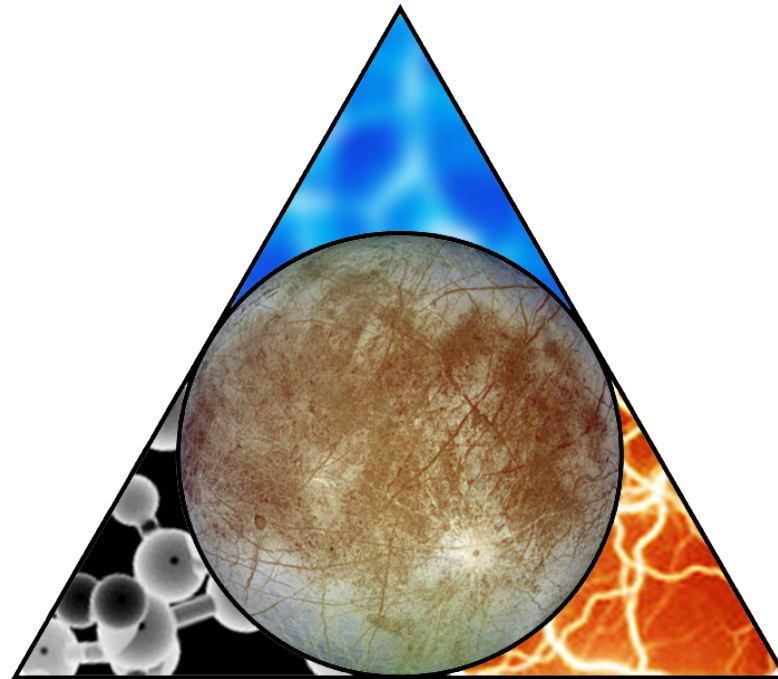
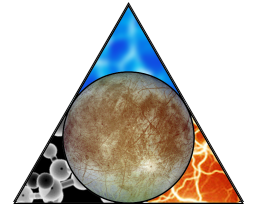


# Europa Study Update

## OPAG meeting, Pasadena, CA

### 10/19/2011

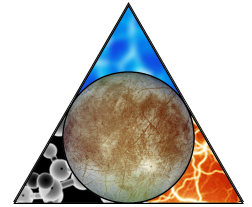


**Science**  
**Bob Pappalardo**  
**Europa Study Scientist and SDT Chair**

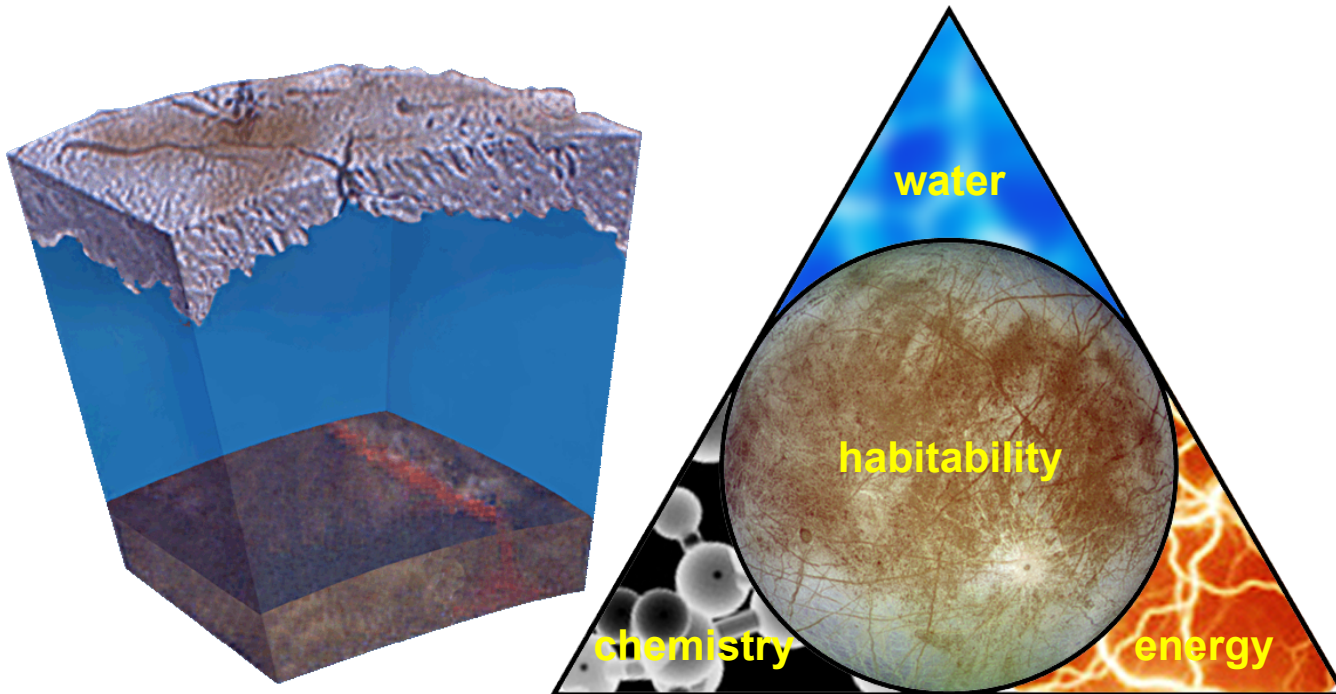
**OPAG meeting, Pasadena, CA**  
**10/19/2011**



# Ingredients for Life



- Three “ingredients” necessary for life:
  - *Water*: Solvent to facilitate chemical reactions
  - *Chemistry*: Constituents to build organic molecules
  - *Energy*: Disequilibrium to drive metabolism

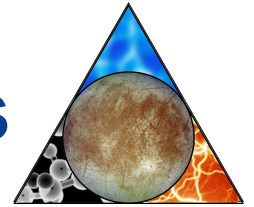


*“Because of this ocean’s potential suitability for life, Europa is one of the most important targets in all of planetary science.”*

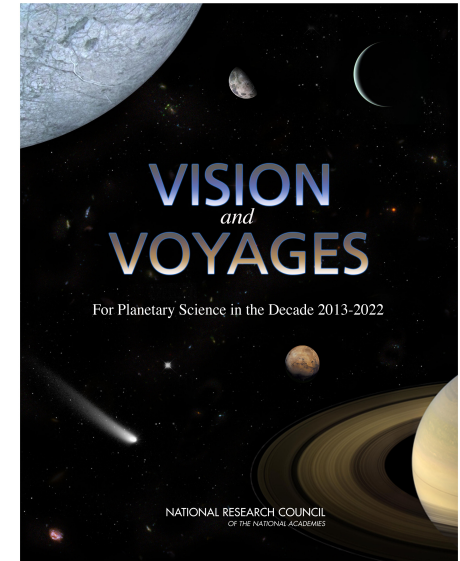
*–NRC Planetary Decadal Survey 2011*



# Decadal Survey Recommendations

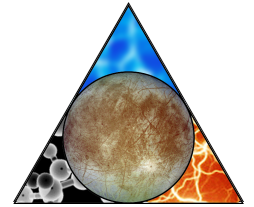


- “[B]oth **JEO and the Mars Sample Return campaign** (beginning with MAX-C) were found to have exceptional science merit.”
- JEO is “**second highest priority Flagship mission**” relative to MAX-C, based on “**pragmatic reasons** associated with the required spending profiles.”
- JEO’s “cost as currently designed is so high that **both a decrease in mission scope and an increase in NASA’s planetary budget** are necessary to make it affordable.”
- “**Possible pathways to lower cost** include use of a **larger launch vehicle that would reduce cost risk** by shortening and simplifying the mission design, and a **significant reduction in the science payload**. Other possible descopes were listed in section 4.1.5 of the 2008 JEO mission study final report.”
- For plutonium savings, **switch from MMRTG to ASRGs**
- MAX-C was chosen over JEO foremost “**to maintain programmatic balance by assuring that no one mission takes up too large a fraction of the planetary budget at any given time.**”

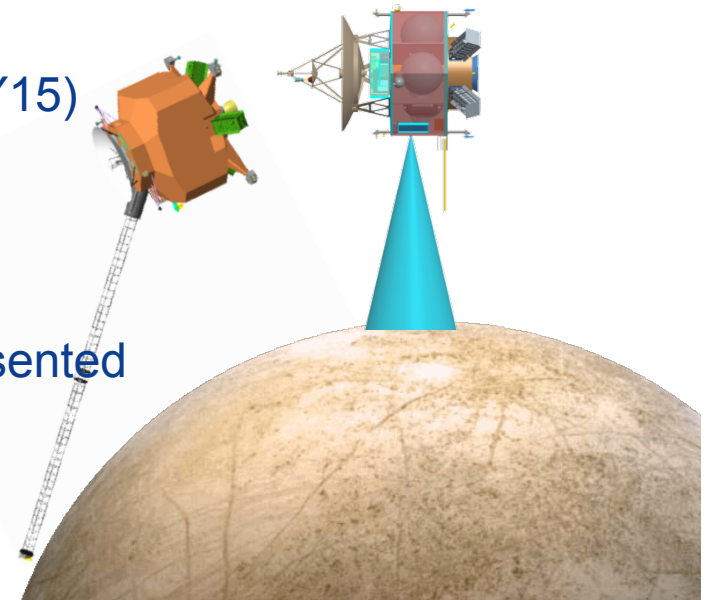




# Europa SDT Summary



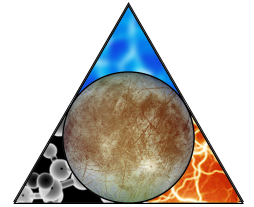
- To achieve the highest priority JEO Europa science objectives, invoke a two-element approach
  - Orbiter element to perform geophysical measurements (“Ocean” science) which can only be achieved from Europa orbit
  - Multiple-flyby element to perform remote measurements (“Chemistry” and “Energy” science) which can be achieved from Jupiter orbit
- Each achieves key science objectives, and each has very high science value of its own
  - Neither science nor element cost (~\$1.5B FY15) is a clear discriminator between elements
- The complementary elements would fly separately, and staggered in time
  - Anticipate the second element would be presented to the next Decadal Survey for consideration
- A landed element is now being studied by the Europa SDT at NASA’s request





# Europa Science Definition Team

## 2011-2012

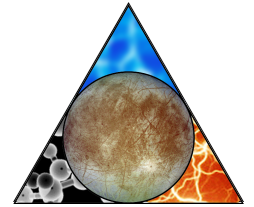


- |                      |                |                         |
|----------------------|----------------|-------------------------|
| • Fran Bagenal       | Univ. Colorado | Plasma                  |
| • Amy Barr           | SwRI           | Geophysics              |
| • Bruce Bills        | JPL            | Geophysics              |
| • Diana Blaney       | JPL            | Composition             |
| • Don Blankenship    | Univ. Texas    | Ice shell               |
| • Will Brinckerhoff* | GSFC           | Astrobiology            |
| • Jack Connerney     | GSFC           | Magnetometry            |
| • Kevin Hand*        | JPL            | Astrobiology            |
| • Tori Hoehler*      | Ames           | Astrobiology            |
| • Bill Kurth         | Univ. Iowa     | Plasma                  |
| • Melissa McGrath    | MSFC           | Atmosphere              |
| • Mike Mellon*       | Univ. Colorado | Ice Physics             |
| • Jeff Moore         | Ames           | Geology                 |
| • Bob Pappalardo     | JPL            | Chair / Study Scientist |
| • Louise Prockter    | APL            | Deputy / Geology        |
| • Dave Senske        | JPL            | Deputy / Geology        |
| • Everett Shock*     | ASU            | Geochemistry            |
| • David Smith        | MIT            | Geophysics              |

\*Recent SDT augmentations



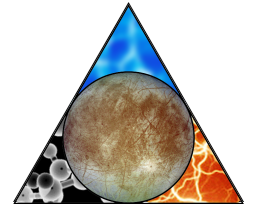
# Europa SDT Meetings



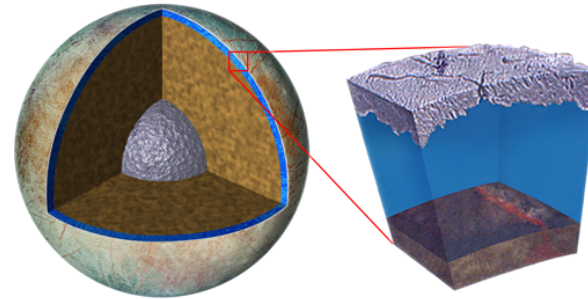
- May 2–3, 2011, JPL
  - Considered Europa objectives and mission design trades, and converged on two-element mission concept
- June 23–24, 2011, JPL
  - Provided feedback on initial orbiter and flyby mission designs, and iterated on model payload and mission requirements
- Aug. 22–23, 2011, JPL
  - Finalized science traceability, model payloads, and mission requirements
- Oct. 17–18, 2011, JPL
  - Developed initial objectives and investigations for a landed element
- *Upcoming*: Nov. 29–30, 2011, Boulder, Colo.
  - Will develop lander element model payload and associated mission requirements



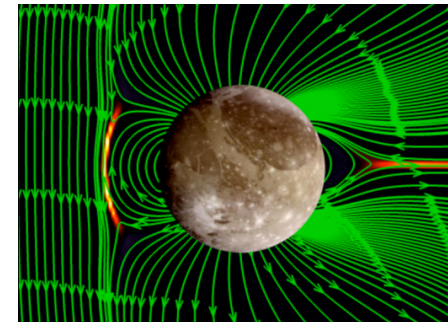
# Reduced from EJSM-JEO to a Europa-Only Focus



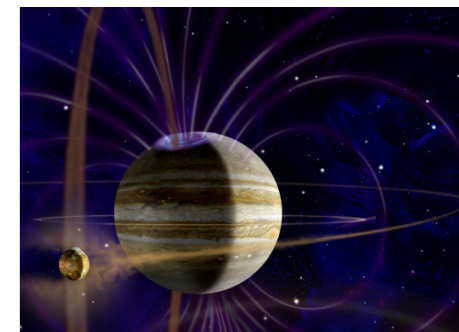
1. Europa:  
Explore Europa to investigate its habitability



2. Ganymede:  
Characterize Ganymede as a planetary object including its potential habitability



3. Jupiter System:  
Explore the Jupiter system as an archetype for gas giants

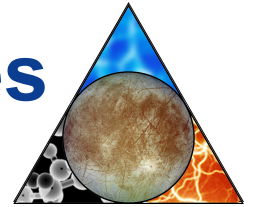






# Science Goal, Objectives, and Themes

Modified and Simplified from JEO

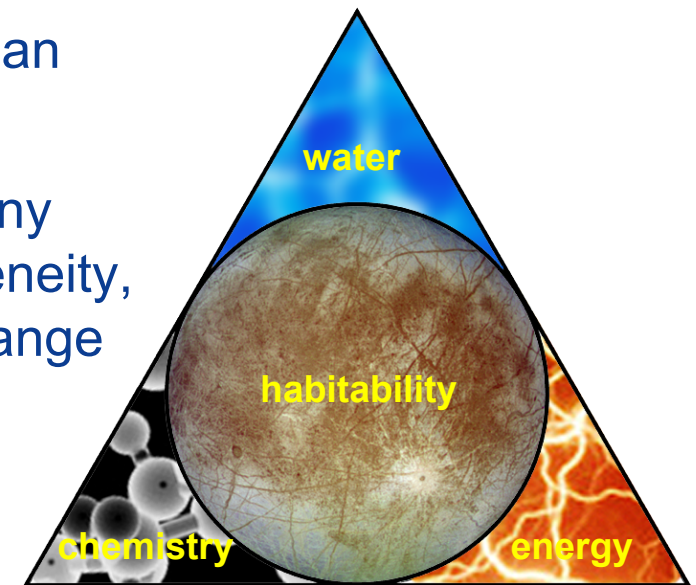


- *Goal:* Explore Europa to investigate its habitability

- *Objectives:*

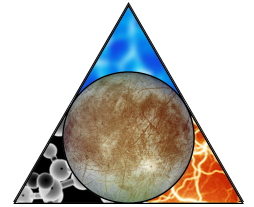
- *Ocean:* Characterize the extent of the ocean and its relation to the deeper interior
- *Ice Shell:* Characterize the ice shell and any subsurface water, including their heterogeneity, and the nature of surface-ice-ocean exchange
- *Composition:* Understand the habitability of Europa's ocean through composition and chemistry
- *Geology:* Understand the formation of surface features, including sites of recent or current activity, and characterize high science interest localities

*Themes:*





# Science as a Driver of Mission Architecture



Science traceability leads to a two element mission concept:



## **Orbiter Element:**

***Geophysical measurements that can be achieved only from orbit***

- Science focused primarily to address “Ocean” objective:
  - Gravity field
  - Tidal amplitude
  - Induction signatures
  - Plasma correction
  - Stratigraphic mapping



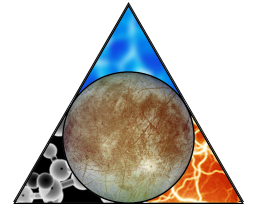
***Flyby Element: Remote measurements that can be accomplished via multiple flybys***

- Science focused primarily to address “Chemistry” and “Energy” themes:
  - Subsurface dielectric horizons
  - Surface constituents
  - Atmospheric constituents
  - Targeted landforms

- Each element achieves key Europa science objectives
- The elements are complementary, and each has very high science value of its own



# Two-Element Approach

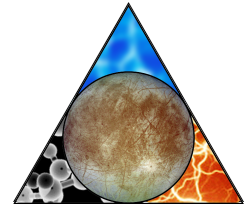


Objective	Europa Science	Orbiter	Multiple Flyby
Ocean	Gravity field	✓	
	Tidal amplitude	✓	
	Induction signatures	✓	
	Plasma correction	✓	
Ice Shell	Subsurface dielectric horizons		✓
Composition	Surface constituents		✓
	Atm. constituents		✓
Geology	Stratigraphic mapping	✓	
	Targeted landforms		✓



# Orbiter Element Traceability

## Ocean Emphasis



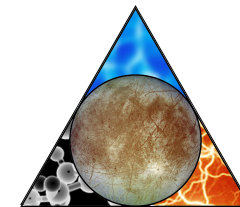
Goal	Objective	Investigation	Model Instruments	Theme		
				W	C	E
Explore Europa to investigate its habitability	Ocean	O.1 Determine the amplitude and phase of gravitational tides.	Radio subsystem, Laser altimeter	✓		
		O.2 Determine Europa's magnetic induction response.	Magnetometer, Langmuir probe	✓		
		O.3 Determine the amplitude and phase of topographic tides.	Laser altimeter, Radio subsystem	✓		
		O.4 Determine Europa's rotation state.	Laser altimeter, Mapping camera	✓		
		O.5 Investigate the deeper interior.	Radio subsystem, Laser altimeter, Magnetometer, Langmuir probe	✓		✓
	Geology	G.1 Understand the formation of surface features, including sites of recent or current activity, and characterize high science interest localities.	Determine the distribution, formation, and three-dimensional characteristics of magmatic, tectonic, and impact landforms.	Mapping camera, Laser altimeter	✓	

**Themes: W= Water, C = Chemistry, E = Energy**



# Flyby Element Traceability

## Chemistry & Energy Emphasis

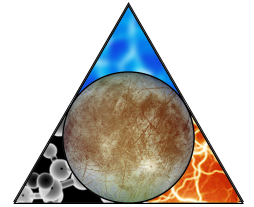


Goal	Objective	Investigation	Model Instr.	Theme				
				W	C	E		
Explore Europa to investigate its habitability	Ice Shell	I.1	Characterize the distribution of any shallow sub-surface water and the structure of the icy shell.	Radar sounder, Topo. Imager	✓		✓	
		I.2	Search for an ice-ocean interface.	Radar sounder, Topo. Imager	✓		✓	
		I.3	Correlate surface features and subsurface structure to investigate processes governing material exchange among the surface, ice shell, and ocean.	Radar sounder, IR spectrometer, Topo. imager	✓	✓	✓	
		I.4	Characterize regional and global heat flow variations.	Radar sounder	✓		✓	
	Composition	Understand the habitability of Europa's ocean through composition and chemistry.	C.1	Characterize the composition and chemistry of the Europa ocean as expressed on the surface and in the atmosphere.	IR spectrometer, INMS	✓	✓	
			C.2	Determine the role of Jupiter's radiation environment in processing materials on Europa.	IR spectrometer, INMS		✓	✓
			C.3	Characterize the chemical and compositional pathway's in Europa's ocean.	IR spectrometer, INMS	✓	✓	
Geology	Understand the formation of surface features, including sites of recent or current activity, and characterize high science interest localities.	G.2	Determine sites of most recent geological activity, and characterize high science interest localities.	Topo. Imager	✓		✓	

**Themes: W= Water, C = Chemistry, E = Energy**



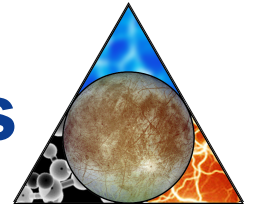
# Traceability Matrix Example



Goal	Objective	Investigation	Measurement	Example instrument	Mission constraints/ requirements	Notional Instrument Specifications	Notional Mission Type
Explore Europa to investigate its habitability.	Characterize the extent of the ocean and its relation to the deeper interior.	Determine the amplitude and phase of gravitational tides.	Measure degree two time dependent gravity field, to recover $k_2$ amplitude at Europa's orbital frequency to 0.003 absolute accuracy, and the phase to 1 degree.	Radio sub-system	(1) Low altitude (~100+ km), circular, near-polar (within 5° to 10° of the pole) orbit, for at least 30 days (baseline), 18 days (floor); (2) Range-rate measurements with an accuracy better than 0.1 mm/s at 60 sec integration time to determine spacecraft orbit to better than 1-meter (rms) in radial direction over several tidal cycles; (3) Several “unperturbed” days for the data arcs (preferably at least one rotation of Europa) for gravity-limit spacecraft momentum dumping or thrusting to an interval of 3 to 4 days, if possible.	Ka and X, both up and down (closed-loop)	Orbiter



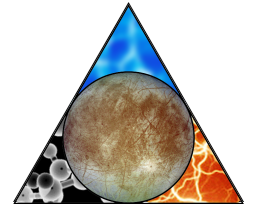
# Key Science Drivers & Requirements



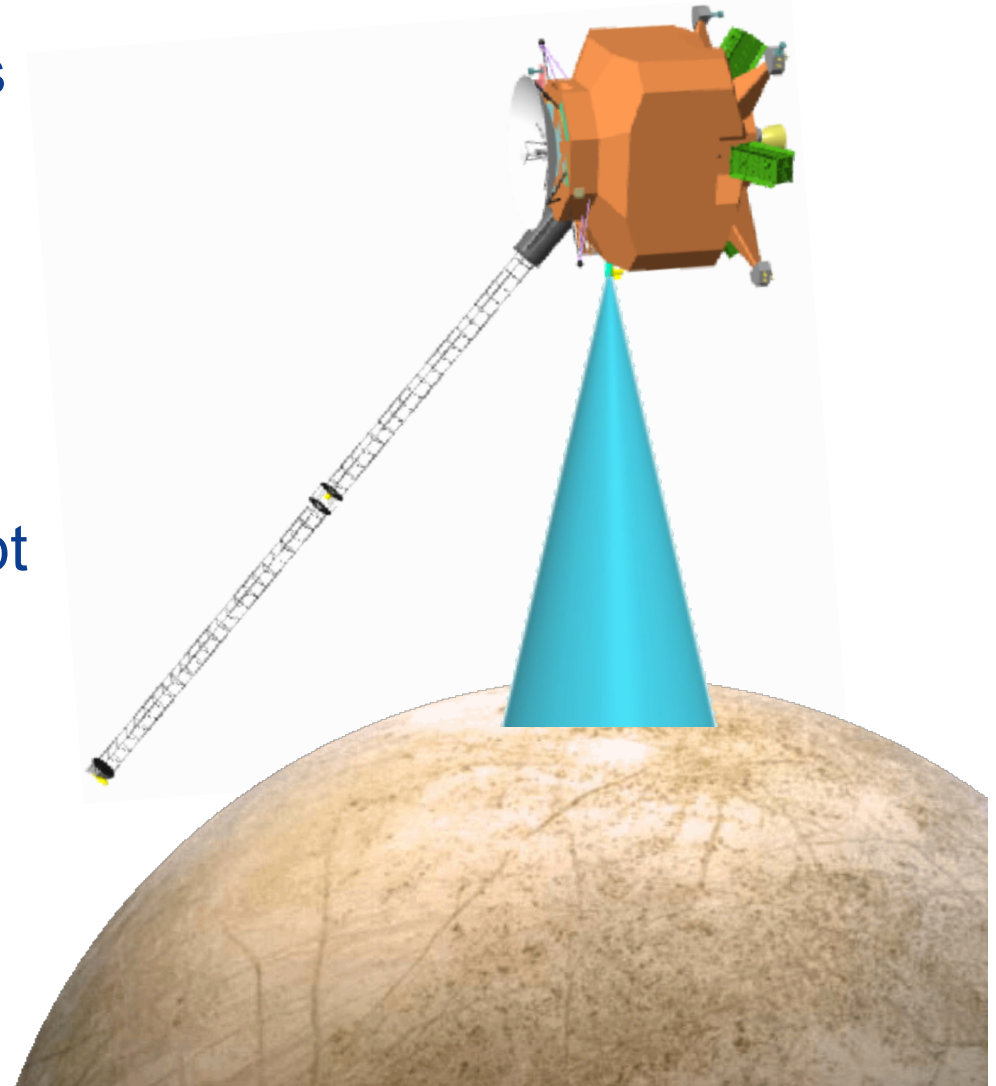
Instrument(s)	Key Accommodation Requirements	Element
<b>Radio Subsystem + Laser Altimeter</b>	Low altitude (~100+ km), circular, near-polar orbit, for at least 30 days. Unperturbed orbital arcs (no thrusting) of at least 3 days.	Orbiter
<b>Magnetometer + Langmuir Probe</b>	Low altitude (~100+ km), circular, near-polar orbit, for at least 30 days. Cover approximately 12 hours of Europa local time.	Orbiter
<b>Mapping Camera</b>	Low altitude (~100+ km), $\geq 80\%$ global coverage under near uniform lighting conditions, solar incidence angle $> 45^\circ$ ( $70^\circ$ preferred).	Orbiter
<b>Ice Penetrating Radar</b>	$\geq 800$ km tracks in 11 of 14 globally distributed regions, intersected by at least 1 other track, with track lengths measured from $\leq 400$ km alt. $\sim 25$ – $100$ km closest approach at $\leq 6$ km/s.	Flyby
<b>ShortWave IR Spectrometer</b>	$\geq 70\%$ coverage at $\leq 10$ km per pixel. Ability to target specific geologic locations with a wide range of surface locations, lighting between 9:00am and 3:00pm. Attitude stability $< \frac{1}{2}$ IFOV over integration time, flyby speed $< 6$ km/s.	Flyby
<b>Topographical Imager</b>	High resolution stereo imagery aligned with IPR coverage; lighting conditions solar incidence angle $> 20^\circ$ ( $70^\circ$ preferred).	Flyby
<b>Ion and Neutral Mass Spectrometer</b>	Low altitude ( $< 200$ km with lower altitudes desired) at $\leq 7$ km/s; long integration times and low altitudes ( $\leq 100$ km) preferred.	Flyby



# Orbiter Model Payload: Outline



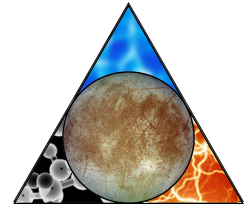
- Model payload instruments
  - Radio Subsystem (RS)
  - Laser Altimeter (LA)
  - Magnetometer (MAG)
  - Langmuir Probe (LP)
  - Mapping Camera (MC)
- Science operations concept
- Example coverage





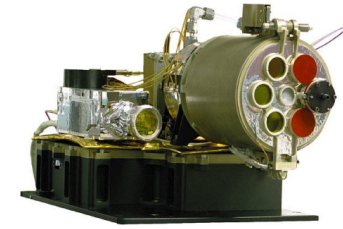


# Orbiter Model Payload Instruments

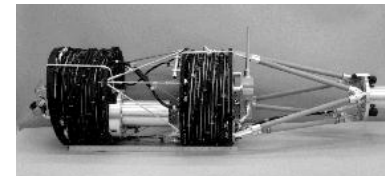


- Radio Subsystem (RS)
  - X-band up and down; Ka-band down only
  - Ka Transponder
- Laser Altimeter (LA)
  - Nadir view, co-boresighted with camera on 2-axis gimbal platform
- Magnetometer (MAG)
  - Dual 3-axis fluxgate
  - Sensors on boom 5 m and 10 m from S/C
- Langmuir Probe (LP)
  - Two 5 cm diameter spheres mounted on 1 m long booms pointed  $> 90^\circ$  from each other
- Mapping Camera (MC)
  - Pushbroom imager; 1024 pixel CMOS or CCD line array
  - 5 separate line arrays in focal plane (radiation shielded)
    - 4 nadir viewing: panchromatic + 3 color bands (color for E/PO)
    - 1 panchromatic viewing  $\sim 40^\circ$  forward or aft for stereo
  - Nadir view, co-boresighted with LA on 2-axis gimbal platform

## Similar instruments



NEAR NLR



Galileo MAG



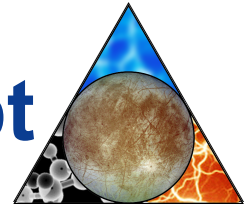
Rosetta LAP



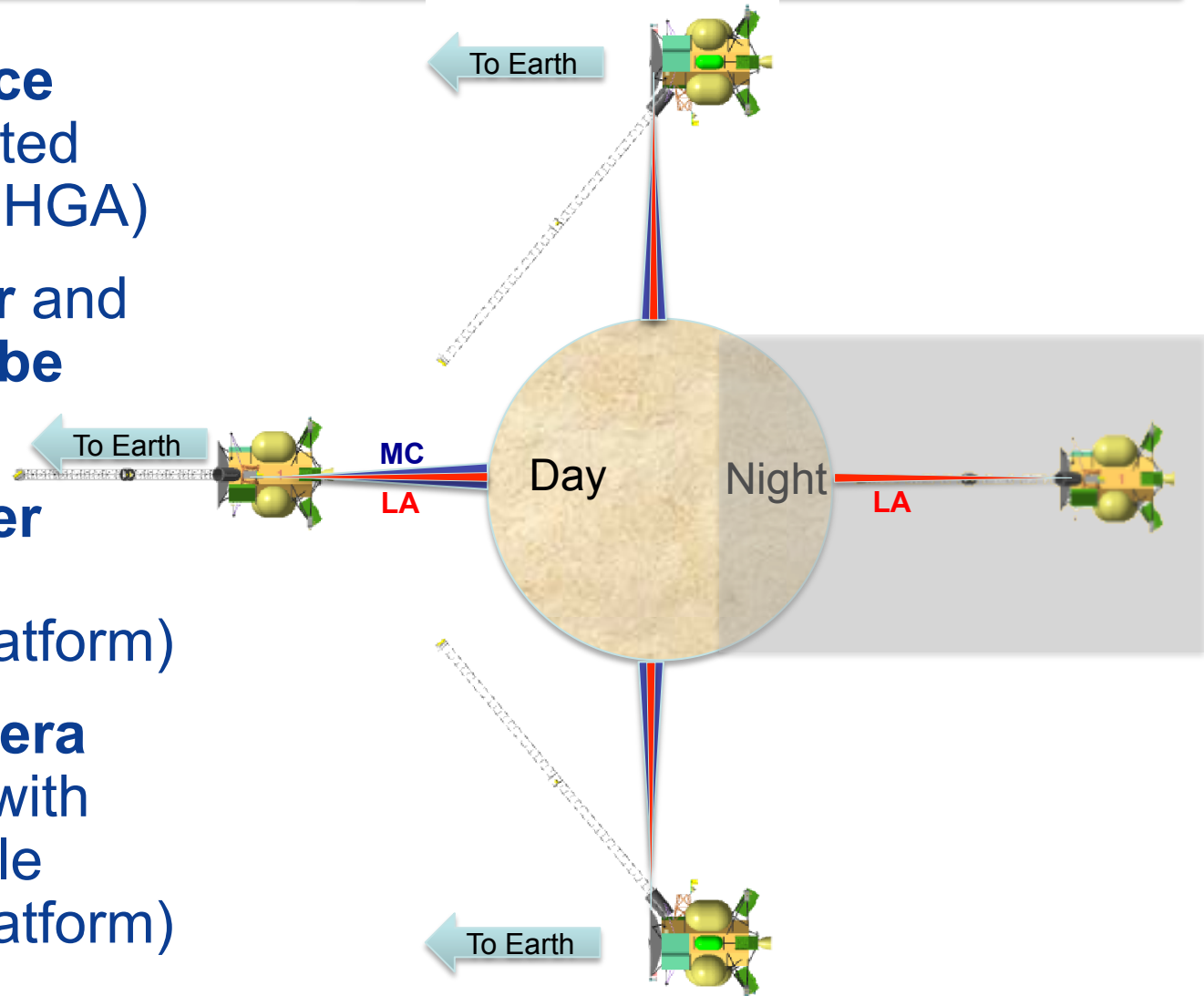
MPL/MSL MARDI



# Orbiter Science Operations Concept



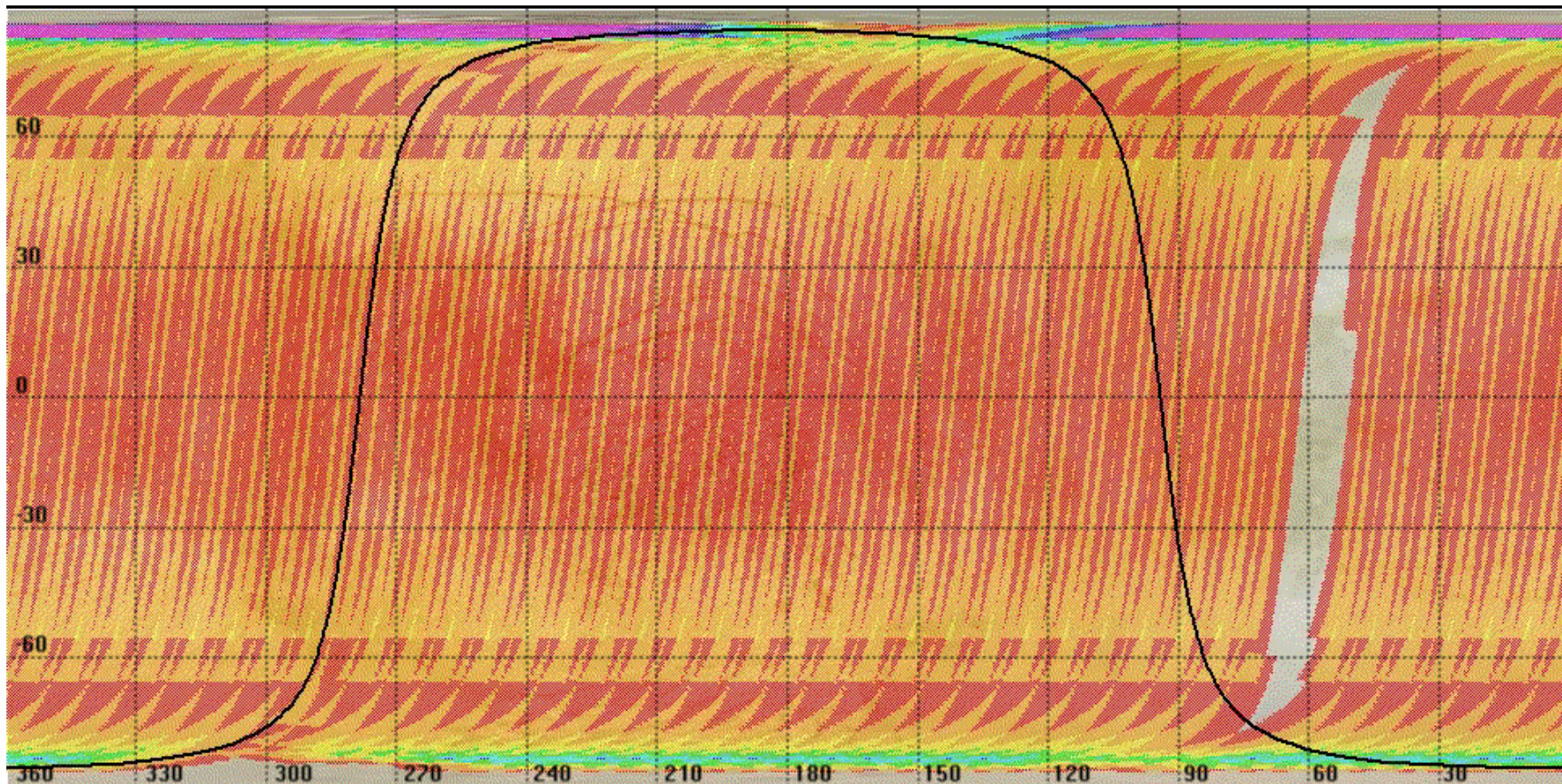
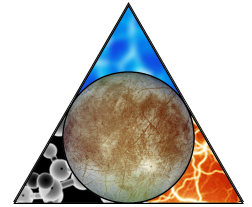
- **Gravity Science**  
while not occulted  
(Earth-pointed HGA)
- **Magnetometer and  
Langmuir Probe**  
on all the time
- **Laser Altimeter**  
on all the time  
(2 axis scan platform)
- **Mapping Camera**  
on during day with  
~45% duty cycle  
(2 axis scan platform)



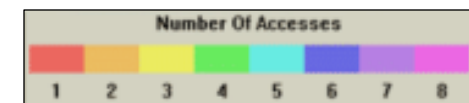


# Mapping Camera

## Global Stereo Map in 3 Eurosols (<11 days)



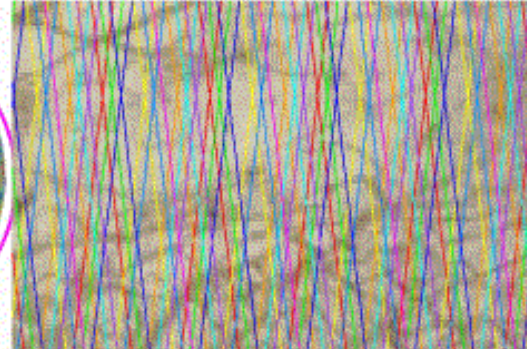
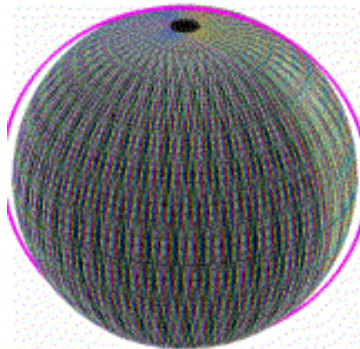
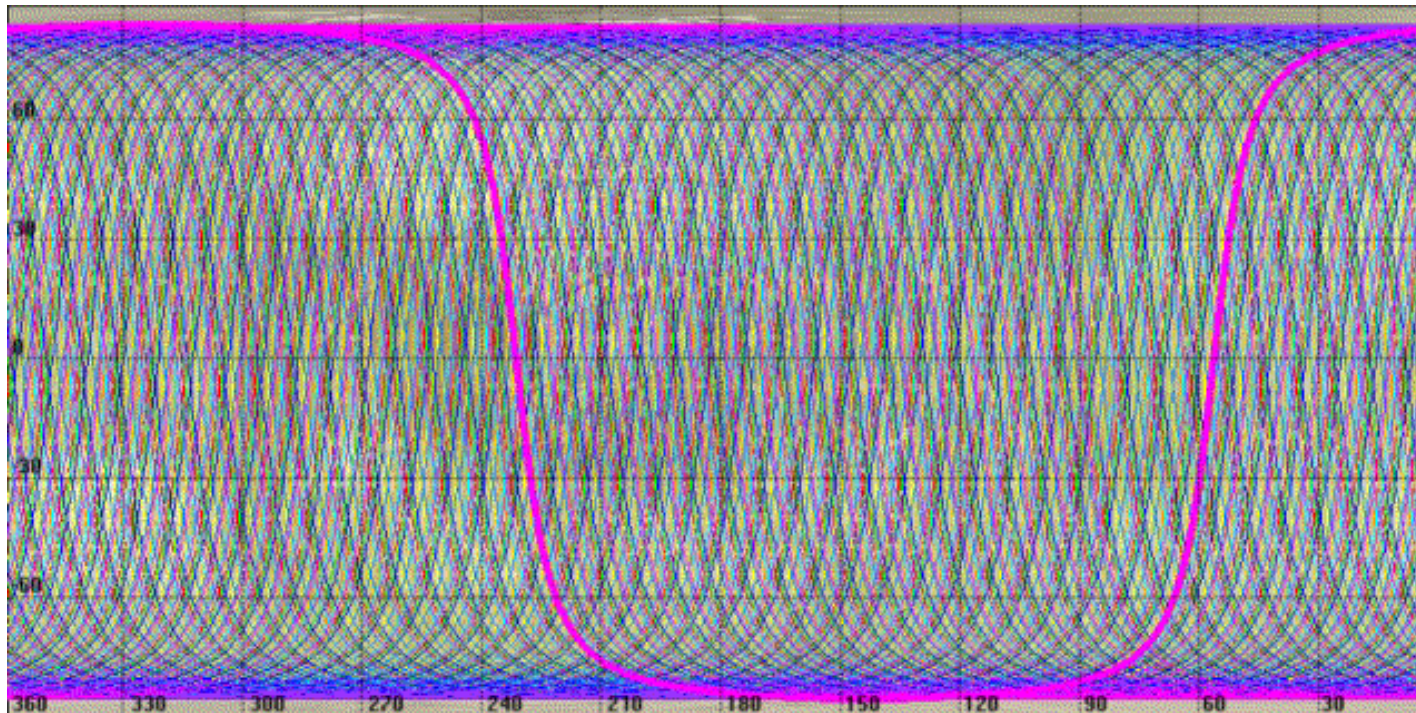
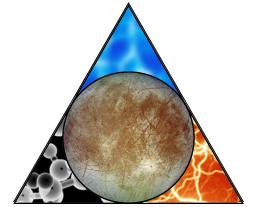
- 10% side-to-side swath overlap
- Orbit altitude of 103 km permits ground track near-repeat
- 1024-pixel line width ~100 m/pixel average resolution across FOV
- Wide gap is due to Jupiter occultation; It could be filled through off-nadir pointing





# Laser Altimeter Nadir Track

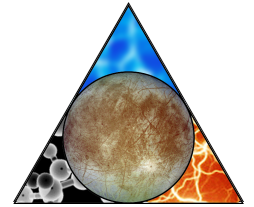
Each color is 1 eurosol (3.55 days)



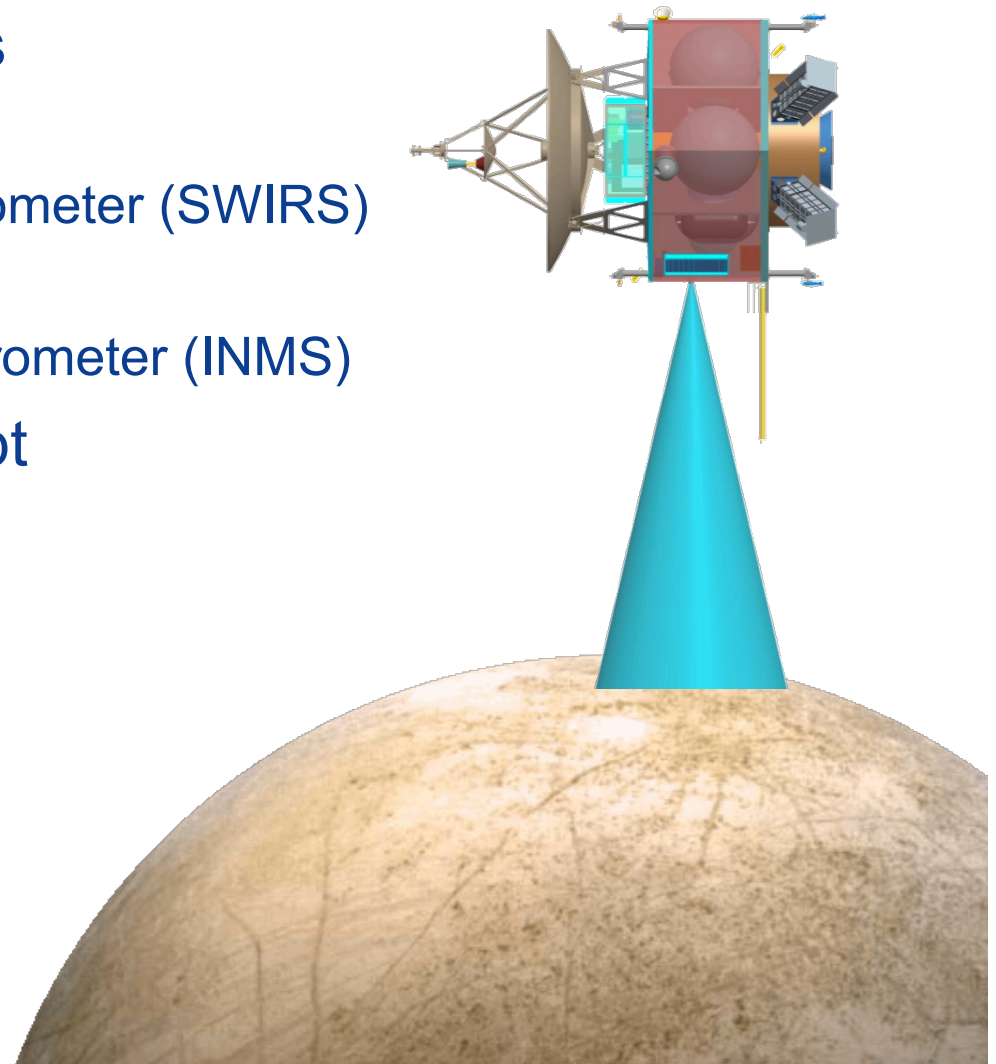
**~25 km equatorial spacing of laser profiles after 30 days**



# Flyby Model Payload: Outline

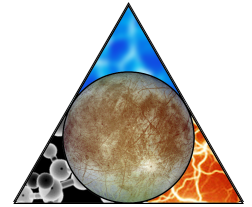


- Model payload instruments
  - Ice Penetrating Radar (IPR)
  - ShortWave InfraRed Spectrometer (SWIRS)
  - Topographical Imager (TI)
  - Ion and Neutral Mass Spectrometer (INMS)
- Science operations concept
- Example coverage



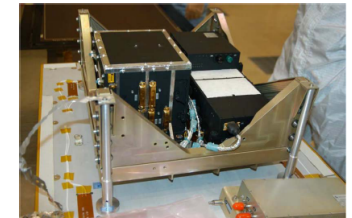


# Flyby Model Payload Instruments

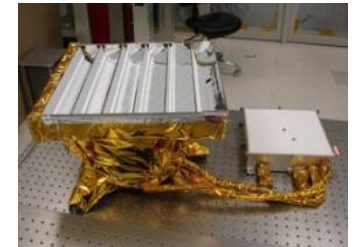


- Ice Penetrating Radar (IPR)
  - Dual-frequency sounder
    - 60 MHz with 10 MHz bandwidth (shallow)
    - 9 MHz with 1 MHz bandwidth (deep)
  - Deployed dipole antenna array on 15 m boom
- ShortWave InfraRed Spectrometer (SWIRS)
  - Spectral Range 0.85 – 5.0  $\mu\text{m}$ ; Spectral Resolution 10 nm
  - Single optic, single grating spectrometer & HgCdTe detector
  - Scan mirror for Target Motion Compensation
- Topographical Imager (TI)
  - Pushbroom, 4096 pixels width
  - Stereo obtained through along-track overlap
- Ion and Neutral Mass Spectrometer (INMS)
  - Mass Range 1 – 300 Da; Mass Resolution  $> 500$ ; Sensitivity 10 particles/cm<sup>3</sup>

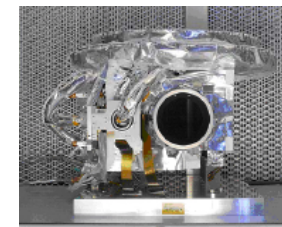
## Similar instruments



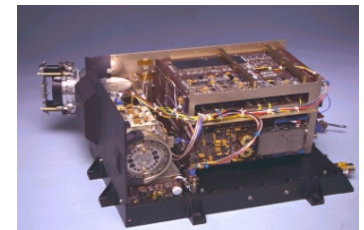
**MRO SHARAD**



**LRO M<sup>3</sup>**



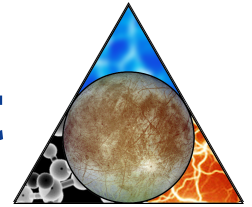
**New Horizons  
Ralph/MVIC**



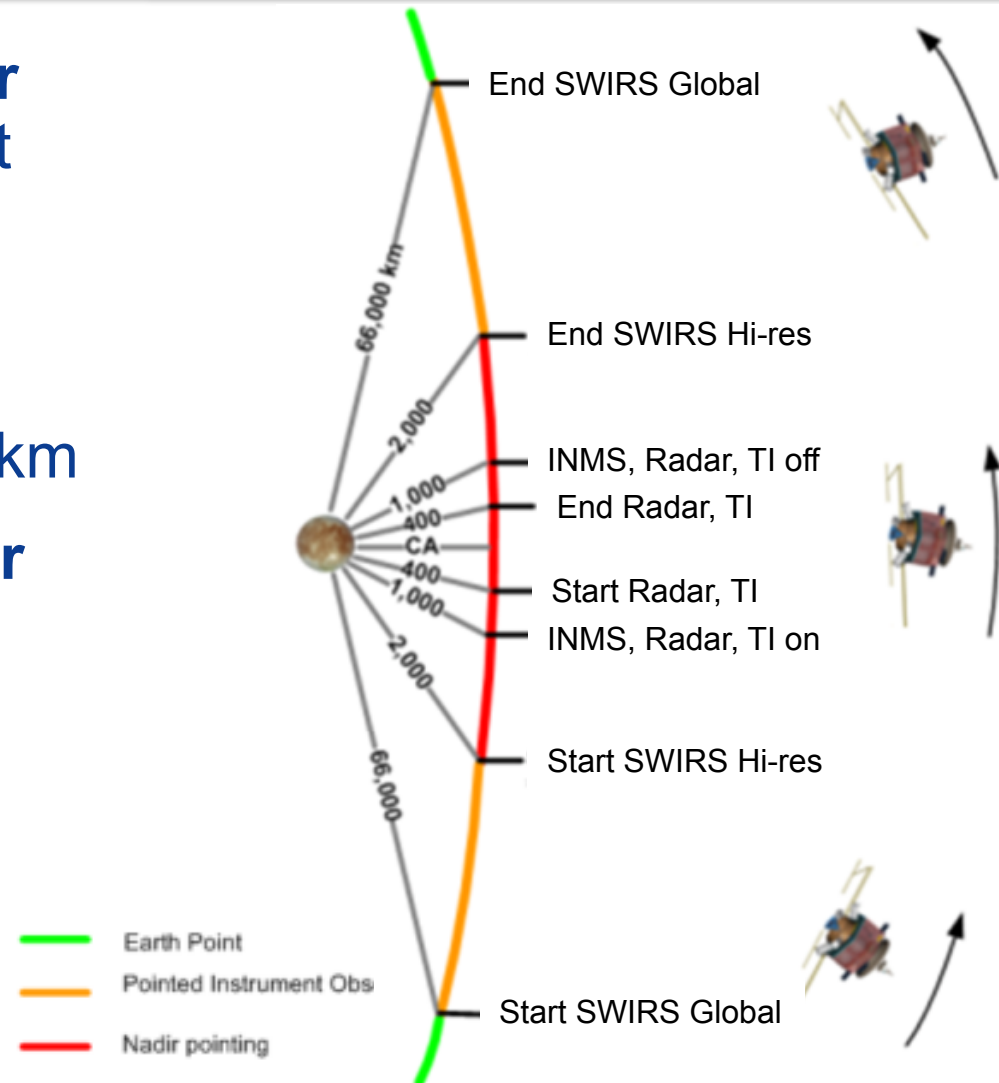
**Cassini INMS**



# Flyby Science Operations Concept



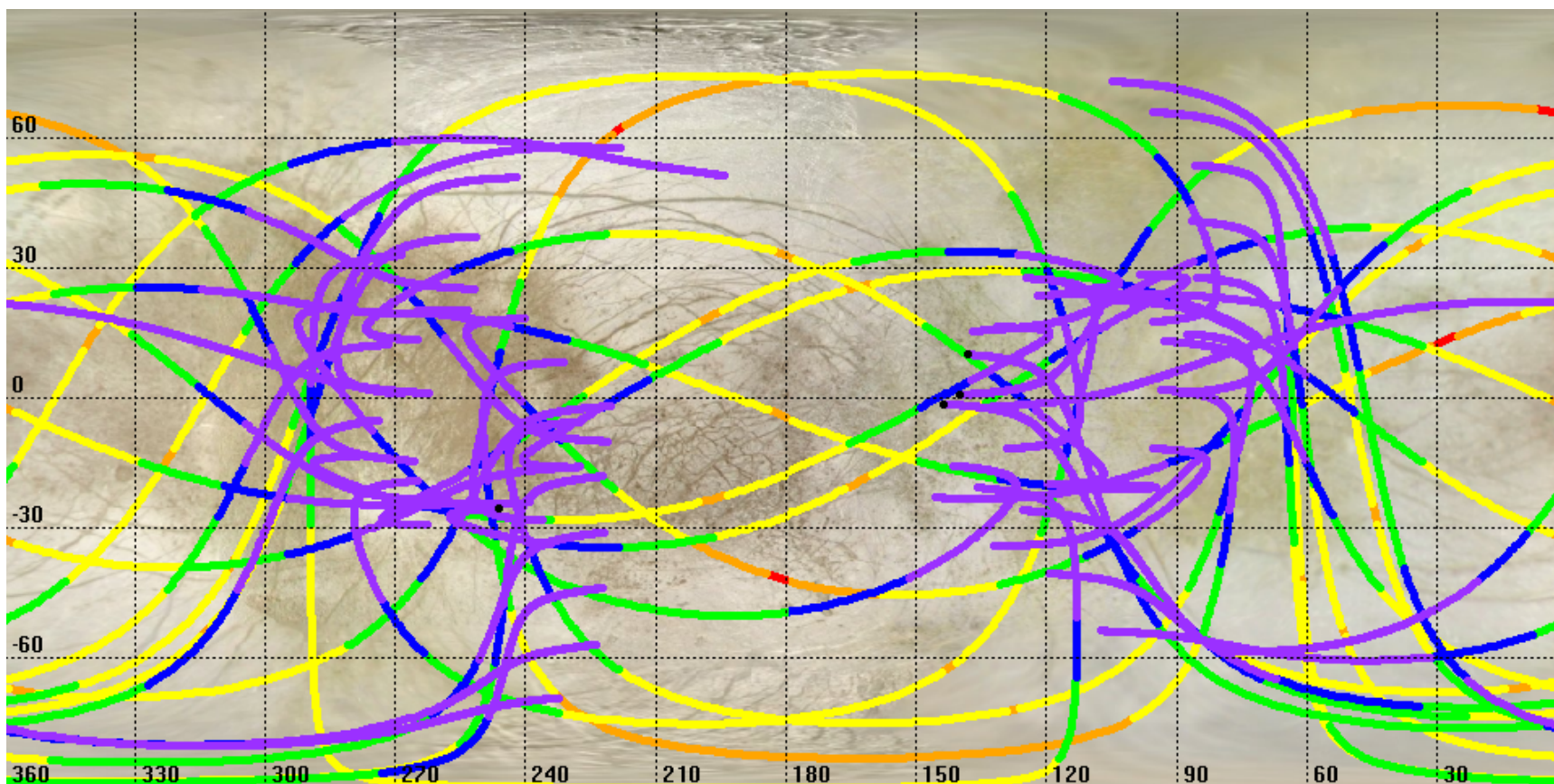
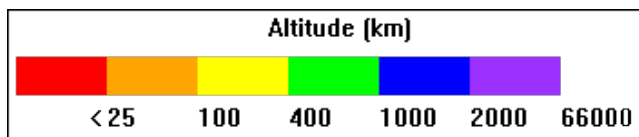
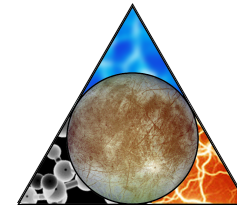
- **Ice Penetrating Radar** obtains primary data at  $\leq 400$  km
- **ShortWave InfraRed Spectrometer** obtains data  $\leq 66,000$  km
- **Topographical Imager** obtains stereo images  $\leq 1,000$  km
- **INMS** obtains data  $\leq 1,000$  km, including several  $\sim 25$  km flybys





# Flyby Ground Tracks

Altitude 25 km – 66,000 km

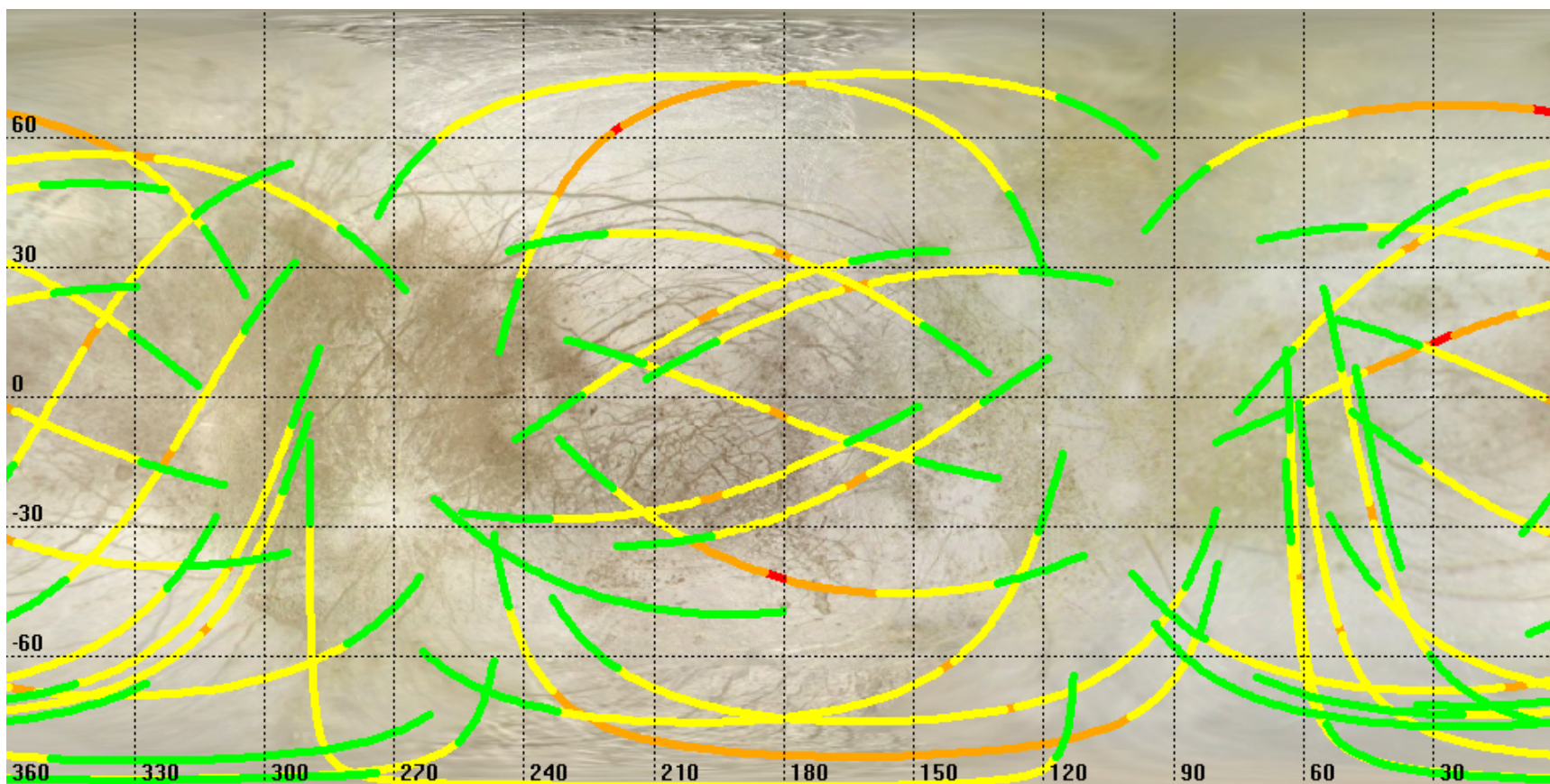
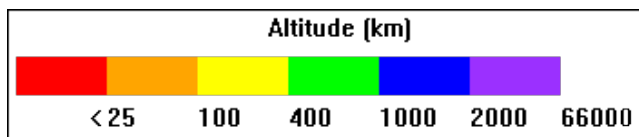
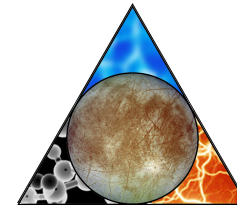


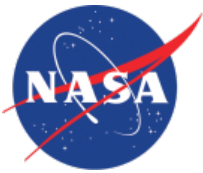




# Flyby Ground Tracks

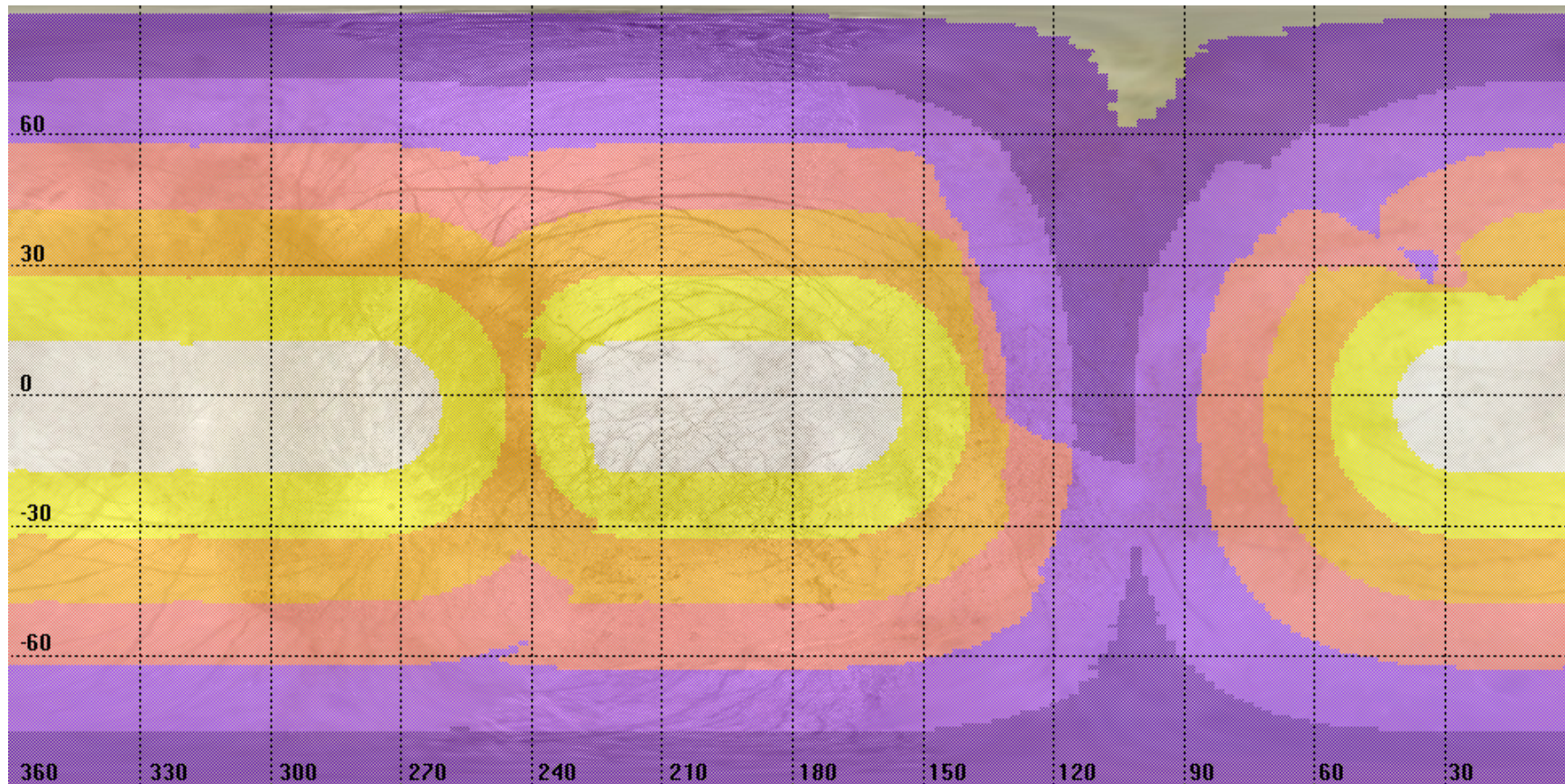
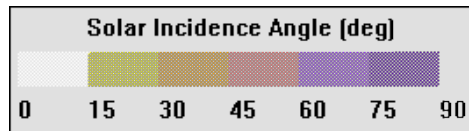
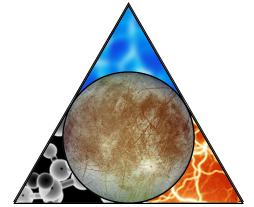
Altitude < 1,000 km





# SWIRS Coverage

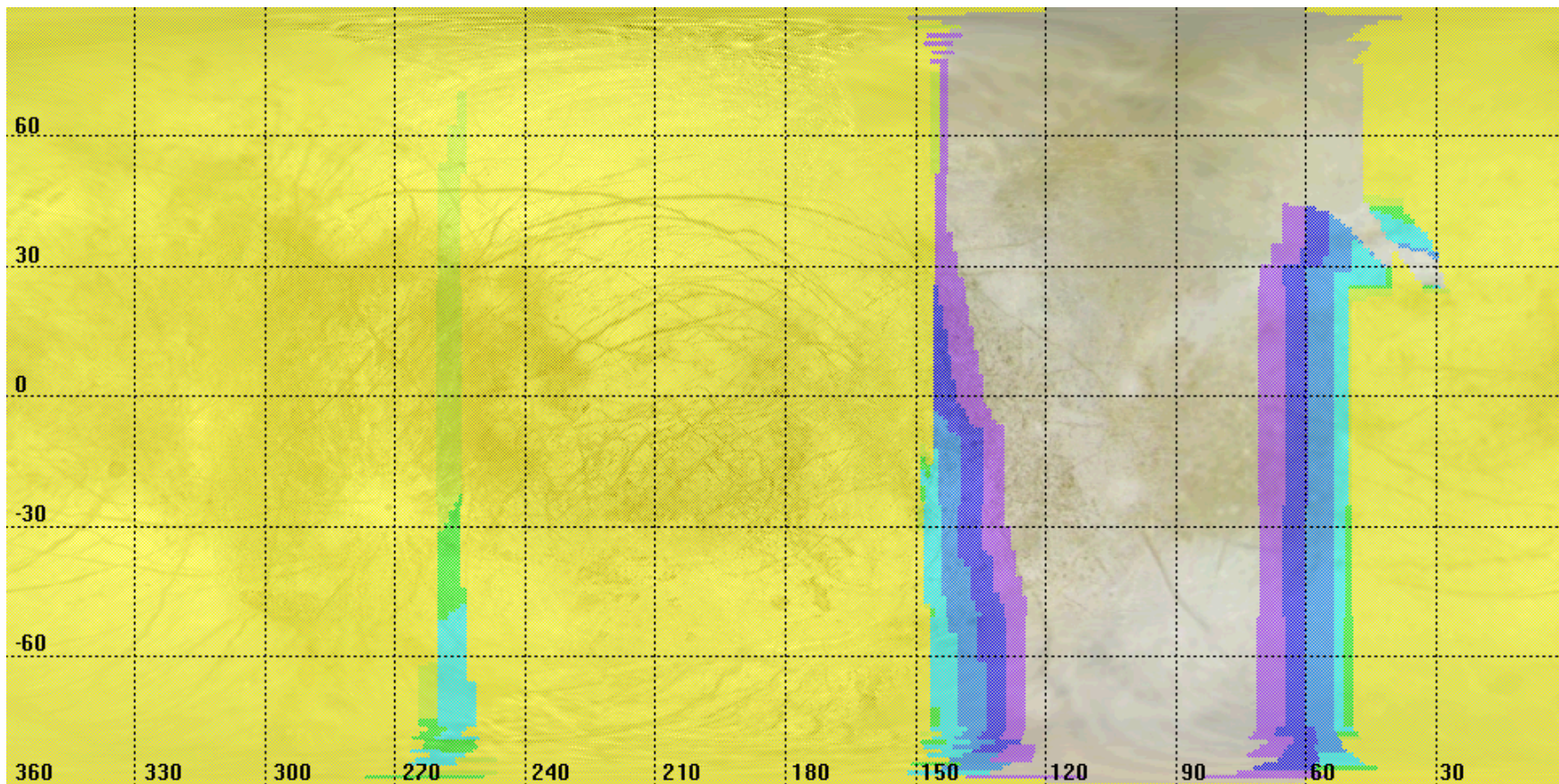
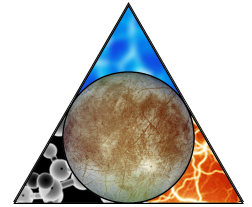
Solar Incidence Angle, Altitude 2,000 km – 66,000 km





# SWIRS Low Resolution Coverage

Altitude 2,000 km – 66,000 km

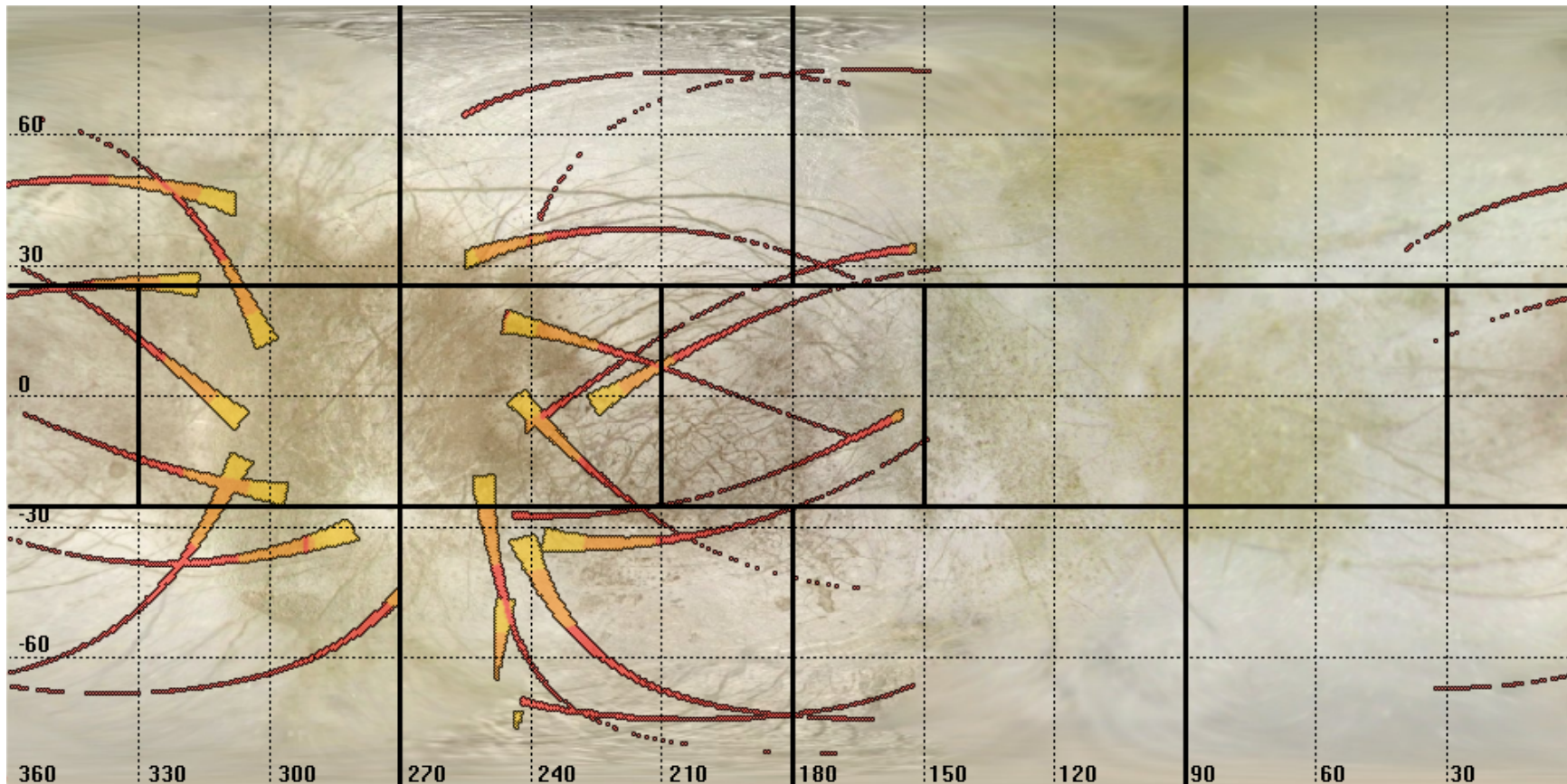
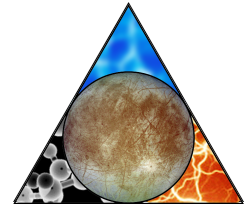


Coverage shown for local true solar times between 9:00am and 3:00pm



# SWIRS High Resolution Coverage

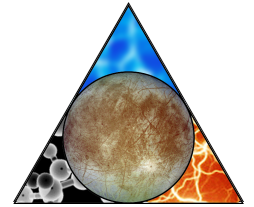
Altitude  $\leq 2,000$  km



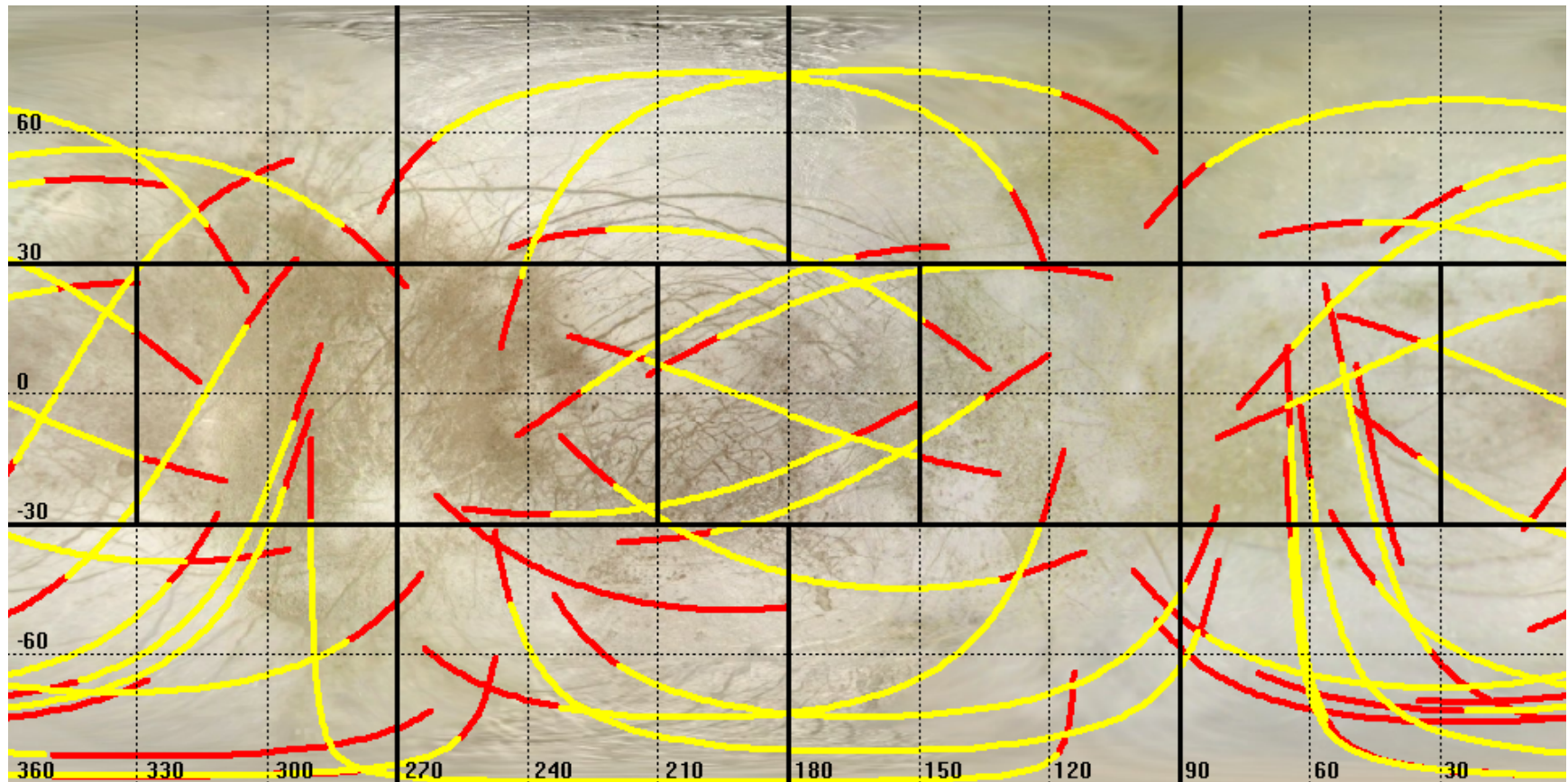
Coverage shown for local true solar times between 9:00am and 3:00pm



# IPR Ground Coverage



- 25 – 400 km (primary data collection)
- 400 – 1,000 km (extended data collection)



Sub-Jovian

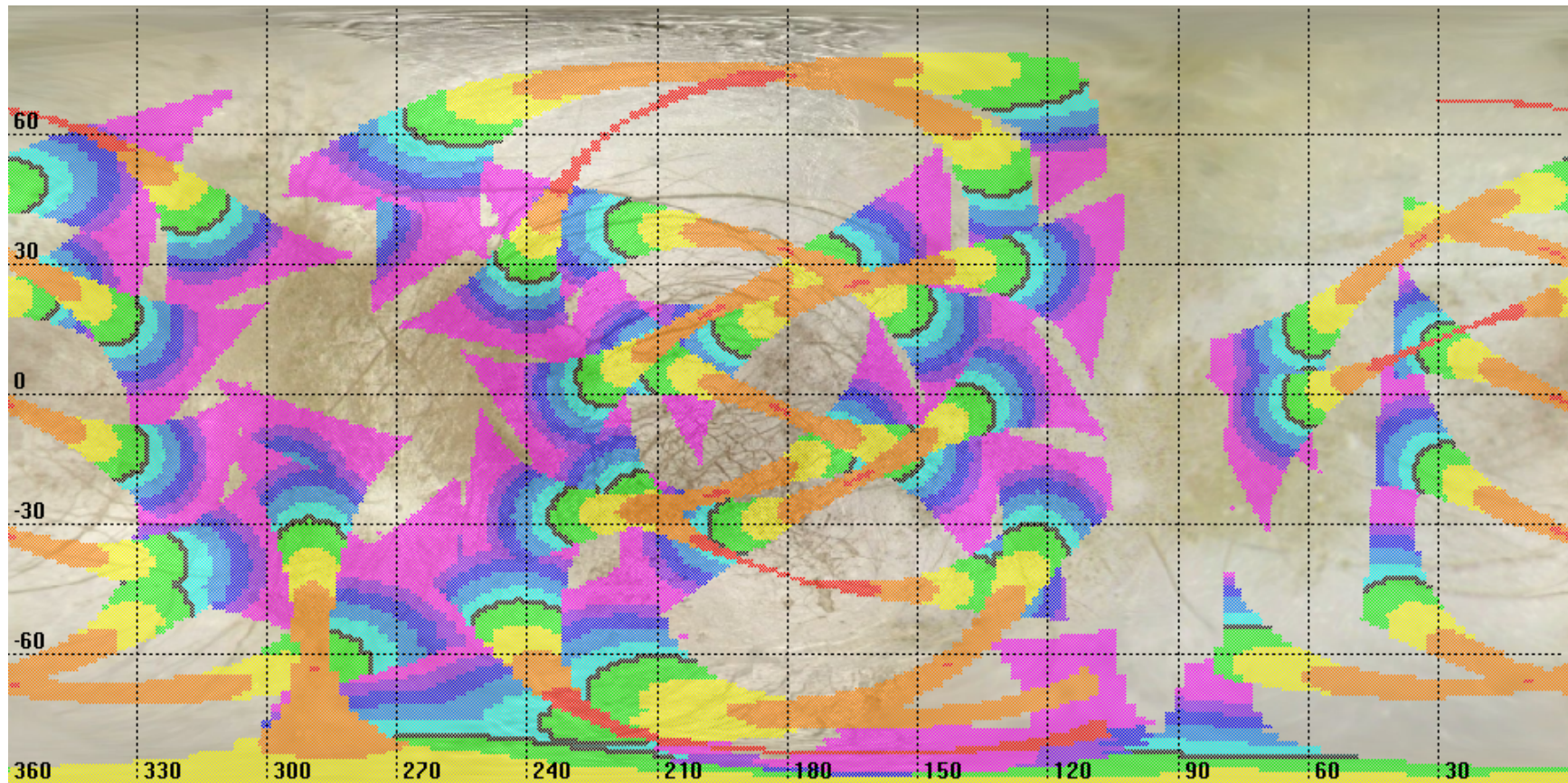
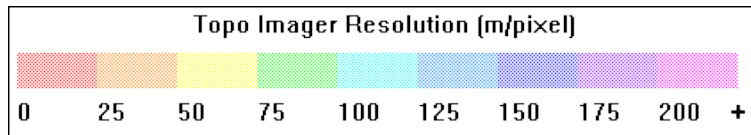
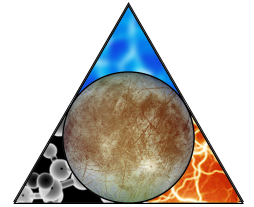
Anti-Jovian

Sub-Jovian



# Topo Imager Coverage

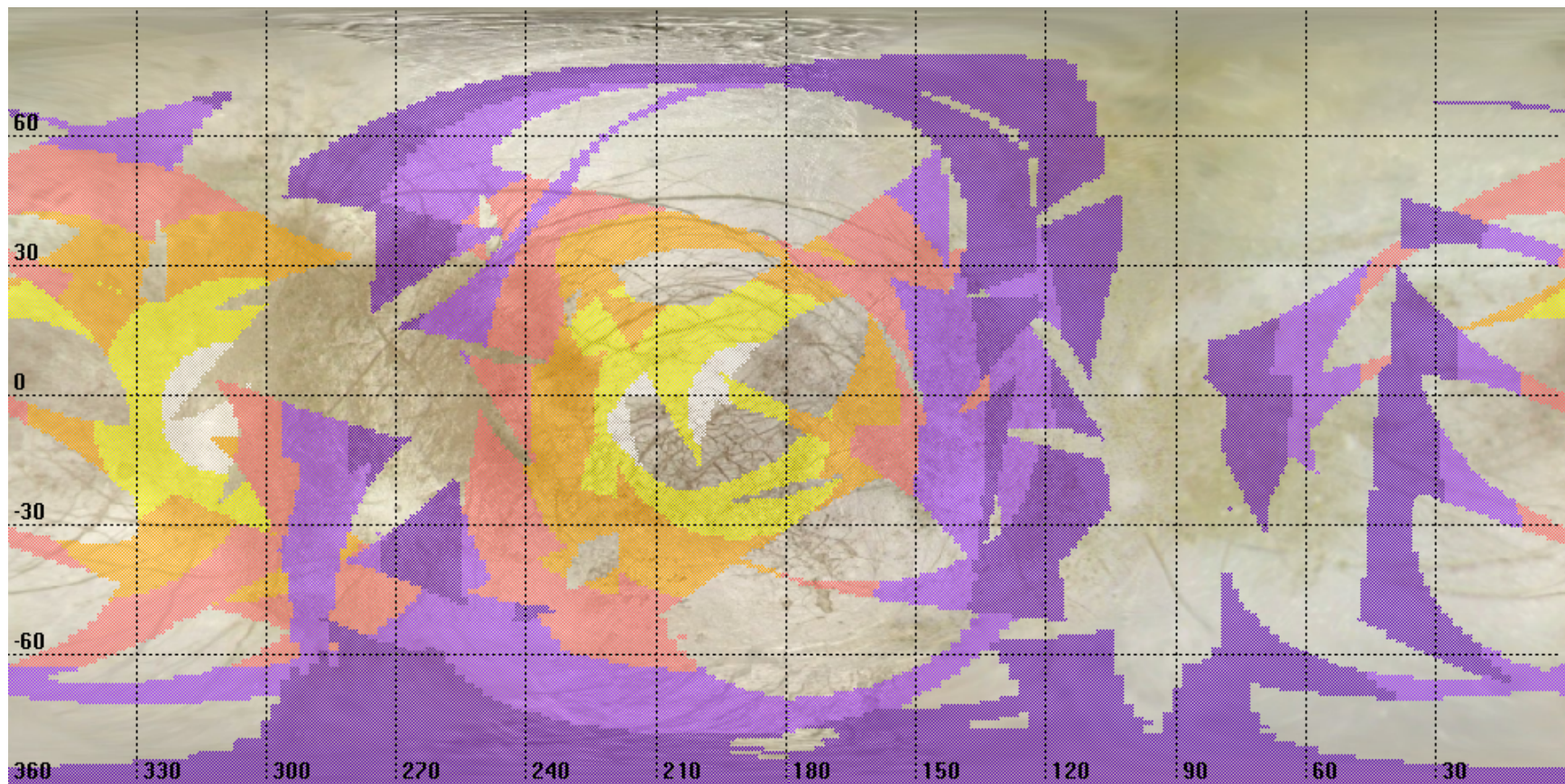
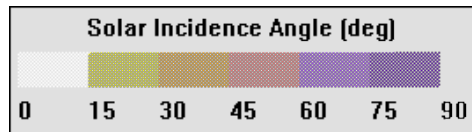
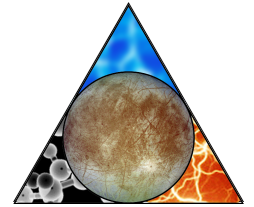
## Resolution for Altitude $\leq 1,000$ km





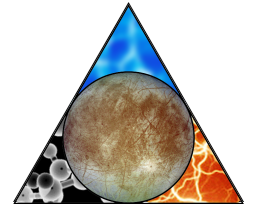
# Topo Imager Coverage

## Solar Incidence Angle for Altitude $\leq 1,000$ km





# Science as a Driver of Mission Architecture



Science traceability leads to a two element mission concept:



## **Orbiter Element:**

***Geophysical measurements that can be achieved only from orbit***

- Science focused primarily to address “Ocean” objective:
  - Gravity field
  - Tidal amplitude
  - Induction signatures
  - Plasma correction
  - Stratigraphic mapping



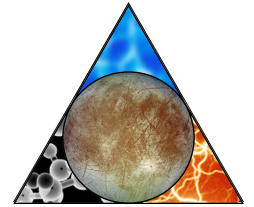
***Flyby Element: Remote measurements that can be accomplished via multiple flybys***

- Science focused primarily to address “Chemistry” and “Energy” themes:
  - Subsurface dielectric horizons
  - Surface constituents
  - Atmospheric constituents
  - Targeted landforms





# Science as a Driver of Mission Architecture



Science traceability leads to a two element mission concept:



## *Orbiter Element:*

*Geophysical measurements that can be achieved only from orbit*

- Payload focused primarily to address “Ocean” objective:
  - Radio Subsystem (RS)
  - Laser Altimeter (LA)
  - Magnetometer (MAG)
  - Langmuir Probe (LP)
  - Mapping Camera (MC)
- Have readily accommodated those instruments that are:
  - Less massive
  - Lower power
  - Lower data rate

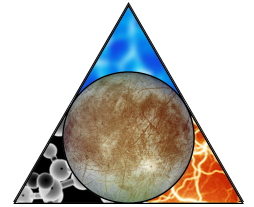


*Flyby Element: Remote measurements that can be accomplished via multiple flybys*

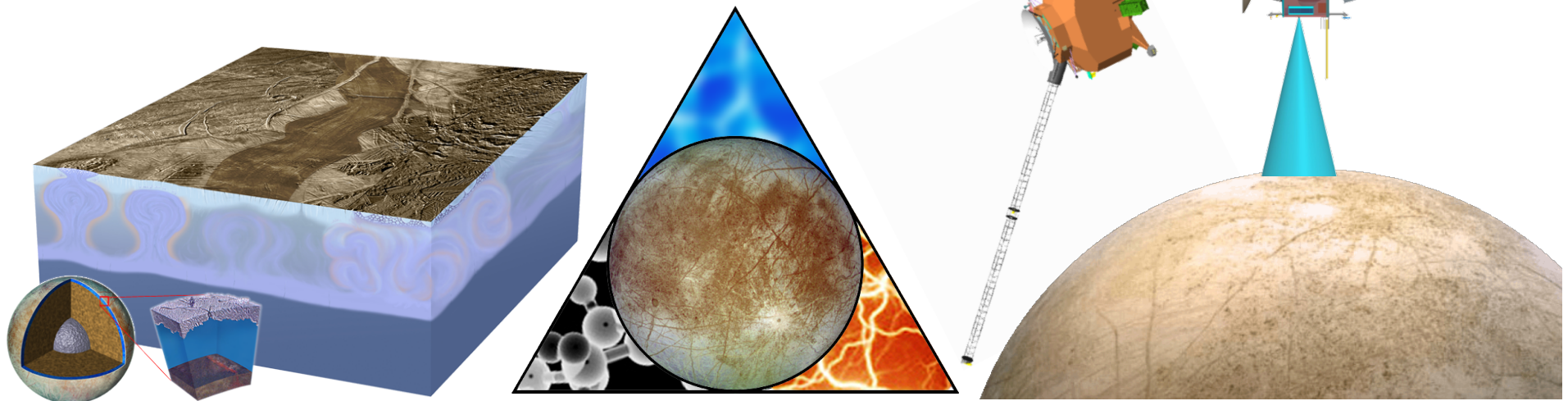
- Payload focused primarily to address “Chemistry” and “Energy” themes:
  - Ice Penetrating Radar (IPR)
  - ShortWave IR Spectrometer (SWIRS)
  - Ion and Neutral Mass Spectrometer (INMS)
  - Topographical Imager (TI)
- Have readily accommodated those instruments that are:
  - More massive
  - Higher power
  - Higher data rate

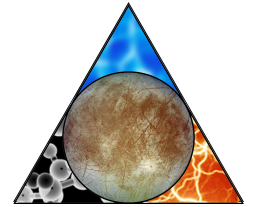


# A Pragmatic Path to Europa Exploration



- Pragmatic two-element architecture would fulfill the highest-priority Europa science objectives
  - Orbiter element concentrates on the “Ocean” objective
  - Multiple-flyby element concentrates on the “Chemistry” and “Energy” themes
- Each element has very high science value on its own
- Directly responsive to Decadal Survey’s recommendation for Europa
- Scientific priorities drive the architecture, permitting low-cost Europa mission options



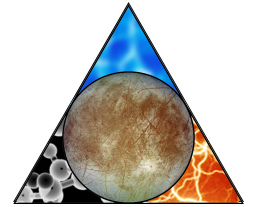


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# Backup Slides: Science



# Europa Exploration Addresses Decadal Survey Themes



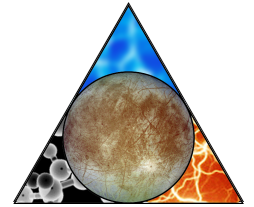
Decadal Survey Theme and Key Questions:		Building New Worlds	Planetary Habitats		Workings of Solar Systems	
		Q2: Planet and satellite accretion and migration	Q4: Organic sources and synthesis	Q6: Modern habitats and life	Q10: Solar system processes and evolution	
Goal	Europa Objective					
Explore Europa to investigate its habitability	Ocean	Characterize the extent of the ocean and its relation to the deeper interior.	✓	✓✓	✓✓✓	✓✓✓
	Ice Shell	Characterize the ice shell and any subsurface water, including their heterogeneity, and the nature of surface-ice-ocean exchange.		✓✓	✓✓✓	✓✓✓
	Comp'tn	Understand the habitability of Europa's ocean through composition and chemistry.	✓	✓✓✓	✓✓✓	✓✓✓
	Geology	Understand the formation of surface features, including sites of recent or current activity, and characterize high science interest localities.		✓✓	✓✓✓	✓✓✓

✓ = good, ✓✓ = very good, ✓✓✓ = excellent



# Europa Science Definition Team

## Subgroup Structure



### Objectives subgroups:

- **Ocean & Ice:**
  - Bills, Barr, Blankenship\*, Connerney, Senske, Smith
- **Composition:**
  - Blaney, Bagenal, Hoehler, Hand, Shock\*, Brinckerhoff, Vance
- **Geology:**
  - Moore, Kurth, McGrath, Mellon, Patterson, Prockter\*

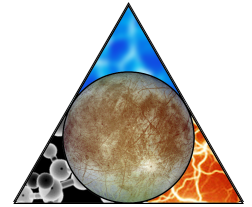
### Cross-Cutting subgroups:

- **Astrobiology:**
  - Hoehler, Blankenship, Hand\*, McGrath, Senske, Shock
- **Instrumentation:**
  - Mellon, Bills, Blaney, Brinckerhoff\*, Connerney, Kurth
- **Landing Sites:**
  - Prockter, Bagenal, Barr\*, Moore, Patterson, Smith, Vance

**Bold = Lead; \* = Deputy Lead**



# Orbiter: Laser Altimeter (LA)



- **Primary Science Investigations**

- Europa global topography and tidal deformation, rotation state, landform topography

- **Measurement Requirements**

- Topographic differences to 1 m vertical accuracy at globally distributed crossover points at varying Europa orbital phases; implies surface ranging to 10 cm accuracy
- Simultaneous ranging with stereo imagery desired
- Near-polar near-circular orbit at  $\leq 200$  km altitude

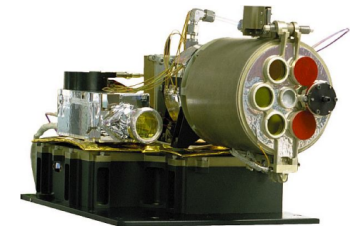
- **Configuration for Model Payload**

- Time-of-flight rangefinder
- Transmits 2.7 mJ pulses at  $1.064 \mu\text{m}$
- 50 m laser spot from 100 km altitude
- Receiver telescope aperture of 12.5 cm diameter provides good SNR from 100 km altitude
- Avalanche photodiode detector (radiation shielded)
- 26 Hz pulse rate yields 50 m spot spacing
- Nadir view co-boresighted with camera; minimize thruster contamination
- Mounted on 2-axis gimbal platform (assuming simultaneous ranging and downlink over fixed-mounted HGA from mid-afternoon orbit)

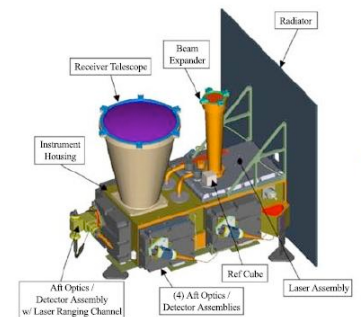
## Similar instruments



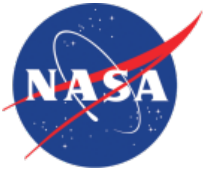
**MESSENGER MLA**



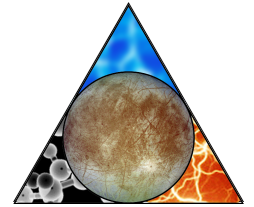
**NEAR NLR**



**LRO LOLA**



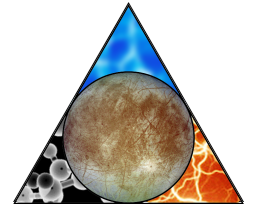
# Orbiter: Radio Subsystem (RS)



- **Primary Science Investigations**
  - Europa static gravity field and gravitational and tides
- **Measurement Requirements**
  - Degree-2 time-dependent gravity field, to recover  $k_2$  amplitude at Europa's orbital frequency to  $\leq 0.003$  absolute accuracy and phase to  $\leq 1^\circ$
  - Europa position wrt Jupiter to  $\leq 10$  m throughout orbital mission
  - Europa static gravity field resolved to degree and order 30
    - Range-rate to  $< 0.1$  mm/s at 60 s integration time
    - S/C orbit determination about Europa to accuracy of  $< 1$  m (rms) in radial direction
  - Near-polar near-circular orbit at  $\leq 200$  km altitude for  $\geq 30$  days
  - Data arcs over several “unperturbed” days
  - Range-rate measurements over several Europa tidal cycles
- **Configuration for Model Payload**
  - X-band up and down; Ka-band down only
  - Ka Transponder/Hybrid Diplexer
  - USO not required



# Orbiter: Magnetometer (MAG)



- **Primary Science Investigations**

- Europa magnetic induction response

- **Measurement Requirements**

- 3-axis magnetic field components at 8 Hz, 8 vectors/s
- Sensitivity of 0.1 nT
- Near continuous measurements
- Near-polar near-circular orbit at  $\leq 200$  km altitude for  $\geq 30$  days

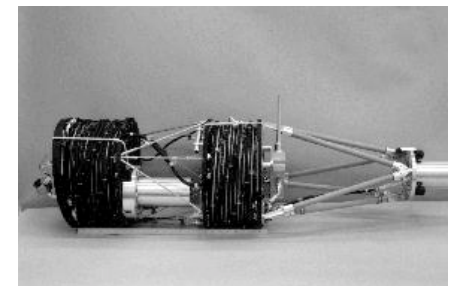
- **Configuration for Model Payload**

- Dual 3-axis fluxgate
- Sensitivity 0.1 nT
- Maximum sampling rate 32 Hz; sampling resolution 0.01 nT
- Sensors on boom 5 m and 10 m from S/C
- Spacecraft cleanliness of 0.1 nT desired, 0.5 nT required at outboard sensor
- Periodic S/C slow spins about two orthogonal axes for calibration

## Similar instruments



**MESSENGER MAG**

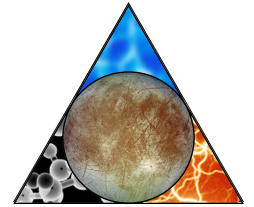


**Galileo MAG**





# Orbiter: Langmuir Probe (LP)

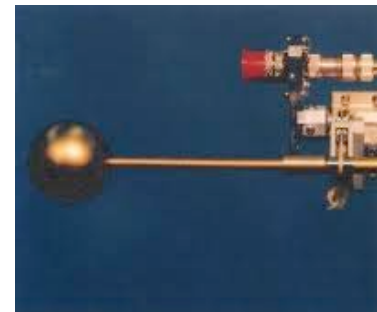


- **Primary Science Investigations**
  - Europa ionospheric plasma effects on magnetic induction response
- **Measurement Requirements**
  - Local plasma density, temperature, and flow
  - Electric field (near DC to 3 MHz)
  - Electron temperature
  - Ion currents
  - $4\pi$  coverage
  - Map approximately 12 hours of Europa local time
  - Near-polar near-circular orbit at  $\leq 200$  km altitude for  $\geq 30$  days
- **Configuration for Model Payload**
  - Two 5 cm diameter spheres mounted on 1 m long booms
  - Booms pointed  $> 90^\circ$  from each other, desire one always free of S/C wake
  - Electronics (including pre-amp) in S/C vault  $< 3$  m from sensors
  - EMI/EMC cleanliness like Rosetta and/or Cassini

## Similar instruments



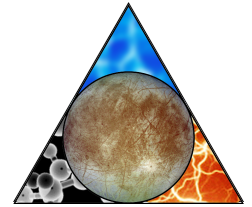
**Rosetta LAP**



**Cassini RPWS**



# Orbiter: Mapping Camera (MC)



- **Primary Science Investigations**

- Landform mapping, Europa rotation state, and surface/subsurface material exchange

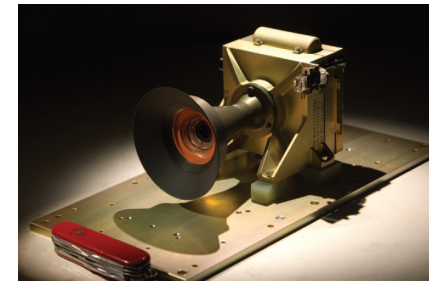
- **Measurement Requirements**

- $\geq 80\%$  global mapping at  $\leq 100$  m/pixel
- 30 m vertical resolution topo,  $\leq 300$  m horizontal footprint resolution
- Incidence angle  $> 45^\circ$  ( $70^\circ$  preferred)
- Near-polar near-circular orbit with consistent lighting

- **Configuration for Model Payload**

- Pushbroom imager; 1024 pixel CMOS or CCD line array
- 5 separate line arrays in focal plane (radiation shielded)
  - 4 nadir viewing: panchromatic, 0.56, 0.76, 0.99  $\mu\text{m}$  bands
  - 1 panchromatic viewing  $\sim 40^\circ$  forward or aft for stereo
- IFOV = 0.85 mrad; FOV =  $50^\circ$ 
  - 85 m/pixel at nadir from 100-km orbit altitude
  - 94 km wide swath ( $\sim 3.4^\circ$  of longitude at the equator)
  - 63 ms for 1 pixel of smear in 100 km orbit, providing good SNR
- Can elect to operate color and stereo bands, or not
- Loose pointing requirements ( $\sim 2^\circ$  accuracy; 5 mrad/s stability)
- Mounted on 2-axis gimbal platform (permits simultaneous imaging and HGA downlink)
- Small radiator with view to dark space

## Similar instruments



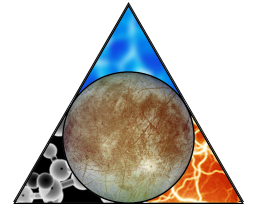
**MRO MARCI**



**MPL/MSL MARDI**



# Flyby: Ice Penetrating Radar (IPR)



- **Primary Science Investigations**

- Characterize distribution of shallow subsurface water and structure of ice shell
- Search for an ice-ocean interface
- Correlate surface features, subsurface structures, and geological processes

- **Measurement Requirements**

- Shallow Mode: 10 m vertical resolution; 100 m to 3 km depth
- Deep Mode: 100 m vertical resolution; 1 km to 30 km depth
- Globally distributed intersecting and adjacent swaths, in 11 of 14 “panels”
- Supporting requirements:
  - Nadir altimetry, 10 m vertical resolution
  - Cross-track surface topography (stereo imaging), 100 m vertical res

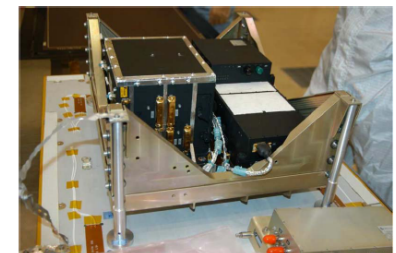
- **Configuration for Model Payload**

- Dual-frequency sounder
  - 60 MHz with 10 MHz bandwidth (shallow)
  - 9 MHz with 1 MHz bandwidth (deep)
- Deployed dipole antenna array on 15 m boom
- Range compression, pre-summing, Doppler filtering, data averaging, resampling in S/C electronics

## Similar instruments



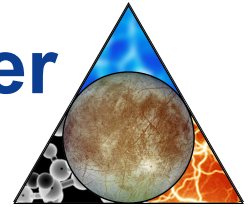
**Mars Express  
MARSIS**



**MRO SHARAD**



# Flyby: ShortWave InfraRed Spectrometer (SWIRS)



- **Primary Science Investigations**

- Characterize surface composition for representative landforms
- Characterize exogenic materials

- **Measurement Requirements**

- Spectral Range                    850 nm – 5.0  $\mu$ m
- Spectral Resolution            10 nm
- Spectral Channels                420
- Spatial Resolution               300 m @ 2000 km (IFOV = 150  $\mu$ rad)
- Image Width                        480 pixels (4.2°)
- Signal to Noise                    ~ 100 at 5  $\mu$ m (TMC 8), 18 at 5  $\mu$ m (TMC 1)
- Exposure time                    ~ 1 s / row (TMC 8) [7.7 min for full image], 0.12 s / row (TMC 1)

- **Configuration for Model Payload**

- Implementation of 4 scans for each flyby: 2 @ 10 km/pixel (s/c slew) and 2 < 300 m/pixel (scan mirror, s/c tracking nadir)
- Single optic, single grating spectrometer & HgCdTe detector; TMC (scan mirror)
- Passive detector and spectrometer cooling
- Detector radiation noise (for TMC 8, 4% pixels are hit @ 9.5 cm tantalum shielding; for TMC 1, 4% pixels are hit @ ~3 cm Ta); calculation for worst case 0° inclination

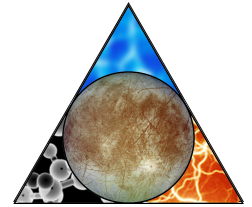
## Similar instruments



LRO M<sup>3</sup>



# Flyby: Topographical Imager (TI)



- **Primary Science Investigations**

- Removal of radar returns from off-nadir surface topography
- High-resolution stereo imaging of landforms for geology

- **Measurement Requirements**

- Spectral Range                    visible
- Spectral Bands                    monochromatic
- FOV                                    58° (for stereo separation)
- Image Width @ C/A            100 km (matches width of radar swath)
- Signal to Noise                     $\geq 100$

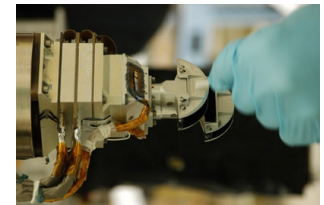
- **Configuration for Model Payload**

- Pushbroom operation
- Stereo obtained through along-track overlap with  $\sim 50$  m vertical resolution
- Image width                        4096 pixels
- Spatial Resolution            25 m @ 100 km alt (250  $\mu$ rad iFOV) (6.2 m @ 25 km alt)
- SNR                                     $\sim 100$  @ 100 km alt (can increase to  $\sim 400$  using TDI)  
 $\sim 27$  @ 25 km alt (with TDI, can get up to  $\sim 100$ )
- Exposure times                    5.5 ms for 1 pixel of smear @ 100 km alt  
1.4 ms for 1 pixel of smear @ 25 km alt
- Radiation noise (% of pixels hit with electrons/1 cm Ta detector shielding) 0.5%

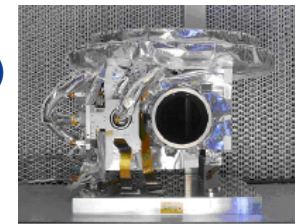
## Similar instruments



**MESSENGER MDIS**



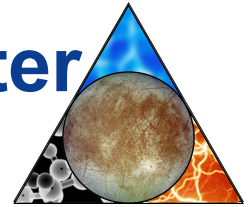
**MRO MARCI**



**New Horizons  
Ralph/MVIC**

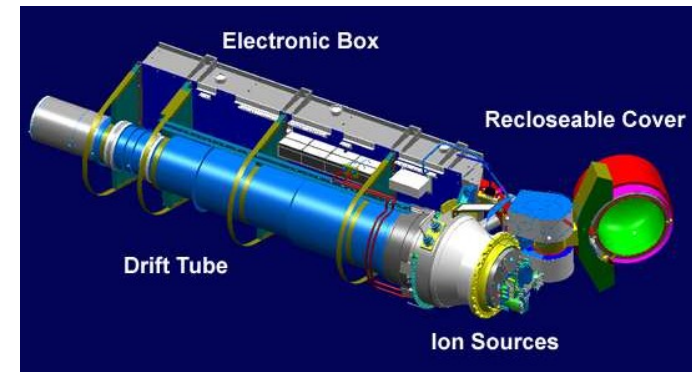


# Flyby: Ion and Neutral Mass Spectrometer (INMS)

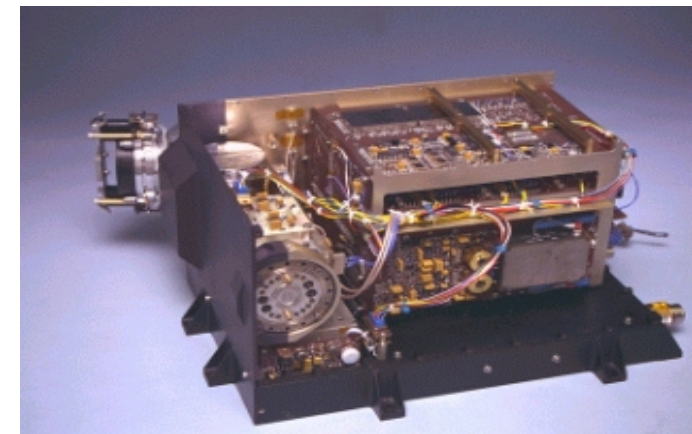


- **Primary Science Investigations**
  - Elemental, isotopic and molecular composition of Europa's atmosphere and ionosphere
- **Measurement Requirements**
  - Mass Range 1 – 300 Daltons
  - Mass Resolution > 500
  - Sensitivity 10 particles/cm<sup>3</sup>
  - Field of View 60 degree clear input
- **Configuration for Model Payload**
  - Ram pointed inlet
  - Remote electronics
  - One-time opening cover
- **Note on Sensitivity and Resources**
  - We solicited input from INMS providers to understand the limits on sensitivity and detection of minor species during fast flybys

## Similar instruments



Rosina RTOF



Cassini INMS