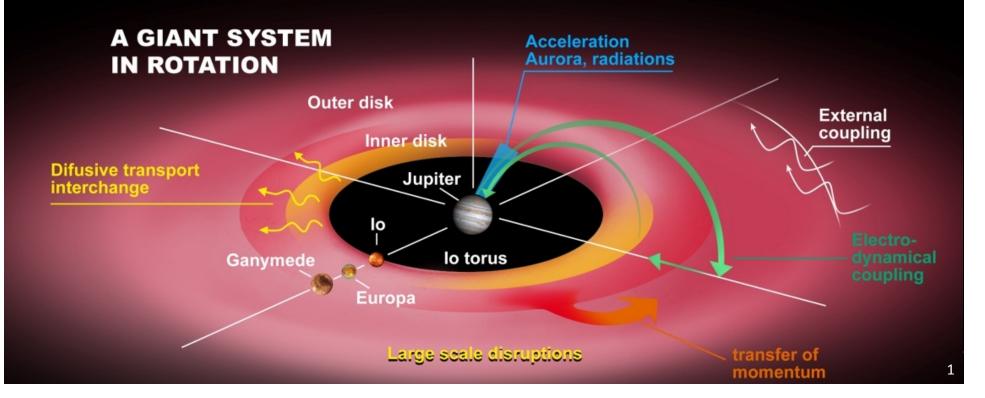
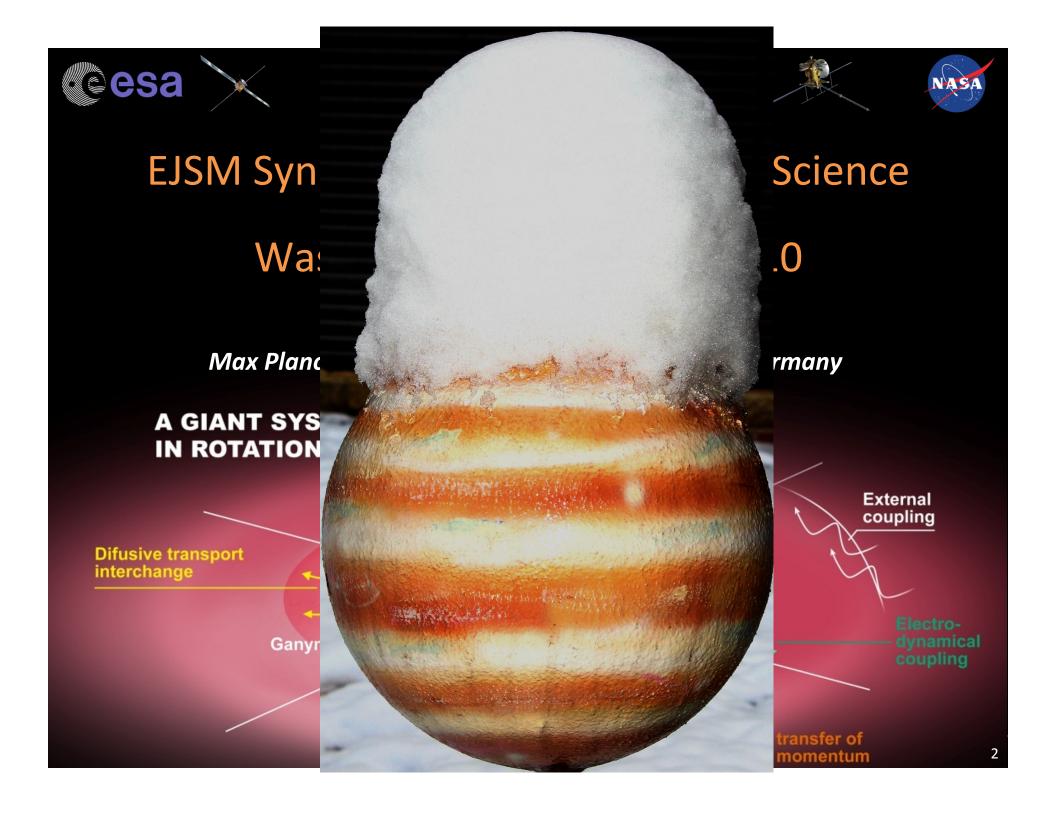




EJSM Synergistic Magnetospheric Science (SN)O(W)PAG Meeting Washington DC, Feb 8/9 2010

Norbert Krupp Max Planck Institute for Solar System Research, Germany







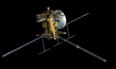






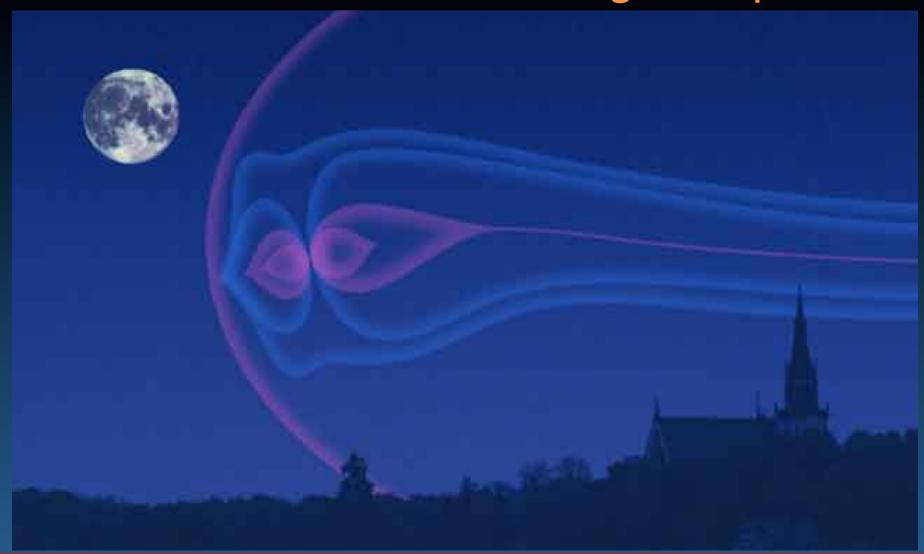








The size of the Jovian magnetosphere





- EJSM can for the first time tackle the global picture of the largest object of the solar system by using global imaging together with multi-point measurements
- Increased time resolution to look for anything of short duration
 - looking for ion pickup (ion cyclotron waves), short-lived interchanging flux tubes
 - resolving the boundary layers of the Ganymede magnetosphere.
- Solar wind monitoring capability with either spacecraft
 - to be able to see the response to change of sector orientation and to crossing of sector boundaries
 - to study the solar wind influence
- Enhanced time resolution coverage beyond 50 RJ
- Magnetospheric coverage capability away from the equatorial plane and beyond 150 RJ in the magnetotail
- Combined plasma data set/magnetometer measurements on both S/C
 - (to better understand the global and temporal variability of the magnetosphere)
- INMS instrument on JEO and JGO, ENA camera on JGO
 - to study the global magnetospheric dynamcis, Europa torus, and sputtered /backscattered neutral particles from the moon surfaces
- Multipoint- and combined remote-sensing/in–situ measurements

















Magnetospheric instrumentation on EJSM

spacecraft	instrumentation	Short cut
JEO	Magnetometer	MAG
JEO	Plasma/Energetic particle package	PPI
JEO	UV spectrometer	UVS
JEO	Ion Neutral Mass spectrometer	INMS
JGO	Magnetometer	MAG
JGO	Plasma/Energetic particle package	PLP/INMS
JGO	Plasma wave instrument	RPWI/LP
JGO	Ion Neutral Mass spectrometer	PLP/ INMS
JGO	UV spectrometer	UVIS















EJSM will provide enhanced Particle Energy Coverage

+

INMS and **ENA** instruments

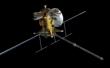
	JEO	JGO	GLL
INMS	Energy Resolution: 10% Reflectron Time-of-Flight Mass Spectrometer Mass Range: 1 to > 300 Daltons Mass Resolution: > 500	Thermal neutrals Mass range: 1-300	
Plasma/ energetic particle instrument	Plasma: Top Hat Analyzer 10 eV to 30 KeV electrons 10 eV to 30 KeV ions with composition Particles: Puck Analyzer 30 KeV to 1 MeV electrons 30 KeV to 10s of MeV ions High Energy Electrons: : >2 MeV, >4 MeV, >8 MeV, >16 MeV	PP: 1-20 keV electrons 1-30 keV ions (mass resolution 1-60, high time resolution) 1-10 keV ions with mass resolution 1-60 3-5000 keV ions with mass resolution 15-1000 keV electrons 10-10000 eV neutrals	PLS: 0.9 eV/q - 52 keV/q EPD: lons > 20 keV Electrons 15-884 keV; > 2 MeV; > 11 MeV lons (p, He, O, S and heavies 10 keV/nuc-15 MeV/nuc)





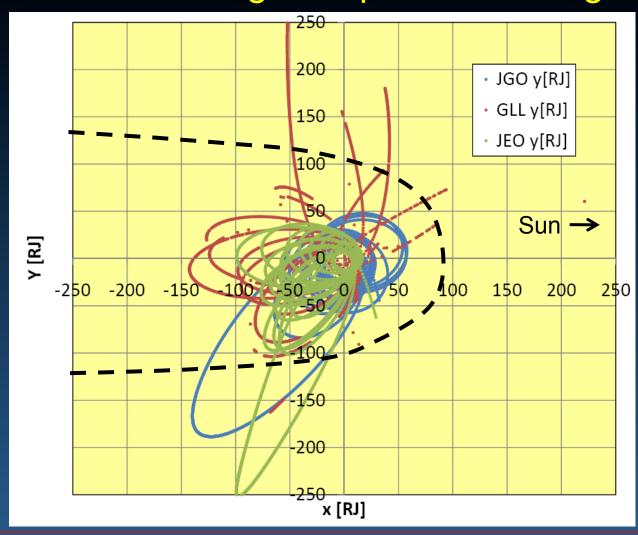








EJSM will provide enhanced magnetospheric coverage









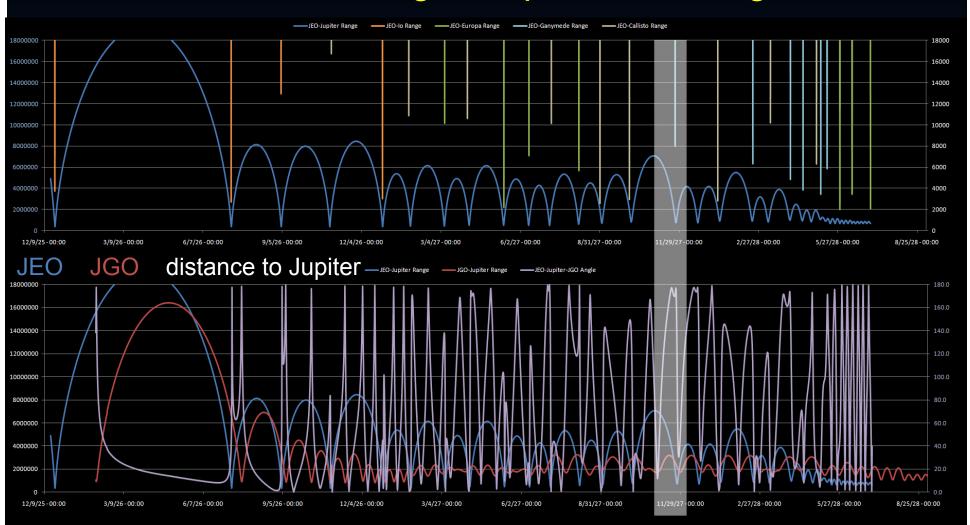








EJSM will provide enhanced magnetospheric coverage















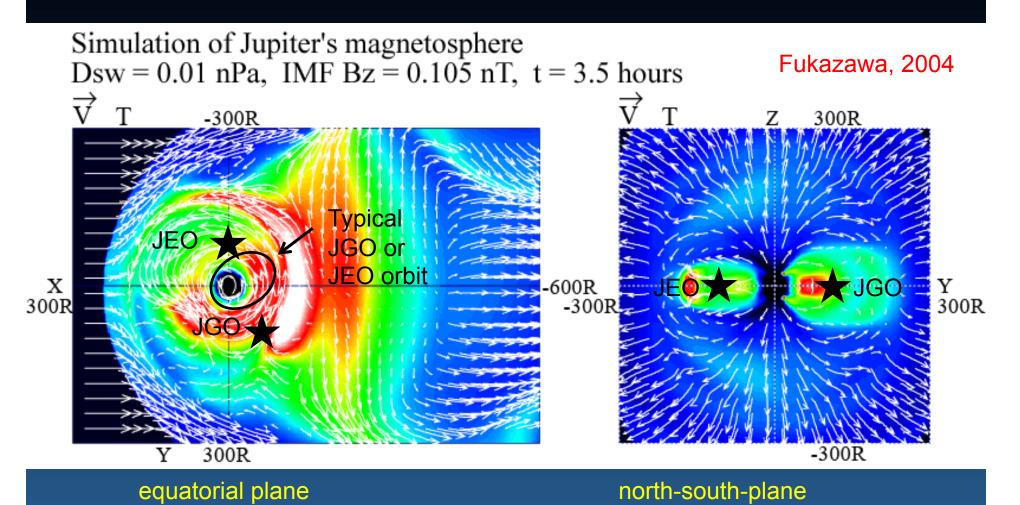


EJSM will be able to perform multipoint measurements

- One spacecraft in solar wind, the other inside the magnetosphere
- One in the outer magnetosphere, the other in the inner magnetosphere
- One spacecraft in the Jovian magnetosphere, the other in Ganymede's magnetosphere
- Two spacecraft in orbit around two moons in parallel

EJSM provides significantly enhanced capability to address magnetospheric dynamics

esa



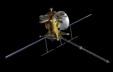














EJSM will provide

local - and global view of Jupiter's magnetosphere

The local view of satellite interaction and plasma transport

JEO +JGO in orbit around moons

The Jupiter/satellite and Jupiter/magnetodisc Coupling. Auroral Processes

JEO/JGO in orbit around moon + JEO/JGO in magnetosphere

Investigation of disc dynamics and large scale transport

JEO +JGO in Jovian magnetosphere







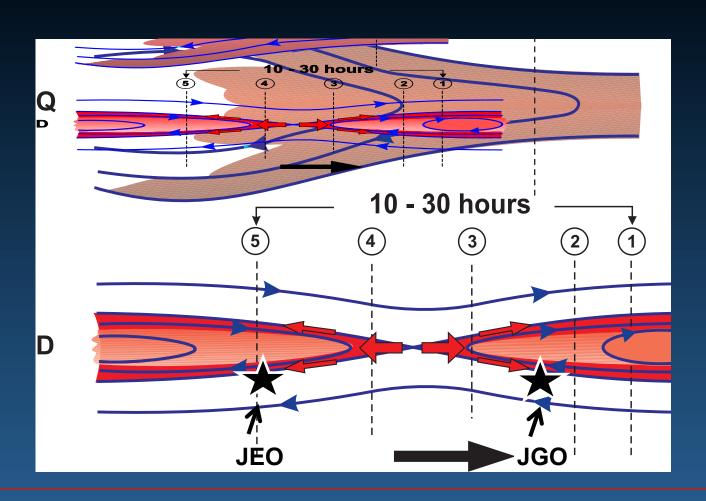






EJSM capability to perform

multipoint measurements in inner and outer magnetosphere in parallel

















EJSM capability to perform remote sensing and in-situ measurements in parallel

- Aurora:
 - Remote sensing of auroral emissions in parallel to particle and fields measurements
- Io and Europa torus:
 - Remote sensing of torus emissions (UV, IR, ENA) with JGO when JEO is inside the tori (particles and fields) and vice versa
- Ring current measurements (ENA+in-situ particles)







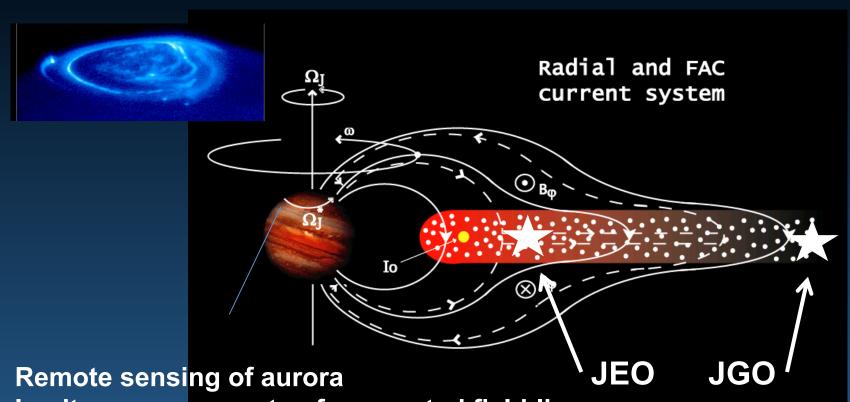








EJSM capability to perform multipoint + remote/in-situ measurements



- In-situ-measurements of connected field-lines
- In-situ multipoint measurements in plasmadisk







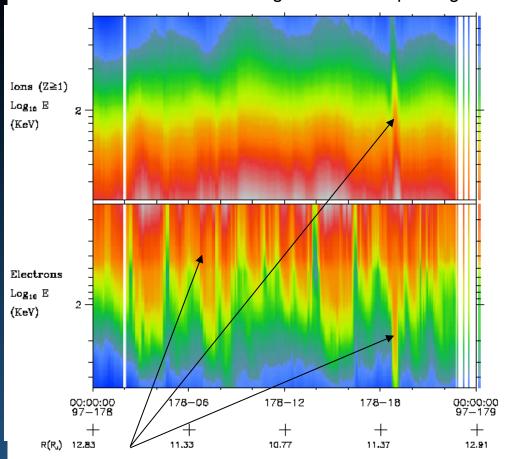


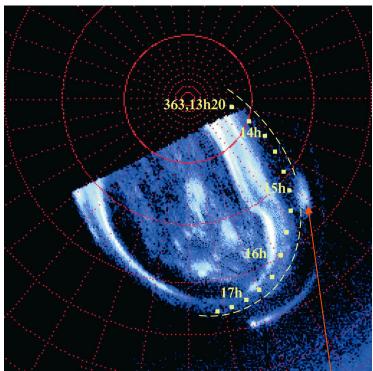






EJSM will enhance the understanding of dynamic processes in the Jovian magnetosphere 24 hour Galileo Energetic Particle Spectrogram HST Image of Jupiter's UV aurora





Mauk et al., 1997; 1999, 2000

Extreme "storm-time" energetic particle dynamics observed in the vicinity of Europa's orbit

Auroral manifestation (HST) of near-Europa storm dynamics



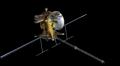










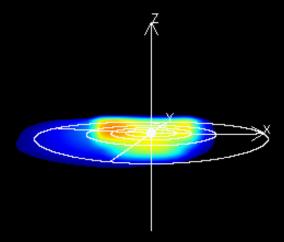


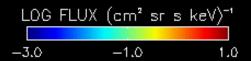


EJSM capability to perform remote sensing (JGO) with in-situ (JEO) measurements

Expected ENA emissions from the EUROPA torus as viewed from JGO orbit (P. Brandt)

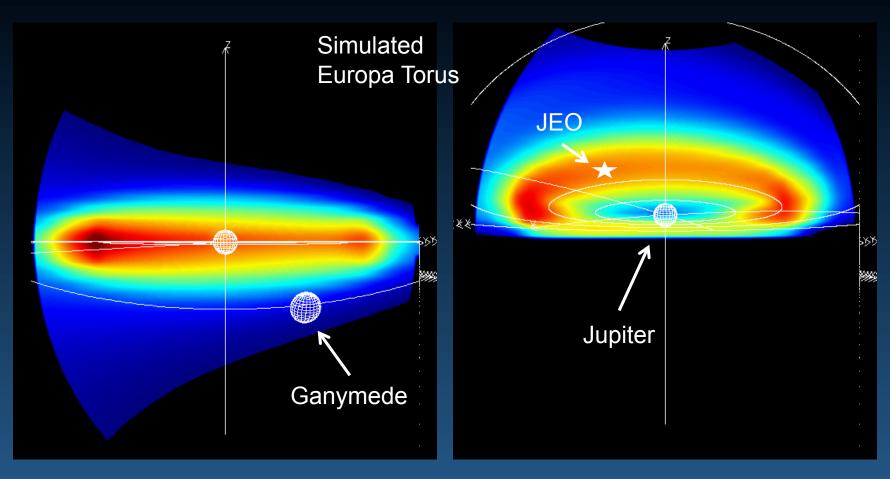
40 keV protons







EJSM capability to perform remote sensing (JGO) with in-situ (JEO) measurements



Expected ENA emissions from the EUROPA torus as viewed from JGO orbit (P. Brandt)







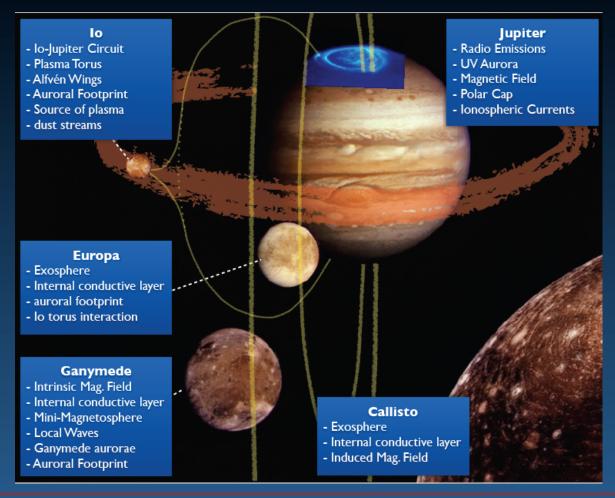








EJSM capability to study the electromagnetic environment of Jupiter



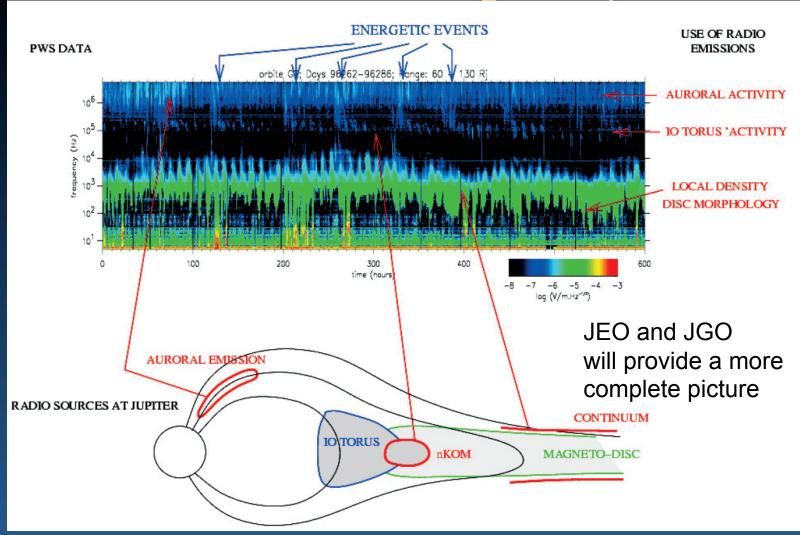
remote and in-situ

plasma wave (JGO only)

UV, particle and field measurements (JGO and JEO)



Radio emissions from various regions inside the Jovian magnetosphere





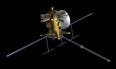








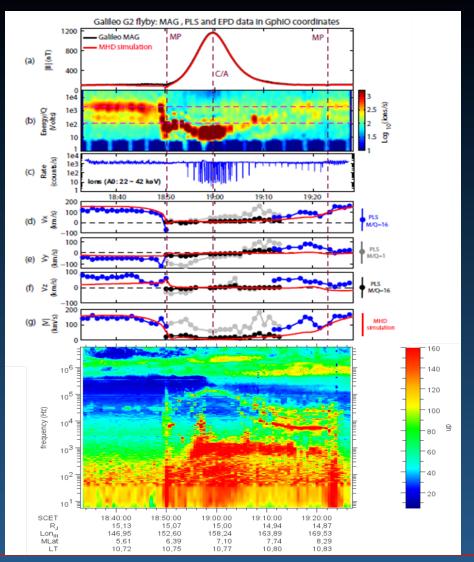






Ganymede environment (Galileo multi instruments)

JGO/JEO
would allow
to do this kind of
combined
measurements
from two
perspectives



Magnetic field

Plasma composition

Energetic particles

Plasma velocity

Plasma waves

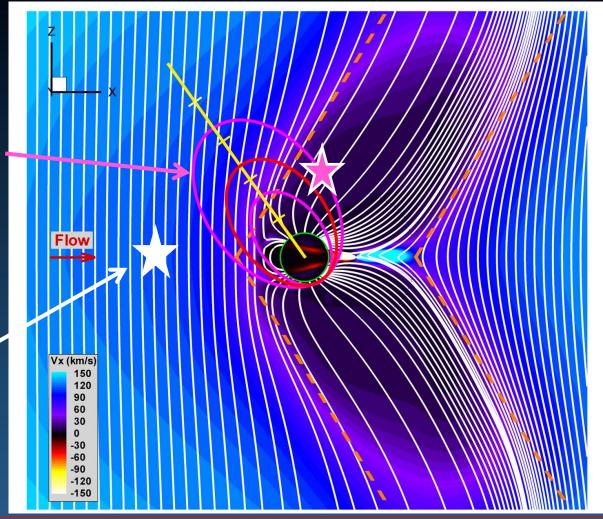


EJSM capability to investigate

the boundaries of Ganymede's magnetosphere in detail

JGO in elliptic / circular orbit around Ganymede

JEO outside Ganymede's magnetosphere providing parameters of Jupiter's magnetosphere



MHD Simulation of Ganymede's magnetosphere from Jia et al. 2009















Jupiter magnetospheric Orbiter (JMO)

- JAXA considers a Jupiter Magnetospheric Orbiter JMO
 - Spin stabilized
 - P/L resource > 40kg (In-situ ~20kg, Remote ~15kg)
 - * Low / Medium / High Energy Particle Packages
 - DC Electric / Magnetic Field & Wave Measurements
 - * Wide Field ENA Imager for Energetic Regions
 - * Wide Field EUV Imager for Io-Torus
 - * Multi-Element Radio Receiver (kHz ~ 50MHz [+ Synchrotron monitor])
 - * Dust
 - * Camera
 - Orbit 10 ~ 100Rj, w/o detailed analysis such as the flyby tour for inclination raising
- Launch schedule :
 - 2022 Launch, ~2028 arrival to Jupiter for joint observation with JEO/JGO

















Unique magnetospheric science from JMO

- JMO would allow 3-point measurements inside the magnetosphere and additional remote/in-situ possibilities
- JMO would allow high-latitude (10-30 deg inclination)
 measurements from large distances (>100 RJ) during phases
 when JGO and JEO are deep in the inner magnetosphere
- If launched slightly later than EJSM JMO could serve as solar wind monitor
- Additional instrumentation such as x-ray spectrometer, dust detector, ENA high-energy instrument would complement EJSM
- JMO as a spinning spacecraft would allow particle and fields measurements from all directions.











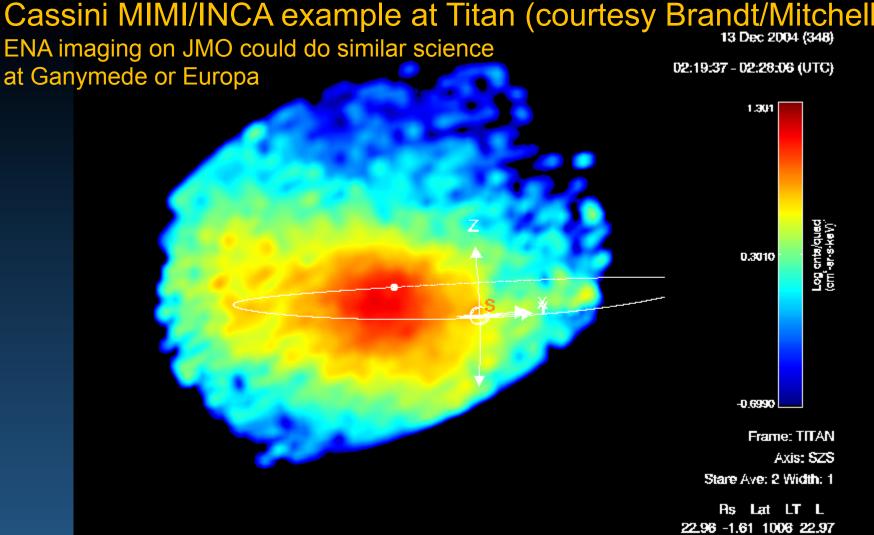






Cassini/MIMI Inca Spatial H+ 20-50 keV

Cassini MIMI/INCA example at Titan (courtesy Brandt/Mitchell)



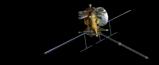
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Summary of EJSM Synergistic Magnetospheric Science

- Multipoint measurements will dramatically enhance the understanding of complex processes in magnetosphere
- Remote and in-situ measurements in parallel will provide new insights into related processes in completely different regions
- The JEO-JGO synergy enables us to address how material from the moons is circulated throughout the magnetosphere, accelerated and reimpact the moons.
- The JEO-JGO-JMO synergy gives us an unprecedented opportunity to reveal the global machinery of the Giant Particle Accelerator of the solar system, providing the ground truth for the planetary environments of other stars.