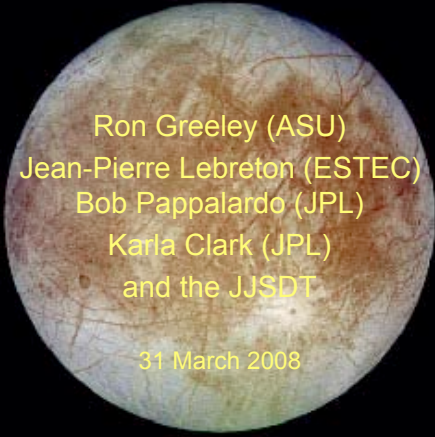


# Preliminary Report of the Joint Jupiter SDT



Ron Greeley (ASU)  
Jean-Pierre Lebreton (ESTEC)  
Bob Pappalardo (JPL)  
Karla Clark (JPL)  
and the JJSdT

31 March 2008



## Joint Jupiter Science Definition Team


- |  |  |
|--|--|
| <p><b>US JJSdT Membership</b></p> <ul style="list-style-type: none"> <li>• Ron Greeley</li> <li>• Bob Pappalardo</li> <li>• Anbar, Ariel</li> <li>• Bills, Bruce</li> <li>• Blaney, Diana</li> <li>• Blankenship, Don</li> <li>• Christensen, Phil</li> <li>• Dalton, Brad</li> <li>• Deming, Jody</li> <li>• Greenberg, Rick</li> <li>• Hand, Kevin</li> <li>• Hendrix, Amanda</li> <li>• Khurana, Krishan</li> <li>• McCord, Tom</li> <li>• McGrath, Melissa</li> <li>• Moore, Bill</li> <li>• Moore, Jeff</li> <li>• Nimmo, Francis</li> <li>• Paranicas, Chris</li> <li>• Prockter, Louise</li> <li>• Schubert, Jerry</li> <li>• Senske, David</li> <li>• Showman, Adam</li> <li>• Sogin, Mitch</li> <li>• Spencer, John</li> <li>• Waite, Hunter</li> </ul> | <p><b>ESA JJSdT Membership</b></p> <ul style="list-style-type: none"> <li>• Jean-Pierre Lebreton</li> <li>• Blanc, Michel</li> <li>• Drossart, Pierre</li> <li>• Grasset, Olivier</li> <li>• Hußman, Hauke</li> <li>• Krupp, Norbert</li> <li>• Prieur, Daniel</li> <li>• Tosi, Federico</li> <li>• Wurz, Peter</li> </ul> |
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- |  |   |
|--|---|
| <p><b>Co-Chair</b></p> <ul style="list-style-type: none"> <li>• Study Scientist</li> <li>• astrobiology</li> <li>• geophysics</li> <li>• composition</li> <li>• radar/geophysics</li> <li>• composition</li> <li>• composition</li> <li>• astrobiology</li> <li>• geophysics</li> <li>• astrobiology</li> <li>• satellites</li> <li>• fields &amp; particles</li> <li>• composition</li> <li>• satellites</li> <li>• geophysics</li> <li>• geology</li> <li>• geophysics</li> <li>• fields &amp; particles</li> <li>• geology</li> <li>• Jupiter</li> <li>• satellites</li> <li>• Jupiter</li> <li>• astrobiology</li> <li>• satellites</li> <li>• fields &amp; particles</li> </ul> | <p><b>Co-Chair</b></p> <ul style="list-style-type: none"> <li>• Lead Scientist</li> <li>• Jupiter</li> <li>• geology</li> <li>• geophysics</li> <li>• fields &amp; particles</li> <li>• astrobiology</li> <li>• origins</li> <li>• origins/instruments</li> </ul> |
|--|---|
- 
- |   |   |
|---|---|
| <p><b>ESA Cross-Disciplinary</b></p> <ul style="list-style-type: none"> <li>• Lorenzo Bruzzone</li> <li>• Paolo Tortora</li> <li>• Ingo Mueller-Wodarg</li> <li>• Frank Sohl</li> <li>• Olga Prieto-Ballasteros</li> <li>• Michele Dougherty</li> </ul> | <ul style="list-style-type: none"> <li>• radar/geophysics</li> <li>• radio science</li> <li>• fields &amp; particles</li> <li>• geophysics</li> <li>• astrobiology</li> <li>• fields &amp; particles</li> </ul> |
|---|---|
- 
- JJSdT Meetings**

  - Feb. 27-29, JPL: Europa Orbiter objectives
  - March 26-28, JPL: US-EU science integration
  - April 23-24, Rome: EO-IPO complementary payloads
  - May 26-29, JPL: Outstanding issues

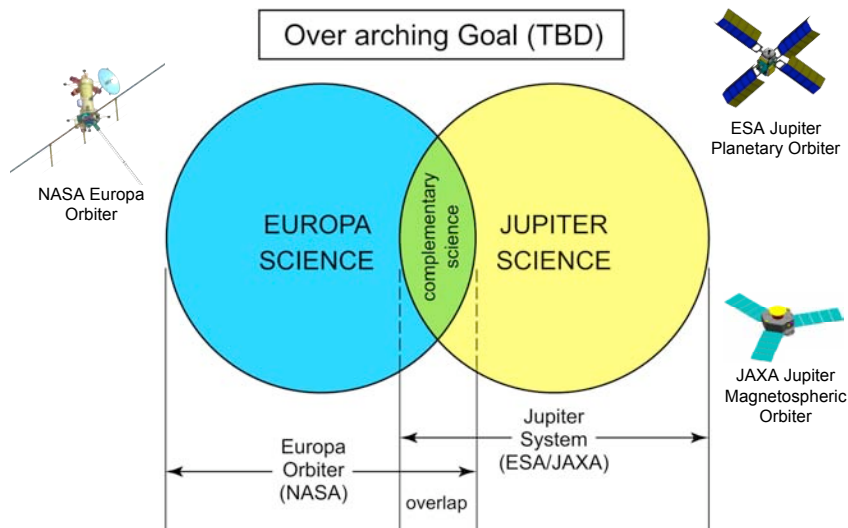
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## Europa-Jupiter System Mission



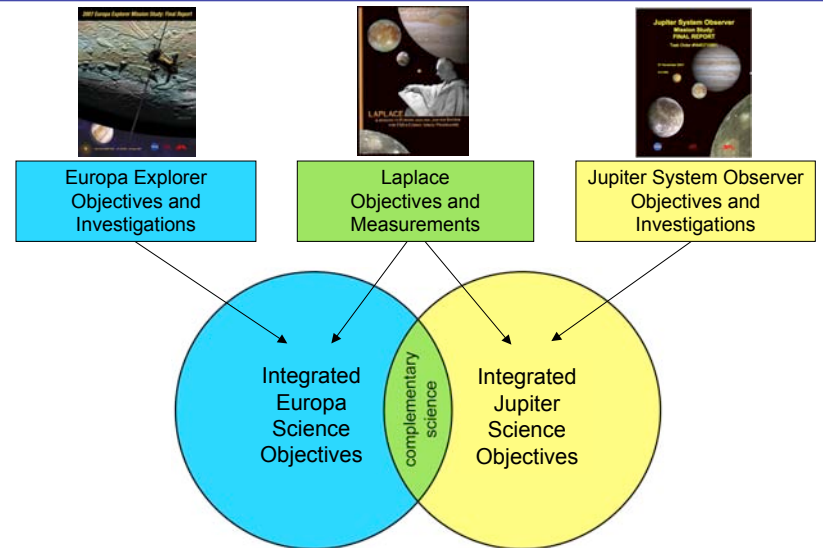
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## JJSdT Science Integration Process



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## Europa Orbiter: Prioritized and Integrated Science



| Goal   | Science Objective  | Science Investigation   |
|--|--|---|
| Explore Europa to investigate its habitability.                              | A. Ocean<br>Characterize the extent of the ocean and its relation to the deeper interior.  | A1. Determine the amplitude and phase of the gravitational tides.   |
|  |  | A2. Characterize the magnetic environment (including ionospheric current plasma effects) to determine the induction response from the ocean over multiple frequencies.                  |
|  |  | A3. Characterize surface motion over the tidal cycle.   |
|  |  | A4. Determine the satellite's dynamical rotation state.   |
|  |  | A5. Investigate the core, rocky mantle, and rock-ocean interface.   |
|  | B. Ice<br>Characterize the ice shell and any subsurface water, including their heterogeneity, and the nature of surface ice-ocean exchange.  | B1. Characterize the distribution of any shallow subsurface water.  |
|  |  | B2. Search for an ice-ocean interface.  |
|  |  | B3. Correlate surface features and subsurface structure to investigate processes governing communication among the surface, ice shell, and ocean.                                       |
|  |  | B4. Characterize regional and global heat flow variations.  |
|  | C. Chemistry<br>Determine global surface compositions and chemistry, especially as related to habitability.  | C1. Characterize surface organic and inorganic chemistry, including abundances and distributions of materials, with emphasis on indicators of habitability.                             |
|  |  | C2. Relate compositions to geological processes, especially communication with the interior.  |
|  |  | C3. Characterize the global radiation environment and the effects of radiation on surface composition, atmospheric composition, albedo, weathering, sedimentation, and redox chemistry. |
|  | D. Geology<br>Understand the formation of surface features, including sites of recent or current activity, and identify and characterize candidate sites for future in situ exploration. | C4. Characterize the nature of exogenic materials.  |
|  |  | D1. Determine the formation history and three-dimensional characteristics of magmatic, tectonic, and impact landforms.  |
|  |  | D2. Determine sites of most recent geological activity, and evaluate future landing sites.  |
| E. Jupiter system<br>Understand Europa in the context of the Jupiter system. | D3. Investigate processes of erosion and deposition and their effects on the physical properties of the surface debris.  |   |
|  | E1.  |   |
|  | E2.  |   |
|  | E3.  |   |
|  | E4.  |   |
| E5.  |  |   |

IN PROGRESS

- Objectives are now prioritized.
- Chemistry objective's traceability is receiving special attention.
- Former "External Environment" objective is folded into the others.
- Jupiter system science is elevated to Level 1 priority.
- US Europa Explorer and Laplace Europa objectives are now integrated.

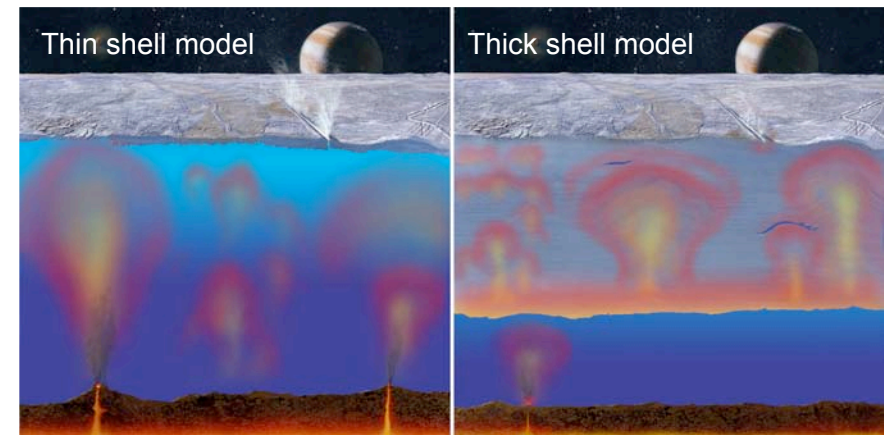
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## Example of Europa Hypothesis Testing: Thin vs. Thick Ice Shell



Data from multiple instruments combine to test fundamental hypotheses:  
Gravity, altimetry, radar sounding, thermal, imaging.

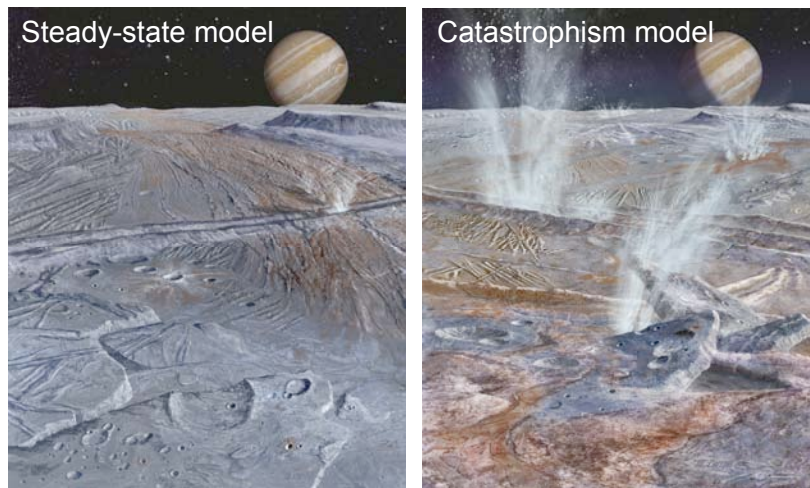
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## Example of Europa Hypothesis Testing: Catastrophic vs. Steady-State Evolution



90% of Europa remains unseen at resolution needed to recognize key units.

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## Jupiter Polar Orbiter: Prioritized and Integrated Science



### III. Study the Jovian magnetodisk/magnetosphere

| Science Objective  | Required Measurements  |
|--|--|
| <b>Jupiter as a fast magnetic rotator</b>  |  |
| 1. Magnetodisk structure, dissipation of rotational energy, transfer of angular momentum | 1.1 Characterization of the 3D properties of the magnetodisk   |
|  | 1.2 Analysis of the influence of the various plasma sources on its structure (including composition measurements)  |
|  | 1.3 Response to mass loading variability (Ionian space weathering)   |
|  | 1.4 Characterization of the different modes of transport   |
|  | 1.5 Determination of the plasma loss processes   |
| 2. Global energy regulation in a complex magnetosphere                                   | 1.6 Boundary layers: Investigate the important boundary layers of the magnetosphere (Bow shock, magnetopause, plasma sheet boundary layer, high energy particle radiation belt boundaries, equatorial plasma disk) |
|  | 2.1 Characterization of the Magnetosphere/Ionosphere/thermosphere coupling processes   |
|  | 2.2 Morphology & modulation of the auroral/radio emissions (multi- $\lambda$ ) and modulation of particles and fields  |
| 3. Satellites/Magnetosphere interactions   | 2.3 Magnetospheric mapping of auroral/radio features/regions   |
|  | 2.4 Response to solar wind variability (Jovian space weather)  |
|  | 3.1 Characterization of the local electrodynamics interaction  |
|  | 3.2 Observations of the moon auroral magnetic footprints   |
|  | 3.3 Study of pick-up & charge-exchange processes in plasma/neutral tori  |
| 4. Origin & effects of its harsh radiation environment                                   | 3.4 Analysis of plasma/surface sputtering processes/surface weathering   |
|  | 3.5 Investigation of Ganymede's magnetosphere  |
| <b>Jupiter as a giant particle accelerator</b>   |  |
| 4. Origin & effects of its harsh radiation environment                                   | 4.1 Characterization of high-energy electron, proton and heavy ion properties  |
|  | 4.2 Identification of high-energy particle acceleration processes (reconnection)   |
|  | 4.3 Determination of high-energy particle loss processes   |
|  | 4.4 Analysis of moon micro-signatures to quantify fundamental processes  |
|  | 4.5 Analysis of high-energy electrons synchrotron emissions  |
|  | 4.6 Radiation effects in the inner magnetosphere on ring system and dust population  |
|  | 4.7 Particles from Jupiter in interplanetary space   |

- US Europa Explorer and Laplace Europa objectives are being integrated.

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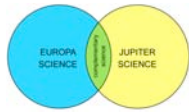
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# Jupiter System Science (in progress)

- Jupiter System Science Objective for EO:
  - Understand Europa in the context of the Jupiter system.
- Overall Jupiter System Science Philosophy:
  - EO adopts investigations which fit under its "Europa habitability" goal.
  - JPO adopts all relevant Jupiter System Science investigations.
  - Each spacecraft might achieve investigations "bookkept" by the other.
- Candidate EO Investigations:
  - Understand the sources and sinks of Io's crustal volatiles and atmosphere.
  - Characterize the composition, variability, and dynamics of Europa's atmosphere and ionosphere.
  - Identify the dynamical processes that cause internal evolution and near-surface tectonics of Ganymede vs. Callisto.
  - Characterize the abundance of Jupiter minor atmospheric species to understand the origin of the Jovian system.
- Possible EO Implementation:
  - Io campaign (3 - 4 Io fly-bys). • Ganymede fly-bys (~8 - 10).
  - Perijove on Jupiter's sunlit side. • Polar Callisto fly-by ( $\geq 1$ ).



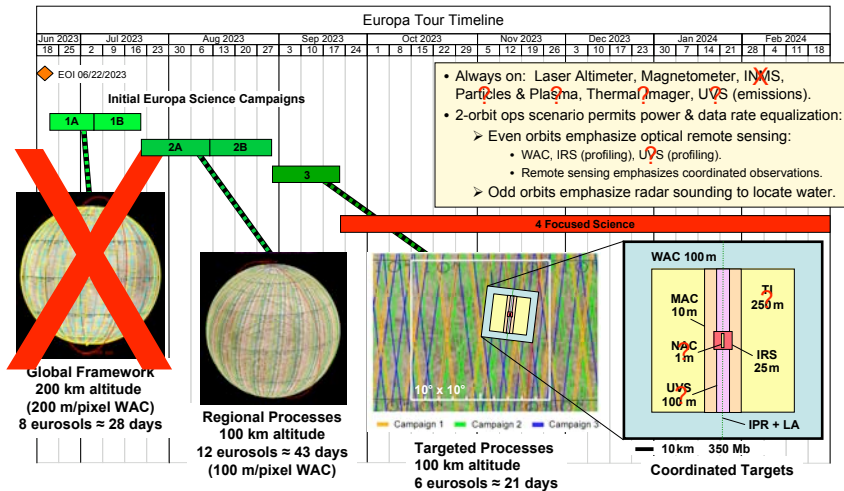
# EO Planning Payload to fit \$2.1B Cost Box

| CORE PAYLOAD                                   | Mass (kg) | Power (W)  | Change from EE07 Floor     |
|--|-----------|------------|----------------------------|
| <b>Ocean</b>                                   |           |            |                            |
| Radio Science (RSS)                            |           |            | Ka + X bands               |
| Laser Altimeter (LA)                           | 7         | 15         |                            |
| <b>Ice</b>                                     |           |            |                            |
| Ice Penetrating Radar (IPR)                    | 31        | 45         | Two dipoles                |
| <b>Chemistry</b>                               |           |            |                            |
| Near-IR Spectrometer (IRS)                     | 20        | 20         | increased mass for realism |
| <b>Geology</b>                                 |           |            |                            |
| Wide- + Med-Angle Camera (WAC+MAC)             | 10        | 12         | Both have color filters    |
| <b>Mag &amp; Plasma</b>                        |           |            |                            |
| Magnetometers + Plasma (MAG + Langmuir Probe?) | 5         | 3          | Dual mags on a 10 m boom   |
| <b>TOTALS:</b>                                 |           |            |                            |
|  | <b>73</b> | <b>95</b>  |                            |
| <b>2007 Floor TOTALS:</b>                      |           |            |                            |
|  | <b>77</b> | <b>106</b> |                            |

| <b>NOT IN CORE PAYLOAD</b>        |    |    |  |
|-----------------------------------|----|----|--|
| Thermal Instrument (TI)           | 5  | 5  |  |
| Particle Instrument (PI)          | 6  | 3  | 10s keV ~ 1.5 MeV (e-)<br>~ 10s MeV (ions) |
| UV Spectrometer (UVS)             | 5  | 10 | 1 channel only                             |
| Narrow Angle Camera (NAC)         | 15 | 12 |  |
| Ion & Neutral Mass Spectr. (INMS) | 15 | 28 |  |



# Implications of Only 60 Days In Europa Orbit

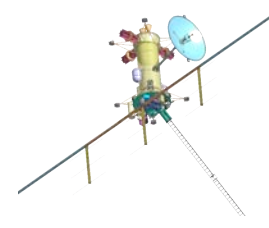


- Campaign 1 is lost, with corresponding loss to science and increased risk.
- Second-order optimizations are possible, e.g. reducing polar overlaps.

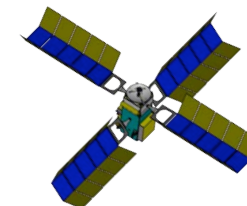


# Europa-Jupiter System Mission

- Joint NASA/ESA Mission Study
  - Multi element architecture
  - Independent launches allow decoupled development schedules
  - Possible JAXA contribution being evaluated



Europa Orbiter - NASA-led



Jupiter Planetary Orbiter/Ganymede Orbiter - ESA-led



Jupiter Magnetospheric Orbiter - JAXA-led

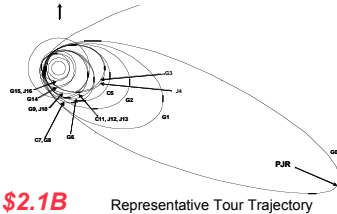
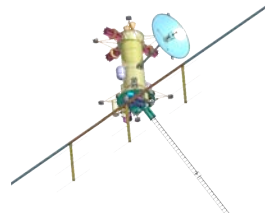
- Rest of discussion is on Europa Orbiter portion of Study



## 2007 Europa Explorer Floor Mission Concept Updated for 2008 Key Ground Rule Changes



- Concept: Europa orbiter with Galilean Satellite Tour
- Launch Vehicle: Atlas 531 => Atlas V551
- Power Source: 5 ASRGs => 5 MMRTG
- Mission Timeline:
  - Launch: ~6/2015 (VEEGA) => 2016/7 (TBD trajectory)
  - Jupiter arrival: ~7/2021 => 2021-2024
  - Galilean satellite tour science: ~2 yr
  - Europa orbital science: ~ 6 months => 60 days
    - 70m stations for 92 days 24/7, then single pass/day => 34m, no arraying, continuous
  - Spacecraft operates until loss of control; final disposition: Europa surface Impact
- Instruments: 8 ; 77 kg, 106 W
- Modified Reserves Base
- Cost: \$2.4B => **\$2.45B**



**The key challenge! Needs to be \$2.1B**



## Europa Orbiter FY08 Mission Specific Tasks



- Incorporate *Jupiter System Science* as Level 1 Objective
- Refine the *chemistry* science objective especially in relation to habitability
- Perform analyses concerning radiation induced effects on instrument measurement quality and mitigation strategies
- Investigate opportunities for *international partnerships* within the \$1B for contributions
- Design and characterize the sensitivity to the design point of a *60 day orbital* mission in terms of cost, mass, science return and other factors
- *Refine radiation plan* in 2007 report
- *Execute* the revised *radiation plan* including:
  - Establishment of all acceptable radiation related lifetime or performance criteria
  - Demonstration of the ability to reach these goals on representative parts including detectors
- Develop a specific *Preferred Parts list* for all hardware and permit only highly justified and well mitigated exceptions



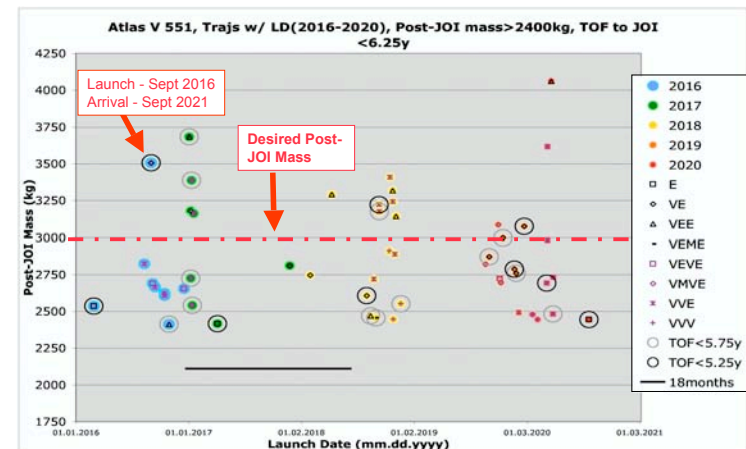
## Challenges for 2008



- Fitting into the \$2.1B box
  - Re-evaluating required management and system engineering workforce
    - Comparison to other missions
    - Evaluating alternate approaches
  - Potential options for Phases B/C/D
    - Simplify spacecraft design
    - Delete MMRTG
  - Potential options for Phases E/F
    - Shorten flight times
    - Simplify operational scenarios to reduce required staff
- Mitigating Radiation Risk
  - Created and began executing radiation plan
  - Identifying, evaluating and testing potential parts and materials
- Preparing potential instrument providers for Announcement of Opportunity (AO)
  - Workshops
  - Design guidelines
  - On-line tutorials



## Chemical Trajectory Performance (by launch year and trajectory type)



**2016 and 2017 launch opportunities exist which would allow arrival at Europa ~ 7 years after launch**



# Arrival Strategy Options



|  | Ganymede |     | Io Once |     | Io Campaign |
|--|----------|-----|---------|-----|-------------|
| Post-JOI Period (days)                               | 200      |     | 200     |     | 200         |
| Number of Io flybys                                  | 0        |     | 1       |     | 4           |
| Perijove @ JOI (Rj)                                  | 12.5     | 9.0 | 5.2     |     | 5.2         |
| Post-PJR Perijove (Rj)                               | 12.5     | 9.0 | 12.0    | 9.0 | 5.2         |
| JOI + PJR ΔV (Rj)                                    | 1010     | 910 | 920     | 820 | 675         |
| ΔTID over EE07 (krad)                                | -        |     | ~50     |     | ~300        |
| ACS sensor high dose-rate accommodation impacts (kg) | -        |     | 10-20   |     | 10-20       |
| ΔV-EGA compatible?                                   | N        |     | N       |     | 2016 only   |

ACS - Attitude Control Subsystem  
 JOI - Jupiter Orbit Insertion  
 PJR - Perijove Raise Maneuver  
 Rj - Jovian radii  
 TID - Total integrated dose

- Major dose-rate increase at Io over that at Europa
- The tour duration is independent of moon assist
- Mass trade looks favorable for conducting an Io campaign



# Major Operations Cost Driver Categories



- Capture relevant lessons learned from the past and present operational missions (Cassini, MRO, New Horizons, MESSENGER), in the area of Phase E cost drivers and operations.
- Each of the following categories are being further expanded with metrics defined for each:
  - Operational processes and architectures
  - Flight system interface complexity (bus and Payload)
  - Ground system interface complexity
  - Flight resource limitations
  - Science observation density and complexity
  - Mission design complexity

Results of analysis will be used to lower operating costs for Europa Orbiter Mission



# Risk Assessment: Early Focus Areas



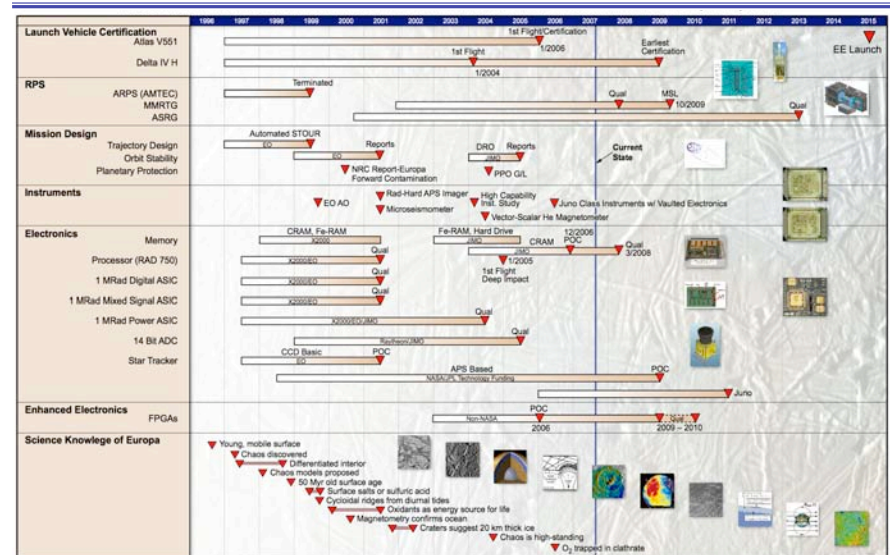
## Key Risk: Impact of radiation and planetary protection on design

| Risk Area            | Components   | Mitigation   | Impact  |
|----------------------|--|--|---|
| Radiation            | a) Dose rate effects<br>b) Sensor impacts (SNR)<br>c) FPGA qualification<br>d) Non-Volatile Memory capability<br>e) Internal Electrostatic Discharge<br>f) Design techniques | a) Quantify dose rate effects<br>b) Use ASICs in place of FPGAs<br>c) FPGA, memory and sensor radiation testing<br>d) Document and disseminate design techniques and guidelines<br>e) Early subject matter expert engagement | a) Reduced cost risk and uncertainty  |
| Planetary Protection | a) Sensor sterilization capability<br>b) Design techniques   | a) Document design techniques and guidelines<br>b) Early subject matter expert engagement  | a) Reduced cost risk and uncertainty  |
| Instrument Maturity  | a) Level of information available for potential providers<br>b) Inexperience of potential providers<br>c) Development schedule   | a) Document design techniques and guidelines<br>b) Instrument provider workshops - early subject matter expert engagement<br>c) Early and streamlined AO with confirmation review  | a) Maximize time instruments can work with experts<br>b) Reduce cost risk and uncertainty at 'commitment' |

Analogous design/development approaches will be assessed and redesigned to meet radiation and planetary protection requirements



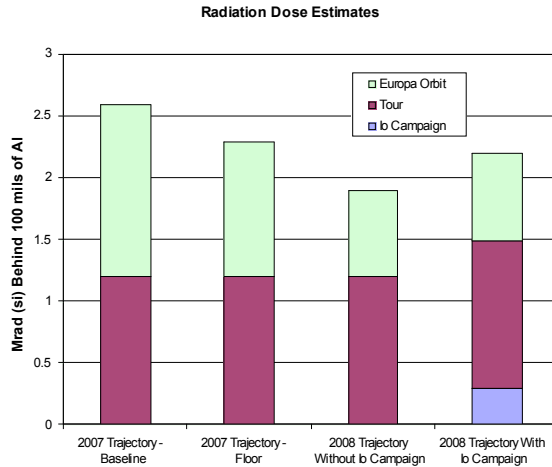
# A Decade of Investment Has Reduced EO Risk





# What's Changed for EE Radiation?

- 2007 baseline mission total => ~2.6 Mrad\*
- ~1.2 Mrad\* when enter Europa orbit
- ~1.4 Mrad\* for @ Europa
- Reducing to 60 days at Europa reduces radiation estimate
- Io campaign would add ~0.3 Mrad\*

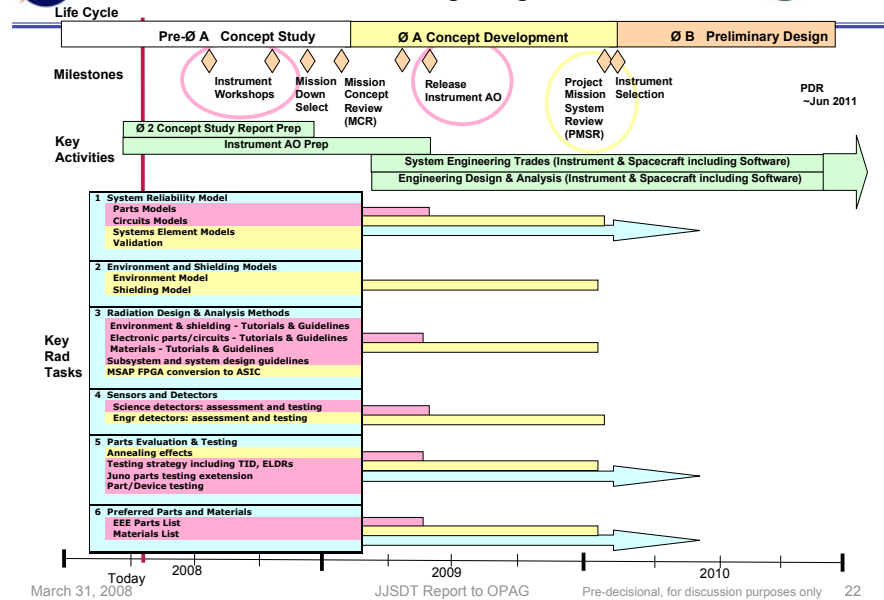


Radiation level estimates will be ~2.2 +/- 0.3 Mrad\*

\* Radiation levels normalized to behind 100 mils of Al  
March 31, 2008



# Near Term Plan for Mitigating Radiation Risk



## Instrument Workshops

- AO release target ~April 2009
- Issues:
  - Europa radiation levels pose significant challenges to the performance of the scientific payloads
  - Planetary protection requirements will demand dry heat microbial reduction
  - Parts/material selection & circuit design for these harsh environments are not business as usual
  - Detectors will suffer performance degradation w/time and are susceptible to transient noise during science operations
- Need to communicate environmental & PP issues before the release of the AO. Benefits include:
  - Proposals with more cost & schedule realism
  - Ensure high-quality science return for the mission

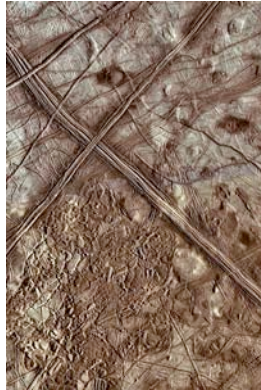


## Instrument Workshops - continued

- Communication strategy
  - Hold community workshops as information becomes available
  - e-tutorials on a community website based on the workshop content
- Current plans are to hold 3 workshops over the next year
- Outer Planet Flagship Instrument Workshop (6/3–5/08)
  - Joint EJSM/TJSM workshop
  - Overviews of mission architectures, science, notional payloads, science scenarios, etc
  - Split sessions that address Europa- and Titan-specific payload issues
  - Poster sessions for instrument providers
- EO Radiation & Planetary Protection Design Guidelines Workshop
  - ITAR restricted to allow for discussion of design details
- Final EO workshop to update design guidelines just prior to AO release



- Fruitful technology investments:
  - Launch vehicles, RPS, Mission Design, Instruments, Electronics.
- Radiation challenges have been met:
  - State-of-the-Art technology and comprehensive system engineering design approach identified, validated and being executed
  - Plan for communicating approaches to potential instrument providers formed
- New Study Guidelines require changes to mission concept and implementation to reduce estimated cost
  - Challenge previous assumptions
  - Simplify mission implementation
  - May require descopes to science as last resort
- Mature and successful operations plan:
  - Updates for shorter Europa Orbital mission being explored
  - Tour science being incorporated.
- Rich science return:
  - For Europa, Jupiter, and the Galilean satellites.
- The first icy satellite orbiter mission is ready to begin!



***The technical capability is mature, and the inspirational goal of the first icy satellite orbiter is within our reach.***