NEW CALCULATIONS AND ESTIMATIONS FOR HYDROTHERMAL ZONES AND IMPACT CONDITIONS ON CHICXULUB, EARTH AND ISIDIS PLANITIA, MARS.
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Synopsis: Chicxulub [1] is among the largest impact crater on Earth and a good analogue for Mars impact processes. Mars's Isidis Planitia [2] is one of the largest impact Planitias on Mars with a diameter of about $1,238 \mathrm{~km}$. Isidis is located at N 14.1 deg and W 271.0 degrees and is the boundary between ancient highlands and the Northern Plains. The exceptionally well-preserved Chicxulub crater is located in the Peninsula of Yucatan in Mexico, and research has identified at least 3 concentric structural rings, which comprise a complex $\sim 200 \mathrm{~km}$ diameter impact basin.

Analytical Method and Results: Our model [3,4] shows for Chicxulub that the asteroid diameter is $\sim 7.01 \mathrm{~km}$, with a velocity and impact angle of $\sim 47.38 \mathrm{~km} / \mathrm{s}$ and $\sim 33.12^{\circ}$ respectively. The number of rings are calculated in $\sim 5.73$ with a crater profundity of $\sim$ 1.31 km and melt volume of $\sim 37,414 \mathrm{~km}^{3}$. The number of ejected fragments are estimated in $\sim 1,903$ millions with sizes of $\sim 5.66 \mathrm{~m}$, the asteroid density is $\sim 5.39 \mathrm{~g} / \mathrm{cm}^{3}$. The total energy in the impact is calculated in $\sim 1.2$ E30 Ergs, i.e., $\sim 571$ millions of Hiroshima. The hydrothermal zone is of $\sim 61.2 \mathrm{~km}$ to 98 km from the nucleus of impact. The lifetimes estimated are of $\sim 1.19 \mathrm{Ma}$ to 1.86 Ma with uncertainties of $\sim+/-0.0076$ Ma to $+/-0.0131 \mathrm{Ma}$. Hydrothermal temperatures for 0.25 years to 1,400 years are estimated in $\sim$ $246.34^{\circ} \mathrm{C}$ to $96.65^{\circ} \mathrm{C}$. The fragments are ejected to $\sim 500.04 \mathrm{~km}$ from the impact center, with velocity of ejection of $\sim 5.87 \mathrm{~km} / \mathrm{s}$, ejection angle of $\sim 4.10^{\circ}$ and maximum height of $\sim 8.97 \mathrm{~km}$.

For Isidis Planitia, the asteroid diameter is $\sim 438.65 \mathrm{~km}$, with a velocity and impact angle of $\sim 19.42 \mathrm{~km} / \mathrm{s}$ and $\sim 74.09^{\circ}$ respectively. The number of rings could be $\sim 206$ with a crater profundity of $\sim 4.8$ km and melt volume of $\sim 5,765,600.6 \mathrm{~km}^{3}$. The number of ejected fragments are estimated in $\sim 3.3 \mathrm{E} 14$ with sizes of $\sim 6.35 \mathrm{~m}$. The total energy in the impact is calculated in $\sim 2.78 \mathrm{E} 33$ Ergs, i.e., $\sim$ 66,110 millions of megatons. The hydrothermal zone is of $\sim 69.69$ km to 617.63 km from the nucleus of impact. The lifetimes estimated are of $\sim 68.09 \mathrm{Ma}$ to 106.28 Ma with uncertainties of $\sim+/-$ 0.88 Ma to $+/-3.74 \mathrm{Ma}$. Hydrothermal temperatures for 0.25 years to 1,400 years are estimated in $\sim 527.63^{\circ} \mathrm{C}$ to $207.16^{\circ} \mathrm{C}$. Finally the fragments are ejected to $\sim 79,276 \mathrm{~km}$ from the impact center, with velocity of ejection of $\sim 24.42 \mathrm{~km} / \mathrm{s}$, ejection angle of $\sim 75.18^{\circ}$ and maximum height of $\sim 74,894 \mathrm{~km}$, these enormous distances could to eject the fragments out of the Mars planet, in a closed orbit.

All the calculations are obtained using a HP 49g, which is a Scientific Programmable Graphing Calculator with 1.5 Mb in RAM memory.

References: [1] Pope, K.O., Ocampo, A.C., Kinsland, G.L., and Smith, R. (1996) Geology., 24, 527-530. [2] Scott, D., and Tanaka, K. (1986) Geological Survey Misc. Inv. Map, I-1802-A. [3] Pope, K.O., Baines, K.H., Ocampo, A.C., and Ivanov, B.A. (1997) Journal of Geophysical Research., 102, 21,645-21,664. [4] Echaurren, J.C., and Ocampo, A.C. (2003) EGS-AGU-EUG Joint Assembly.

