

## HELIUM LOSS AND SHOCK PRESSURE IN MARTIAN METEORITES – A RELATIONSHIP

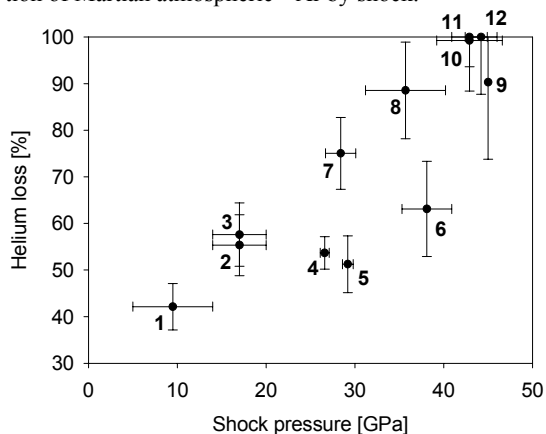
S. P. Schwenzer<sup>1</sup>, J. Fritz<sup>2</sup>, A. Greshake<sup>2</sup>, S. Herrmann<sup>1</sup>, K. P. Jochum<sup>1</sup>, U. Ott<sup>1</sup>, D. Stöffler<sup>2</sup>, B. Stoll<sup>1</sup>

<sup>1</sup>Max-Planck-Institut für Chemie, J.-J. Becher Weg 27, D-55128 Mainz, Germany, schwenze@mpch-mainz.mpg.de; <sup>2</sup>Humboldt-Universität zu Berlin, Museum für Naturkunde, Institut für Mineralogie, D-10099 Berlin, Germany.

**Introduction:** Impact on a parent body causes shock effects, which have been shown to influence the inventory of noble gases and easily volatilized elements in the affected rocks [1, 2]. A well-known example is the resetting of the Ar-“clock” by a major degassing event and therefore a pronounced clustering of Ar-ages around ~0.5 Ga for the L-chondrites [2]. The influence of shock on the noble gas content in chondrites has been demonstrated for <sup>40</sup>Ar [3] and <sup>4</sup>He [4], with a complete loss resulting from shock pressures in excess of ~35 GPa.

**Methods:** Here we investigate the effects on the radiogenic <sup>4</sup>He inventory of Martian meteorites. We combine data from three studies: Radiogenic <sup>4</sup>He from noble gas mass spectrometry, U and Th contents measured by SSMS [5], and peak shock pressures based on the shock induced reduction of the refractive index of plagioclase [6]. The loss of <sup>4</sup>He is inferred from the difference between the amount of <sup>4</sup>He produced by radioactive decay of U and Th since closure of the magmatic system (as deduced from ages obtained by Rb/Sr and/or Sm/Nd dating [7, 8]) and the measured <sup>4</sup>He amount.

**Results:** Our data show a correlation between the <sup>4</sup>He loss and the shock pressure as determined with the method of [6] in the range of ~10–45 GPa, indicating that the amount of energy deposited by shock is correlated with the loss of <sup>4</sup>He. In addition to our data it is known [9] that the heavily shocked LEW88516 (~44 GPa [6]) also suffered complete loss of <sup>4</sup>He. For <sup>40</sup>Ar the situation is more complicated, as there is both, loss and implantation of Martian atmospheric <sup>40</sup>Ar by shock.



Helium loss vs. shock pressure. 1: Lafayette, 2: Nakhla, 3: Gov. Valadares, 4: Chassigny, 5: Zagami, 6: EETA79001, 7: Shergotty, 8: ALHA84001, 9: QUE94201, 10: SaU005, 11: DaG476, 12: ALHA77005.

**References:** [1] Keil K et al. (1994) *Planet Space Sci*, 42: 1109–1122. [2] Bogard DD (1995) *Meteoritics*, 30: 244–268. [3] Stöffler D et al. (1991) *GCA*, 55, 3845–3867. [4] Schultz L & Stöffler D (1993) *MAPS*, 28, 432. [5] Jochum KP et al. (2001) *MAPS*, 36, A90–A91 [6] Fritz J et al. (2003) *LPSC*, XXXIV, #1335. [7] Meyer C (2001) Mars Meteorite Compendium; Houston. [8] pers. comm. E. Jagoutz, Mainz. [9] Ott U & Löhr HP (1992) *MAPS*, 27, 271.