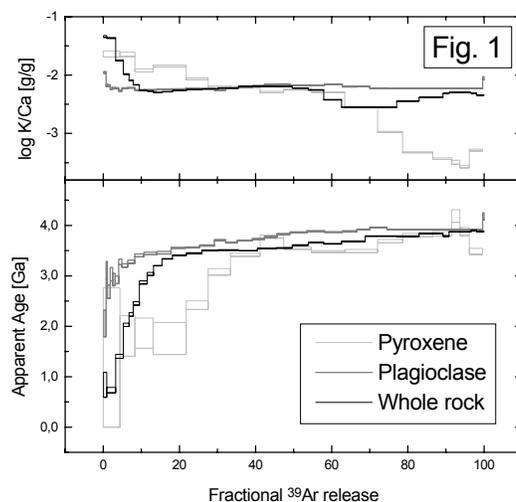


**$^{40}\text{Ar}$ - $^{39}\text{Ar}$  AGES OF MINERAL SEPARATES OF EUCRITES DHO300 AND DHO007.** E. V. Korotchantseva<sup>1,2</sup>, A. I. Bouikine<sup>1,2</sup>, M. A. Nazarov<sup>1</sup> and M. Trierloff<sup>2</sup>, <sup>1</sup>Vernadsky Institute of Geochemistry and Analytical Chemistry, Moscow 119991, Russia, <sup>2</sup>Mineralogisches Institut, Universität Heidelberg, D-69120 Heidelberg, Germany, trierloff@min.uni-heidelberg.de.

**Introduction:** We report new  $^{40}\text{Ar}$ - $^{39}\text{Ar}$  measurements on two brecciated eucrites recently discovered in the Oman desert: Dho007 and Dho300. Dho007 was classified as cumulate eucrite [1], though it has very high bulk Ni and Co contents in comparison to other eucrites [2]. Dho300 belongs to the noncumulate type of eucrites [3]. To better separate and quantify the various Ar components, high-resolution  $^{40}\text{Ar}$ - $^{39}\text{Ar}$  stepwise heating analyses were carried out on whole rock samples, as well as handpicked mineral separates (plagioclase, pyroxene and vein glass).

**Results and discussion:** All samples display distinct release peaks of trapped argon at 400°C and 700°C. This trapped component is of atmospheric composition, as revealed by three isotope correlation diagrams. Correction of the low temperature extractions results in less complex age spectra, as shown in Fig. 1 for Dho300.  $^{40}\text{Ar}$ - $^{39}\text{Ar}$  age spectra of Dho300 and Dho007 are similar to other typical eucrite age spectra and are characterized by stepwise increasing apparent ages. Judging from the release temperature of K-derived  $^{39}\text{Ar}$  and the K/Ca spectra, the main K carrier in all samples is plagioclase, but of presumably differing grain size: While for the plagioclase separate, preferably large grains were picked, in the pyroxene separates there remained some % of small plagioclase grains that could not be successfully removed. However, these dominate the K budget and, hence, the age and release spectrum. Release of K-derived  $^{39}\text{Ar}$  that correlates with Ca-derived  $^{37}\text{Ar}$  release from the pyroxene lattice at high temperatures of >1200°C was hardly observed. The Dho300 age spectra (Fig. 1) can be explained by a recent thermal event that induced varying degrees of secondary  $^{40}\text{Ar}$  loss (weakest for large plagioclase grains). Apparent ages in the high temperature extractions indicate that the last total reset occurred ~3.9 Ga ago (Fig.1), coeval with the intense cratering period on the moon [4].



**References:** [1] Afanasiev S.V. et al. (2000) *MAPS*, 35, A19. [2] Yamaguchi A. et al. (2003) *LPSC XXXIV*, 1377. [3] Meteoritical Bulletin (2001) *MAPS*, 36, A293-A322. [4] Kunz J. et al. (1995) *Planet. Space Sci.*, 43, 527-543.