

**Sm-Nd AND Rb-Sr AGES FOR NORTHWEST AFRICA 2977, A YOUNG LUNAR GABBRO FROM THE PKT.** L. E. Nyquist<sup>1</sup>, C.-Y. Shih<sup>2</sup>, Y. D. Reese<sup>3</sup>, and A. J. Irving<sup>4</sup>. <sup>1</sup>NASA Johnson Space Center, Houston, TX 77058. E-mail: [laurance.e.nyquist@nasa.gov](mailto:laurance.e.nyquist@nasa.gov). <sup>2</sup>JE-23, ESCG Jacobs-Sverdrup, Houston, TX 77058. <sup>3</sup>JE-23, ESCG/Muniz Eng., Houston, TX 77058. <sup>4</sup>Dept. Earth&Space Sci. Univ. Washington, Seattle, WA 98195.

**Introduction:** Northwest Africa (NWA) 2977 is an olivine gabbro cumulate equivalent to one of the lithologies in lunar mare breccia NWA 773 [1,2,3]. The <sup>39</sup>Ar-<sup>40</sup>Ar age is 2.77±0.04 Ga based on the last ~57% of the gas release [4], similar to results for NWA 773 [5]. A Sm-Nd age (T) of 2.865±0.031 Ga and  $\epsilon_{Nd} = -7.84 \pm 0.22$  for the NWA 773 gabbro reported by [6] has been revised to T = 2.993±0.032 Ga,  $\epsilon_{Nd} = -4.5 \pm 0.3$  [7].

**<sup>147</sup>Sm-<sup>143</sup>Nd isochron for NWA 2977:** Whole rock, pyroxene, olivine, plagioclase, whole rock leachate (~phosphate) and the combined leachates from the mineral separates yield a well defined Sm-Nd isochron for an age T = 3.10±0.05 Ga and  $\epsilon_{Nd,CHUR} = -3.74 \pm 0.26$  [8], or  $\epsilon_{Nd,HEDR} = -4.61 \pm 0.26$  [9].

**<sup>87</sup>Rb-<sup>87</sup>Sr isochron:** NWA 2977 contains only a modest amount of Rb and/or Sr contamination. The Sr-isotopic composition of the contaminant closely resembles that of seawater. The whole rock residue after leaching combined with leach residues for plagioclase and pyroxene define an isochron age of 3.29±0.11 Ga for initial <sup>87</sup>Sr/<sup>86</sup>Sr = 0.70287±18. The olivine residue, with lower Sr abundance of ~ 1.5 ppm, is only slightly displaced from the isochron. The relatively small uncertainties of the Rb-Sr isochron parameters and near-concordancy with the Sm-Nd age indicate that both the Rb-Sr and the Sm-Nd ages are reliable.

**Discussion:** The somewhat older Rb-Sr and Sm-Nd ages of NWA 2977 compared to its <sup>39</sup>Ar-<sup>40</sup>Ar age suggest some radiogenic <sup>40</sup>Ar\* loss from the sample analyzed by [4]. Initial  $\epsilon_{Nd}$  for NWA 2977, when combined with  $\epsilon_{Nd}$  for KREEP-rich samples, defines a line that when extrapolated passes between the CHUR and HEDR values at 4.568 Ga ago. Both types of samples lie within error limits of Nd-isotopic evolution for <sup>147</sup>Sm/<sup>144</sup>Nd ~0.177 characteristic of KREEP-rich sources. Also,  $\epsilon^{142}Nd$  for NWA 2977 is the same within error limits as for KREEP melt rock 14310 [10]. The <sup>142,143</sup>Nd isotopic evolution can be modeled for: (a) 50 Ma with chondritic <sup>147</sup>Sm/<sup>144</sup>Nd ( $\mu_{CHUR}$ ) = 0.1967, (b) 168 Ma with  $\mu_{moon} = 0.205$ , (c) 1250 Ma with source region  $\mu_S = 0.173$  (present-day values). A two-stage model for Sr-isotopic evolution gives <sup>87</sup>Rb/<sup>86</sup>Sr ~0.207 in the NWA 2977 source, compared to ~0.18 in the rock and ~0.14-0.17 in the sources of A14 and A15 KREEP basalts. The NWA 2977 source evidently contained highly evolved Trapped Instantaneous Residual Liquid [11]. The age, Sr-and Nd-isotopes, and trace element abundances suggest an origin in the PKT (Procellarum KREEP Terrane).

**References:** [1] Bunch T. E. et al. 2006. *LPS XXXVII*, #1375. [2] Jolliff B. L. et al. 2007. *LPS XXXVIII*, #1489. [3] Zeigler R. A et al. 2007. *LPS XXXVIII*, #2109. [4] Burgess R. et al. 2007. *LPS XXXVIII*, #1603. [5] Fernandes V. A. et al. 2003. *Meteoritics & Planet. Sci.* 38: 555-564. [6] Borg L. et al. 2004. *Nature* 432: 209-211. [7] Borg L. et al. 2009. *Geochim. Cosmochim. Acta*, in press. [8] Jacobsen S. B. and Wasserburg G. J. 1984. *Earth & Planet. Sci. Lett.* 67: 137-150. [9] Nyquist L. et al. 2006. *Geochim. Cosmochim. Acta*, 70: 5990-6015. [10] Touboul M. et al. 2009. *LPS XL*, #2269. [11] Snyder G. A. and Taylor L. 1993. *Proc. NIPR Symp. Antarctic Met.* 6: 246-267.