



Jet Propulsion Laboratory  
California Institute of Technology

# Cassini Science Update

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Outer Planets Assessment Group

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# Cas

- Four Cassini p
- Buratti et al., 2019, Pandora, and Ep <https://science.>
- Iess et al., 2019, and ring mass
- Tiscareno et al., 2019, Cassini's ring-gr finale <https://s>
- Spilker, 2019, of discovery <htt>

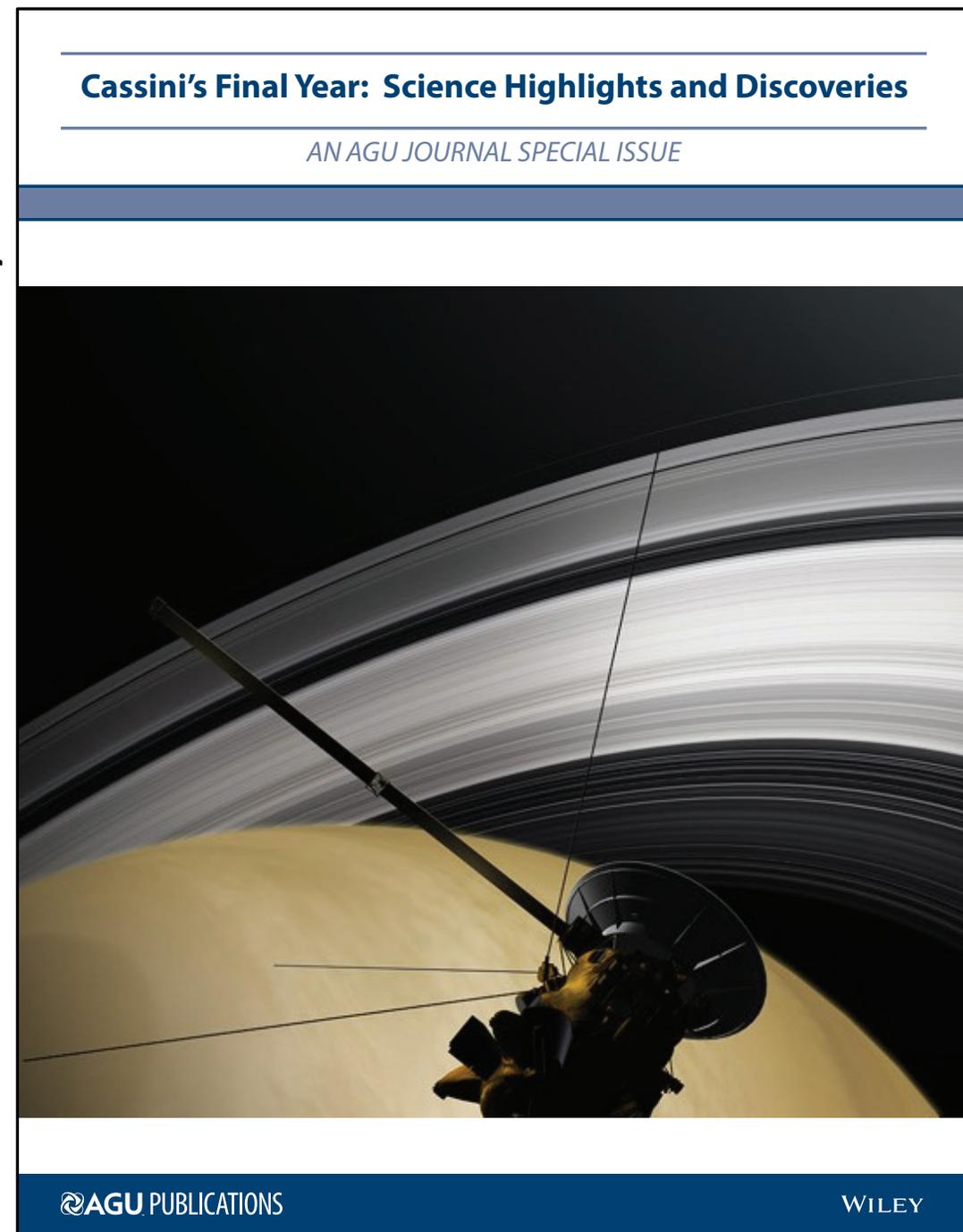


# Science

- of Science magazine
- Pan, Daphnis, Atlas, <https://science.> 2349
- Jupiter's gravity field [/364/6445/eaat2965](https://doi.org/10.1126/science.1236464/6445/eaat2965)
- Saturn's rings during [/45/eaau1017](https://doi.org/10.1126/science.1236464/6445/eaau1017)
- Jupiter system: 13 years [/64/6445/1046](https://doi.org/10.1126/science.1236464/6445/1046)

# Cassini GRL Special Issue: 40 papers

- 40 Cassini papers are on line and will appear soon in the print version of Geophysical Research Letters (GRL)
- Topics include: Saturn's interior, atmosphere, stratosphere, ionosphere, hexagon, aurora, auroral hiss, radio emissions, ring-gap phenomena (electric currents, radiation belts, gap particle composition and behavior, ring rain), rings, icy moons, Titan, and magnetosphere
- Cassini Icarus Special issue will be published in early 2020, 26 papers submitted



# Cassini Legacy: Future Plans

- Continue to share new Cassini results at OPAG meetings and other venues
  - Valuable for community to continue to hear about new Cassini results
  - New results may motivate future mission discussions
- Cassini researchers requested to continue to submit accepted papers to Linda Spilker and Gretchen McCartney (Cassini media lead: [gretchen.p.mccartney@jpl.nasa.gov](mailto:gretchen.p.mccartney@jpl.nasa.gov)) for consideration for press releases
  - JPL media infrastructure exists to continue to support Cassini press activity longer term even though Cassini funding ends in September
- Encourage younger scientists to help continue Cassini's legacy
  - Contact Linda if interested in more details

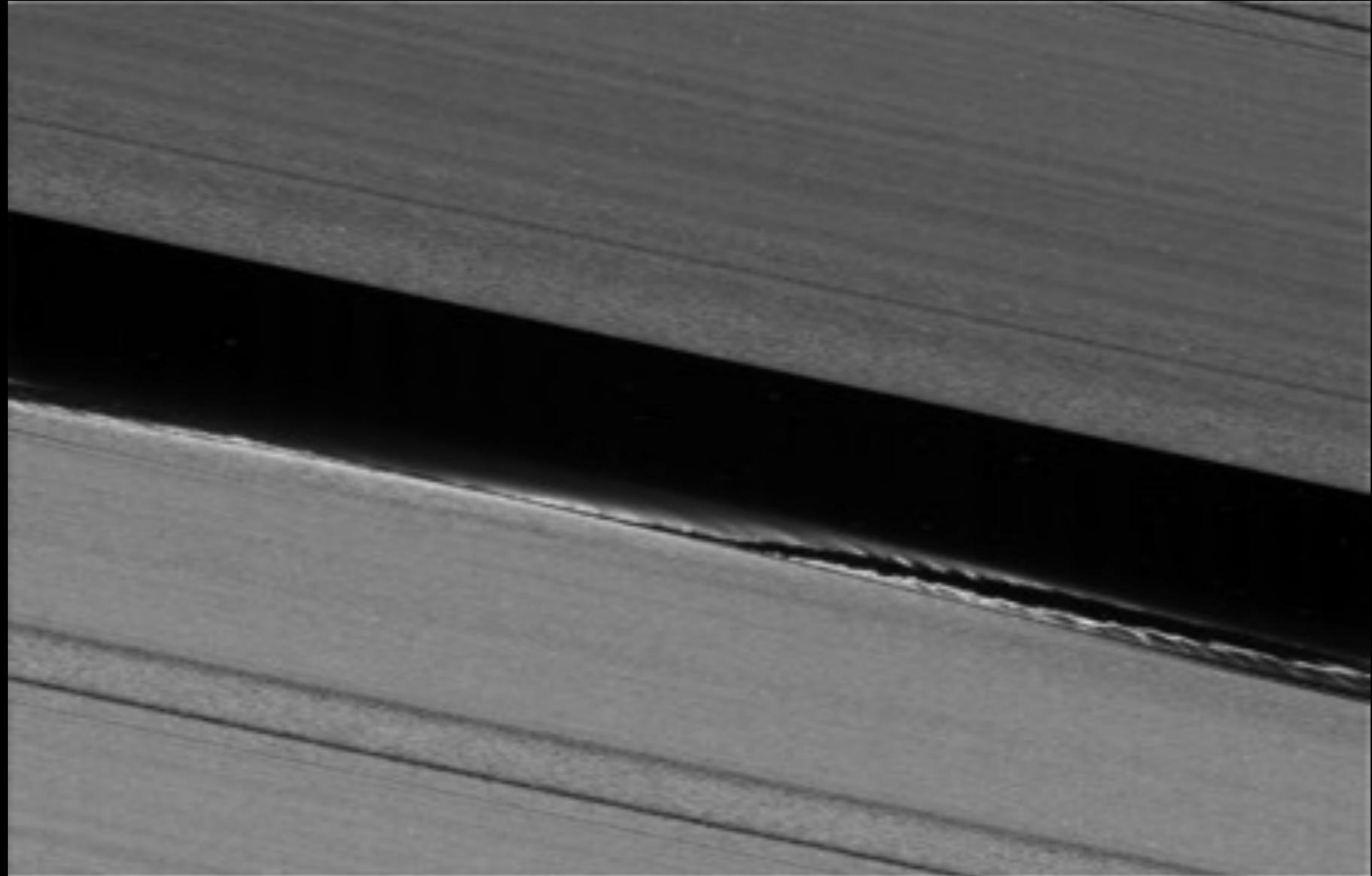
# Embedded moon, Daphnis, sculpts rings



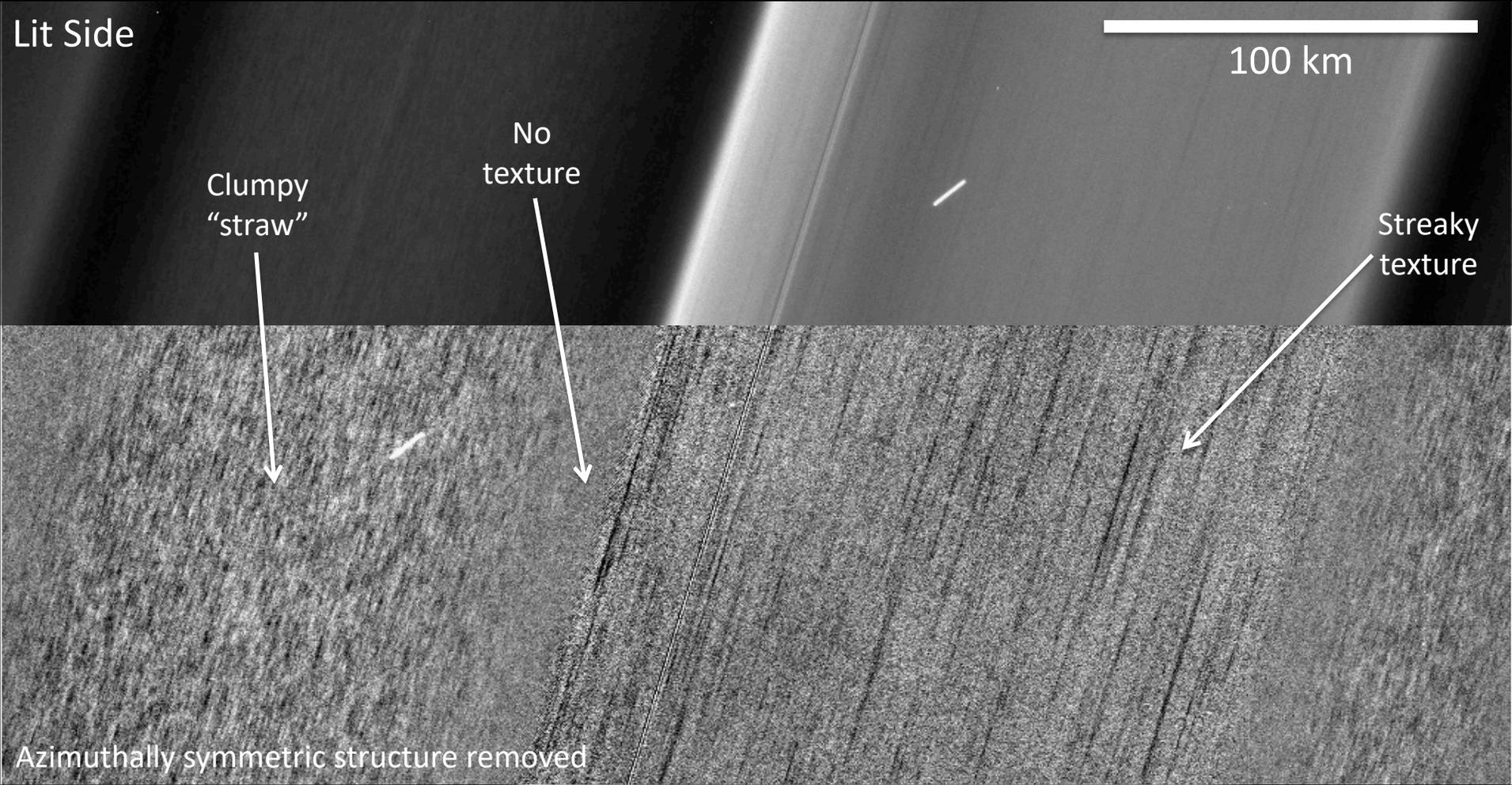
- Tiny Daphnis creates edge waves
- Color differences seen on either side of Keeler gap in enhanced mosaic
- Color change probably has more to do with particle properties including size rather than composition changes

# Multiple strands in 3<sup>rd</sup> wave crest

- Ribbon of material soars above and below ring plane
- Ribbon dives below rings on left-hand side and disappears
- Resolution: 230 m/pixel



# Plateau in the C ring



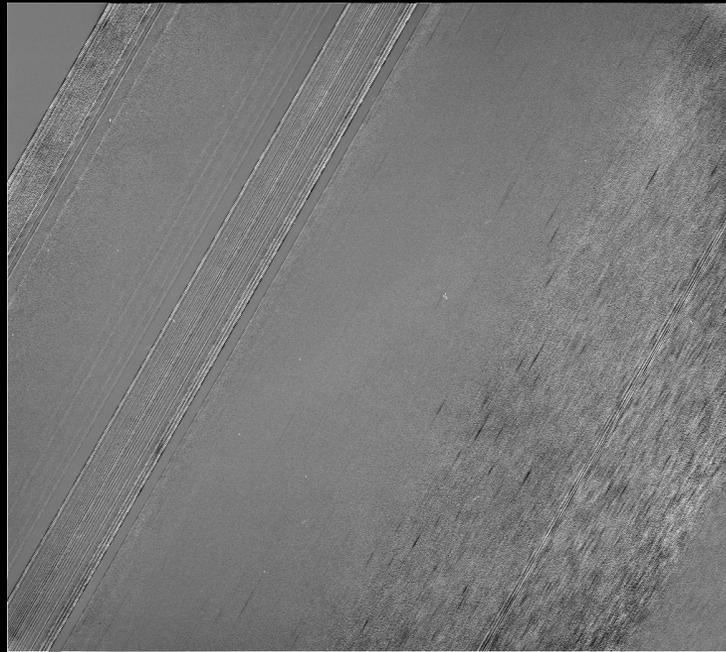
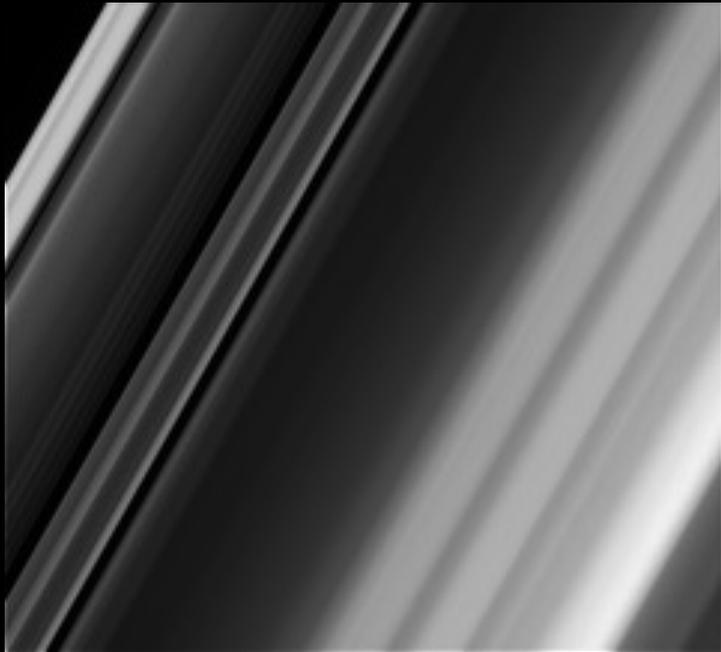
Resolution 330 m/pixel

Plateau P1



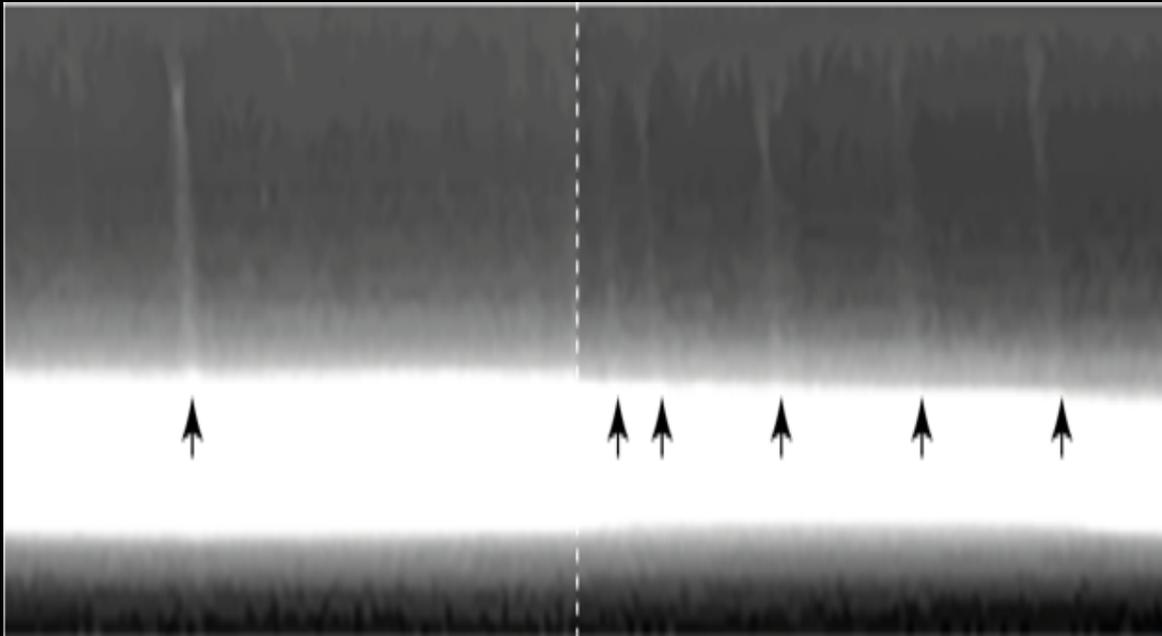
Tiscareno et al.,  
Science, 2019

# Texture in Outer Cassini Division



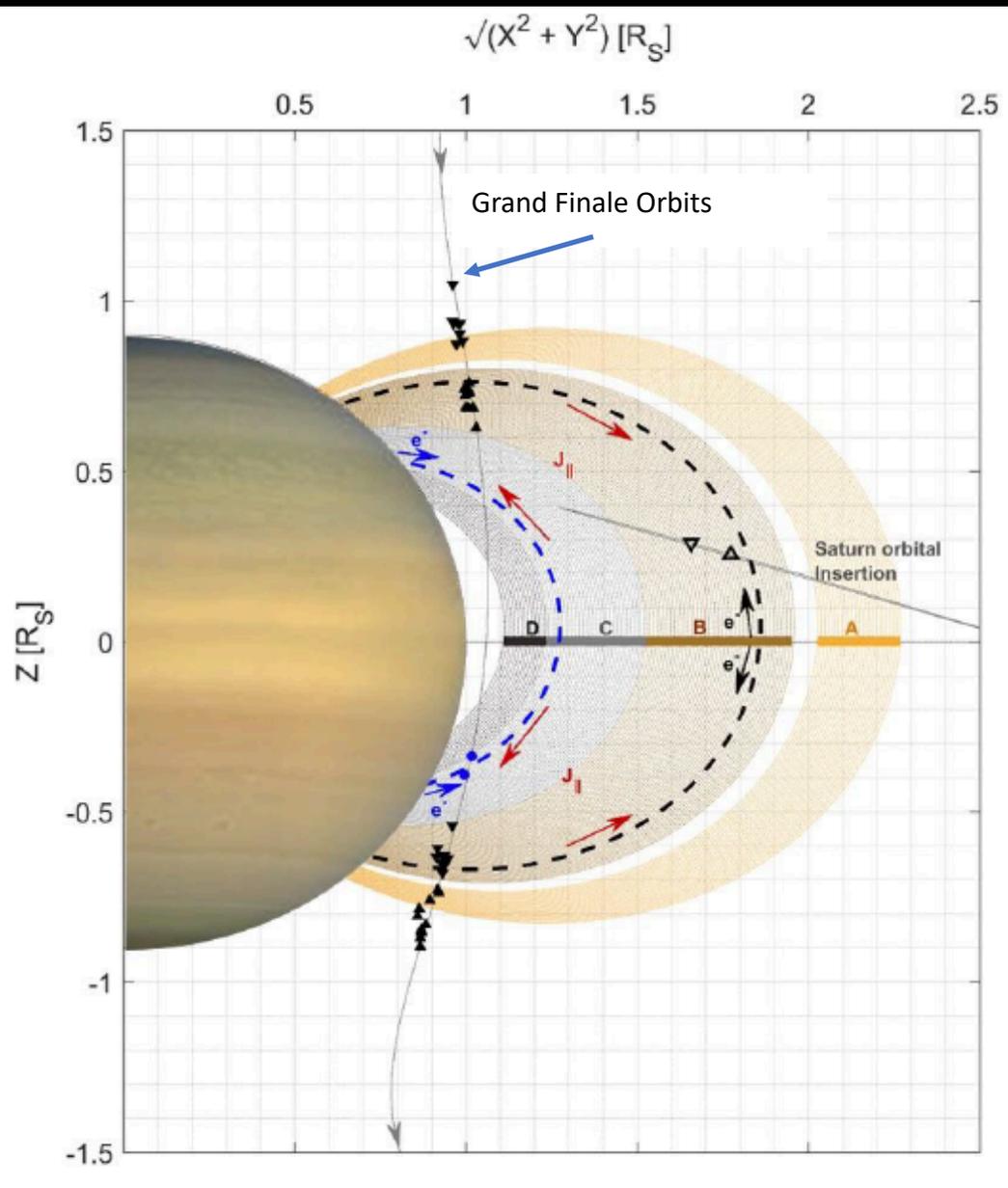
- Feathery texture apparent in 3 brightest regions of Cassini Division, but not in surrounding regions
- Right image filtered to enhance feathery texture
- Resolution 690 m/pixel

# Mini-Jets in Saturn's F ring Formed by Flock of Impactors



- Bright core of F ring is visible and overexposed to make streaks visible
- Streaks of material (see arrows) show evidence for a flock of small objects that impacted F ring at same time, leaving trails about 30 km long.
- Ring is shaped by streams of material that orbit Saturn itself rather than, e.g., cometary debris
- 2 images spliced together along dashed line

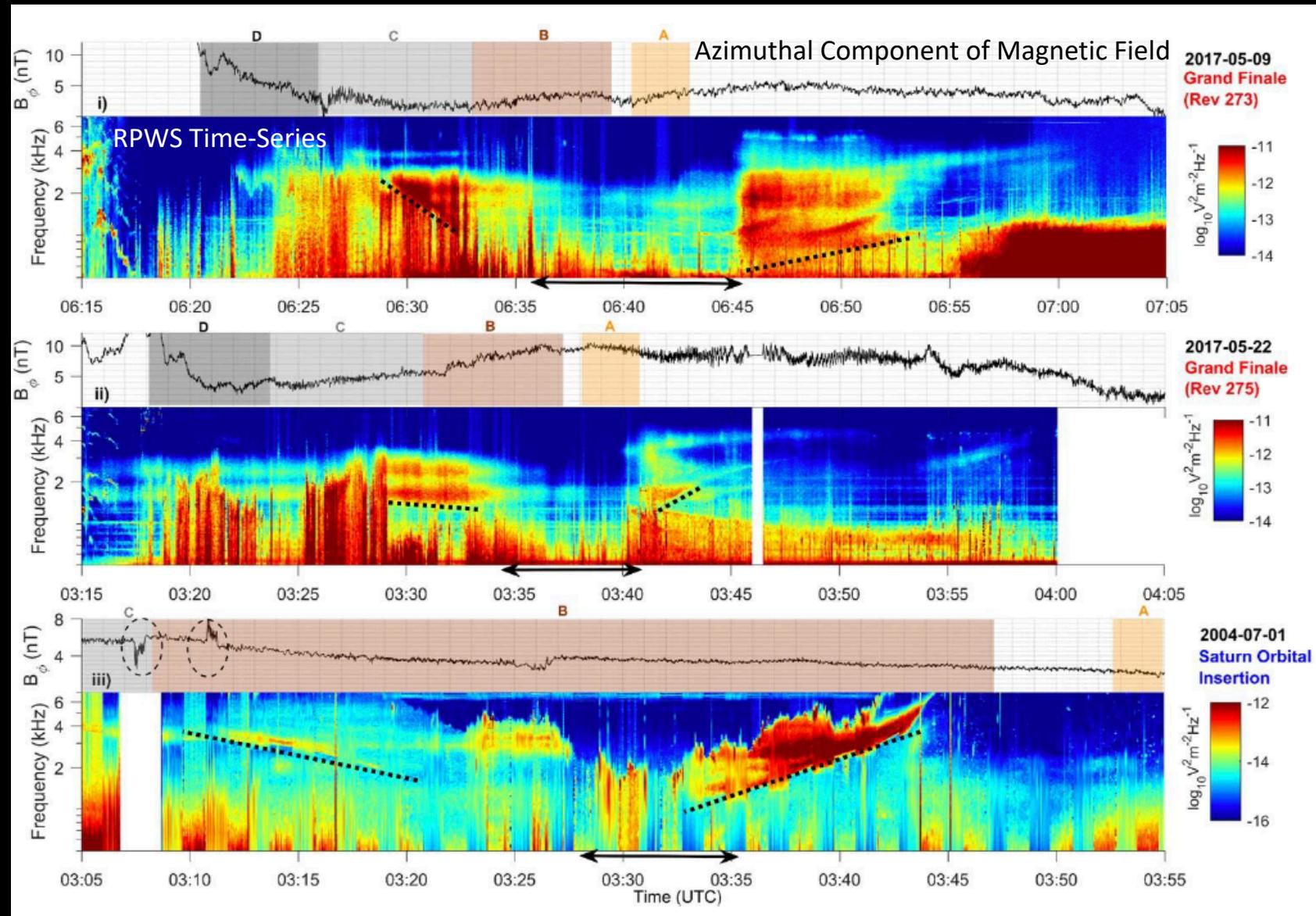
# Persistent large-scale electrodynamic coupling between Saturn and its rings

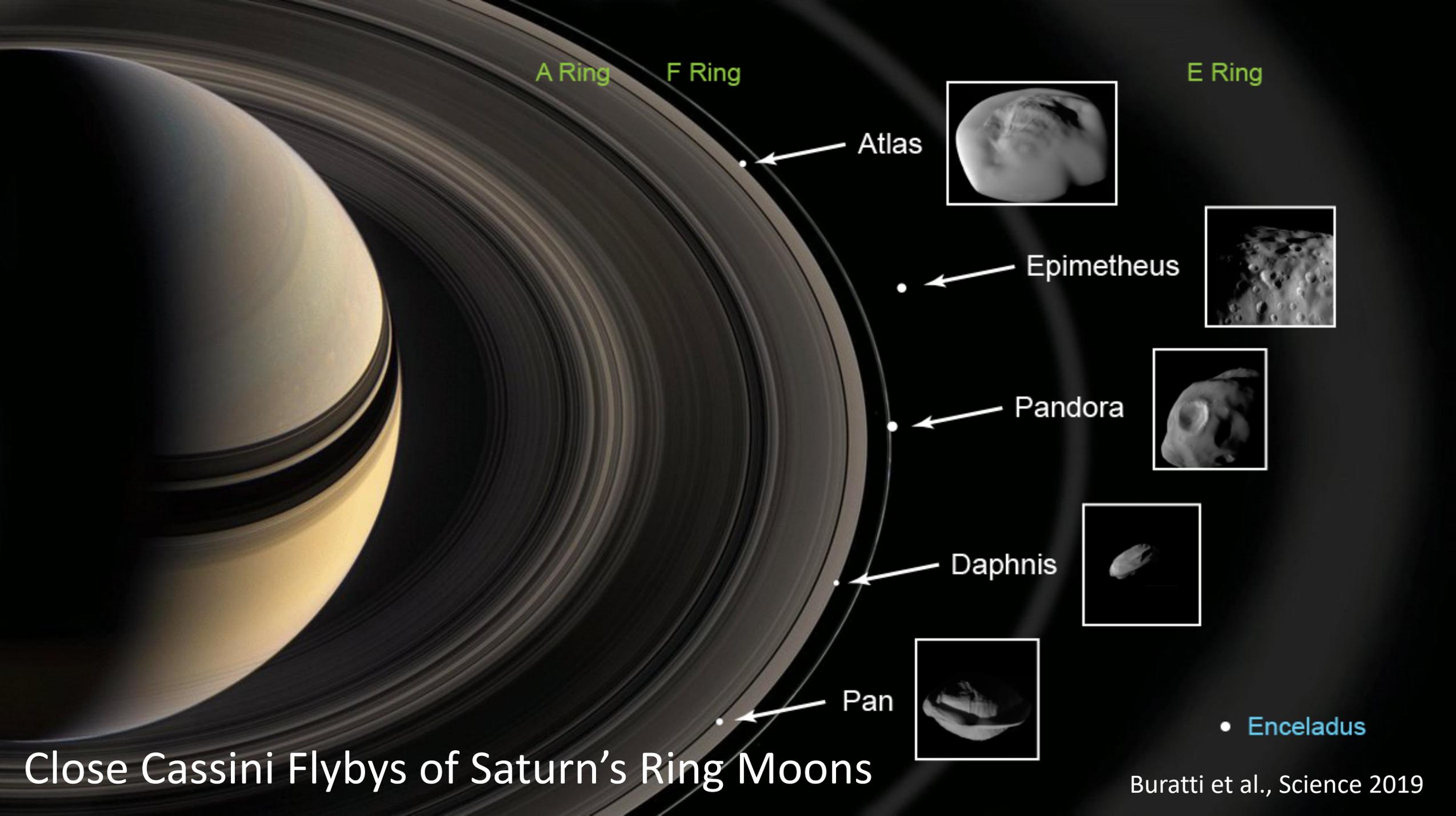


- Saturn's rings are a dynamic system in electromagnetic coupling with Saturn
- During all 22 Grand Finale orbits, the Radio and Plasma Wave Spectrometer (RPWS) repeatedly detected "auroral hiss" emissions on magnetic field lines connected to Saturn's rings
- North and south footprints of ring-connected field lines are at different N/S latitudes, indicating a small northward shift ( $0.047 R_S$ ) of magnetic equator
  - Blue dots are locations of source magnetic field lines

# Whistler-mode “auroral hiss” emissions in RPWS spectrograms

- Black solid line is  $B_\phi$
- Shaded areas correspond to L-shells connected to A, B, C and D rings
- Emissions propagate in whistler mode and exhibit characteristic funnel or V-shape in frequency-time spectrograms (dashed black lines)
- Vertex point (lowest frequency) is coincident with source of magnetic field line



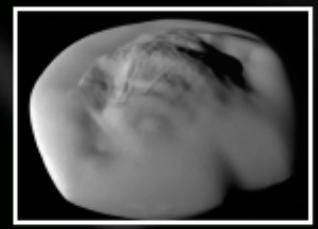


A Ring

F Ring

E Ring

Atlas



Epimetheus



Pandora



Daphnis



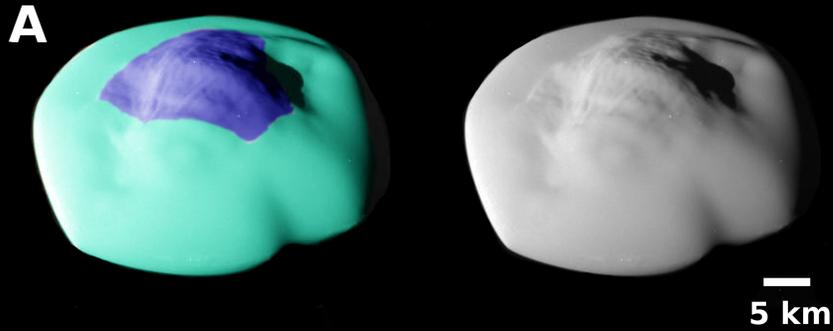
Pan



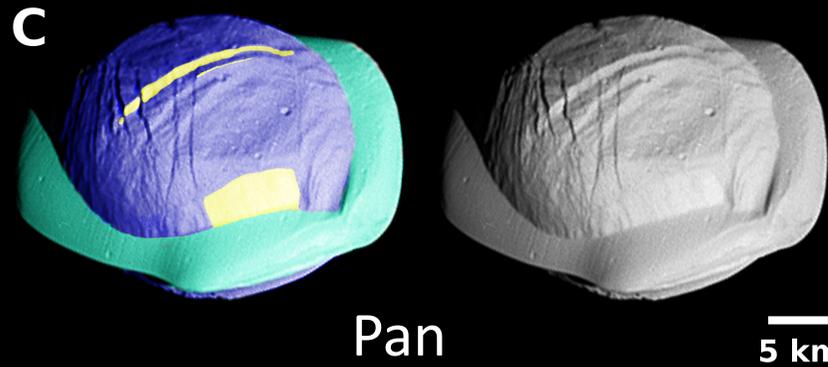
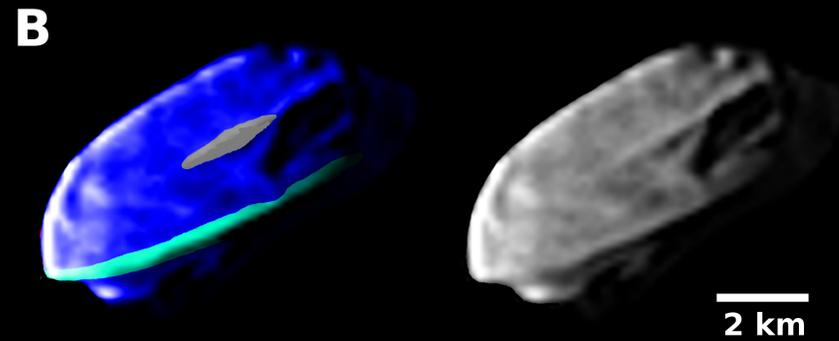
• Enceladus

# Close Cassini Flybys of Saturn's Ring Moons

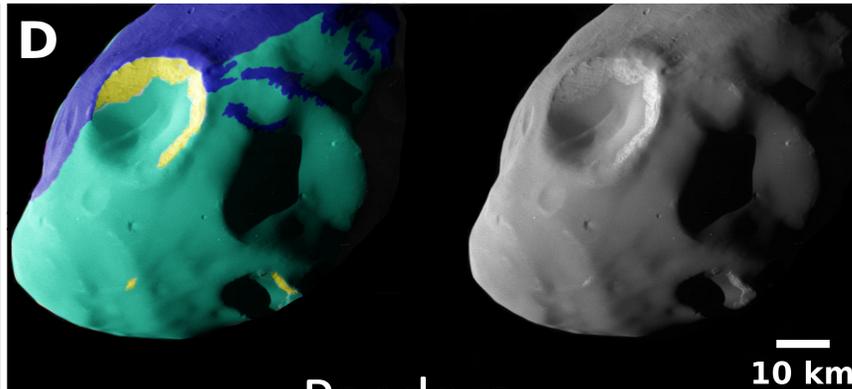
Atlas



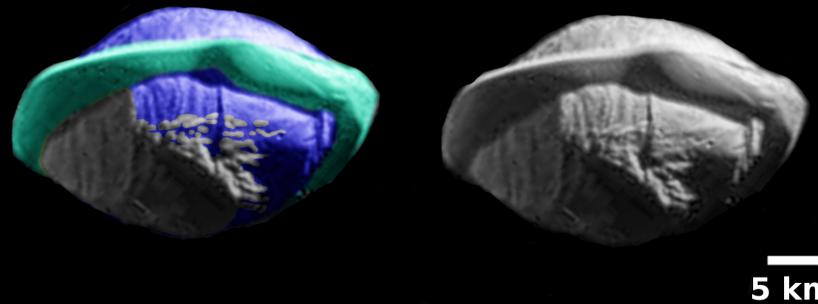
Daphnis



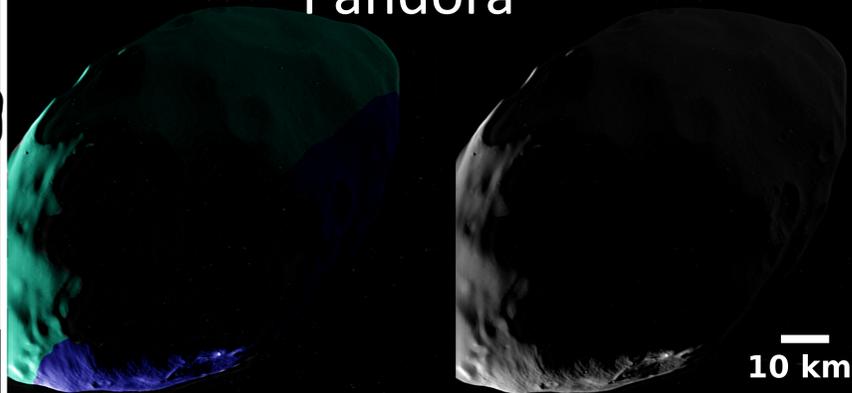
Pan



Pandora



5 km



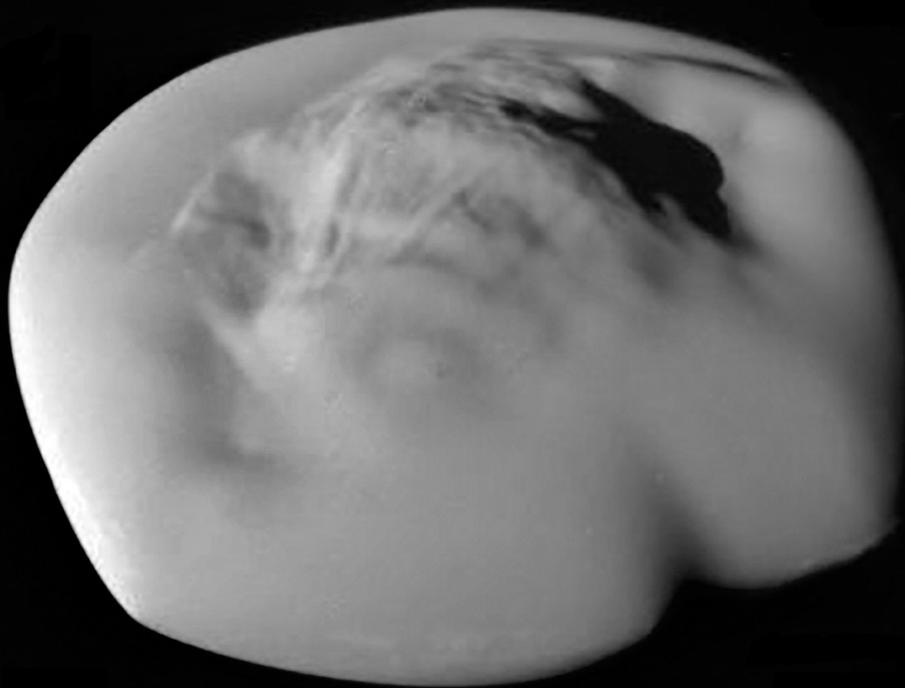
10 km

■ UNMODIFIED SURFACE MATERIALS    ■ EQUATORIAL BAND MATERIAL  
■ EXPOSED SCARP OR CRATER WALL    ■ UNCLASSIFIED

- Moons are highly porous
- Moons near rings have accreted equatorial ridges of ring material
- Ring moons closest to Saturn appear reddest, similar to main rings
- Three main geologic units shown in color
  - Blue: Cratered, unmodified surface
  - Cyan: Smooth terrain
  - Yellow: Exposed substrates

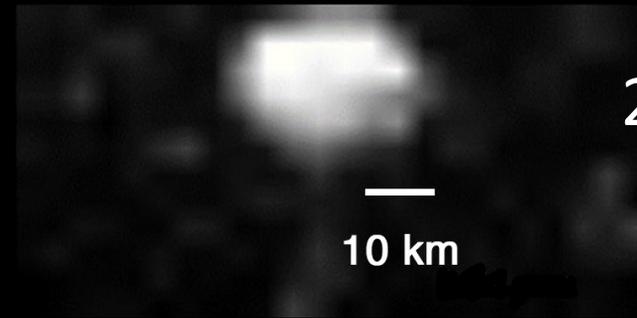
# Atlas

## A



10 km

## B

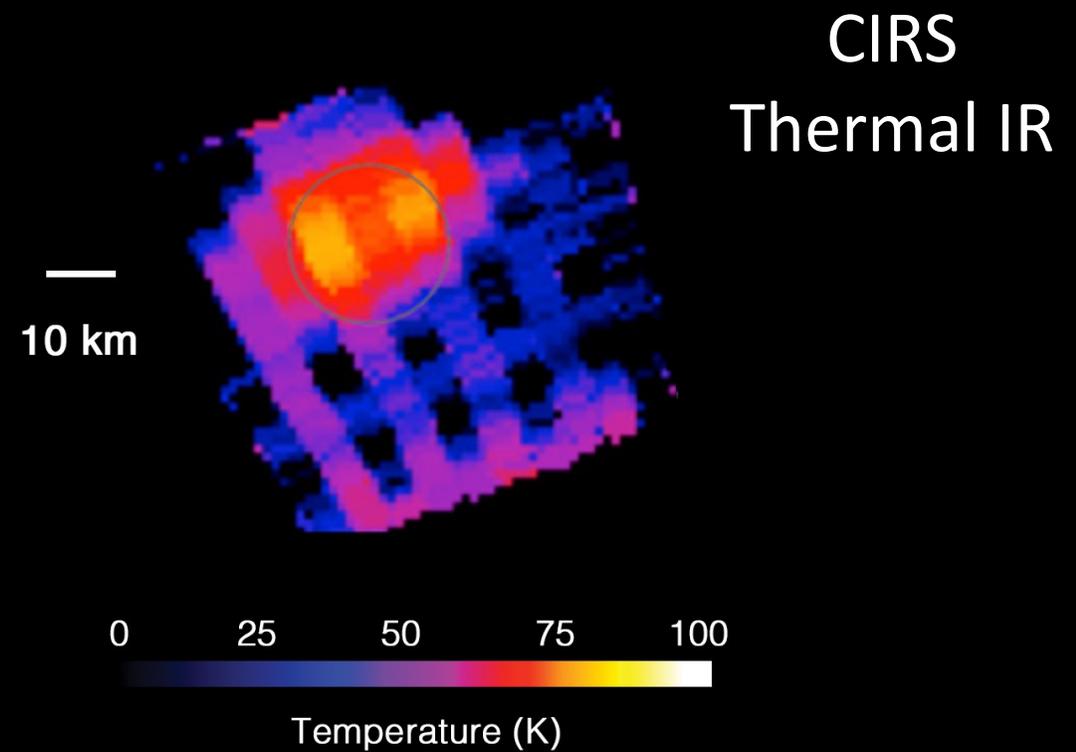


VIMS

2.01 microns

10 km

## C



CIRS

Thermal IR

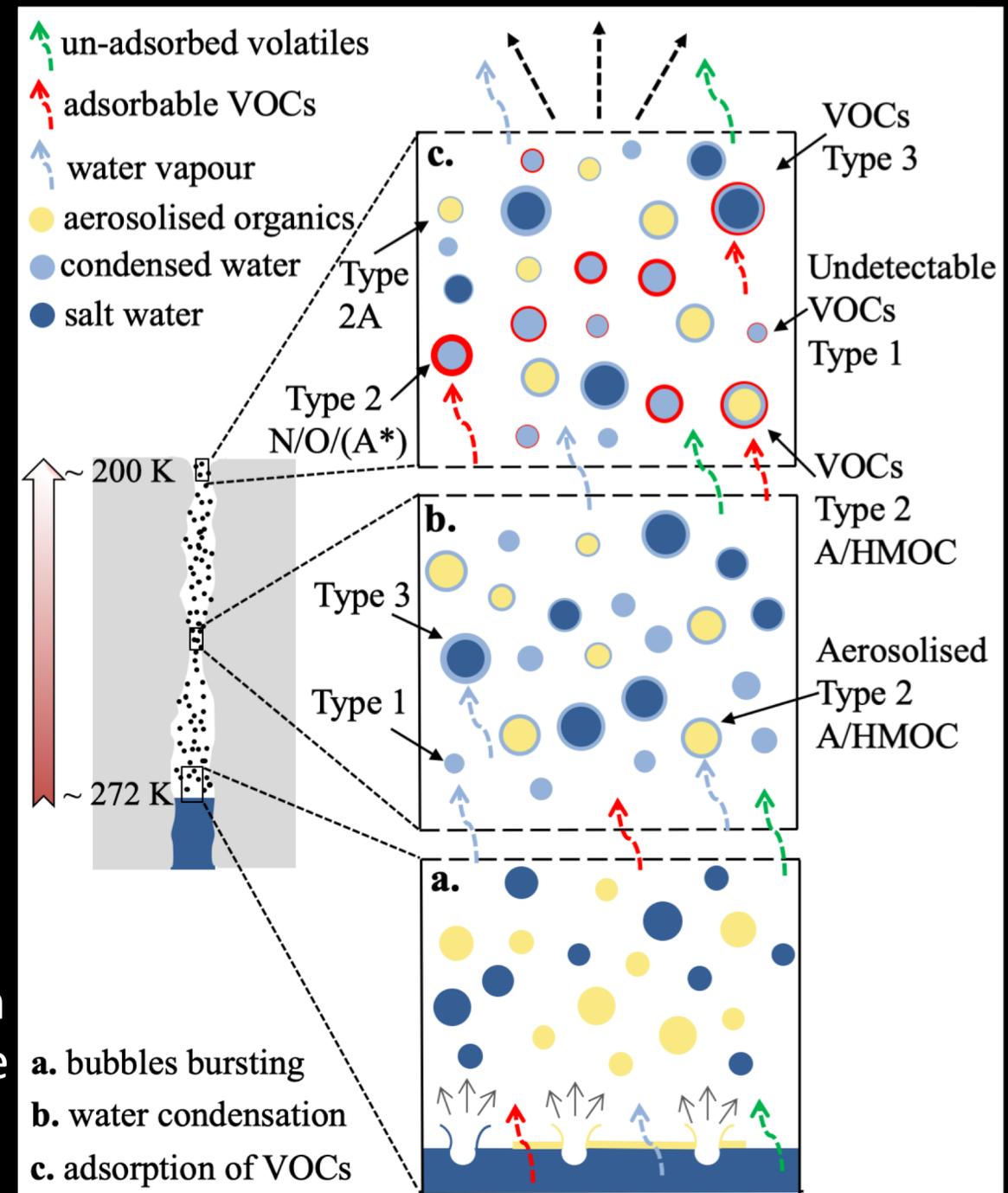
10 km

0 25 50 75 100

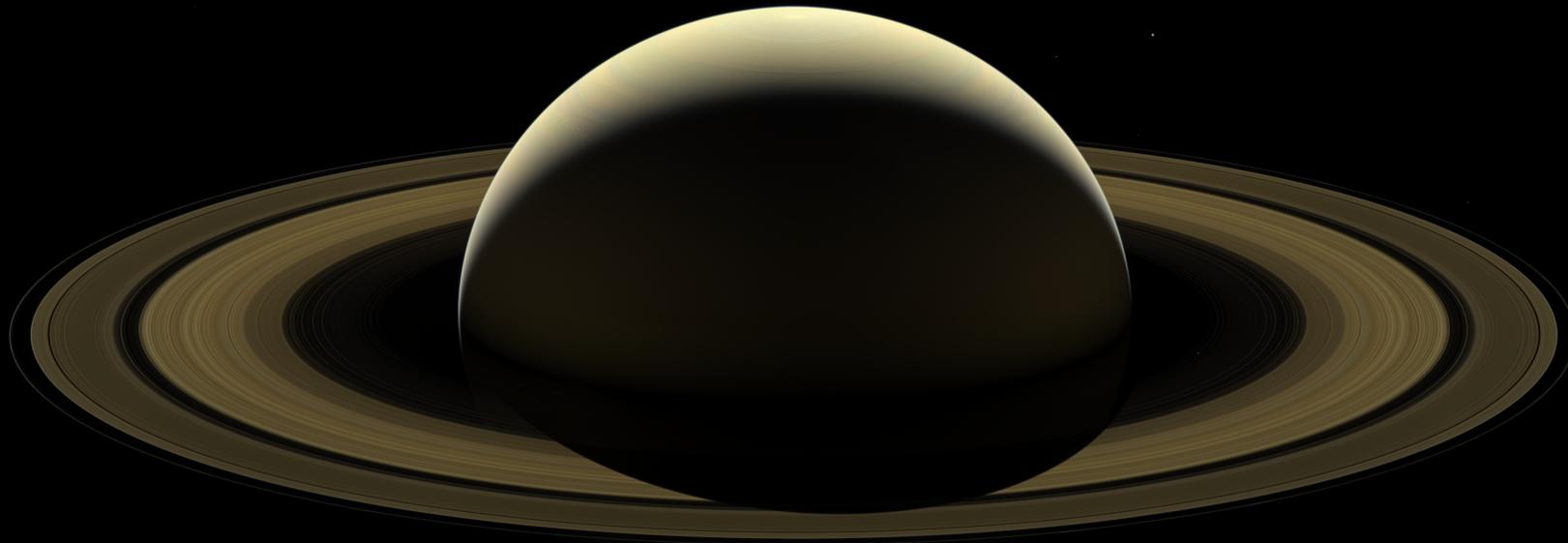
Temperature (K)

# Ice Formation Mechanisms in Enceladus Vent

- Gaseous organic molecules condense onto Enceladus ice grain
- Bubble bursting (a): Creates spray of salt-water aerosols & insoluble organics
- Water condensation (b): Vapor becomes supersaturated; condenses to form almost pure water ice grains & coats other grains
- Adsorption of volatile organic compounds (VOCs): N- & O-bearing VOCs adsorb onto pre-existing grains
  - Volatile gasses with low binding energy & high condensation temperatures leave in gas phase along with uncondensed water vapor



Questions?



Movie