

**SIMULATION OF THE MASSIVE FRACTIONATION OBSERVED IN CO PHOTODISSOCIATION EXPERIMENTS.**

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**Introduction:** The observation of massive fractionation (~10,000 ‰) in CO<sub>2</sub> produced in laboratory CO photolysis experiments [1] provides a superb illustration of the process of self-shielding. Self-shielding results from the saturation of CO isotopologue absorption lines, and yields an abundance-dependent fractionation upon dissociation. Because <sup>12</sup>C<sup>16</sup>O is ~ 500 times more abundant than <sup>12</sup>C<sup>18</sup>O, saturation of <sup>12</sup>C<sup>16</sup>O lines occurs at much lower gas column densities than for the rare isotopologues, creating a region of <sup>17</sup>O and <sup>18</sup>O-enriched O atoms. CO self-shielding is essential to understanding the distribution of oxygen isotopes in the solar system [2],[3],[4]. For the experiments performed in [1] the typical column densities were ~ 10<sup>18</sup> cm<sup>-2</sup>, implying <sup>12</sup>C<sup>16</sup>O optical depths > 1000 [5]. Thus, self-shielding was unavoidable in these experiments. In order to elucidate the origin of the large fractionations measured in [1], we have performed detailed modeling of the photolysis of the CO E(0) – X(0), E(1) – X(0), and C(1) – X(0) bands at 107.6, 105.17, and 106.3 nm, respectively. To the cross sections computed in [6] we have included the perturbation of the E(1) state by the k(6) triplet state using the formalism of [7].

**Results:** For simulation of the experiments at 105.17 nm, the model δ<sup>x</sup>O(CO<sub>2</sub>) values are ~ 6,000-12,000 ‰ (comparable to measured values), but have a δ<sup>17</sup>/δ<sup>18</sup> slope ~ 1.05, lower than the measured value of 1.3. δ<sup>x</sup>O(CO<sub>2</sub>) decrease with CO column density and become negligible when optical depth is << 1, as expected for a self-shielding effect. Low optical depth experiments are needed to confirm this result. The δ<sup>17</sup>/δ<sup>18</sup> slope can be accounted for by a 30% stronger band oscillator strength for <sup>12</sup>C<sup>17</sup>O versus <sup>12</sup>C<sup>18</sup>O. However, we do not expect such large isotopic variation in band oscillator strengths, and doubt that this is a satisfactory explanation.

Simulation results [6] for the synchrotron beam at 107.61 nm (E(0) band) show that a slope ~ 1.4, comparable to the measured values, occurs at the highest column densities (as observed), but is not present at lower densities. It is important to note that self-shielding does occur in this band, contrary to the claims of [1]. The non-unity δ<sup>17</sup>/δ<sup>18</sup> slope arises from the pattern of closely overlapping lines among the isotopologue spectra.

**Conclusions:** The key results here are that massive fractionation is expected from self-shielding alone, and that the fractionation depends on wavelength. The δ<sup>17</sup>/δ<sup>18</sup> slope associated with CO photolysis must be determined from the summation over all CO bands for comparison with the CAI line (0.95-1.0).

**References:** [1] Chakraborty S. et al. 2009. *Science* 321:1328-1331. [2] Clayton R. N. 2002. *Nature* 415:860-861. [3] Yurimoto H. and Kuramoto K. 2004 *Science* 305:1763-1766. [4] Lyons J. R. and Young E. D. 2005. *Nature* 435:317-320. [5] Lyons J., Lewis R., Clayton R. (submitted) [6] Lyons J. R. 2009. Abstract #2377 40th Lunar and Planetary Science Conference. [7] Ubachs W., I. Velchev, P. Cacciani 2000. *Journal of Chemical Physics* 113:547-560.