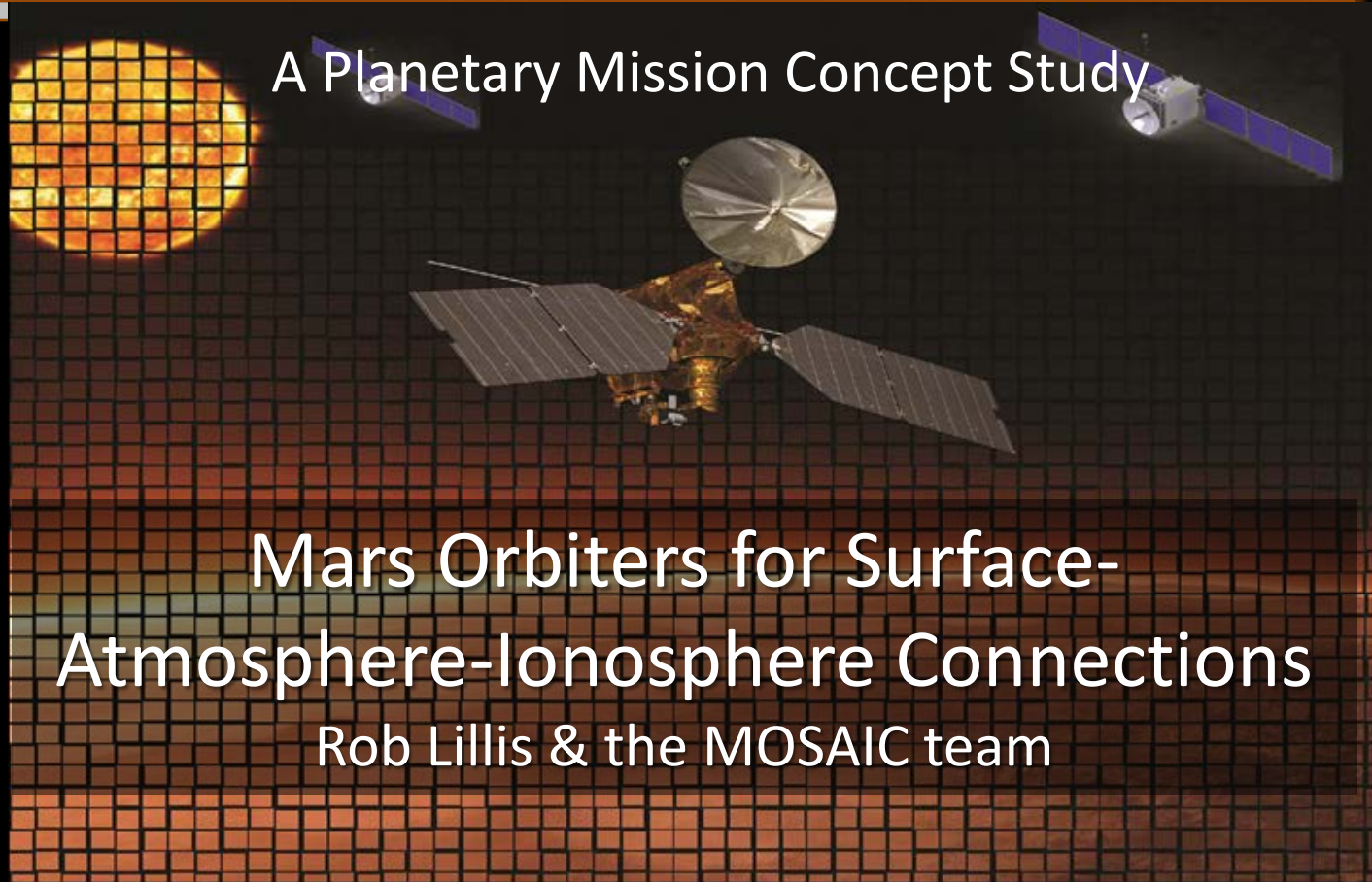


MOSAIC



A Planetary Mission Concept Study



Mars Orbiters for Surface- Atmosphere-Ionosphere Connections

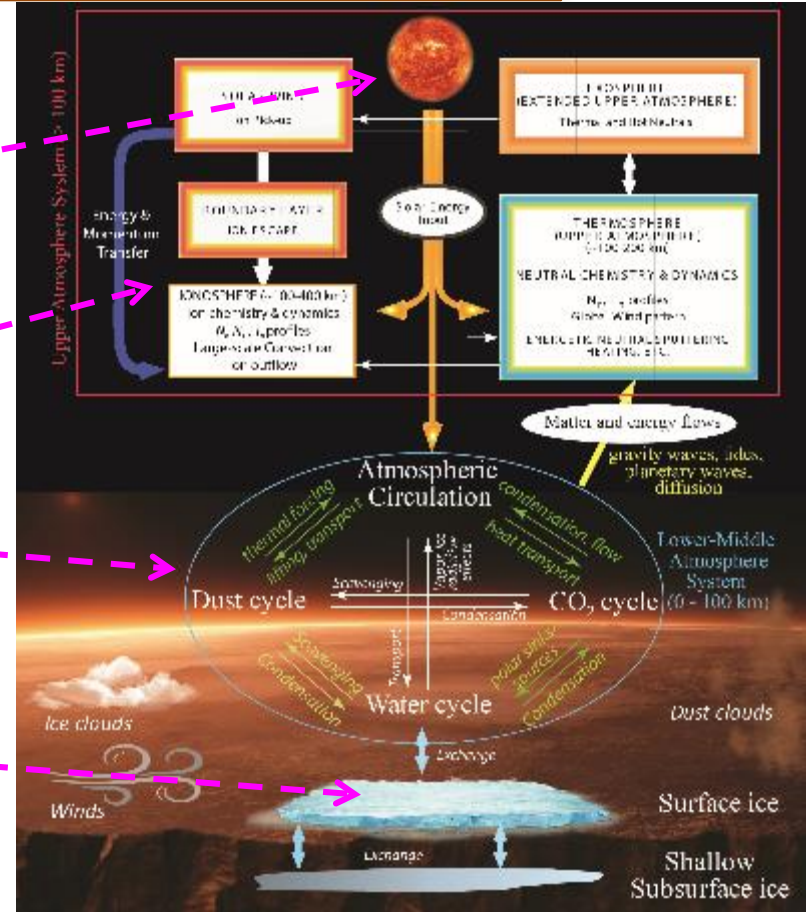
Rob Lillis & the MOSAIC team

NOTE ADDED BY JPL
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MOSAIC Goals & Objectives: Science



MOSAIC GOALS	Mission Objectives	
<p><i>Understand Mars's present day climate processes and their inter-connections, from the sub-surface to the solar wind</i></p> <p><i>(i.e. how do the different parts of the Mars climate system talk to each other?)</i></p>	I.D Characterize fields and plasma flows in the upstream solar wind and throughout the magnetosphere and upper ionosphere, separating spatial from temporal variability.	
	I.C: Correlate variability in the thermosphere, ionosphere, and escape rates to:	Conditions in the lower-middle atmosphere.
		The space weather environment
	I.B: Characterize the structure and dynamics of the Martian lower-middle atmosphere on meso- and global scales, and its geographic, diurnal, and seasonal variability.	
I.A: Characterize volatile cycling between the subsurface, surface and atmospheric reservoirs.		

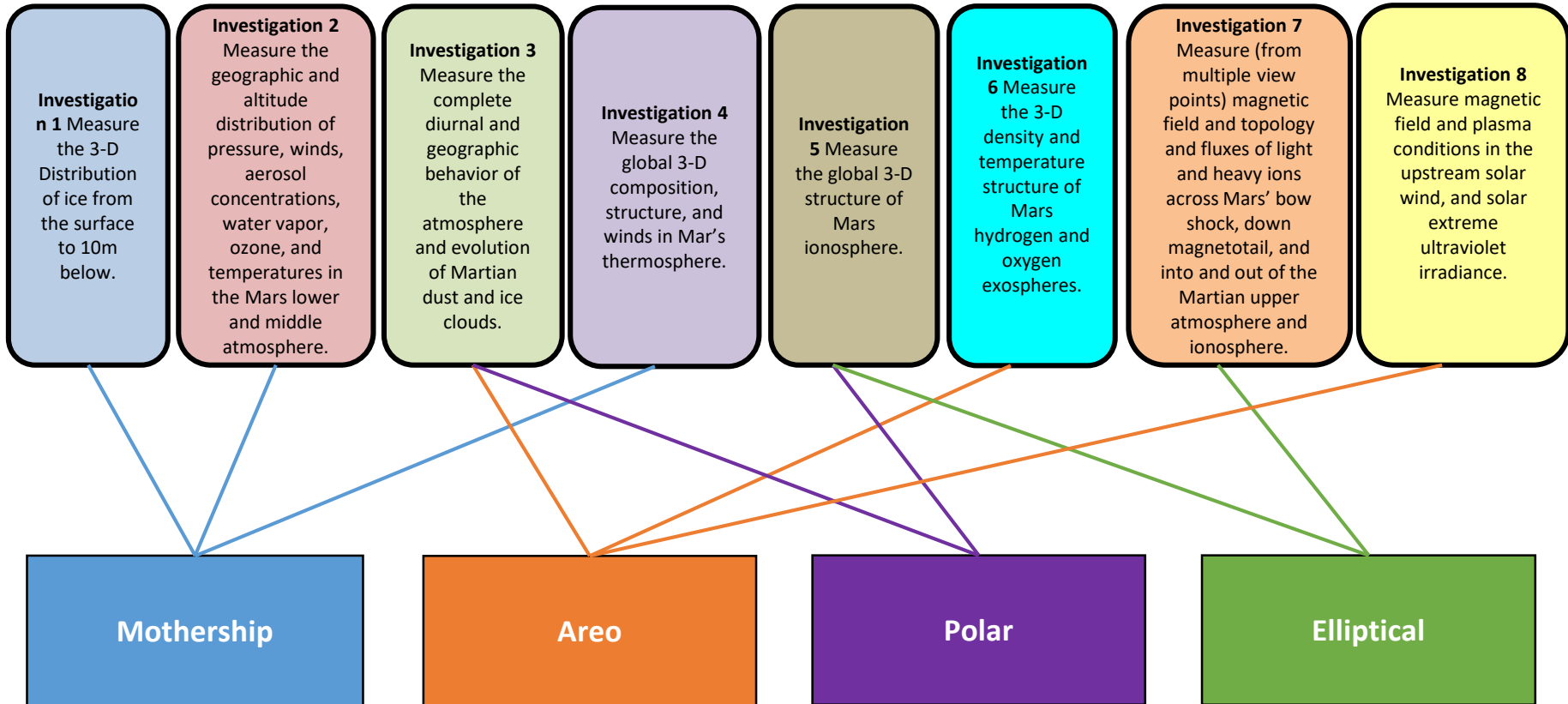


Objectives to Investigations



Mission Objectives		Investigations:	1. Ice distribution	2. Atmosphere structure	3. Atm. Diurnal behavior	4. Thermosphere	5. Ionosphere	6. Exosphere, neutral escape	7. Plasma & ion escape	8. Space weather
I.A: Characterize volatile cycling between the subsurface, surface and atmospheric reservoirs.			✓	✓	✓					
I.B: Characterize the structure and dynamics of the Martian lower-middle atmosphere on meso- and global scales, and its geographic, diurnal, and seasonal variability.				✓	✓					
I.C: Correlate variability in the thermosphere, ionosphere, and escape rates to:	Conditions in the lower-middle atmosphere.			✓	✓	✓	✓	✓	✓	
	The space weather environment					✓	✓	✓	✓	✓
I.D Characterize fields and plasma flows in the upstream solar wind and throughout the magnetosphere and upper ionosphere, separating spatial from temporal variability.							✓	✓	✓	✓

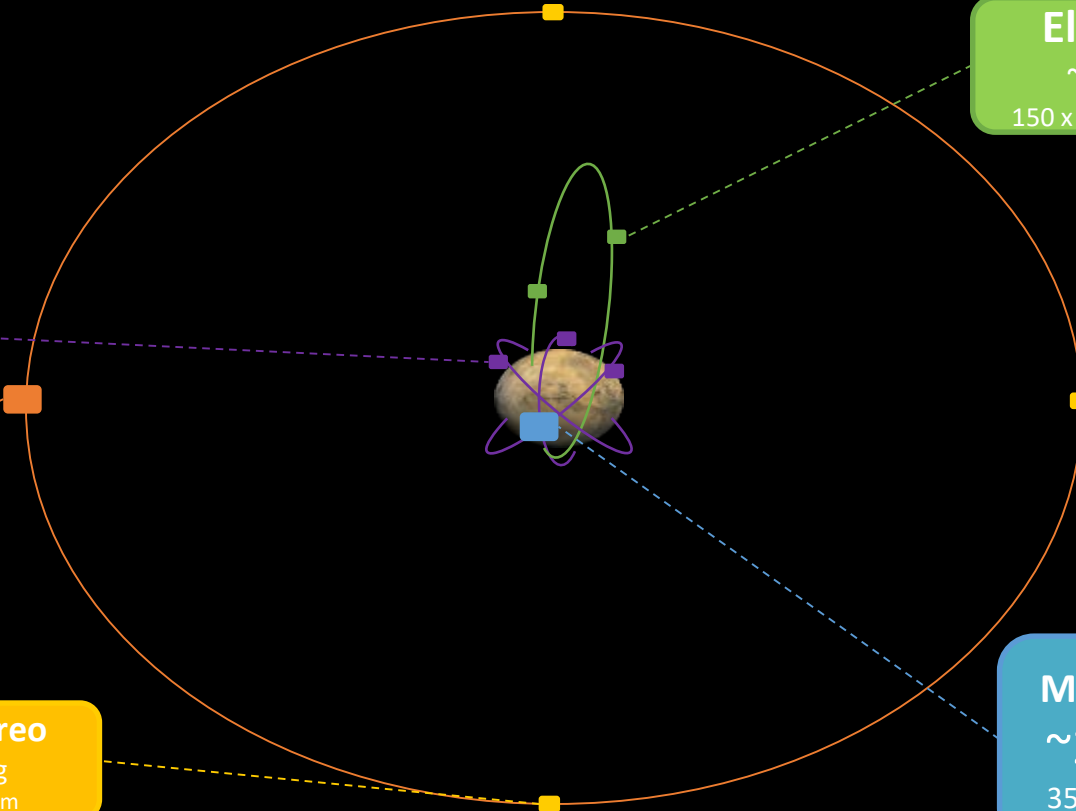
Connections between science and architectures



The MOSAIC Constellation



1. LMO Mothership [1 S/C] ←
2. SSO Polar [3 S/C]
3. Elliptical [2 S/C]
4. Areostationary [1 S/C]
5. Mini-Areo [3 S/C]



Elliptical ★
~220 kg
150 x 6000 km x 75°

Polar
~90 kg
350km x 92.8
(Spread in LST)

Areo ★
~500 kg
17,000 km

Mini-Areo
~50 kg
17,000 km

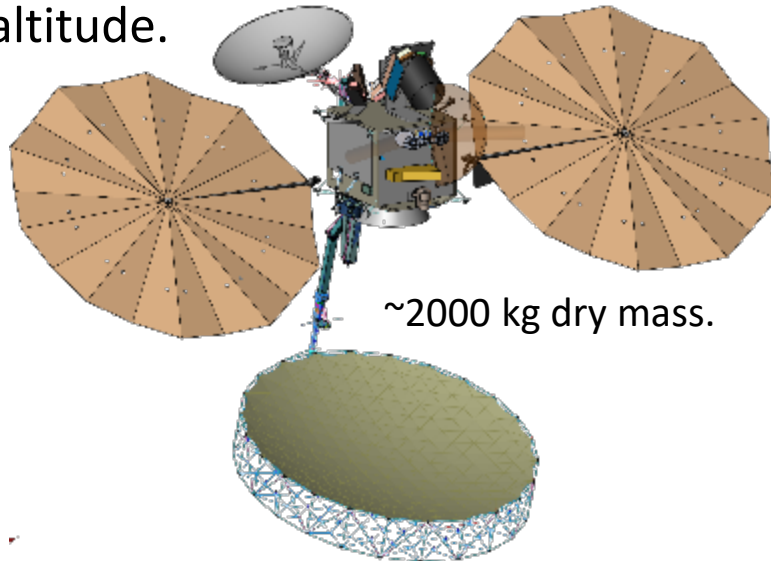
Mothership
~2000 kg
350 km x 92.8°



Mothership Platform



- Measures vertical profiles at fixed local time of:
 - Subsurface ice
 - Lower/middle atmosphere: Temperature, pressure, H₂O, aerosols and [wind](#)
 - Upper atmosphere: neutral composition & density, ion density, [neutral wind](#)
- Large SEP Class B spacecraft. Low circular sun-synchronous polar orbit, ~350 km altitude.

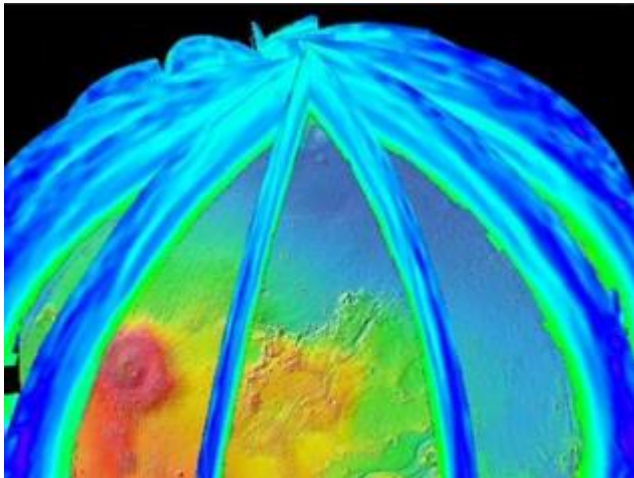
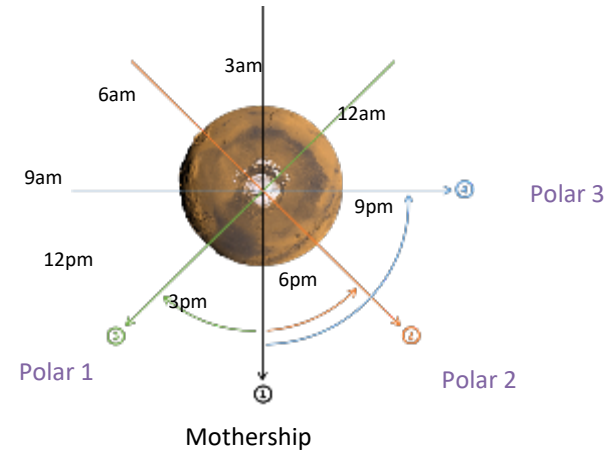


Instrument	Inv	Priority		Mass	Power	FY20 Cost
P-band SA radar	1	T : 1	B : 1	125 kg	200 W	\$170 M
Visible camera	1	T : 1	B : 1	3.4 kg	10 W	\$5 M
Thermal IR radiometer (MCS)	2	T : 1	B : 1	9 kg	18 W	\$25 M
Wind LIDAR (MARLI)	2	T : 1	B : 1	45 kg	91 W	\$40 M
Sub-mm sounder	2	T : 1	B : 1	35 kg	39 W	\$35 M
Near IR spectrometer	3	T : 0	B : 1	0.3 kg	2.5 W	\$0.34 M
Wind doppler interferometer	4	T : 1	B : 1	40 kg	20 W	\$40 M
FUV/MUV spectrometer	4	T : 1	B : 1	27 kg	28 W	\$30 M
Radio occultation	3,5	T : 1	B : 1	1.5 kg	3 W	\$2 M
Total				286 kg	412 W	\$347 M

Polar Orbit Smallsat



- **Measures: vertical profiles of pressure, temperature, water vapor, aerosols, ice and ionospheric plasma density.**
- Sun-synchronous low ~ 350 km polar orbit, 3-axis stabilized
- 3 copies of this spacecraft, each at a different local solar time
- Dropped off by the Mothership in a low polar orbit, and then moves itself into a sun-sync orbit with the correct local time

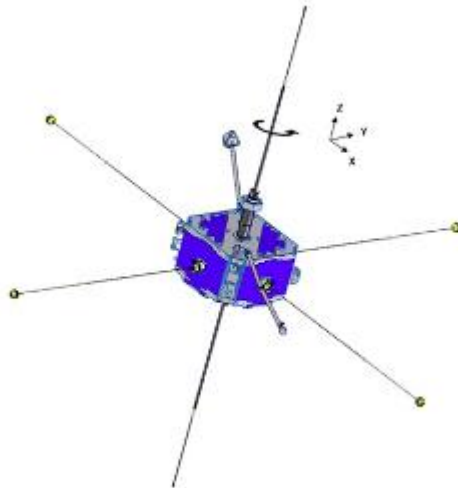
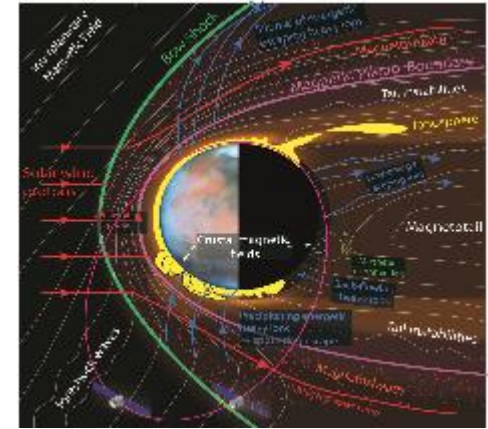


Instrument	Inv.	Priority	Mass	Power	Cost
Mini therm. IR radiometer	3	T : 1 B : 3	3.5 kg	8 W	\$10 M
Mini near IR spectrometer	3	T : 0 B : 3	0.3 kg	2.5 W	\$0.34 M
Radio occultation	3,5	T : 1 B : 3	1.5 kg	3 W	\$2 M
Total			5.3 kg	13.5 W	\$12.3 M

Elliptical Orbit Smallsat



- **Measures: ion and sputtering escape, magnetic morphology and topology, thermal plasma, and [electric fields](#). Also neutral temperature/pressure via RO.**
- 2 copies, pearls-on-a-string orbit allows disambiguation of spatial/temporal variability and measurement of real-time response to solar storms.
- Elliptical orbit, 7000 km x 150km, 4.5 hours. Spinner.
- Booms, including four 20m centrifugally stabilized wire booms
- Dropped off by the Mothership in a near-polar elliptical orbit.
- Smaller cheaper non-spinning version without electric fields also examined.

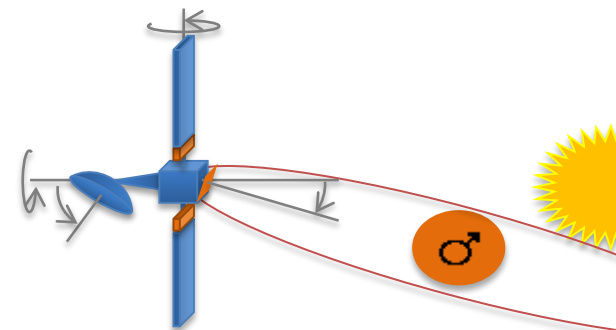


Instrument	Inv	Priority	Mass	Power	Cost
Fluxgate magnetometer	7	T : 2 B : 2	1.3 kg	4.9 W	\$4 M
Ion energy/angle/mass	7	T : 2 B : 2	3.3 kg	4.2 W	\$4 M
Electron energy/angle	7	T : 2 B : 2	1.8 kg	1.6 W	\$3 M
Electric fields	7	T : 0 B : 2	12 kg	0.24 W	\$2 M
Search coil magnetometer	7	T : 0 B : 2	1.8 kg	0.1 W	\$3 M
Langmuir probe	5	T : 2 B : 2	0.45 kg	1.5 W	\$0.5 M
Radio occultation	3,5	T : 2 B : 2	1.5 kg	3 W	\$2 M
Total			22.2 kg	15.5 W	\$18.5 M

Areostationary Spacecraft



- **Measures: 1) synoptic view of lower atmosphere (dust & ice column, ozone, temperature profile), 2) neutral oxygen and hydrogen escape rates 3) upstream space weather environment,**
- Separates shortly after launch, goes to Areostationary orbit using SEP
- Solar arrays co-pointed with solar instruments.
- Mars-pointed instruments: Camera to Mars center; rest need to scan.



IR 9.3 micron absorption CDOD normalized at 610 Pa

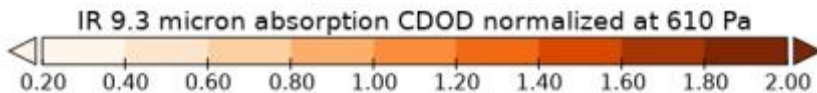
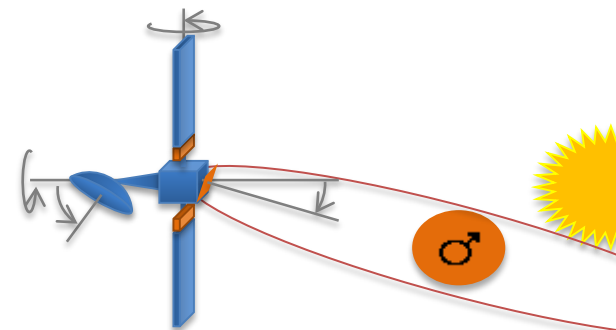


Instrument	Inv	Priority	Mass	Power	Cost
Visible camera (MSSS)	3	T : 1 B : 4	1.6 kg	7 W	\$0.5 M
Mini therm. IR radiometer	3	T : 1 B : 4	3.5 kg	8 W	\$10 M
Mini near IR spectrometer	3	T : 1 B : 4	3 kg	15 W	\$2 M
FUV/EUV spectrograph	6	T : 1 B : 1	21 kg	12 W	\$20 M
Fluxgate magnetometer	8	T : 1 B : 2	1.3 kg	4.9 W	\$4 M
Ion energy/angle	8	T : 1 B : 2	2.6 kg	2.1 W	\$3 M
Electron energy/angle	8	T : 0 B : 2	1.8 kg	1.6 W	\$3 M
Energetic ion/electron	8	T : 1 B : 2	0.9 kg	5.5 W	\$0.8 M
Extreme UV monitor	8	T : 1 B : 2	1.1 kg	0.7 W	\$2 M
Total			46.8 kg	57 W	\$43.6 M

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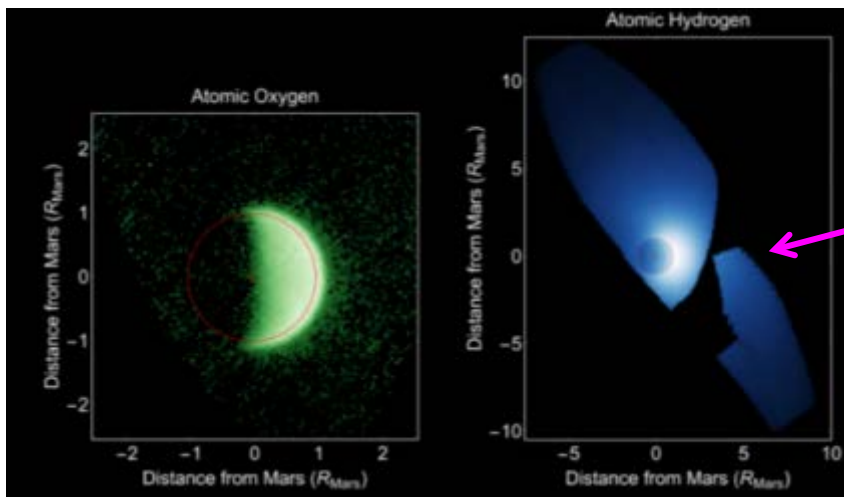
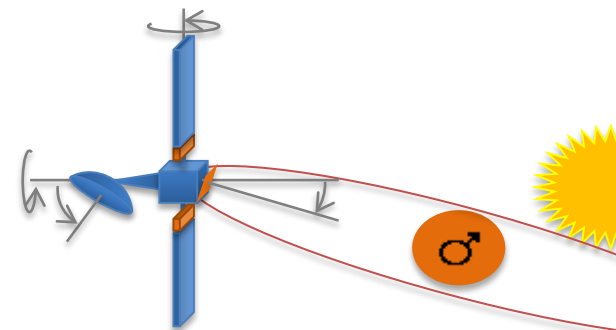


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Areostationary Spacecraft



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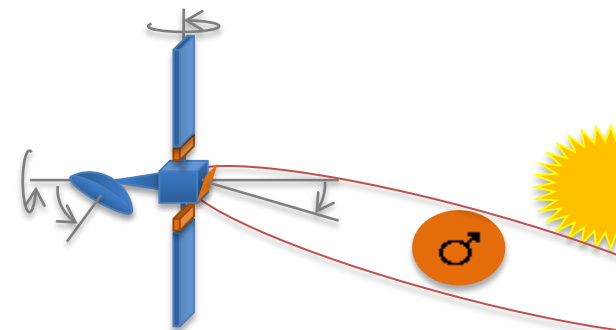


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FUV/EUV spectrograph	6	T : 1 B : 1	21 kg	12 W	\$20 M
Fluxgate magnetometer	8	T : 1 B : 2	1.3 kg	4.9 W	\$4 M
Ion energy/angle	8	T : 1 B : 2	2.6 kg	2.1 W	\$3 M
Electron energy/angle	8	T : 0 B : 2	1.8 kg	1.6 W	\$3 M
Energetic ion/electron	8	T : 1 B : 2	0.9 kg	5.5 W	\$0.8 M
Extreme UV monitor	8	T : 1 B : 2	1.1 kg	0.7 W	\$2 M
Total			46.8 kg	57 W	\$43.6 M

Areostationary Spacecraft

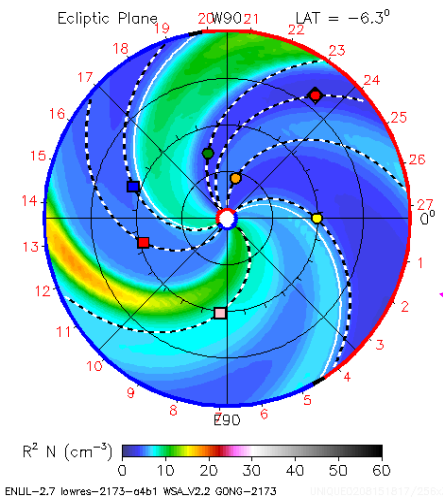


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2016-02-07T00:00

● Earth ● Mars ● Mercury ● Venus ◆



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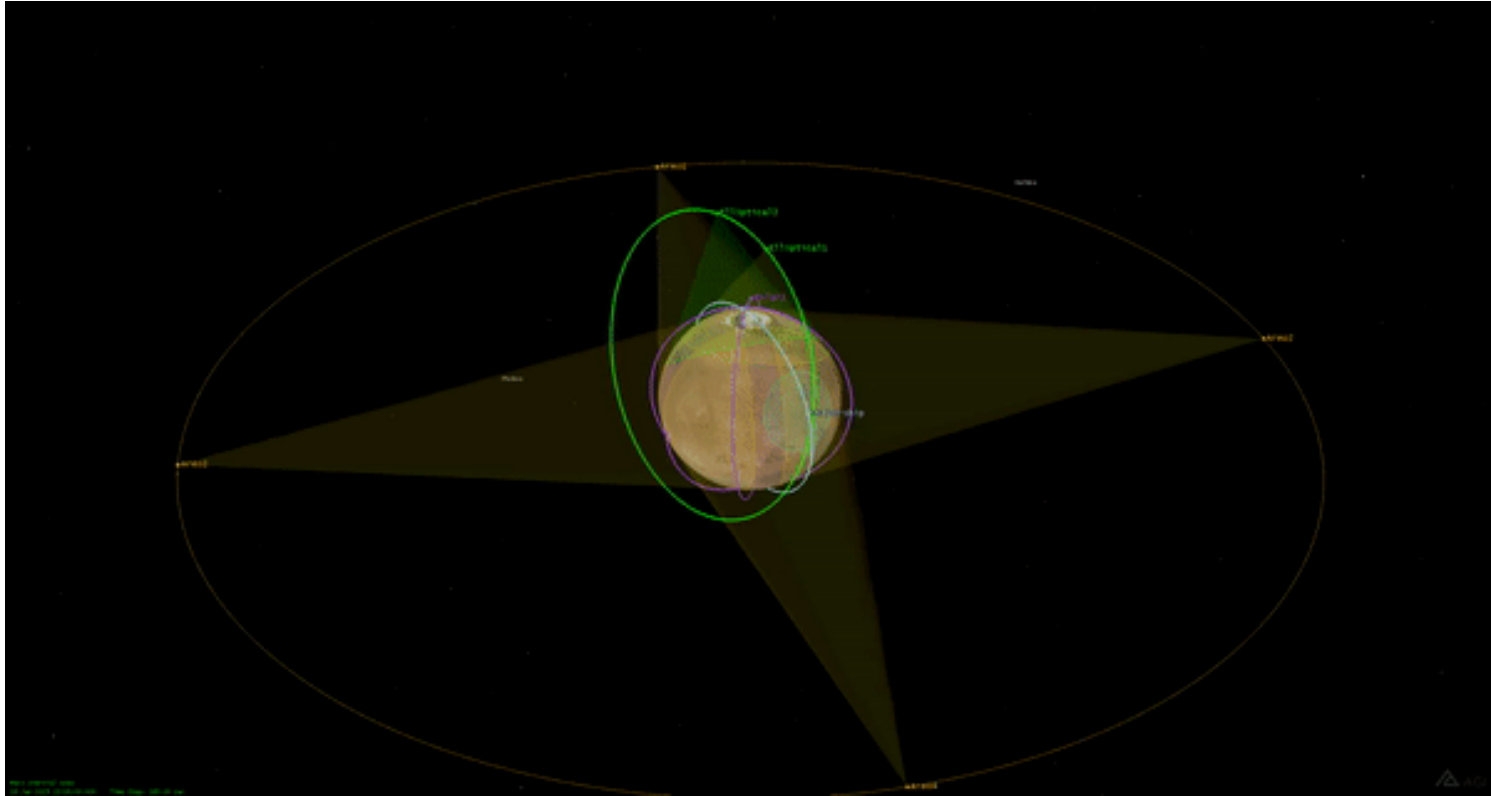
Descoped Mothership option



- Preserves constellation aspect, global/diurnal coverage, and lower-upper atmosphere connection.
- Loses all the newest measurements: wind and ice. But ***significantly*** cheaper.

Instrument	Inv	Priority		Mass	Power	FY20 Cost
Visible camera	1	T : 1	B : 1	3.4 kg	10 W	\$5 M
Thermal IR radiometer (MCS)	2	T : 1	B : 1	9 kg	18 W	\$25 M
Near IR spectrometer	3	T : 0	B : 1	0.3 kg	2.5 W	\$0.34 M
FUV/MUV spectrometer	4	T : 1	B : 1	27 kg	28 W	\$30 M
Radio occultation	3,5	T : 1	B : 1	1.5 kg	3 W	\$2 M
Total				40 kg	62 W	\$62 M

MOSAIC Constellation Movie



Note: embedded movie wouldn't play. Had to resort to lower resolution GIF.

MOSAIC Goals & Objectives: Exploration



MOSAIC GOALS	Mission Objectives
<i>II. Identify hazards, characterize resources, and demonstrate technologies to enable the Human Exploration of Mars.</i>	II.A: Characterize potentially extractable water ice resources to support in situ resource utilization
	II.B: Characterize the Mars atmospheric state with sufficient spatial sampling and cadence to allow accurate data assimilation and weather forecasting.
	II.C: Characterize the Mars ionospheric state and variability sufficiently to determine its likely disruptive effect on communications and positioning
	II.D: Demonstrate delay-tolerant networking and relay communication between Mars surface, Mars orbit, and Earth.
	II.E: Demonstrate high-bandwidth deep space communication between Earth and Mars.



MOSAIC 30-second summary



- **What?** *“MOSAIC will unveil the complete Martian climate system, from meso- to planetary scales and from subsurface to exosphere, while enabling hazard forecasting, communication, and resource prospecting for human exploration.”*
- **Why?**
 - A) *The interconnections between the surface, lower and upper atmospheres, and ionized environment, are stronger than previously thought and poorly understood.*
 - B) *Martian explorers will need water, air, fuel, weather forecasts and comms.*
- **How?** *5-10 coordinated satellites making simultaneous measurements of Mars’ climate system, many for the first time.*
- **When?** *post-2028*
- **How much?** *We’ll tell you soon but \$1-3B*
- **Are you crazy?** *Quite possibly. We’ll see.*



Thank you

Backup slides



MOSAIC org chart



Science Team

Thermosphere - Inv. 4, 8
Lead: **S. England**
J. Deighan, S. Bougher, A. Brecht

Magnetosphere/Escape -
Inv. 7,8
Lead: **S. Curry**
J. Luhmann, **D. Mitchell**, F. LeBlanc,
J. Halekas, D. Brain, X. Fang, J.
Espley, H. Opgenoorth, O. Vaisberg

Comms, Radio Science - Inv. 2, 3, 5, 9
Lead: **C. Ao**
P. Withers, **J. Vander Hook**, D. Hinson, O.
Karatekin, S. Asmar, M. Van Woerkum

Upper-Lower
Atmosphere
Connections - Inv.
2-8
Lead: **M. Chaffin**
S. England, **B. Jakosky**, **A.
Brecht**, **J. Deighan**,
F. Forget, **S. Bougher**
Leads: **R. Lillis**, **D. Mitchell**

Subsurface & Surface Ice - Inv. 1
Lead: **T. Harrison**
C. Neish, I. Smith, G. Osinski, C. Stuurman, S.
Spencer.

Lower-Middle Atmosphere – Inv. 2, 3
Leads: **L. Montabone**, **S.
Guzewich**
M. Kahre, **N. Heavens**, M. Smith, A. Spiga,
M. Mischna, M. Wolff, A. Kleinboehl, D.
Hinson, F. Forget, L. Tamppari, B. Cantor

Ionosphere - Inv. 5, 7, 8, Lead: **P. Withers**
R. Lillis, D. Andrews, M. Paetzold, S. Tellmann, K. Peter,
C. Fowler, M. Lester, B. Sanchez-Cano

JPL Study Team

Lead: **S. Matousek**, Systems: **Nathan Barba**, Mission design: **Ryan Woolley**
Plus lots of JPL folks from A-Team, Team-X, and Team-Xc

PMCS Study Plan/schedule



1. Define Science Requirements (10-11/19) ✓
2. Define Instrument Requirements (11-12/19) ✓
3. JPL A-Team Session I: building blocks (12/17/19) ✓
4. Science Team meeting Berkeley (01/27/20) ✓
5. JPL A-Team Session II: narrow architecture (02/5-6/20) ✓
6. JPL Team Xc Smallsat Designs & Cost (02/18-26/20) ✓
7. LPSC Progress Report (03/15/20) ✓
8. JPL Team X Point Design & Cost (03/31 – 04/03/20) ✓
9. Writing.
10. Final report July 2020.