

**GROWTH OF FERROUS OLIVINE IN THE OXIDIZED CV CHONDRITES DURING FLUID-ASSISTED THERMAL METAMORPHISM.** A. N. Krot<sup>1</sup>, M. I. Petaev<sup>2</sup>, and P. A. Bland<sup>3</sup>  
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The CV carbonaceous chondrites are subdivided into the oxidized [Allende-like (CV<sub>oxA</sub>) and Bali-like (CV<sub>oxB</sub>)] and reduced subgroups [1], which reflect complex alteration histories experienced by CVs [2]. Although recent mineralogical [3-5] and isotopic [6-8] studies show that CV alteration occurred in the presence of aqueous solutions and lasted for at least several Ma, the origin of secondary ferrous olivine (Fa<sub>40-60</sub>) remains controversial. Two types of models have previously been proposed: (i) gas-solid condensation from an oxidized nebular gas [9] and (ii) dehydration of phyllosilicates during thermal metamorphism [10, 11]. Here we argue that ferrous olivine in the oxidized CVs formed during fluid-assisted metamorphism by several mechanisms (1-4) described below:

(1) Replacement of opaque nodules in CV<sub>oxB</sub> chondrules by fayalite (Fa<sub>>90</sub>) [2, 12] is a two-stage process involving oxidation of metal to magnetite:  $3\text{Fe}_{(s)} + 4\text{H}_2\text{O}_{(l,g)} = \text{Fe}_3\text{O}_{4(s)} + 4\text{H}_{2(g)}$ , and replacement of magnetite by fayalite:  $2\text{Fe}_3\text{O}_{4(s)} + 3\text{SiO}_{2(aq)} + 2\text{H}_{2(g)} = 3\text{Fe}_2\text{SiO}_{4(s)} + 2\text{H}_2\text{O}_{(l,g)}$ . In MET00430, which is intermediate between CV<sub>oxA</sub> and CV<sub>oxB</sub>, the opaque nodules are replaced by ferrous olivine (Fa<sub>60-80</sub>), whereas those in Allende are replaced by less Fe-rich olivine (Fa<sub>40-50</sub>). This probably reflects higher alteration temperatures experienced by the CV<sub>oxA</sub> compared to the CV<sub>oxB</sub> [3].

(2) Direct precipitation of ferrous olivine from a supersaturated fluid in pore space:  $\text{Fe}^{2+}_{(aq)} + \text{Mg}^{2+}_{(aq)} + \text{SiO}_{2(aq)} = (\text{Fe,Mg})_2\text{SiO}_{4(s)}$ , is indicated by the presence of euhedral ferrous olivines overgrowing Mg-rich olivines and pyroxenes in chondrules and AOAs in oxidized CVs, and by the <sup>16</sup>O-poor compositions of the ferrous olivines in Allende AOA [13]. Large compositional variations of the ferrous olivines (Fa<sub>50-80</sub>) within individual AOAs in Kaba and complex Fe-Mg zoning (Fa<sub>50-80</sub>) of individual ferrous olivines in MET00430 probably reflect fluctuations of fluid compositions on a local scale.

(3) Replacement of low-Ca pyroxene by ferrous olivine is commonly observed in the CV<sub>oxA</sub> and CV dark inclusions (DIs) [11, 14]. The presence of phyllosilicates coexisting with these olivines [14, 15] suggests that this process occurred in the presence of aqueous solutions:  $\text{Fe}^{2+}_{(aq)} + \text{MgSiO}_{3(s)} + \text{H}_2\text{O}_{(l,g)} = (\text{Fe,Mg})_2\text{SiO}_{4(s)} + \text{H}_{2(g)}$ .

(4) Formation of ferrous olivine by dehydration of phyllosilicates is inferred from fibrous textures of some of the ferrous olivines in the Allende DIs [10], and from the presence of inclusions of pentlandite and poorly-graphitized carbon in the CV<sub>oxA</sub> matrix olivines [16]. However, the Mg-rich compositions of phyllosilicates in CVs, the rarity of fibrous olivines, and the lack of O-isotopic evidence for extensive hydration-dehydration of CVs [17] suggest that a direct substitution of olivine for phyllosilicates has played a minor role in the origin of ferrous olivine. Growth of ferrous olivine in the presence of fluid released during dehydration of CVs seems more likely.

**References:** [1] Weisberg et al. (1997) *MAPS* 32, A138; [2] Krot et al. (1998) *MAPS* 33, 1065; [3] Krot et al. (2001) *Geochem. Int.* 36, 351; [4] MacPherson et al. (2002) *MAPS* 37, A91; [5] Cozzirinsky et al. (2003) *LPS* 34, #1043; [6] Hutcheon et al. (1998) *Science* 282, 1865; [7] Pravdivtseva et al. (2001) *MAPS* 36, A168; [8] Krot et al. (2002) *MAPS* 37, A82; [9] Weisberg and Prinz (1998) *MAPS* 33, 1087; [10] Kojima and Tomeoka (1996) *GCA* 60, 2651; [11] Krot et al. (1997) *Meteoritics* 32, 31; [12] Choi et al. (2000) *MAPS* 35, 1239; [13] Imai and Yurimoto (2003) *GCA* 67, 765; [14] Krot et al. (1999) *MAPS* 34, 67; [15] Brearley (1997) *Science* 276, 1103; [16] Brearley (1999) *Science* 285, 1380; [17] Clayton and Mayeda (1999) *GCA* 63, 2089.