

Magnetic properties of primitive achondritic meteorites GRA 06128 and GRA 06129. J. Geissman¹ and C.K. Shearer². ¹Dept. of Earth and Planetary Sciences, Univ. of NM, Albuquerque, NM 87131 (jgeiss@unm.edu), ² Institute of Meteoritics and Dept. of Earth and Planetary Sciences, Univ. of NM, Albuquerque, NM 87131.

Introduction: Newly discovered achondritic meteorites GRA 06128 and 06129 (GRA) are somewhat unusual meteorites that contain a high percentage of sodic feldspar ($\sim \text{An}_{15}$) with Fe-rich silicates (olivine, high-Ca and low-Ca pyroxene), two phosphates, oxides, sulfides, and Fe,Ni-metal [1,2,3,4,5]. ²⁶Al-²⁶Mg isotope measurements and a variety of chemical-mineralogical characteristics indicate that these meteorites represent products of very early planetesimal melting (4565.9 ± 0.3 Ma) of an unsampled geochemical reservoir from an asteroid that has characteristics similar to the brachinite parent body [1]. Superimposed upon the initial crystallization event are several other events that represent subsolidus reequilibration-metamorphism and fluid interaction. Our intent in this abstract is to examine the magnetic properties of GRA 06128 to gain insights into its evolving mineral assemblage and the environment under which GRA obtained its magnetism.

Analytical approach: Magnetic properties of GRA were measured at the University of New Mexico using a DC-SQUID based superconducting rock magnetometer (2G Enterprises), housed in a low magnetic induction space. The specimen mass used for all magnetic measurements was about 120 mg. First, the specimen was subjected to repeated measurements of the natural remanent magnetization (NRM) to test for viscous remanent magnetization (VRM) acquisition. Alternating field (AF) demagnetization was utilized to a peak field of 120 mT, to examine the character of the remanence carried by relatively low coercivity phases. An anhysteretic remanent magnetization (ARM) was acquired in a DC field of 0.1 mT and a peak AF of 98 mT. Acquisition of an isothermal remanent magnetization (IRM) to near saturation (SIRM) utilized a pulse DC impulse magnet to about 3.0 T. Backfield pulse DC demagnetization of near saturation IRM was conducted until IRM changed sign. AF demagnetization of IRM imparted in a pulse DC field of 98 mT was carried out to 100 mT. The specimen was subsequently mounted in pure alumina cement to make a thermally resistant specimen of cubic shape, and subjected to progressive thermal demagnetization to over 670°C of three-component IRM, acquired in orthogonal fields of 3.0 T, 0.3 T, and 0.03 T, following the method of Lowrie [6].

Results: The natural remanent magnetization (NRM) moment of the UNM specimen is 1.94×10^{-6} emu (intensity of 16.08×10^{-6} Am²/kg) (Figure 1). Progressive alternating field demagnetization of the GRA specimen measured indicates that the meteorite NRM

may contain at least two components. Alternating-field (AF) demagnetization to 95 mT randomizes less than 50 percent of the NRM; at higher peak fields demagnetization behavior is more erratic. A possible higher coercivity component only partially responds to AF demagnetization and thermal demagnetization of NRM. Intensity of anhysteretic remanent magnetization (ARM) is about 81.0×10^{-6} Am²/kg and over a factor of 6 higher than NRM. The intensity of ARM as well as AF demagnetization response to ARM indicates that a considerable fraction of moderate coercivity phases is present that does not contribute to the NRM (Figure 2). Acquisition of an isothermal remanent magnetization shows nearly complete saturation at about 3.0T, and backfield DC demagnetization of near-saturation IRM yields an estimate of coercivity of remanence of about 280 mT. The near saturation IRM intensity is about 2970×10^{-6} Am²/kg, and yields an ARM/SIRM ratio of approximately 0.027. Comparison of AF demagnetization of ARM and near-saturation IRM suggests the presence of fine, single domain to pseudo single domain ferrimagnetic particles in the specimen (Figures 2 and 3). In addition, the very low crossover value ($R < 0.1$) for IRM acquisition and AF demagnetization of near saturation IRM at least suggests the possibility of strongly interacting magnetic phases [7,8]. Thermal demagnetization of three component IRM suggests that the magnetization is carried by high-coercivity phases with laboratory unblocking temperature intervals that center on $\sim 350^\circ\text{C}$, $\sim 570^\circ\text{C}$, and $\sim 675^\circ\text{C}$ that are likely pentlandite, a Fe-Ni metal, and hematite, respectively (Figure 4). Thermal demagnetization of both the high (>3.0 T) and intermediate ($0.3 < H < 3.0\text{T}$) coercivity components also suggests the possibility of the thermal destruction of iron hydroxide phases between about 100 and 200°C . The intermediate and lowest field (coercivity) ($< 0.03\text{T}$) components show a \sim complete unblocking below about 570°C , and this response is consistent with the observed occurrence of Fe,Ni metal. Notably, the response to thermal demagnetization of three component IRM does not suggest the generation of abundant new magnetic phases that could be reflected in an increase in the intensity of one of the components during progressive heating.

Discussion: Based on the available rock magnetic and demagnetization data for GRA 06129, the origin of the remanence present in this sample is difficult to determine with confidence. The available data are not inconsistent with the remanence being a weak field thermoremanent magnetization (TRM) blocked over a range of temperatures below about 680°C . Progressive

thermal demagnetization of NRM, on a separate sample, is required to more adequately evaluate this hypothesis. Alternatively, the remanence could be an elevated temperature chemical remanent magnetization (TCRM). The paragenesis of the remanence carrying magnetic phases is critical in assessing this possibility.

