

**CARBON DIOXIDE ON THE SATELLITES OF SATURN.** Dale P. Cruikshank, NASA Ames Research Center, Mail Stop 245-6, Moffett Field, CA, 94035. Dale.P.Cruikshank@nasa.gov

The spectral signature of solid CO<sub>2</sub> (the C-O asymmetric stretching mode at 4.27 μm) is seen in the Cassini VIMS data for several satellites of Saturn [1,2,3,4]. It often appears strongest in geographic regions of low albedo, compared to the higher albedo regions dominated by H<sub>2</sub>O ice; this may be an effect of band contrast rather than true spatial association. In all cases, although to varying degrees, the CO<sub>2</sub> band is shifted to shorter wavelengths than is seen in pure CO<sub>2</sub> at 80 K in the laboratory, suggesting that CO<sub>2</sub> is complexed in one or more ways with another molecule or material. Lab spectra and *ab initio* calculations of molecular complexes of CO<sub>2</sub> with 1, 2, and more H<sub>2</sub>O molecules [5] show that the wavelength shift seen on some satellites can be explained in this way. The CO<sub>2</sub> band on Hyperion is different from that on the other satellites in that its band is further shifted and has a different profile. Both of these spectral parameters are matched by the type II clathrate of CO<sub>2</sub> in H<sub>2</sub>O [6]. Although CO<sub>2</sub> clathrate has been postulated to occur on Mars and Enceladus, its spectroscopic signature has not been seen until this identification on Hyperion. The implications of CO<sub>2</sub> clathrate on Hyperion are far reaching in terms of its propensity to decompose explosively [7], Hyperion's low bulk density, the distribution of low-albedo (organic rich) material on its surface, and other features of this satellite.

**References:** [1] Clark, R. N., et al. (2005) *Nature* 435, 66, [2] Buratti, B. J., et al. (2005) *Ap. J. Lett.* 622, L149, [3] Brown, R. H., et al. (2006) *Science* 311, 1425, [4] Cruikshank, D. P., et al. submitted, [5] Chaban, G. M., Bernstein, M. P., and Cruikshank, D. P. (2007) *Icarus* 187, 592, [6] Prasad, P. S. R., Shiva Prasad, K., and Thakur, N. K. (2006) *Current Science* 90, 1544, [7] Kieffer, S. W., et al. (2006) *Science* 314, 1764.