

ON THE MODELS OF COMETARY'S ORBITS EVOLUTION. D. G. Modulina¹ and E. N. Tikhomirova^{2, 1}
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Introduction: The interest to integrated problems of celestial mechanics increases now [1]-[5]. In [1] the model of comets motion (with a zero mass m) like Kreutz ones in the gravitational fields of the Sun with the mass m_s and planets with the masses m_{pi} is considered. We estimated the interval of comet motion time from a perihelion to an aphelion equal $t=255.5288$ of years - for the comet C/1843 D1 (the Great Comet of 1843). This size corresponds (closely) to a half of the motion period of the considered comet ($T=515.2126$ of years). The perturbations from Jupiter are taken into account.

The aim of this work is forecasting of comets appearance near the Earth and the estimation of their danger degree.

Fundamental Equations: Let's consider movement of a variable mass point at the isotropic radiation law in the Newtonian gravitation field of n – motionless concentric circles.

The equation of point motion is

$$m \frac{dv}{dt} = -v \cdot \frac{dm}{dt} + F \quad (1)$$

The mass changes under the law

$$m(t) = m_0 \cdot (1 - \alpha \cdot t) \quad (2)$$

$$\frac{d^2 z}{dt^2} + \frac{dz}{dt} \cdot \frac{1}{m} \cdot \frac{dm}{dt} - \sum_{i=1}^n \frac{m_i \cdot z(t)}{(R_i^2 + z(t)^2)^{3/2}} = 0 \quad (3)$$

The equations (1)-(3) are reduced to the following quadrature

$$U(t) = \frac{dz}{dt}$$

$$\frac{dU}{dt} = U(t) \cdot \left(\frac{\alpha}{1 - \alpha \cdot t} \right) + \sum_{i=1}^3 \frac{G \cdot m_i \cdot z(t)}{(R_i^2 + z(t)^2)^{3/2}}$$

As a result of its decision (Runge-Kutta method of the 4th order), we received the following relation graphs:

Fig.1 A comet movement with an increasing mass

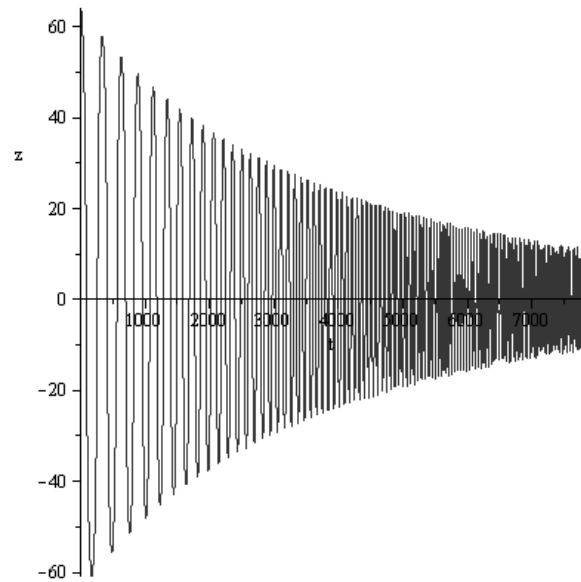


Fig.2 A comet movement with a decreasing mass

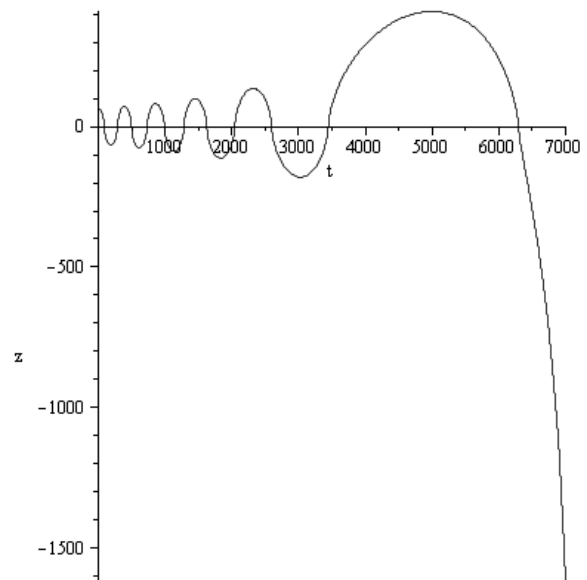
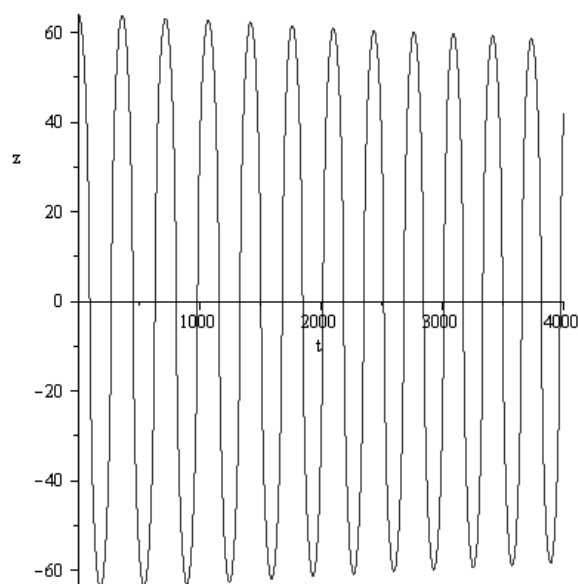


Fig.3 A comet movement with a constant mass

**Conclusion:**

1. The model allows to estimate the period of movement of comets taking into account perturbations (from quadratures).
2. From observation and calculations: at the Saturn and the Jupiter conjunction the center of mass of system the Sun-the Jupiter-the Saturn is always behind the Sun (the natural protection of Earth against the comets).
3. At mass loss a comet is jumped out of the Solar system (according to the model).
4. At absorption mass a comet comes nearer to the center of the Solar system (asymptotically).

References: [1]. Modulina D. G. (2012) *Astronautics and society: the problems and the decisions*, Yaroslavl, P. 48-51. [2]. Kondrat'iev B. P. (2012) *Astronomical Vestnik RAS*, V 46, № 5, P. 380-391. [3]. Vashkov`ayk M. A., Vashkov`ayk S. A. (2012) *Astronomical Vestnik RAS*, V. 46. №1, P.72-80. [4]. Perov N. I. (2010) *LAP LAMBERT Academic Publishing GmbH & Co. KG. Saarbrücken*, P. 300. [5]. Perov, N. I. et al. (2011) *Yaroslavl: YSPU*, 208 p.