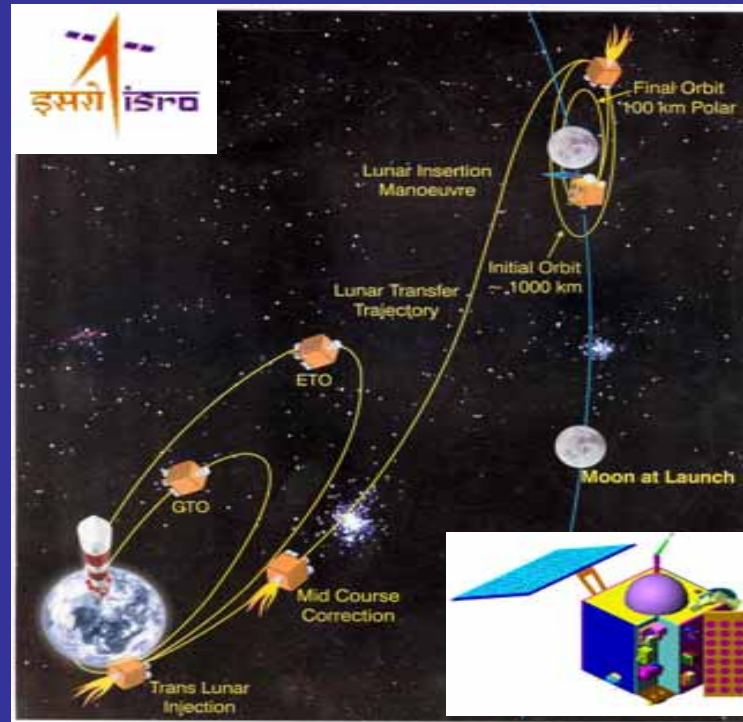


# Chandrayaan Mission Objectives and future lunar programs



*Paul Spudis, channeling J. N. Goswami*

*Principal Scientist, Chandrayaan-1 Mission*

*The Clementine & Lunar Prospector Missions to Moon raised*

## **New Questions about the Evolution of Moon**

I. Origin of the Moon: Chance or Destiny?

II. Melting of the Moon: The Magma Ocean Hypothesis

III. The bulk chemistry of Moon, particularly the abundance of iron, magnesium and uranium

IV. Lunar asymmetries & formation of SPA basin

V. The nature and structure of the lunar crust and mantle

VI. Water on Moon and the nature of Volatile transport



And, the List is growing.....

# *Understanding the origin and Evolution of the Moon*

## **Physical Properties of the Moon**

Topography

Gravity

Magnetic Field

Radiation Environment



## **The bulk chemistry of Moon**

**Nature of the Lunar Crust**

**The lunar far side: rock types, chemistry**

## ***Special Regions of Interest:***

Polar Regions ,

South Pole-Aitken Basin Region,

Selected Basins and Craters as windows into the crust

**Nature of the Magma Ocean and Lunar Interior**

***Nature of Volatile Transport on Moon (Water on Moon?)***



# Objectives of the Chandrayaan-1 Mission



**Simultaneous Mineralogical, Chemical and Photogeological Mapping at resolutions better than previous and currently planned lunar missions**

**Direct estimation of lunar surface concentration of the elements Mg, Al, Si, Ca, Ti and Fe with high spatial resolution (20 km)**

**High resolution UV-VIS-NIR mapping of the lunar surface to identify mineralogy and selected elements (Fe, Al, Mg, Ti)**

**3D mapping of lunar surface at very high spatial resolution (~5 m)**

**Volatile Transport to colder polar regions (using Pb-210 as tracer)**



# Chandrayaan-1 Mission



**Configuration :** 100 km polar orbiter

**Observation Period :** 2 years

## Baseline Payloads:



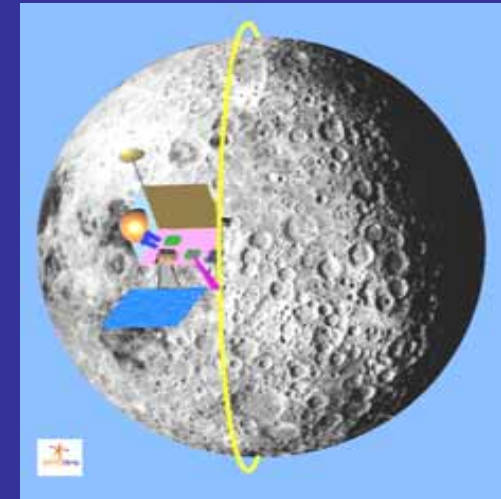
Hyper Spectral Imager (HySI) (0.4-0.9 $\mu$ m)

Terrain Mapping Camera (TMC)

Laser Ranging (LLRI)

High energy X- $\gamma$  ray spectrometer (HEX) (10-200KeV)

Moon Impact Probe (MIP), camera and altimeter hard lander





# Chandrayaan-1 Mission



## *Imaging Payloads on Chandrayaan-1*

### I. Terrain Mapping Camera (TMC)

*Stereoscopic instrument in Panchromatic band for  
3D Topographic mapping with high spatial  
and altitude resolution.*

### II. Hyper Spectral Imager (HySI)

*Imager for Mineralogical mapping in UV-VIS-NIR  
with high spectral resolution.*

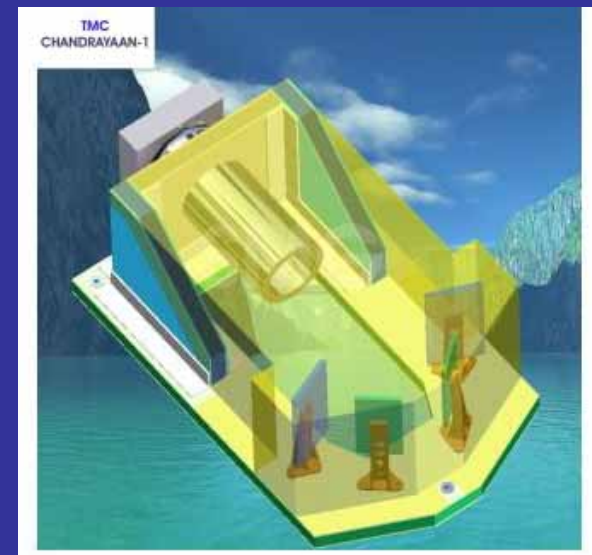
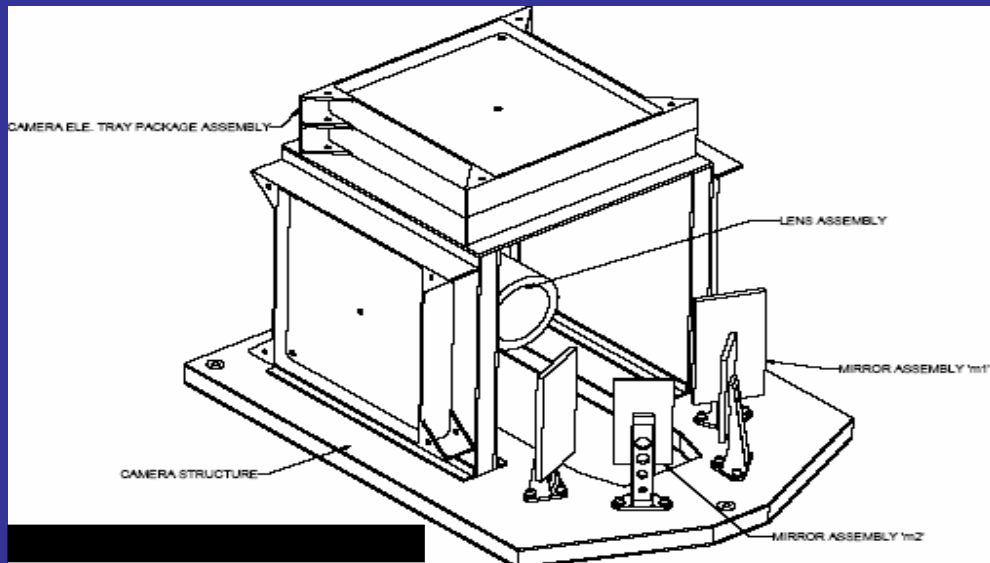
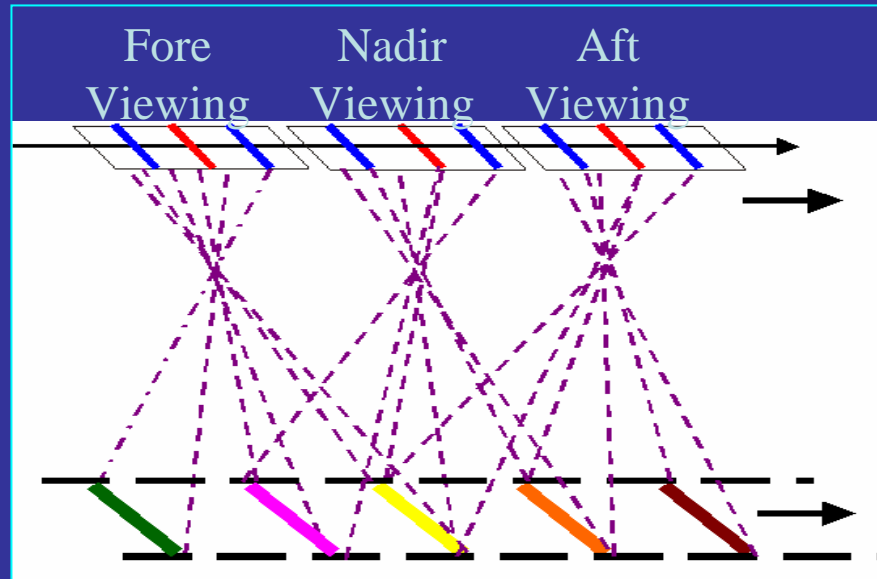
[Designed and Fabricated at the Space Application Center, ISRO]



# Chandrayaan-1 Mission



## TMC concept & Instrument configuration





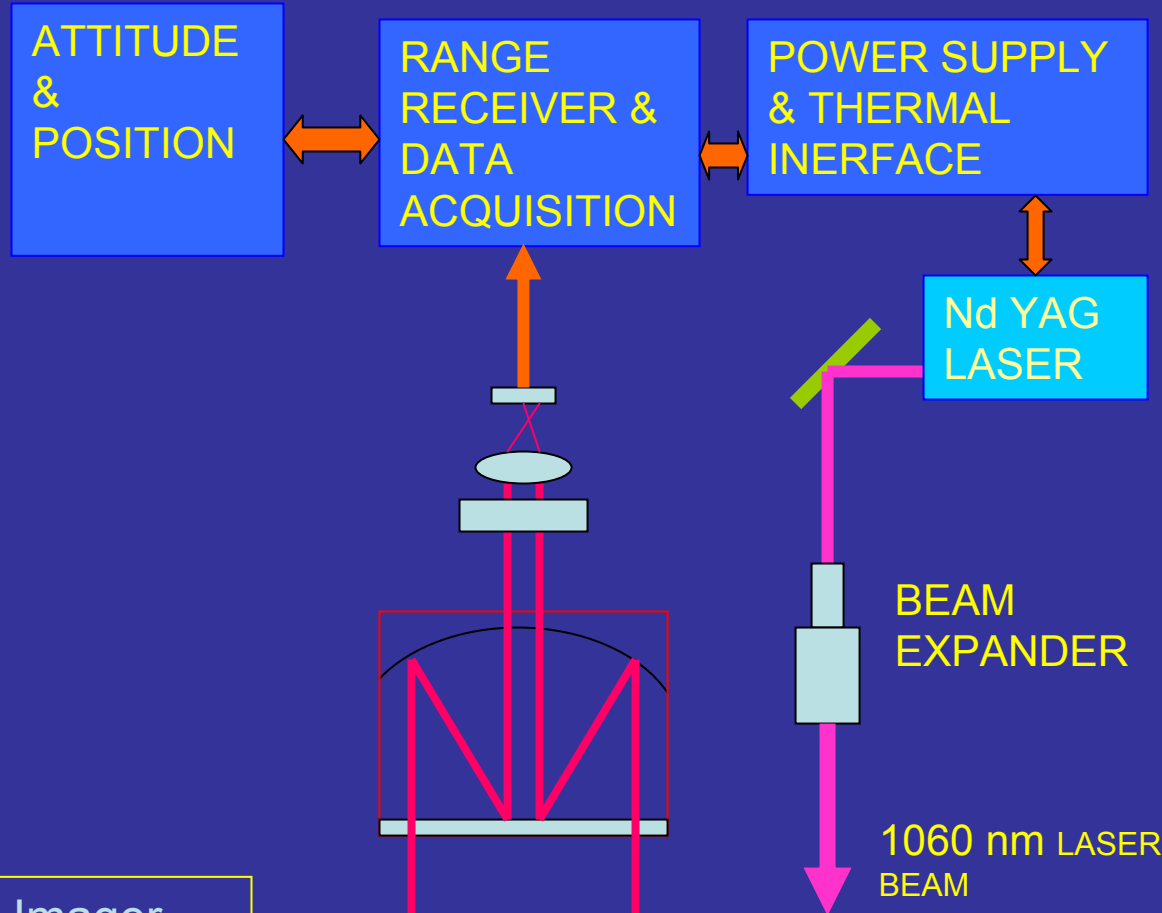
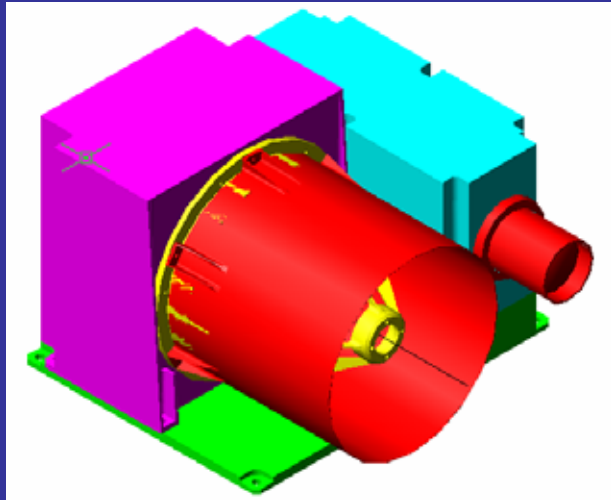
# Chandrayaan-1 Mission



## Lunar Laser Ranging Instrument

### Primary Objective:

**Determine Global Topographic Figure of Moon**



- Supplement TMC and HySI Imager
- Improved model of lunar gravity field

Developed at LEOS, ISRO

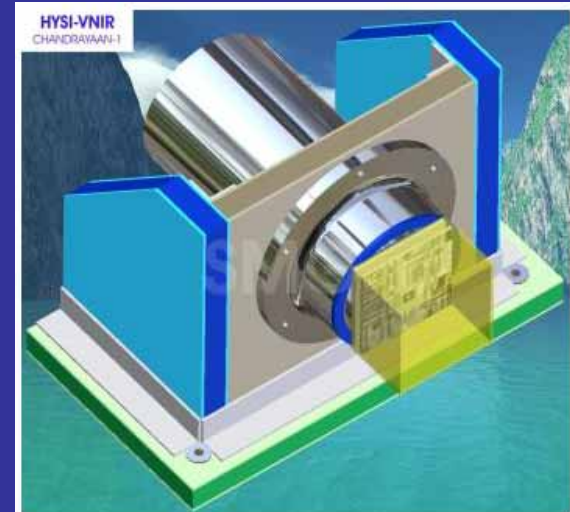
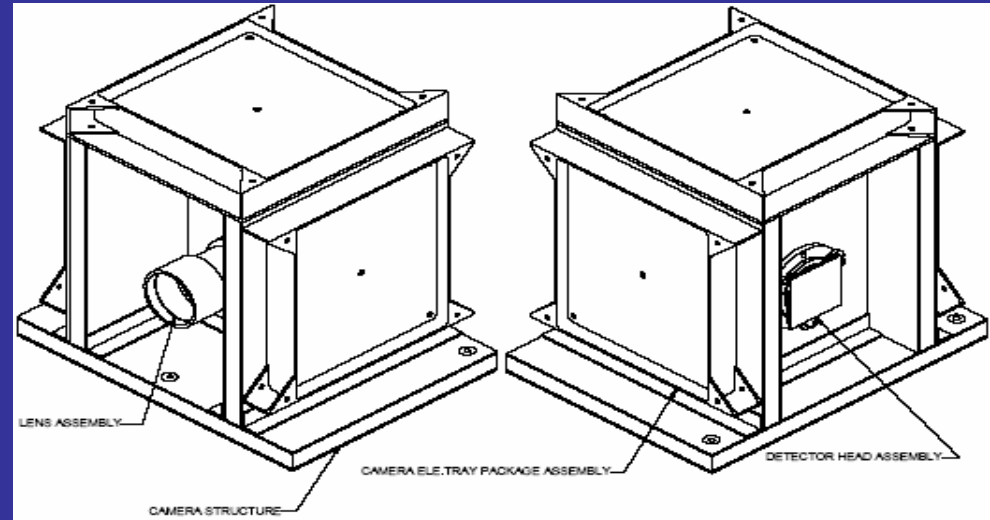
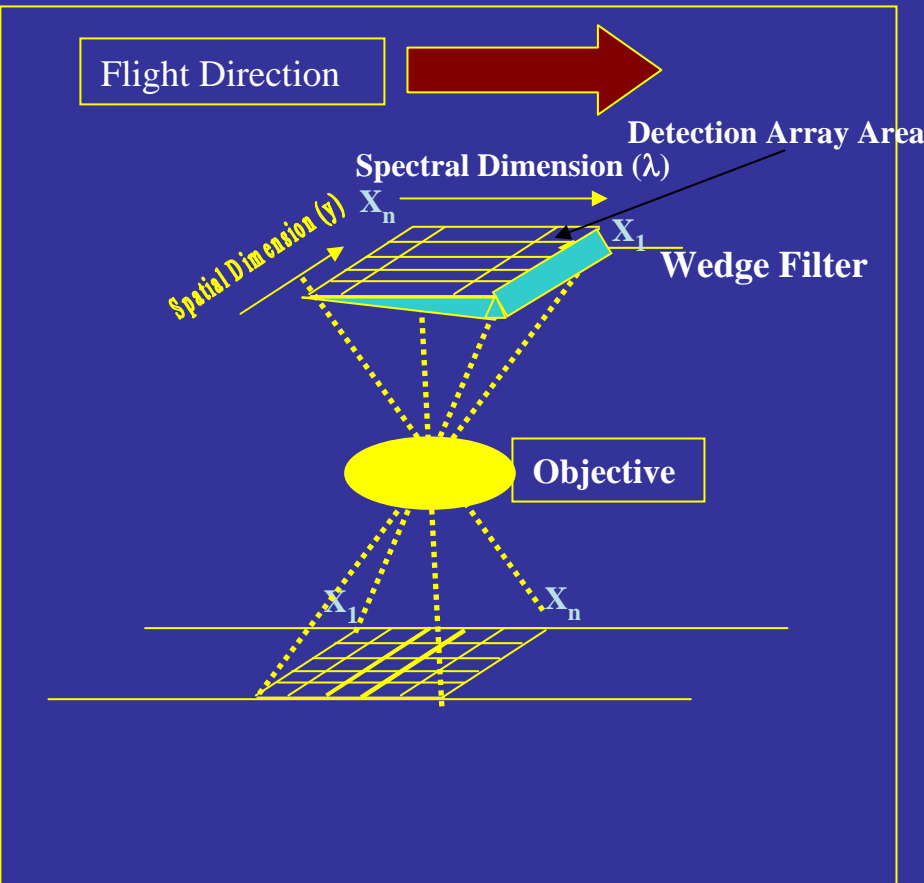




# Chandrayaan-1 Mission



## HySI concept & Instrument configuration

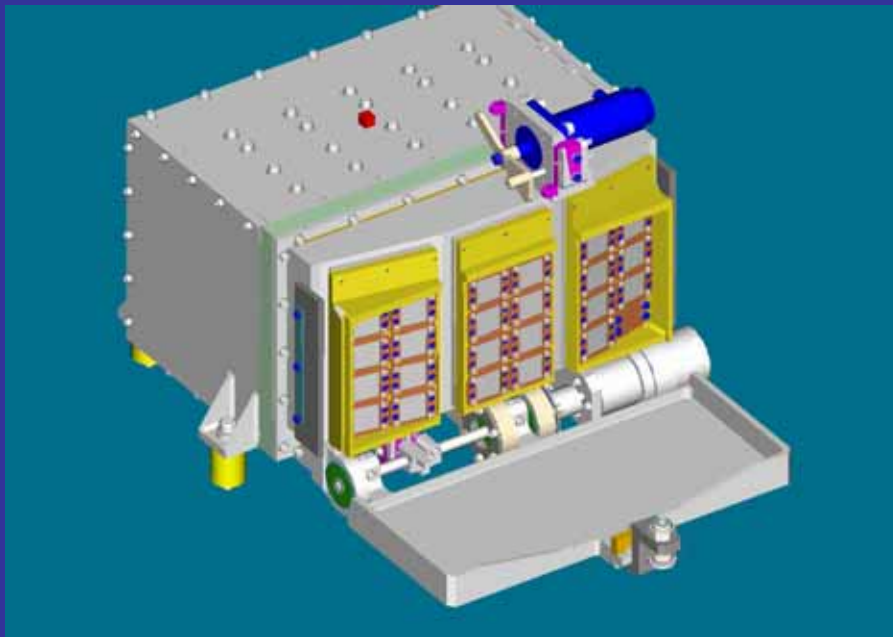


400-900 nm, 32 bands, resolution 80 m



## Compact Imaging X-ray Spectrometer (CIXS)

A modified version of D-CIXS, the Low Energy X-ray Spectrometer, flown on SMART-1



New generation Swept-charge X-ray Detector

*Chemical (Elemental) Mapping of Lunar Surface based on Solar X-ray induced fluorescence emission*

[Mg, Al, Si : Solar quiet time]

Ca, Ti, Fe : Solar Flare time]

**High resolution mapping of Fe and estimation of Mg#**

P. I. Prof. Manuel Grande,  
Rutherford Appleton Laboratory, UK  
Indian Co-I : Dr. P. Sreekumar, ISAC



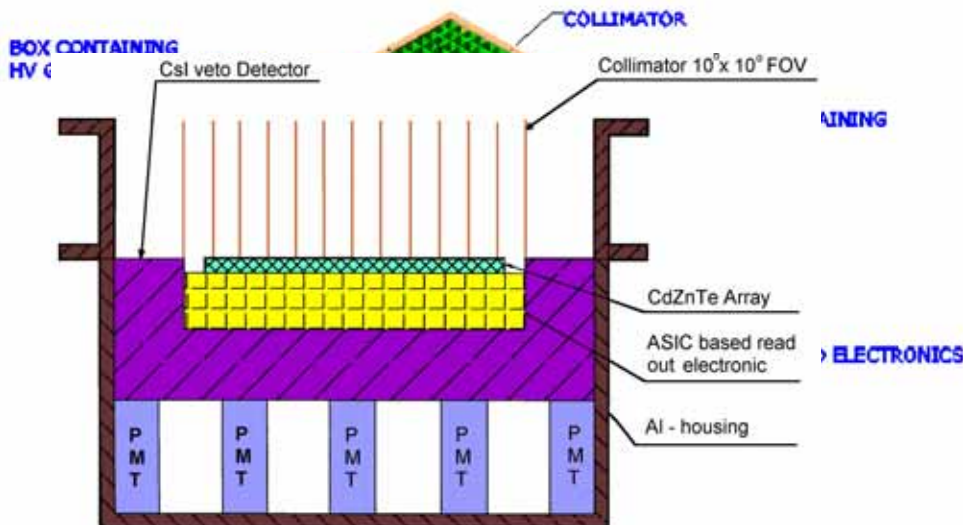
# Chandrayaan-1 Mission



## High Energy X- $\gamma$ Ray (HEX) PAYLOAD

- Volatile Transport on Moon through detection of 46.5 keV line from Pb-210
- Th and U map of Polar and U-Th-rich regions

### High Energy X- $\gamma$ Ray Spectrometer (10-200 KeV)



CdZnTe Array Detector with Collimator

### Basic Features

Detector: Cd-Zn-Te Array

Effective Area: 100 cm<sup>2</sup>

Energy Range: 20-250 keV

Energy Resolution:  
≤6% @ 60 keV

FOV: 40km x 40km

Active Anticoincidence  
Shielding: CsI(Tl)+PMT

Collimator: Tantalum

Weight: 15 kg

Power: 23 W

Developed by Physical Research Laboratory & ISRO Satellite Center

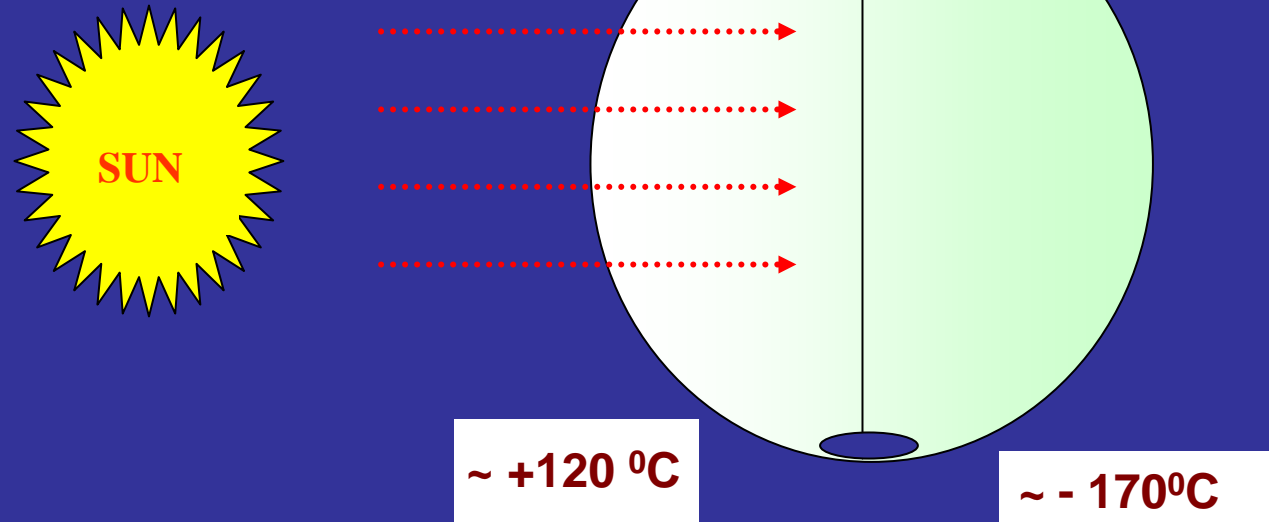


# High Energy X- $\gamma$ Ray (HEX)

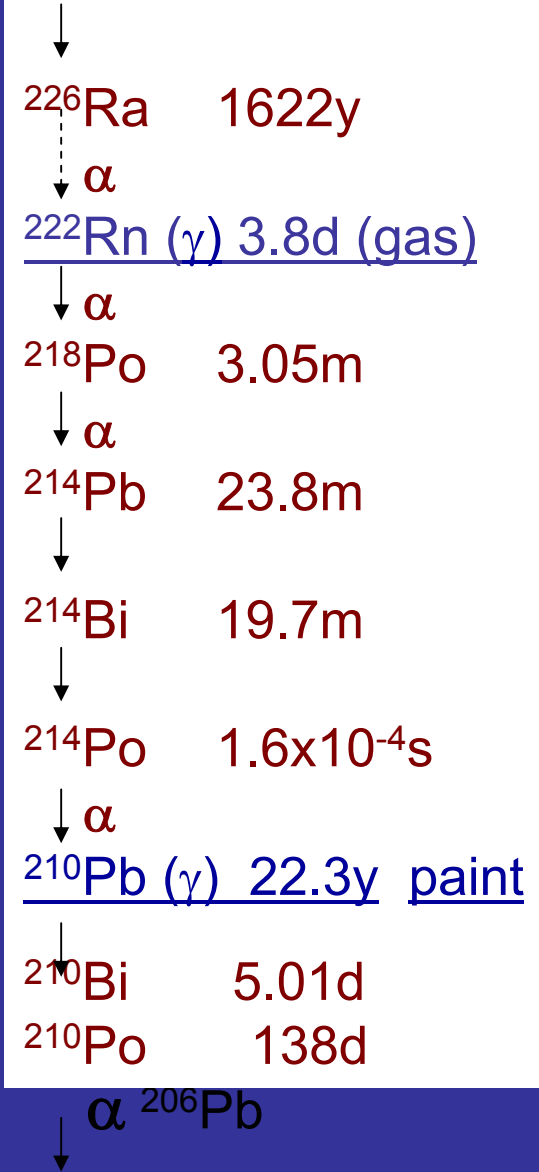


First attempt to detect low-energy (20-250 keV) gamma ray emission from a planetary surface

Mapping of U, Th and Pb-210



Source: U-238 in Moon



## Water on Moon:

Movement of volatile ( $^{222}\text{Rn}$ ) to permanently shadowed polar regions using  $^{210}\text{Pb}$  (46 keV) as tracer



# Chandrayaan-1 Mission



## Payload Configuration & Details

Payload	Configuration	Range	Resolution	Objective
Hyper Spectral Imager (HySI)	Wedge filter pixelated imager	0.4-0.9 $\mu\text{m}$	Spatial - 80m Spectral-15nm 32 channels	Mineralogical mapping
Terrain Mapping Camera (TMC)	Three stereo cameras with pixelated imagers	Panchromatic; 20 km swath	Spatial - 5m Vertical - 5m	High resolution Atlas of the whole moon
Laser ranging (LLRI)	Pulsed Nd-Yag laser	1064 nm	Vertical - 10 m or better	Topography & Gravity model
Compact X-ray Spectrometer (CIXS)	X-ray SCXD type detector 25 sq. cm area	0.5-10 keV	20 km	Elemental mapping Si, Al, Mg, Ca, Fe, Ti
High energy X-ray spectrometer (HEX)	CdZnTe detector 100 sq. cm. area	10-200 keV	40 km	$^{210}\text{Pb}$ , Volatile transport
Solar X-ray Monitor (SXM)	Si-Pin Diode 2 or 3 detectors viewing orthogonally	2-10 keV		Solar X-ray flux monitoring



# Chandrayaan-1 Mission



## *A new era of International Cooperation*

Based on science objectives and spacecraft resources, several proposals were accepted from international community; they complement/add to the Indian experiments to meet the basic science goals of the mission.

- I. IR spectrometers for mineral mapping (SIR-2 and MMM)*
- II. An experiment to detect neutral atoms (SARA)*
- III. An experiment to map the poles and search for water ice (mini-SAR)*
- IV. An experiment to monitor energetic particle environment (RADOM)*

*There will be Indian collaboration in analyzing data from all these experiments*

# Chandrayaan-1 Mission (AO Payloads)



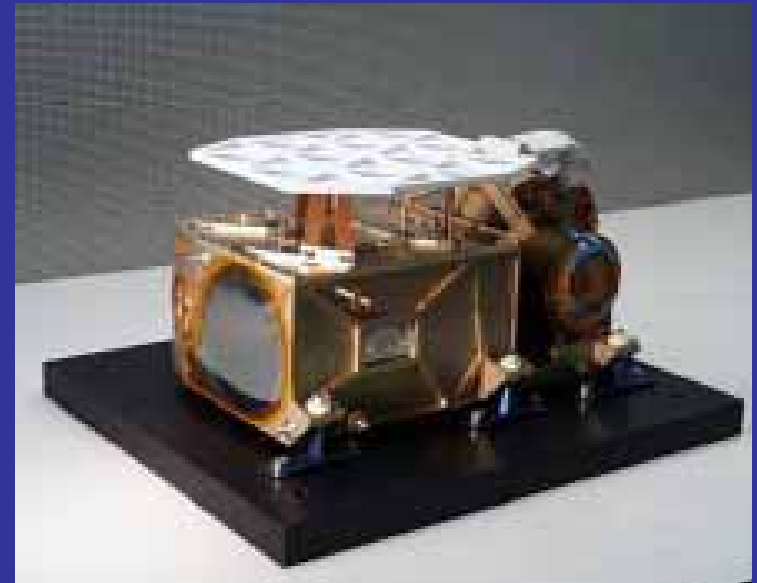
## **SIR-2: A Near-infrared Spectrometer**

P. I. Dr. Urs Mall,  
Max-Planck Institute for Solar System Research  
Katlenburg-Lindau, Germany

The first version of SIR  
orbited the Moon onboard  
ESA SMART-1 Mission

Spot spectrometer 1.11  
milliradian FOV (~110 m  
surface spot size from 100  
km orbit)

930-2400 nm spectral range  
with 6 nm spectral resolution

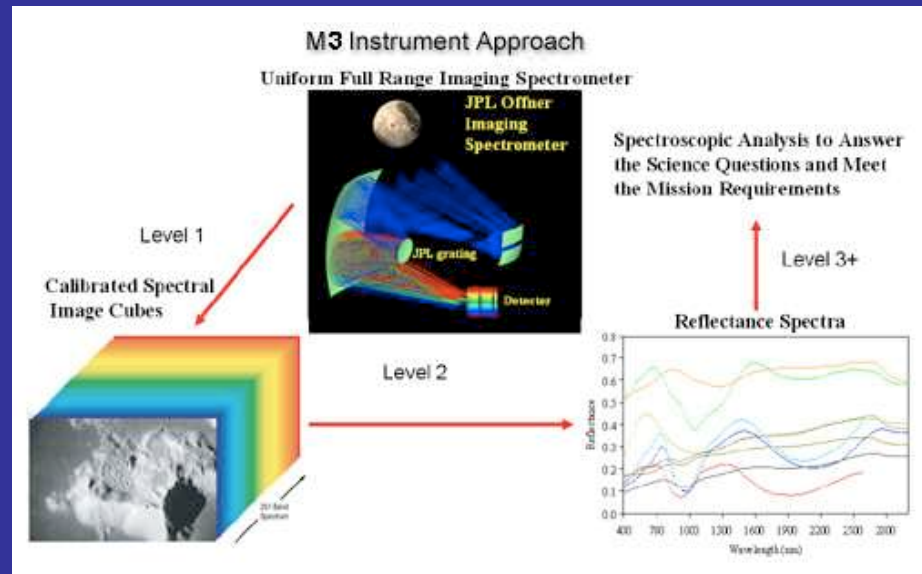


# Chandrayaan-1 Mission (AO Payloads)

## Moon Mineralogy Mapper

An Imager in VIS-NIR band

P. I. Dr. C. Pieters, Brown University, USA



700 to 3000 nm range  
Sampling : 10 nanometers  
Spatial resolution: 70 m/pixel  
[from 100 km orbit]  
Field of View: 40 km [from 100  
km orbit]  
Weight: ~7 kg  
Power average: ~13 W

*Mineral Mapping of Lunar Surface*

This Instrument, together with SIR-2, HySI, CIXS and SARA, provide detailed mapping of surface composition of the Moon



# Chandrayaan-1 Mission (AO Payloads)



## **SARA: Sub-KeV Atom Reflecting Analyzer**

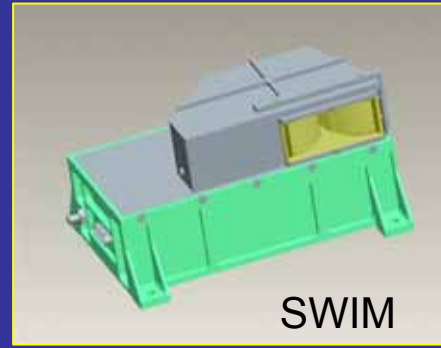
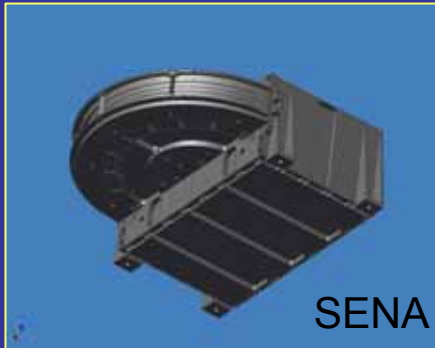
P. I. Dr Stas Barabash, Swedish Institute of Space Physics &  
Dr. A. Bharadwaj, Space Physics Laboratory, India

**Basic Science Objective :**

*Imaging of:*

*(i) Moon's surface composition*

*(ii) lunar surface magnetic anomalies*



## **Anticipated Highlights:**

- ▶ Surface composition of permanently shadowed areas and complement data from other payloads on surface composition
- ▶ Surface magnetic anomalies: magnitudes and plausible causes

# Chandrayaan-1 Mission (AO Payloads)



## **RADOM: Radiation characterization**

Bulgarian Academy of Sciences

***Basic Science Objective : Characterization of lunar radiation environment and efficacy of shielding effects by lunar materials***



### **Anticipated Highlights:**

Particle flux, deposited energy spectrum, accumulated absorbed dose rates in Lunar orbit

Estimate dose around Moon at different altitudes and latitudes

Evaluate the shielding characteristics (if any) of the near-Moon environment towards galactic and solar cosmic radiation and solar particle events

# Chandrayaan-1 Mission (AO Payloads)



## Mini-SAR: A Miniature Synthetic Aperture Radar

P. I. Dr. Paul D. Spudis: Applied Physics Laboratory,  
Johns Hopkins University

**Basic Science Objective:** Map terrain of lunar poles and determine surface scattering properties of deposits to search for ice in the permanently shadowed regions of the poles

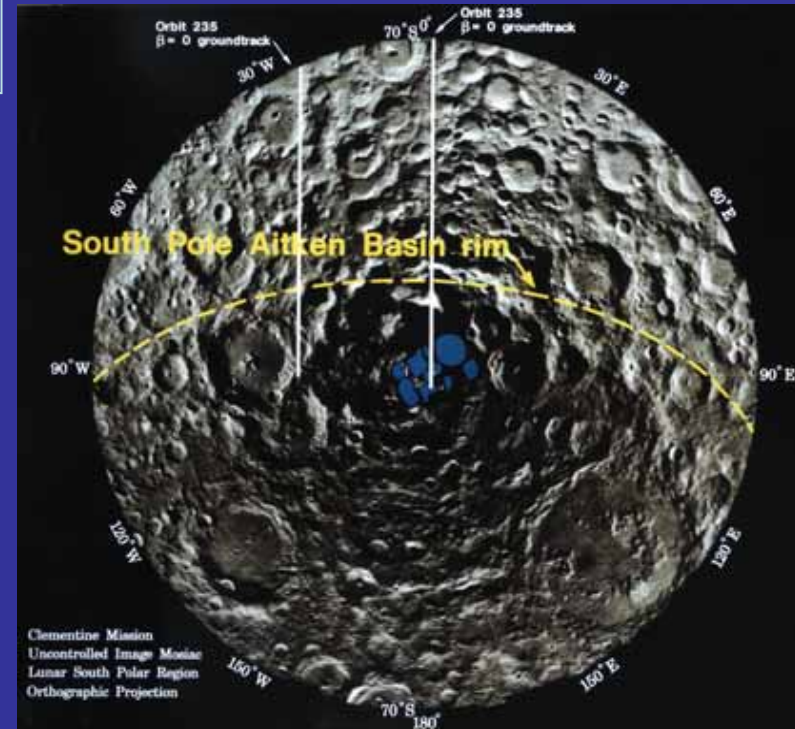
Uses unique hybrid architecture to reconstruct Stokes matrix scattering parameters

## Clementine Mission

Radar reflections mimicking volume-scattering behavior of ICE

## Prospector Mission

Neutron Spectrometer suggests presence of excess hydrogen (water?) in polar regions





# Chandrayaan-1 Mission



## Salient features of Chandrayaan-1 Mission

### *High resolution Global Chemical Mapping*

Elemental mapping of the moon (Mg, Al, Si, Ca, Ti and Fe) using both X-ray Fluorescence and UV-VIS-IR spectroscopy;

*First set of detail data for Fe (from XRF) and mineralogy (IR region)*

*Timing of the mission: Mid-level to high level of solar flare activity*

Prospector data for Fe from gamma-ray observation suffers from low spatial resolution (150 km) and Al interference

The Mg\* [Mg/(Mg+Fe)] is one of the MOST IMPORTANT diagnostic parameters to understand lunar bulk composition

*LEX, HySI, SIR-2, MMM (also SARA; shaded area)*



# Chandrayaan-1 Mission



## Salient features of Chandrayaan-1 Mission

*High resolution global topographic map of the Moon*

3D mapping with spatial resolution of 5 meter : **TMC and LLRI** - Study of small (10-100 m) features [e.g., crater distributions on Moon]

*First attempt to detect emission of low energy (<300keV) gamma-rays from a planetary surface (HEX) & detailed study of the permanently shadowed polar regions*

**Volatile Transport on Moon:**

*Enhanced Pb-210 emission in such region (HEX)*

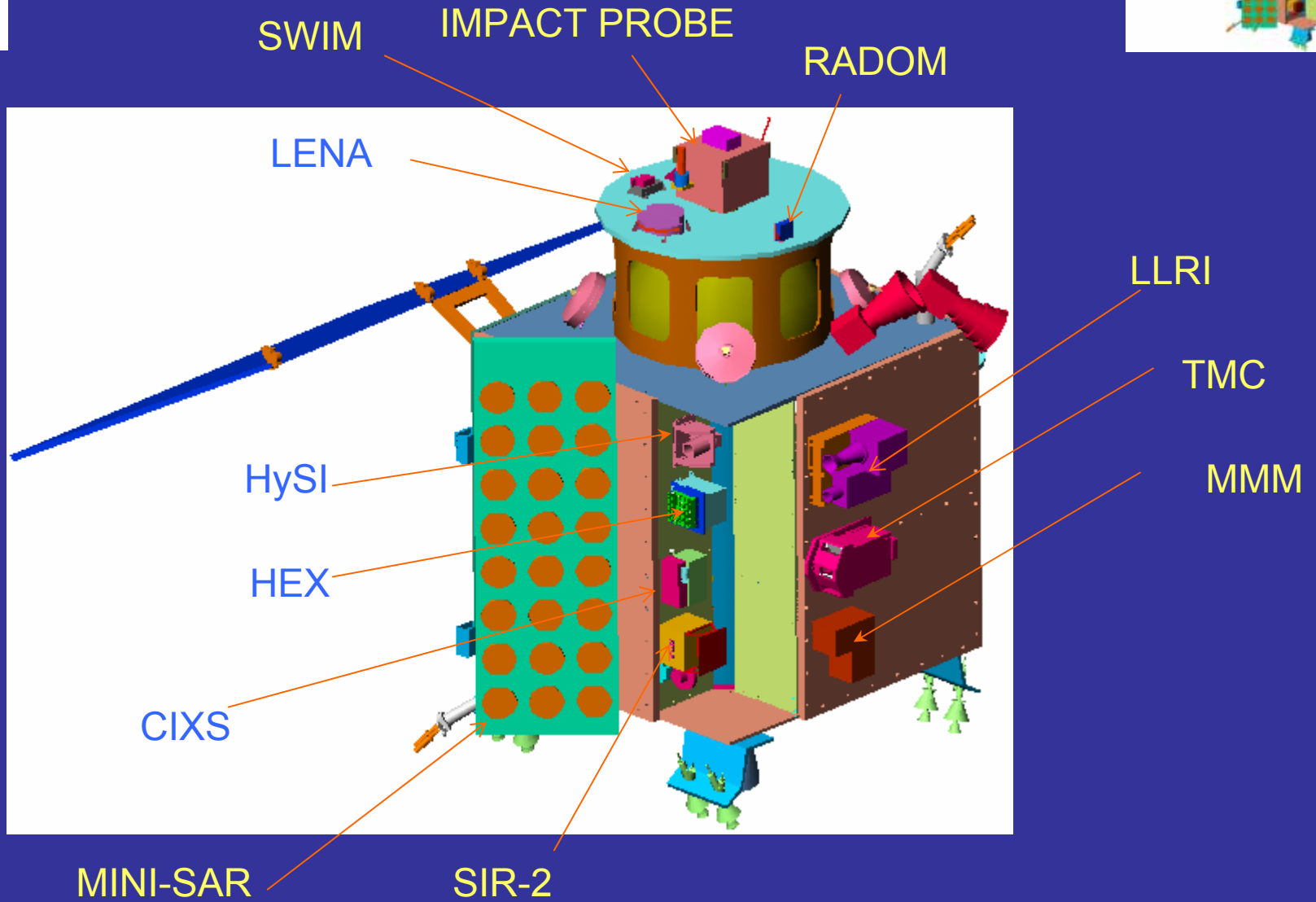
**Water on Moon in shadowed polar regions via volatile transport:**

*Detection, mapping, and study using mini-SAR*

**Magnetic anomaly and composition of shaded area (SARA)**

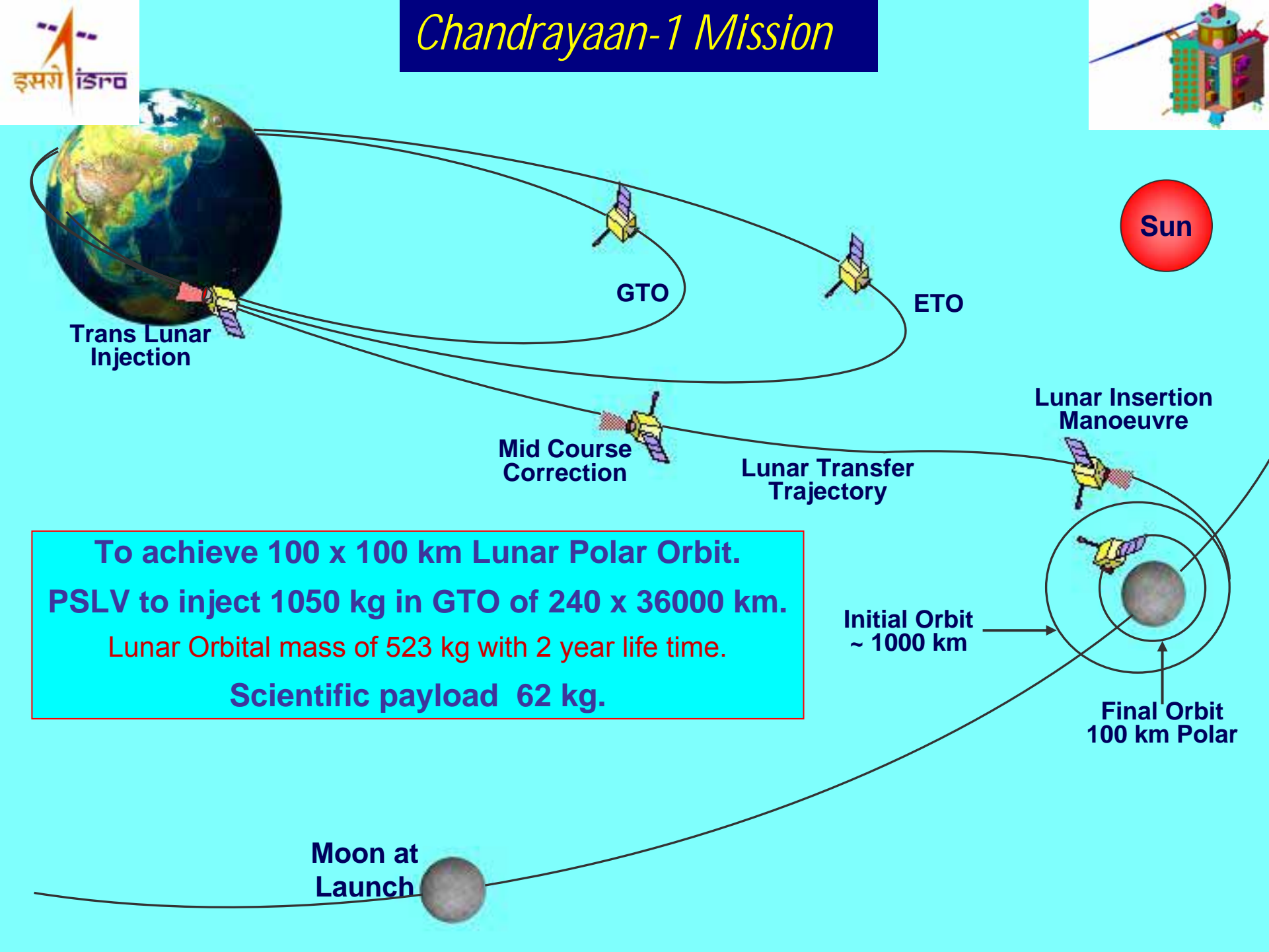


# Chandrayaan-1 Mission



## PAY LOAD ACCOMMODATION

# Chandrayaan-1 Mission



**To achieve 100 x 100 km Lunar Polar Orbit.**  
**PSLV to inject 1050 kg in GTO of 240 x 36000 km.**  
**Lunar Orbital mass of 523 kg with 2 year life time.**  
**Scientific payload 62 kg.**

Sun

GTO

ETO

Trans Lunar Injection

Mid Course Correction

Lunar Transfer Trajectory

Lunar Insertion Manoeuvre

Initial Orbit ~ 1000 km

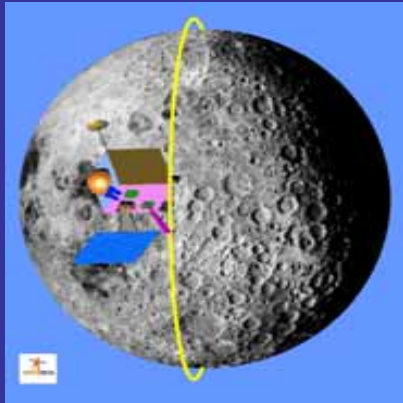
Final Orbit 100 km Polar

Moon at Launch



# Launch Of Chandrayaan - I by PSLV

(Target Date: April, 2008 )

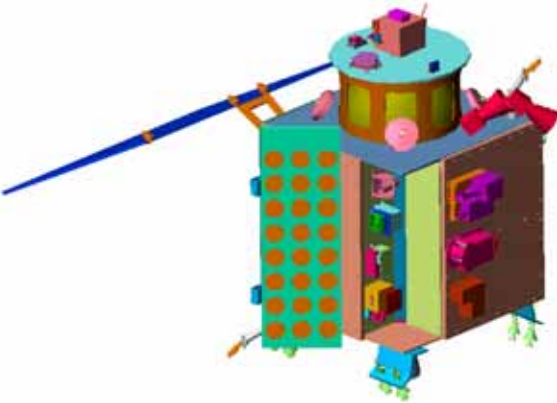


Altitude = 100km;

Inclination =  $90^\circ$

Period = 117.6 min.

Track Repeat = 28 days







## *Chandrayaan-1 Mission*



Instrument Calibration: Laboratory Calibration to cover ALL aspects. Additional Calibration during Tran lunar cruise as well as in lunar orbit, if needed

A National Science Data Center will be established to host data from Chandrayaan-1 and other Science Missions (e.g., ASTROSAT-1); Instrument Specific Software Packages for generation of Data Products and Analysis will be available at the Data Center

Spacecraft and Instrument data will be available to PI and Team on real time

Final Data Products and format for archiving have been formulated

Data sharing between various PI groups for enhancing Science return



# Chandrayaan 2



- Proposed launch 2010-2011
- Orbital spacecraft and surface rover
- Rover 30-100 kg; semi-hard or soft landing
- Solar powered; may be designed to hibernate during lunar night for a second two weeks of operation
- Instruments TBD, but will conduct compositional analysis of surface materials