

### Mn-Cr SYSTEMATICS OF PALLASITE OLIVINE BY SIMS: CONTRASTING Cr IN BRAHIN AND BRENHAM.

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**Introduction:** Pallasite meteorites are mixtures of Fe-Ni metal and olivine and would appear to be good candidates for  $^{53}\text{Mn}$ - $^{53}\text{Cr}$  dating ( $t_{1/2}$  3.7 Ma). However, conflicting ion-probe results have been reported for these rocks, with some workers finding elevated  $^{53}\text{Cr}/^{52}\text{Cr}$  correlated with  $^{55}\text{Mn}/^{52}\text{Cr}$  [1, 2] while others do not [3]. Recently, it has been suggested that a statistical effect at low count rates arising from the reduction of data as a series of ratios could introduce positive bias into isotope measurements for some applications [4], leading [5] to retract earlier findings of elevated  $^{53}\text{Cr}/^{52}\text{Cr}$  in pallasite olivine [1, 2]. To investigate whether in-situ accumulated radiogenic  $^{53}\text{Cr}$  exists in these meteorites, we have re-analysed olivine from two petrographically contrasting pallasites: Brahin (which has a fragmental texture), and Brenham (with rounded olivine).

**Methods:** *In situ* isotopic analysis was performed using Sensitive High-mass Resolution Ion Microprobe Reverse Geometry (SHRIMP-RG) on olivine and chromite from Brahin and olivine from Brenham. The primary sputtering beam was run at high current (17-66 nA as a mixture of  $\text{O}^-$  and  $\text{O}_2^-$ ) and focused to a ~30 micron diameter spot, giving strong Cr secondary ion beams and hence minimising the effect described by [4].  $^{52}\text{Cr}^+$ ,  $^{53}\text{Cr}^+$  and  $^{55}\text{Mn}^+$  were measured with an electron multiplier, and San Carlos olivine was used as an external standard to correct for instrumental fractionation on  $^{53}\text{Cr}^+/^{52}\text{Cr}^+$  and  $^{55}\text{Mn}^+/^{52}\text{Cr}^+$ .

**Results:** We find that olivine and chromite in Brahin has normal Cr isotopic composition within analytical error (~1 ‰ 1 $\sigma$ ). In contrast, Brenham olivine shows Cr-isotope excesses (up to 6 ±1‰ 1 $\sigma$ ) within a few  $10^3$ 's of microns of olivine rims. These are well correlated with  $^{55}\text{Mn}/^{52}\text{Cr}$  and imply an initial  $^{53}\text{Mn}/^{55}\text{Mn}$  of  $2.55 (\pm 0.62) \times 10^{-5}$  at the time of isotopic closure of Cr in Brenham olivine rims.

**Discussion:** Our finding of isotopically normal Cr in Brahin is in agreement with [3], while our initial  $^{53}\text{Mn}/^{55}\text{Mn}$  for Brenham is even higher than originally reported by [1, 2]. Previous ion-probe studies had 30-250 counts per cycle for  $^{52}\text{Cr}^+$  over 150 cycles [5], but with higher beam currents and improved ion transmission we have achieved  $1.4 \times 10^4$  counts per cycle, with 80-200 cycles for each analysis. We estimate bias arising from the effect highlighted by [4] to be ~0.1‰ for our measurements.

The contrasting Mn-Cr systematics found here is good evidence for localised Cr-isotope variation in some, but not all, pallasites. The initial  $^{53}\text{Mn}/^{55}\text{Mn}$  found for Brenham is not compatible with current understanding of extinct nuclide chronology (with Solar System initial  $^{53}\text{Mn}/^{55}\text{Mn}$  of  $9.1 (\pm 1.7) \times 10^{-6}$ ; [6]) and probably represents concentration of radiogenic  $^{53}\text{Cr}$  at olivine rims. This appears to be related to the arrangement of olivine, with the interconnected networks in Brenham [7] exhibiting anomalous Cr while isolated olivine fragments in Brahin do not.

**References:** [1] Hsu et al. 1997 28<sup>th</sup> LPSC, Abstract # 1783. [2] Hsu W. 2005 Geochemical Journal 39, 311. [3] Tomiyama T. et al. 2007 38<sup>th</sup> LPSC, Abstract # 2007. [4] Oglione et al. 2011 42<sup>nd</sup> LPSC, Abstract # 1592. [5] Huss et al. 2011 42<sup>nd</sup> LPSC, Abstract # 2608. [6] Nyquist et al. 2009 Geochimica et Cosmochimica Acta, 73, 5115-5136. [7] Spinsby et al. 2008 Computers & Geosciences 34, 1-7.