

BASALTIC SHERGOTTITE, JIDDAT AL HARASIS 479

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Introduction: Martian meteorite, Jiddat Al Harasis 479, (JaH 479) was found in the desert of Oman in 2008 [1]. We have studied its mineralogy, petrography, chemistry and oxygen isotopic compositions. In addition, results of its noble gases investigation are presented in this volume [2].

Results and discussion: The meteorite is unbrecciated, coarse-grained subophitic rock, consisting mainly of pyroxene and plagioclase. Pyroxene is presented by differently oriented, anhedral pigeonite and subhedral, elongated, augite grains of 1-3 mm in size. Several pigeonite grains are long lathes up to 10 mm. Pigeonite crystals are twinned and show undulatory extinction. Feldspar is isotropic (maskelynite), but several grains demonstrate a weak anisotropy. Maskelynite is presented by lathes of 0.04-0.06 x 1.5 mm in size, and by prismatic grains of smaller size, growing between pyroxene crystals. The minor phases are olivine, ilmenite, silica, orthoclase, Ca-phosphates, zircon, pyrrhotite and troilite. The fine-grained aggregates of clinopyroxene, olivine and silica occur on the outer parts of pyroxene grains. The rare black melt veins of 0.1-0.2 mm wide are crossing the rock.

Pigeonite is normally zoned, from $\text{En}_{45.8}\text{Wo}_{23.5}$ to $\text{En}_{15.3}\text{Wo}_{19.9}$ ($\text{Fe}/\text{Mn} = 37$). Augite is $\text{En}_{14.7-46.0}\text{Wo}_{20.4-31.4}$ ($\text{Fe}/\text{Mn} = 28$). Maskelynite is not stoichiometric and irregularly varies in composition from $\text{An}_{76}\text{Ab}_{18}\text{Or}_6$ to $\text{An}_{66}\text{Ab}_{31}\text{Or}_3$. Olivine is $\text{Fo}_{3.85}$ ($\text{Fe}/\text{Mn} = 54$). Pyroxene in fine-grained aggregates with olivine and silica is $\text{En}_{4.5}\text{Wo}_{40.4}$. These aggregates are similar to those that found in several shergottites [3]. They probably were formed due to breakdown of metastable FeO-rich pyroxene or pyroxferroite. Ca-phosphates are F-Cl apatite and Cl-poor phosphate - $\text{Ca}_{9.2}(\text{Fe}_{0.9}\text{Mg}_{0.2})\text{Na}_{0.4}(\text{PO}_3\text{F}_{0.5})(\text{PO}_4)_6$, that could be close to Fe-rich bobdownsite ($\text{Ca}_9\text{Mg}(\text{PO}_4)_6(\text{PO}_3\text{F})$) [4]. According to [5-6], the JaH 479 was shocked to approximately 32 GPa. Concentrations of REEs in JaH 479, determined by INAA, are (ppm): La 1.4, Ce 3.36, Nd 2.18, Sm 0.77, Eu 0.24, Tb 0.21, Yb 1.27, Lu 0.25. Oxygen isotopic compositions of JaH 479 is: $\delta^{17}\text{O} = 2.951\text{‰}$, $\delta^{18}\text{O} = 5.070\text{‰}$, $\Delta^{17}\text{O} = 0.315\text{‰}$.

Conclusions: The Mn/Fe ratios in the pyroxene and oxygen isotopes composition of JaH 479 are in the range of those of Martian meteorites [7]. Based on mineralogy and petrology, JaH 479 is a basaltic shergottite. The Cl-normalized REEs pattern of JaH 479 shows moderate depletion in LREE relatively HREE and it is close to that of Zagami [8].

References: [1] Meteoritical Bulletin 95. 2009. MAPS 44:429-462. [2] Cartwright J. A. et al. 2010. This volume. [3] Ikeda Y. et al. 2006. Antarctic Meteorite Research 19:20-44. [4] Gnos E. et al., 2002. Meteoritics & Planetary Science 37:835-854. [5] Stoeffler D. et al. 1988. In "Meteorites and the early Solar System" (Ed. Kerridge J.F.), Univ. Arizona Press:165-204. [6] Fritz J. et al. 2005. Meteoritics & Planetary Science 40:1393-1411. [7] McSween H. Y., Treiman A. H. 1998. Reviews in mineralogy 36:6-1-6-53. [8] Warren P. H. and Bridges J. C. 2005. 36th Lunar & Planetary Science Conference. Abstract #2098.