

SIX MODELS FOR THE FORMATION OF THE EARTH'S ATMOSPHERE. M. Maurette. CSNSM, Bat.104, 91405 Orsay–Campus, France. E–mail: maurette@csnsm.in2p3.fr.

Introduction: The chemical composition of planetary atmospheres can be defined by the mass mixing ratios, $R(A)$, of their constituent species, A, relatively to N_2 . Moreover, in the case of a pure meteoroid atmosphere released by meteoroids upon atmospheric entry, these ratios would scale as the corresponding concentration ratios measured in Antarctic micrometeorites [1]. Previously, we only considered Ne, N, C (trapped in carbonates) and H_2O . We recently added S locked in the upper mantle to this list, because the predicted meteoroid burdens of S, Ir, Os and Ru strikingly well fit the corresponding values observed in the upper mantle [2], i.e., within a factor 1.05x (Ru) to 1.25x (S). Like in classical basalt magma volcanism, C and S were first likely oxidized into CO_2 and SO_2 , by a high temperature in–situ oxidation that would be related to the oxygen fugacity of their assemblage of minerals.

Comparison of 6 models of atmosphere: Table I lists the mixing ratios of Ne, H_2O , CO_2 and SO_2 . The successive lines from top to bottom refer to: **#Earth**, *observed* composition; **Model #1**, volcanic atmosphere defined by Rubey [2]; **Model #2**, atmosphere released by geysers and fumaroles that are heavily contaminated by surface water [4]; **Model #3**, outgassing of Hawaiian volcanoes, which would both sample the most pristine gases from the degassed upper mantle [4], and give a model for the Tharsis bulge volcanism on Mars; **Model #4**, atmosphere expected from a cometary impact [5]; **Model #5**, "steam" atmosphere from a CI–type asteroid impact [6]; **Model #6**, "pure" meteoroid atmosphere (up–dated ratios from Table 3.1, in Ref. 1).

Table I: Mass mixing ratios (relatively to N_2)

Type	$R(Ne)$	$R(H_2O)$	$R(CO_2)$	$R(SO_2)$
#EARTH	1.6×10^{-5}	350	83	75–150
#1 (Rubey)	?	380	21	1.1
#2 (Fumaroles)	?	2000	6.6	1.2
#3 (Hawaii)	?	10	4	4.4
#4 (Comet)	?	17	2.6	0.06
#5 ("Steam")	10^{-8}	55	54	0.3
#6 (Meteoroid)	3×10^{-5}	140	130	140

Verdict: From the 6 models of atmosphere considered, yet, only the meteoroid atmosphere (**#6**) reasonably fits the measured composition of the Earth's atmosphere (2nd line). The very low Ne mixing ratio that invalidates model **#5** right away, was derived in Ref. 1 (section 3.3.2).

References: [1] Maurette M. 2006. In *Comets and the Origin and Evolution of Life*, edited by Thomas P.J. et al., pp. 69–111. [2] Maurette M. Forthcoming. Hydrous-carbonaceous meteoroids in the Hadean Aeon. *ASP Conf. Series*. [3] Rubey W.W. 1955. In *Crust of the Earth*. New-York: Geol. Soc. America, pp. 630–650. [4] Hartmann W.K. 1999. *Moon and Planets*. Belmont: Wadsworth. 428p. [5] Delsemme A. 2006. In *Comets and the Origin and Evolution of Life*, edited by Thomas P.J. et al., pp. 29–68. [6] Fegley B. and Schaefer L. 2008. *Meteoritics & Planetary Science* 43: A42.