

**THE VIGARANO CAI REFERENCE SUITE, II. HIGH-PRECISION AL-MG ISOTOPIC STUDIES.**

G. J. MacPherson<sup>1</sup>, N. T. Kita<sup>2</sup>, E. S. Bullock<sup>3</sup>, T. Ushikubo<sup>2</sup>, A. M. Davis<sup>3,4</sup>.

<sup>1</sup>US National Museum of Natural History, Smithsonian Institution, Washington, D.C., USA 20560. E-mail: macphers@si.edu.

<sup>2</sup>Wisc-SIMS, Univ. of Wisconsin/Madison, Madison, WI 53706.

<sup>3</sup>Dept. of the Geophysical Sciences, Univ. of Chicago, Chicago, IL 60637, USA. <sup>4</sup>Enrico Fermi Institute, Univ. of Chicago, Chicago, IL 60637, USA.

**Introduction:** Vigarano (CV3) CAIs have far less secondary alteration than do Allende CAIs, and have not experienced the intense (presumably shock-induced) flattening observed in Efremovka and Leoville. Thus they may be the most pristine representatives of CV3 CAIs. Preliminary Mg isotope data are given here for high-Al/Mg phases in 4 Vigarano CAIs extracted as part of the CAI Reference Suite Project [1] – one fluffy Type A (FTA), one amoeboid olivine aggregate (AOA), and 2 Type B2s.

**Samples and Techniques:** Vigarano 3138-F1 and -F2 are Type B2, both showing minimal secondary alteration except for minor nepheline within anorthite. Inclusion 3138-F5 is an exceptionally unaltered AOA, consisting of unzoned forsteritic olivine (Fo<sub>97-100</sub>), numerous small nodules of spinel that are rimmed by pyroxene, and a 350 μm diameter spherule consisting entirely of melilite (Åk<sub>5-15</sub>) and spinel, with a thin outer pyroxene rim. Inclusion 3138-F8 is an unaltered FTA that contains abundant melilite (Åk<sub>0-17</sub>) and hibonite, with lesser spinel. The Cameca IMS-1280 ion microprobe at the Univ. of Wisconsin was used for Al-Mg isotope analyses in EM-monocollection mode for high (5–1000) Al/Mg phases using a 5 μm-diameter primary <sup>16</sup>O<sup>-</sup> beam.

**Results:** Analyses of Al-rich melilite in the AOA and melilite+hibonite in the FTA yielded undisturbed isochrons corresponding to initial <sup>26</sup>Al/<sup>27</sup>Al ratios of (5.5±0.4)×10<sup>-5</sup> (AOA) and (5.3±0.3)×10<sup>-5</sup> (FTA), both close to the assumed canonical value of 5×10<sup>-5</sup> [2] and the recently determined “whole-CAI” isochron [3] value of 5.2×10<sup>-5</sup>. Melilite in both Type B2s yielded reasonably well-defined isochrons of ~5×10<sup>-5</sup>, although with relatively large uncertainties. The Al-Mg system in anorthite in these Type B2s has been almost completely reset by some later event. Anorthite in 3138-F1 retains <sup>26</sup>Mg excesses that do not correlate positively with Al/Mg, and no resolvable <sup>26</sup>Mg excesses were found in anorthite in 3138-F2.

**Discussion:** The two primitive CAIs (FTA, AOA) have initial <sup>26</sup>Al/<sup>27</sup>Al ratios consistent with the recently updated “canonical” ratio of 5.2×10<sup>-5</sup>. The behavior of the Al-Mg system in the two Type B2s is quite different from that in Allende: the CAIs in the latter seem to have experienced near-closed-system partial isotopic exchange between high- and low-Al/Mg phases [e.g. 4], but the resetting of the Al-Mg system in anorthite in the two Vigarano Type B2s may be the result of fluid-controlled isotopic exchange.

**References:** [1] Bullock E. S. & MacPherson G. J 2008. *Meteoritics & Planetary Science* 43:A29. [2] MacPherson G. J., Davis A. M. & Zinner E. K. 1995. *Meteoritics & Planetary Science* 30:365–386. [3] Jacobsen B. et al. 2008. *Earth and Planetary Science Letters* 272:353–364. [4] Podosek F. A. et al. 1991. *Geochimica et Cosmochimica Acta* 55:1083–1110.