ON THE POSSIBILITY OF FORMATION OF THE CHONDRITE DIAMONDS IN THE NON-EQUILIBRIUM PROCESS. A.V. Fisenko, L.F. Semjonova, V.F. Tatsy, G.V. Baryshnikova, and A.K. Lavrukhina. V.I. Vernadsky Institute of Geochemistry and Analytical Chemistry, USSR Academy of Sciences, Moscow, USSR

The diamonds of chondrites have presolar origin, but the process of their formation is discussed. We report the results of X-ray and differential thermal analyses (DTA) of the Efremovka chondrite diamond-rich fraction (DE-4). Diamonds in this fraction are not contamination origin since they contain the isotopically anomalous Xe-X/1/. Besides the fraction DE-4 diamonds of different genesis have also been analyzed: natural diamonds ND 3/2 (particles range 2-3 µm); synthetic diamonds SD 1/0, obtained under static conditions (particles range 0-1 µm); submicro-and micropowders of UDD and DDG diamonds synthesized by the detonation method from carbon of explosive in gaseous phase and graphite, respectively.

The sizes of the coherent scattering region (CSR) of the investigated diamonds were calculated (Table) accor-

ding to extendsion of lines on X-rayogram.

Table.	Parameters	of	the	diamond	samples

Parameters		Diamond samples						
		ND 3/2	SD 1/0	DDG	מכט	DE-4		
Size of CSR, A	Calculated values	_	860	110	3 5	30		
	References /2,3/	2000	n.d.x	7 5	40	n.d.		
Mean size	U	6100	1150	630	50	26 –3 0		

x not determined

Mean size of diamond particles calculated according to their specific surface /4/ are also given in the Table with the exception of DE-4. For latter we used the same size of diamonds as for Murchison chondrite /5/.

It is seen from Table that for Efremovka chondrite diamonds the sizes of CSR and the mean size of particles agree with each other. From this follow chondrite diamonds are monocrystalline. DTA-curves of diamonds air oxidation were recorded in range from 20°C to 950°C. Intensive oxidation of diamond particles was recorded in the form of sharp single peak at ~583°C. Not less than 90% of the diamonds was oxidized over the range from 350°C to 630°C. On the basis of the oxidation exotherm follow the sample DE-4 is homogeneous on grain size powder. The Efremovka

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chondrite diamonds agrees to a considerable degree with UDD in particle sizes, their monocrystalline and of the thermal oxidative stability. Based on this, we are proposed that the Efremovka chondrite diamonds could be formed from carbon in a gaseous phase at the fast non-equilibrium process as well as the UDD. The condensation of diamond particles may be initiated and enhanced in shock waves, e.g. in the star wind emitted from O,B-star in early stages of core He-burning. In this stage of the star evolution a gas phase had C/O>1 /6/. The isotopic data on carbon and nitrogen for the diamonds of chondrites correspond also to the nuclear process proceeding in the mentioned stage. The possibility of the formation of the diamond nucleations in the star atmosphere in shock waves was pointed out earlier /7/. However, the authors of this work are supposed diamond nucleus served as substrates for the further epitaxial growth of interstellar diamonds. From our data it follow that the fast non-equilibrium process has played a main role in the formation of the chondrite diamonds. By confirm of this one may indicate, e.g. that the diamonds of chondrites have a considerable amount of nitrogen (up to ~0.9 wt.%), pointing to the non-equilibrium process of nitrogen incorporation into diamonds /8/. It is may be proposed that the considerable amount of carbon atoms on the non-equilibrium formation of diamond particles is replaced by nitrogen that gives its large content in diamonds.

References. 1. Fisenko A.V. et al. Meteoritica. 1987,

V.46. P.58 (in Russ.).

2. Staver A.M. et al. Physics of Combustion and Explosion. 1984, N5. P.100 (in Russ.).

3. Drobyshev V.N. Physics of Combustion and Explosion.

1983. N5. P.158 (in Russ.).

4. Breusov O.N. et al. Superhard Materials. 1989. N1. P.25 (in Russ.).

5. Tang M., Anders E. Geoch. Cosm. Acta. 1988. V.52. P.1235. 6. Prantzos N. Preprint. CEN-Sacla-Service ol Astrophysi-

que. 1985. 6P.

7. Saslaw W.C., Gaustad J.E. Nature. 1969. V.221. N5176.

P.160.

8. Russell S.S., Pillinger C.T. L.P.S. XXI. 1990. P.1051.