

# DIFFRACTION MODEL OF THE NEGATIVE POLARIZATION OF LIGHT SCATTERING BY ATMOSPHERELESS COSMIC BODIES

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Existing models of the negative polarization (NP) based on the geometrical optics approach [1,2] do not explain increasing of NP with decreasing of particle size up to wavelength of incident light for glass samples [3]. This fact and a lot of other ones are explained in a frame of a new NP model [3,4]. The model takes into account polarization of single and double scattering beams. Moreover, interference of double scattered rays interacting with the same scatters is taken into account, e.g. (see Fig.1): the first path is (source → particle # 1 → particle # 2 → observer), the second path is (source → particle # 2 → particle # 1 → observer). For a phase angle ( $\alpha$ ) dependence of polarization ( $P$ ) a following approximate formula is obtained:

$$P = \frac{G}{18} (1+2\sqrt{1-w})^2 \left[ \sin^2 \alpha + 2 \frac{\mu w (1-\sqrt{1+\beta^2 \sin^2 \alpha})^2}{\beta^2 \sin^2 \alpha \sqrt{1+\beta^2 \sin^2 \alpha} \ln(1-\mu)} \right]$$

$$\beta = 8\pi r / 3 \ln(1-\mu);$$

$$w = 24A (32A+15+4\sqrt{24A+9}) / (9+32A)^2.$$

Used parameters:

$A$  - a visible surface albedo ( $w$  - a single scattering albedo)

$r$  - an average particle radius divided by wavelength;

$\mu$  - a porosity (a number of particles in the particle volume);

$G$  - a polarimetric ability of scatterers ( $0 < G \leq 1$ );

Two theoretical  $P(\alpha)$  dependencies are plotted in Fig.2, where they are fitted to observed lunar data [5] (crosses) and to laboratory measurements of MgO sample [6] (dots). Curves 1 and 2 (for the Moon and for MgO) are characterised by following numerical values of the model parameters:

$A = 12\%$ ,  $\mu = 0.45$ ,  $r = 0.5$ ,  $G = 0.2$  and

$A = 97\%$ ,  $\mu = 0.07$ ,  $r = 0.4$ ,  $G = 0.2$  respectively.

Relation of the polarimetric parameters  $P_{min}$ ,

$\alpha_{inv}$ ,  $\alpha_{inv}/2\alpha_{min}$ ,  $h$  to the model parameters

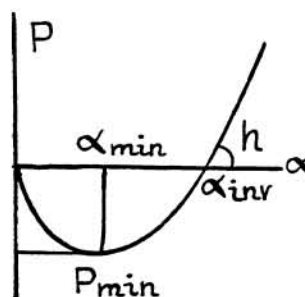
$G$ ,  $\mu$ ,  $A$ ,  $r$  are shown in Fig.3. The main numerical means of the model parameters are:  $G = 0.2$ ,

$\mu = 0.4$ ,  $A = 0.12$ ,  $r = 0.3$ .

**Comments.** Decreasing  $\alpha_{inv}$  with increasing  $r$  attracts attention. It is quantitatively close to laboratory data [6]. A dependence between  $P_{min}$  and  $A$  has two branches, that corresponds to experimental data [7]. The same situation is for  $h$  -  $A$  dependence in the low albedo case.

## References.

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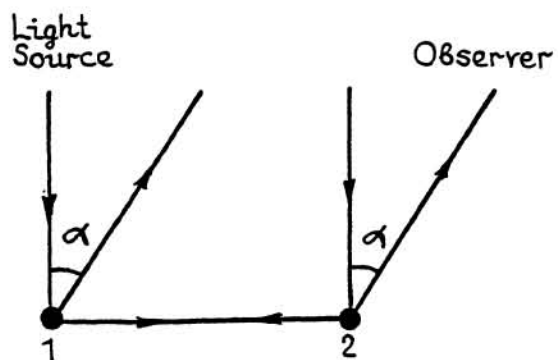
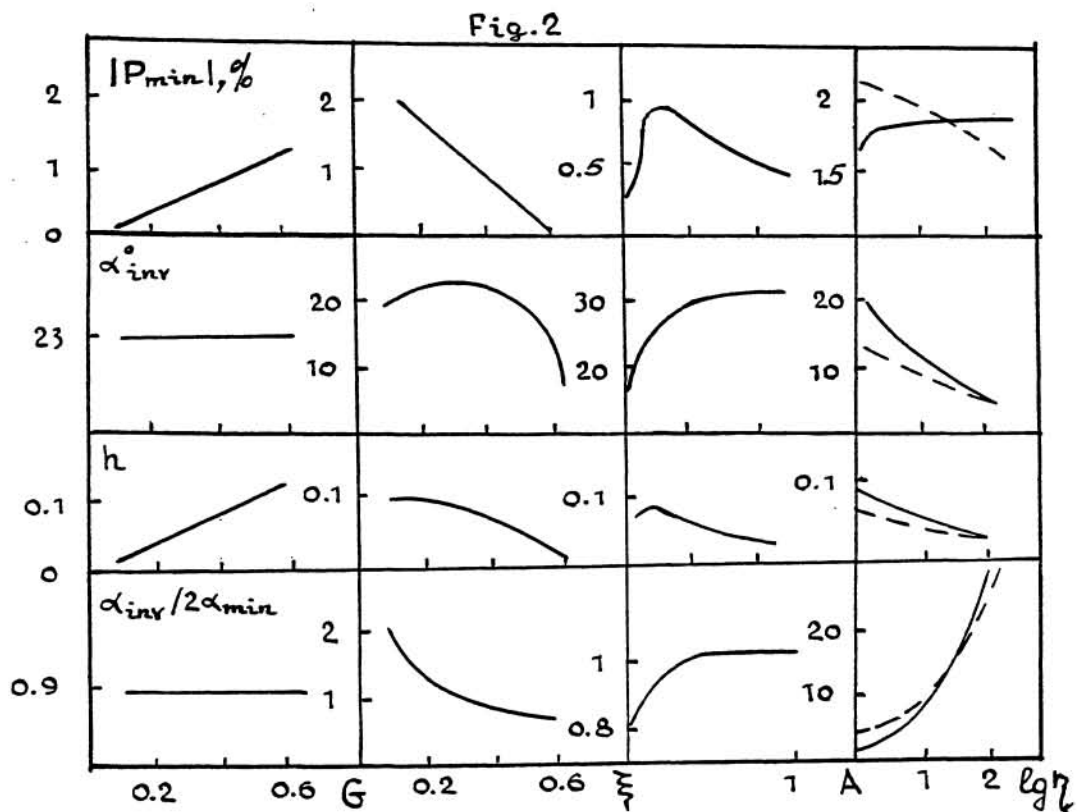


Fig. 1