

THE LUNAR PHOTOELECTRON SHEATH: A CHANGE IN TRAPPING EFFICIENCY DURING A SOLAR STORM

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Abstract: On the lunar dayside, photoelectron emission is quasi-constantly occurring and this electron emission acts to charge the dayside lunar surface a few volts positive. In arriving at an equilibrium surface potential, the surface will charge to balance the two primary currents: the outgoing photoelectron flux, J_p , against the incoming solar wind electron thermal flux, J_e . In nominal solar wind conditions, $J_p > J_e$ and the surface charges positive, trapping most (all but 6%) of the photoelectrons in the first few Debye lengths from the surface. We demonstrate via particle-in-cell plasma simulations that the morphology of the lunar photoelectric sheath during a coronal mass ejection (CME) passage by the Moon is very different than in the case of the nominal solar wind. Specifically, during the passage of the plasma associated with a CME, the solar wind electron thermal flux will greatly increase (by a factor of 5-10), making $J_e > J_p$. We find that during these times, there is substantially less near-surface electrostatic trapping of the photoelectrons. The photoelectron population has almost direct access to upstream regions with photoelectron escape in excess of 50%.

References. [1] Poppe, A., and M. Horányi (2010), *J. Geophys. Res.*, 115, A08106.

