulitic 70-80% plagioclase fragments, one crushed feldspar or anorthositic fragment, three intersertal plagioclase-pyroxene-olivine fragments, one crushed olivine or dunite, one poikilitic 50-60% plagioclase, two crushed spinel-olivine fragments, one crushed troctolite fragment and three aphanitic feldspathic fragments. McGee et al. (1979) one basalt clast (1 mm) with acicular plagioclase and subhedral olivine in section 76315,95 and one dunite clast (2 mm) with polygonal olivine and symplectite in 76315,97.

Norman et al. (1993) reported on the chemical composition of mineral clasts in the matrix of 76315, with the hope that they could determine the precursor rock type(s), concluding that “troctolites” were the source of the mineral fragments. However, this doesn’t explain the relatively high trace element content (figure 9). No mineralogical evidence for “low-K Fra Mauro basalt” could be found.

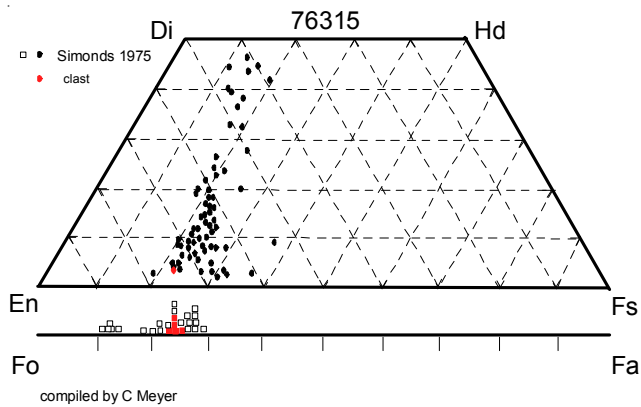


Figure 6: Pyroxene and olivine composition of matrix and clast in 76315 (from Simonds 1975).

Significant Clasts

Clast 1: Feldspathic Granulite (1, 3, 22, 52)

Figure 3 shows a large patch of crumbly “pink-grey” material. This thin brecciated clast is aligned with the direction of foliation in the matrix forming a zone of weakness along which the rock was broken during sampling from the boulder. According to Simonds (1975) this clast is granulitic in texture with ~70% plagioclase (An_{95}), ~15% low-Ca pyroxene ($Wo_{2-3}En_{83}Fs_{12}$) and ~15% olivine (Fo_{82}). Table 1 and figure 8 include an analysis (52). However, there may be more than one rock type present in this patch.

Clast 2: Anorthositic breccia (46, 61, 62, 95)

Figure 3 shows a cm-sized light-grey clast, which also turned out to be thin (figure 12a). It has a poikilitic texture with ~70% plagioclase (An_{96}), ~17% low-Ca pyroxene ($Wo_{3-5}En_{78}Fs_{18}$) and ~13% olivine (Fo_{75}). The minerals in this clast were found to be homogeneous in composition. An analysis (62) showed it to be Al-rich and the Ar plateau (61) showed that it may be significantly older than the matrix (figure 10). McGee et al. (1979) reported that thin section 76315,95 has a seriate size distribution of subangular plagioclase with less abundant pyroxene and olivine. The largest plagioclase grains have olivine necklaces with 50 micron overgrowths (figure 13).

Clast xyz:

Internal clast exposed by slab cut (see figure 12b).

Mineralogy

Olivine: Olivine in 76315 has a narrow range of composition Fo_{70-76} .

Pyroxene: Simonds (1975) determined the composition of pyroxene in matrix and clasts (figure

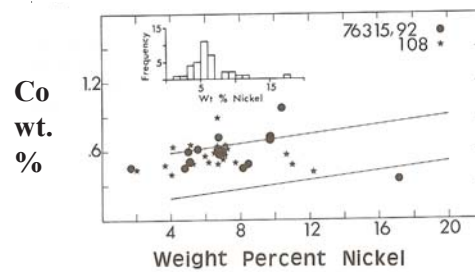


Figure 7: Ni and Co content of metal grains in 76315 (Misra et al. 1976).

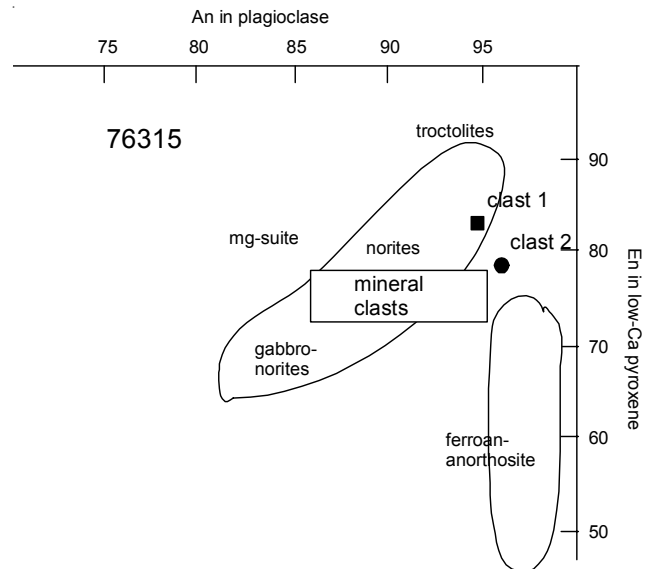


Figure 8: Composition of plagioclase and low-Ca pyroxene in 76315.

6). The low-Ca pyroxene is $Wo_4En_{60-73}Fs_{19-26}$ while the high-Ca pyroxene is $Wo_{20-30}En_{44-57}Fs_{12-15}$.

Feldspar: Plagioclase in 76315 is variable (An_{81-97}).

Pink Spinel: Simonds (1975) reports the occurrence of minor amounts of pink spinel in some areas.

Metallic Iron: Misra et al. (1976) determined Ni and Co in iron grains (figure 7).

Chemistry

Hubbard et al. (1974), Rhodes et al. (1974) and Wiesmann and Hubbard (1976) reported numerous analyses of matrix and clasts in 76315 (Table 1 and figure 9). Morgan et al. (1974), Gros et al. (1976) and Norman et al. (2002) have determined the siderophile and trace element content of the matrix of 76315. Jovanovic and Reed (1975) and Allen et al. (1975) determined F, Cl, I, Li, U, Ru, Os and Pb.

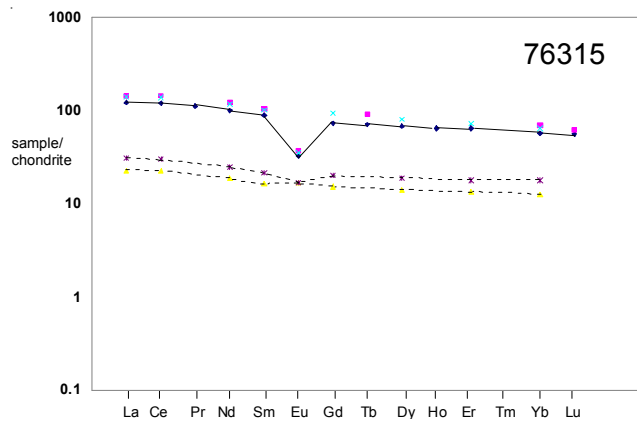


Figure 9: Normalized rare-earth-element diagram for 76315 matrix and clasts.

Radiogenic age dating

Turner and Cadogen (1975) determined the age of the station 6 boulder (figure 10). This has been confirmed by Dalrymple and Ryder (1996) who obtained an age of 3.9 b.y. (figure 11). That's the age of Serenitatis!

Nyquist et al. (1974) found the Rb/Sr systematics of the granulite clast were aligned with other Apollo 17 granulites and gave old model ages. Silver (1974) studied U, Th and Pb isotopes in 76315.

Cosmogenic isotopes and exposure ages

Turner and Cadogen (1975) determined the cosmic ray exposure age of 20, 19 and 17 m.y. by the ^{38}Ar method while Crozaz et al. (1975) determined $^{81}\text{Kr} = 22$ m.y. (see discussion in Arvidson et al. 1975). Crozaz et al. (1975) also studied cosmic ray tracks in 76315.

Other Studies

Crozaz et al. (1975) and Hohenburg et al. (1980) reported Xe analyses of 76315. Bogard (1974) also analysed rare gas content.

Mayeda et al. (1975) determined the isotopic composition of oxygen in plagioclase and olivine in 76315.

Nagata et al. (1975) reported magnetic data and discussed Ni/Fe phases and Brecher (1976) studied the alignment of magnetization with foliation. Magnetization of this rock was also studied by Stephenson et al. (1974), Pearce et al. (1974) and Gose et al. (1976). Housley et al. (1975) determined the ferromagnetic resonance and Huffman et al. (1975) determined the iron distribution by Mossbauer.

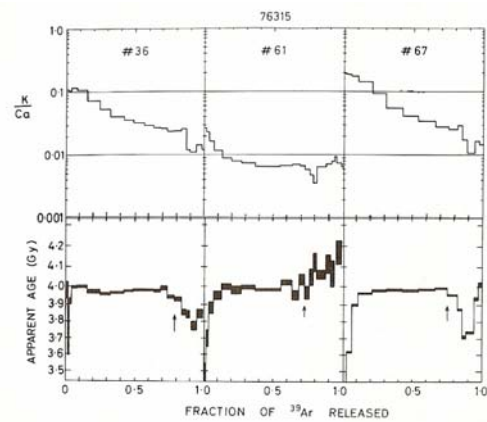


Figure 10: Argon release diagram (Turner and Cadogen 1975). Split ,61 is the white clast.

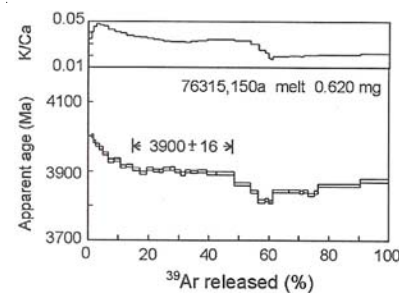


Figure 11: Ar/Ar plateau age diagram for 76315 (from Dalrymple and Ryder 1996).

Summary of Age Data for 76315

	Ar/Ar
Turner and Cadogen 1975	3.98 ± 0.04 b.y.
Dalrymple and Ryder 1996	3.900 ± 0.016

Caution: Beware changes in decay constant.

Adams and Charette (1975) compared the reflectance spectra with other samples (but didn't discuss the effect of patina).

Processing

76315 was studied as part of the Station 6 boulder consortium (Phinney 1981).

A slab was cut through the middle of 76315 (figure 12). There are 22 thin sections.

There were about 22 grams of dirt and chips in the bag with 76315 - which might include additional chips of the white clast (someone needs to look). In addition, chips of the white clast may have fallen off during sawing of the slab (sawing is a rough operation).

Table 1. Chemical composition of 76315.

reference weight	Norman2002	Dalrymple96 melt	,62c Hubbard74 Wiesmann76 Rhodes 74	,2 Wiesmann76 Rhodes 74	,30M	C3	,35M	,52c	Gnos 76	Morgan74				
SiO2 %	46.6	(b)			45.64	46.45	46.21	48.57	(e)					
TiO2	1.4	(b)		(e)	1.5	1.43	1.5	0.32	(e)					
Al2O3	18.3	(b)		(e)	17.53	18.8	18.14	17.91	(e)					
FeO	8.02	(b)	8.4	(c) 5.29	(e) 9.53	8.83	8.95	7.66	(e)					
MnO	0.1	(b)		(e) 0.07	(e) 0.13	0.13	0.12	0.13	(e)					
MgO	12	(b)		(e) 7.46	(e) 12.5	12.34	12.02	13.84	(e)					
CaO	11	(b)		(e) 15.12	(e) 10.97	11.3	11.32	10.36	(e)					
Na2O	0.65	(c) 0.69	(c) 0.47	(e) 0.61	0.7	0.63	0.81	0.47	(e)					
K2O	0.21	(c) 0.24	(c) 0.097	(d) 0.27	0.27	0.22	0.26	0.15	(d)					
P2O5				(e) 0.06	(e) 0.3	0.29	0.29	0.12	(e)					
S %				(e) 0.04	(e) 0.08	0.07	0.07		(e)					
sum														
Sc ppm	18.2	(a)	16.3	(c)										
V	46	(a)												
Cr	1355	(a)	1317	(c) 770	(d) 1228	1302	1257	1389	813	(d)				
Co	31.2	(a)	16	(c)										
Ni	283	(a)	83	(c)					423	256	260	(f)		
Cu	12.1	(a)												
Zn	13.8	(a)							2	3.1	3.4	(f)		
Ga	5.1	(a)												
Ge ppb									58	346	354	(f)		
As														
Se									71	100	107	(f)		
Rb	7.7	(a)		2.336	(d) 5.88	6.56	3.85	5.78	3.73	(d) 2.73	5.91	5.9	(f)	
Sr	182	(a)	179	(c) 153.1	(d) 180	175	172	174	115	(d)				
Y	113	(a)												
Zr	493	(a)	330	(c) 95	(d) 477	485	465		105	(d)				
Nb	35.1	(a)												
Mo														
Ru	6.77	(a)												
Rh														
Pd ppb	7.87	(a)								23				
Ag ppb										0.7	0.8	0.9		
Cd ppb										12	5	6		
In ppb										4.6				
Sn ppb														
Sb ppb										0.8	1.5	1.5	(f)	
Te ppb										3.4	4	5.1	(f)	
Cs ppm	0.12	(a)	0.22	(c)						0.11	0.25	0.25	(f)	
Ba	355	(a)	373	(c) 72.8	(d) 359	349	366	337	129	(d)				
La	28.6	(a)	34	(c) 5.41	(d) 30.1	32.9	24.7	31.6	7.33	(d)				
Ce	72.5	(a)	88.3	(c) 13.7	(d) 84.6	84	78.6	82.3	18.4	(d)				
Pr	9.99	(a)												
Nd	45.8	(a)	56	(c) 8.6	(d) 53.5	53.5	50.2	52.7	11.5	(d)				
Sm	13.1	(a)	15.4	(c) 2.42	(d) 15.1	15.1	14.1	14.8	3.2	(d)				
Eu	1.83	(a)	2.09	(c) 0.94	(d) 2	1.97	1.88	1.95	0.971	(d)				
Gd	14.6	(a)		2.99	(d) 18.9	18.5	17.6	18.8	3.93	(d)				
Tb	2.59	(a)	3.3	(c)										
Dy	16.5	(a)		3.39	(d) 19.9	19.7	18.3	19.1	4.59	(d)				
Ho	3.61	(a)												
Er	10.3	(a)		2.14	(d) 11.7	11.5	11	11.4	2.91	(d)				
Tm														
Yb	9.43	(a)	11.3	(c) 2.07	(d) 11	10.6	10	10.4	2.98	(d)				
Lu	1.37	(a)	1.5	(c) 0.3	(d)				0.455	(d)				
Hf	10.1	(a)	12	(c) 5.3	(d) 12.5					(d)				
Ta	1.5	(a)	1.56	(c)										
W ppb	0.86	(a)												
Re ppb	0.36	(a)								1.85	0.507	0.575	(f)	
Os ppb										21			(f)	
Ir ppb	3.45	(a)								18.6	5.42	5.97	(f)	
Pt ppb	7.76	(a)												
Au ppb										6.4	3.2	3.5	(f)	
Th ppm	5.87	(a)	5.3	(c) 1.234	(d) 5.2	5.36	5.23	5.69	1.34	(d)				
U ppm	1.54	(a)	1.62	(c) 0.343	(d) 1.52	1.47	1.36	2.52	0.34	(d)	0.355	1.54	1.49	(f)

technique: (a) ICP-MS, (b) fused-bead, e-probe, (c) INAA, (d) IDMS, (e) XRF, (f) RNAA

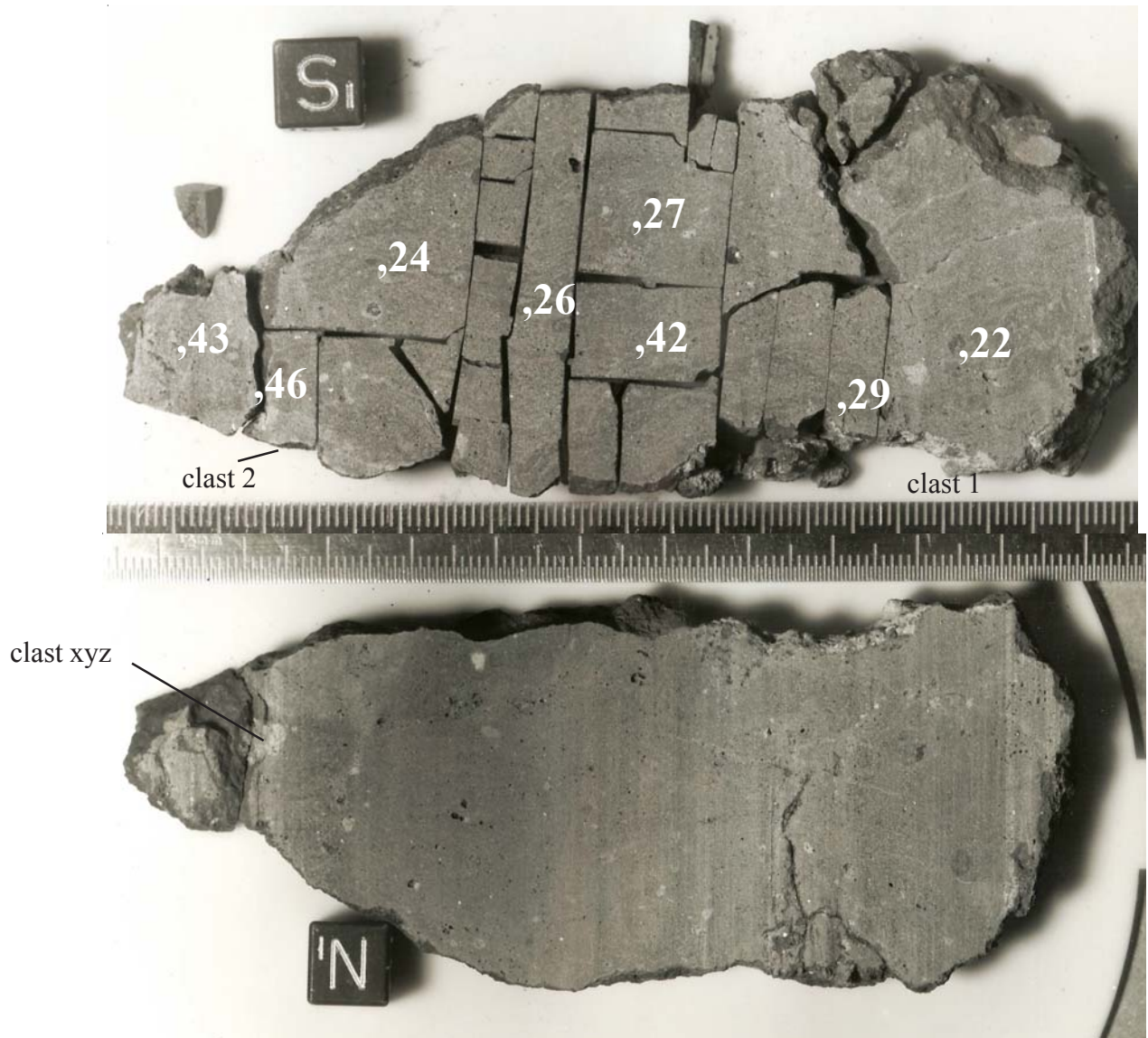
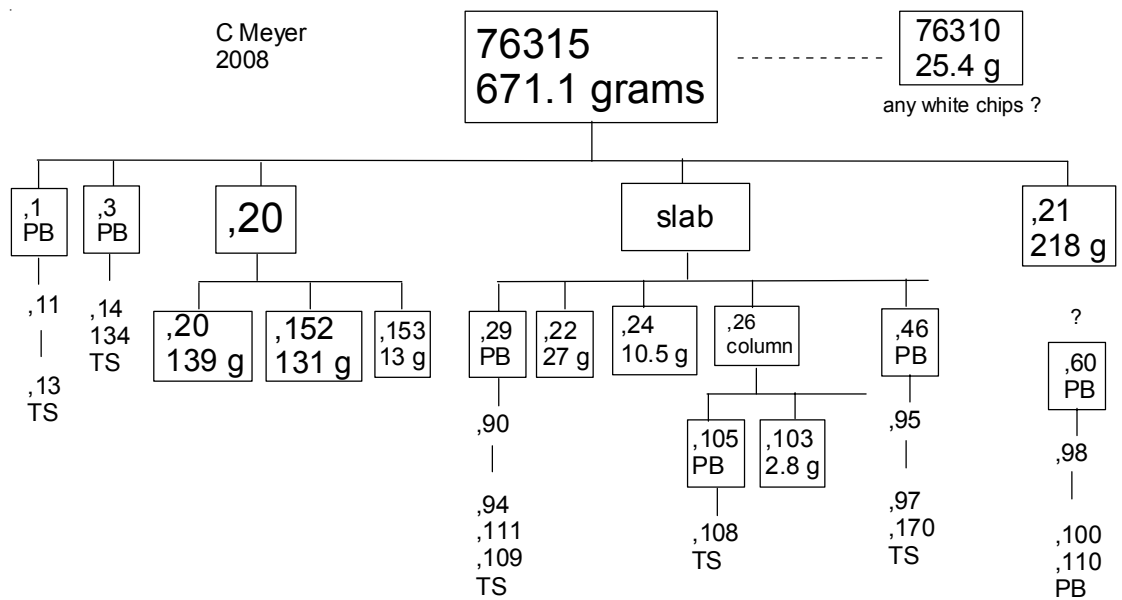


Figure 12a, b: Front and back sides of slab cut through the middle of 76315. Cube is 1 cm. Top is S73-35833; bottom is 34145. Note how thin the “white clast” is.



References for 76315

- Adams J.B. and Charette M.P. (1975) Spectral reflectance of highland rock types at Apollo 17: Evidence from Boulder 1, Station 2. *The Moon* 14, 483-489.
- Arvidson R., Crozaz G., Drozd R.J., Hohenberg C.M. and Morgan C.J. (1975) Cosmic ray exposure ages of features and events at the Apollo landing sites. *The Moon* 13, 259-276.
- Bogard D.D., Nyquist L.E. and Hirsch W.C. (1974) Noble gases in Apollo 17 boulders and soils (abs). *Lunar Sci.* V, 73-75. (unpublished data is available in Phinney 1981)
- Brecher A. (1977a) Interrelationships between magnetization directions, magnetic fabric and oriented petrographic features in lunar rocks. *Proc. 8th Lunar Sci. Conf.* 703-723.
- Butler P. (1973) **Lunar Sample Information Catalog Apollo 17.** Lunar Receiving Laboratory. MSC 03211 Curator's Catalog. pp. 447.
- Crozaz G., Drozd R., Hohenberg C., Morgan C., Ralston C., Walker R. and Yuhas D. (1974a) Lunar surface dynamics: Some general conclusions and new results from Apollo 16 and 17. *Proc. 5th Lunar Sci. Conf.* 2475-2499.
- Dalrymple G.B. and Ryder G. (1996) $^{40}\text{Ar}/^{39}\text{Ar}$ laser step heating ages of some Apollo 17 melt rocks and the age of the Serenitatis impact (abs). *Lunar Planet. Sci.* XXVII, 285-286. Lunar Planetary Inst., Houston.
- Dalrymple G.B. and Ryder G. (1996) Argon-40/argon-39 age spectra of Apollo 17 highlands breccia samples by laser step heating and the age of the Serenitatis basin. *J. Geophys. Res.* 101, 26069-26084.
- Gibson E.K. and Moore G.W. (1974a) Sulfur abundances and distributions in the valley of Taurus-Littrow. *Proc. 5th Lunar Sci. Conf.* 1823-1837.
- Gose W.A., Strangway D.W. and Pearce G.W. (1976) Origin of magnetization in lunar breccias: An example of thermal overprinting (abs). *Lunar Sci.* VII, 322-324. Lunar Planetary Institute, Houston
- Gros J., Takahashi H., Hertogen J., Morgan J.W. and Anders E. (1976) Composition of the projectiles that bombarded the lunar highlands. *Proc. 7th Lunar Sci. Conf.* 2403-2425.
- Heiken G.H., Butler P., Simonds C.H., Phinney W.C., Warner J., Schmitt H.H., Bogard D.D. and Pearce W.G. (1973a) Preliminary data on boulders at Station 6, Apollo 17 landing site. NASA TMX-58116, pp. 56.
- Hohenberg C.M., Hudson B., Kennedy B.M. and Podosek F.A. (1980) Fission xenon in troctolite 76535. In *Proc. Conf. Lunar Highlands Crust. Geochim. Cosmochim. Acta, Suppl.* 12. Pergamon Press. 419-439. Lunar Planetary Institute, Houston.
- Housley R.M., Cirlin E.H., Goldberg I.B., Crowe H., Weeks R.A. and Perhac R. (1975) Ferromagnetic resonance as a method of studying the micrometeorite bombardment history of the lunar surface. *Proc. 6th Lunar Sci. Conf.* 3173-3186.
- Hubbard N.J., Rhodes J.M., Wiesmann H., Shih C.Y. and Bansal B.M. (1974) The chemical definition and interpretation of rock types from the non-mare regions of the Moon. *Proc. 5th Lunar Sci. Conf.* 1227-1246.
- Huffman G.P., Schwerer F.C., Fisher R.M. and Nagata T. (1974) Iron distribution and metallic-ferrous ratios for Apollo lunar samples: Mossbauer and magnetic analyses. *Proc. 5th Lunar Sci. Conf.* 2779-2794.
- James O.B. (1994) Siderophile and volatile elements in Apollo 17 impact melts (abs). *Lunar Planet. Sci.* XXV, 617-618. Lunar Planetary Institute, Houston.
- Jovanovic S. and Reed G.W. (1974a) Labile and nonlabile element relationships among Apollo 17 samples. *Proc. 5th Lunar Sci. Conf.* 1685-1701.
- LSPET (1973) Apollo 17 lunar samples: Chemical and petrographic description. *Science* 182, 659-672.
- LSPET (1973) Preliminary Examination of lunar samples. Apollo 17 Preliminary Science Rpt. NASA SP-330. 7-1 – 7-46.
- Mayeda T.K., Shearer J. and Clayton R.N. (1975) Oxygen isotope fractionation of Apollo 17 rocks. *Proc. 6th Lunar Sci. Conf.* 1799-1802.
- McGee P.E., Simonds C.H., Warner J.L. and Phinney W.C. (1979) Introduction to the Apollo Collections: Part II Lunar Breccias. Curators Office.
- Meyer C. (1994) Catalog of Apollo 17 rocks. Vol. 4 North Massif Curator's Office
- Misra K.C., Walker B.M. and Taylor L.A. (1976a) Textures and compositions of metal particles in Apollo 17, Station 6 boulder samples. *Proc. 7th Lunar Sci. Conf.* 2251-2266.
- Morgan J.W., Ganapathy R., Higuchi H., Krahenbuhl U. and Anders E. (1974a) Lunar basins: Tentative characterization of projectiles, from meteoritic elements in Apollo 17 boulders. *Proc. 5th Lunar Sci. Conf.* 1703-1736.

Muehlberger et al. (1973) Documentation and environment of the Apollo 17 samples: A preliminary report. *Astrogeology* 71 322 pp superceded by *Astrogeology* 73 (1975) and by Wolfe et al. (1981)

Muehlberger W.R. and many others (1973) Preliminary Geological Investigation of the Apollo 17 Landing Site. *In Apollo 17 Preliminary Science Report*. NASA SP-330.

Norman M.D., Taylor G.L., Spudis P. and Ryder G. (1993) Lithologies contributing to the clast population in Apollo 17 LKFM basaltic impact melts. *In Workshop on Geology of the Apollo 17 Landing site*. Lunar Planetary Institute, Houston Tech. Rpt. 92-09. 42-44.

Norman M.D., Bennett V.C. and Ryder G. (2002) Targeting the impactors: highly siderophile element signatures of lunar impact melts from Serenitatis. *Earth Planet. Sci. Lett.* 202, 217-228.

Nagata T., Fisher R.M., Schwerer F.C., Fuller M.D. and Dunn J.R. (1975a) Effects of meteorite impact on magnetic properties of Apollo lunar materials. *Proc. 6th Lunar Sci. Conf.* 3111-3122.

Nyquist L.E., Bansal B.M., Wiesmann H. and Jahn B.-M. (1974a) Taurus-Littrow chronology: some constraints on early lunar crustal development. *Proc. 5th Lunar Sci. Conf.* 1515-1539.

Pearce G.W., Strangway D.W. and Gose W.A. (1974a) Magnetic properties of Apollo samples and implications for regolith formation. *Proc. 5th Lunar Sci. Conf.* 2815-2826.

Phinney W.C. (1981) Guidebook for the Boulders at Station 6, Apollo 17. Curatorial Branch Publication 55, JSC- 17243 pp. 125.

Rhodes J.M., Rodgers K.V., Shih C., Bansal B.M., Nyquist L.E., Wiesmann H. and Hubbard N.J. (1974a) The relationships between geology and soil chemistry at the Apollo 17 landing site. *Proc. 5th Lunar Sci. Conf.* 1097-1117.

Silver L.T. (1974) unpublished in Phinney (1981)

Simonds C.H., Phinney W.C. and Warner J.L. (1974) Petrography and classification of Apollo 17 non-mare rocks with emphasis on samples from the Station 6 boulder. *Proc. 5th Lunar Sci. Conf.* 337-353.

Simonds C.H. (1975) Thermal regimes in impact melts and the petrology of the Apollo 17 Station 6 boulder. *Proc. 6th Lunar Sci. Conf.* 641-672.

Simonds C.H., Phinney W.C., Warner J.L. and Heiken G.H. (1975) Thermal regimes in crater debris as deduced from

the petrology of the Apollo 17 Station 6 boulder and rake samples (abs). *Lunar Sci.* VI, 747-749. Lunar Planetary Institute, Houston.

Stephenson A., Collinson D.W. and Runcorn S.K. (1974) Lunar magnetic field paleointensity determinations on Apollo 11, 16, and 17 rocks. *Proc. 5th Lunar Sci. Conf.* 2859-2871.

Turner G. and Cadogan P.H. (1975a) The history of lunar bombardment inferred from ⁴⁰Ar-³⁹Ar dating of highland rocks. *Proc. 6th Lunar Sci. Conf.* 1509-1538.

Warner J.L., Phinney W.C., Bickel C.E. and Simonds C.H. (1977) Feldspathic granulitic impactites and pre-final bombardment lunar evolution. *Proc. 8th Lunar Sci. Conf.* 2051-2066.

Wiesmann H. and Hubbard N.J. (1975) A compilation of the Lunar Sample Data Generated by the Gast, Nyquist and Hubbard Lunar Sample PI-Ships. Unpublished. JSC

Wolfe E.W., Bailey N.G., Lucchitta B.K., Muehlberger W.R., Scott D.H., Sutton R.L and Wilshire H.G. (1981) The geologic investigation of the Taurus-Littrow Valley: Apollo 17 Landing Site. US Geol. Survey Prof. Paper, 1080, pp. 280.

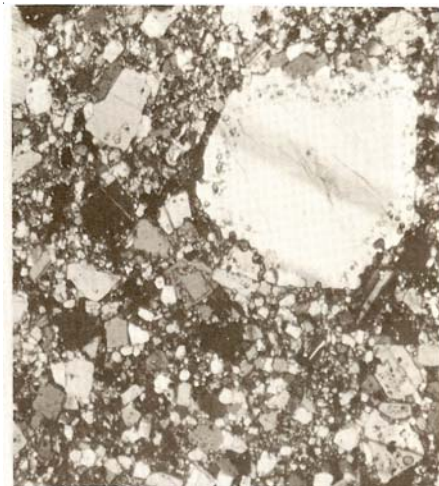


Figure 13: Plagioclase-rich clast with olivine necklace in large plagioclase grain (0.7 mm). 76315,95 (clast 2?). From McGee et al. (1979).