NASA JPL

L S T I C E

Linda Spilker Cassini Project Scientist Outer Planets Assessment Group 15 July 2013

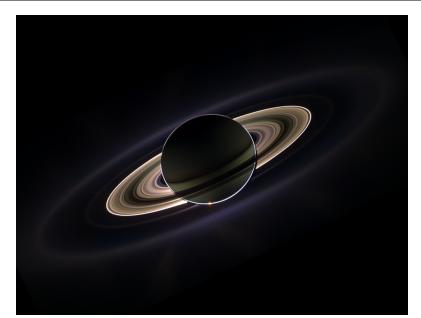


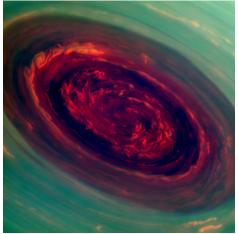
Cassini Solstice Mission



Cassini Solstice Mission: Introduction

- Key Cassini events
 - New Cassini Program Manager
 New Participating Scientists
 Recent Flybys
- Latest Science Highlights
 - ≻Last Rhea Flyby
 - ≻Titan Ice Floats
 - ≻Meteor Strikes in Rings
 - Kronoseismology
 - ≻Polar Hurricane
 - ≻Auroral Campaign
 - Enceladus Plume Variability
- Upcoming Cassini Senior Review
- Wave at Saturn campaign







Cassini Solstice Mission



New Program Manager: Earl Maize





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New Cassini Participating Scientists

Caitlin Griffith



Alex Hayes



Henrik Melin





Francis Nimmo



Christopher Parkinson



Mark Perry

Joe Spitale



Matt Tiscareno



Ann Verbiscer





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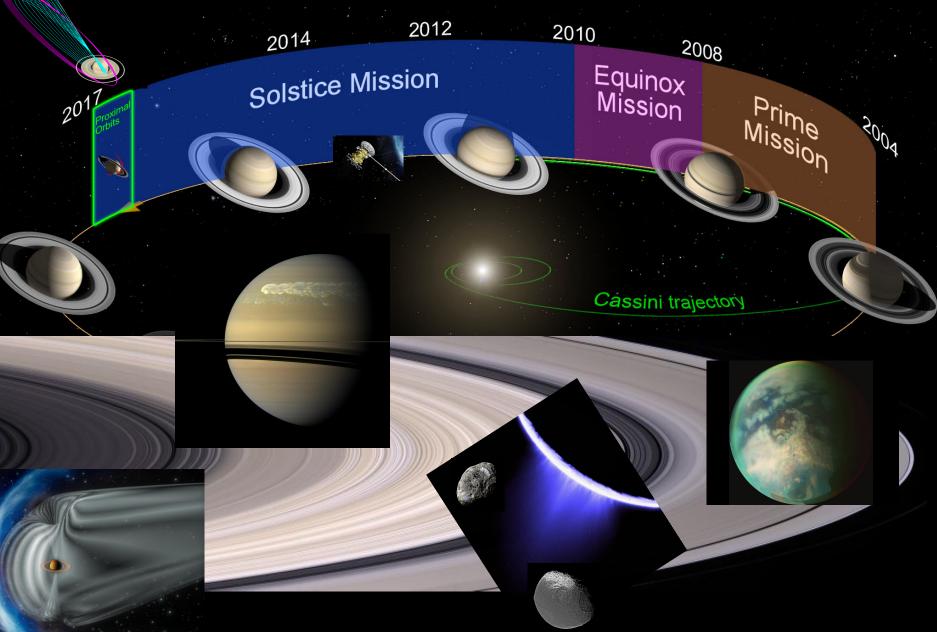


New Participating Scientists

PI	Instruments/DWG	Institution	Title	Discipline	E-mail
Caitlin Griffith	primary: VIMS occasional: CIRS, RADAR /TWG	University of Arizona	Titan's Methane Cycle: An analysis of Cassini Data	Titan	Griffith@lpl.arizona.edu
Alex Hayes	RADAR / TWG (VIMS, ISS through proposal Co-Is)	Cornell University	Seas, Lakes, Channel Networks and Hillslopes: A Coupled Analysis to Explore the Evolution of Titan's Polar Landscapes	Titan, Icy Satellites	hayes@astro.cornell.edu
Henrik Melin	UVIS, VIMS	Space Environment Technologies	Simultaneous infrared and ultraviolet observations of Saturn's aurora using Cassini VIMS and UVIS	Saturn, Magnetosphere	h.melin@gmail.com
Francis Nimmo	RSS, RADAR, NAV / ISWG	University of California, Santa Cruz	Integrating shape and gravity data to investigate Saturnian satellite structure and evolution	Icy Satellites	fnimmo@es.ucsc.edu
Christopher Parkinson	UVIS (main), CIRS, VIMS, ISS / SWG	University of Michigan	Analysis of Extreme and Far Ultraviolet Observations of Saturn ¹ s Atmosphere	Saturn (main), Icy Satellites	theshire@umich.edu
Mark Perry	INMS / MAPS WG	John Hopkins Univ., APL	Analysis of INMS observations of ions and neutrals in Saturn's inner magnetosphere	Magnetosphere, Icy Satellites	mark.perry@jhuapl.edu
Joe Spitale	ISS / ISWG	Planetary Science Institute	Instantaneous Jet Source Locations on Enceladus: Testing the Tidal Control Hypothesis	Icy Satellites: Enceladus Jets	jnspitale@psi.edu
Matt Tiscareno	ISS / RWG (comm. with ISWG)	Cornell University	Dynamics of Saturn's Rings and Moons	Rings (mainly), Icy Satellites	matthewt@astro.cornell.edu
Anne Verbiscer	ISS, VIMS, CIRS / ISWG	University of Virginia	Spectrophotometric Analysis of Thermally Anomalous Terrain on Icy Saturnian Satellites	Icy Satellites	verbiscer@virginia.edu

Cassini-Huygens: Mission to Saturn





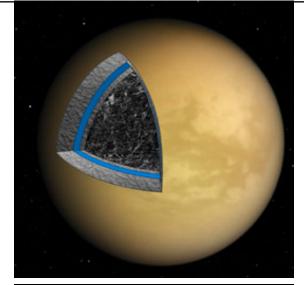


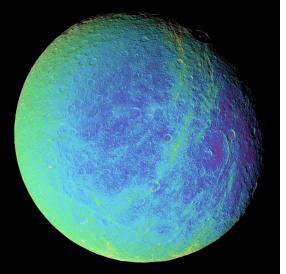
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Flybys

- T89 Titan flyby (1978 km) Feb. 17, 2013
 RSS gravity flyby (1 of 4 in CSM) to understand possible ocean and soft interior
- R4 Rhea flyby (997 km) Mar. 9, 2013
- Determined gravity to understand internal structure, measured dust environment, and imaged surface in UV, IR, and visible
- T90 Titan flyby (1400 km) Apr. 5, 2013
- ORS flyby: CIRS searched for atmospheric seasonal change and VIMS looked for surface evolution in southern fall





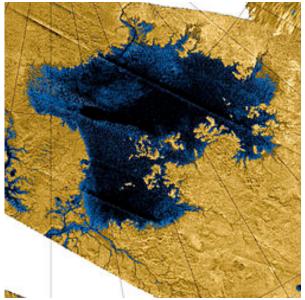


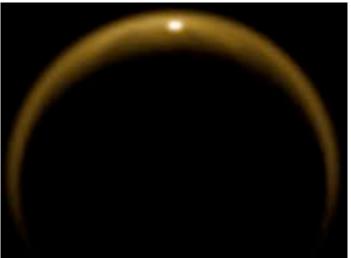
Cassíní Solstice Míssíon



Flybys

- T91 Titan flyby (970 km) May 23, 2013
 RADAR flyby to addresses Titan lake depths and seasonal change and to understand
 - relationship between Titan's gravity and surface topography
- T92 Titan flyby (964 km) Jul. 10, 2013
- RADAR flyby: stereo pair (T91 and T92) to look for lake change and surface porosity
- T93 Titan flyby (1400 km) Jul. 26, 2013
 ORS flyby: Search for lake specular reflections and monitor mid-latitude cloud formation



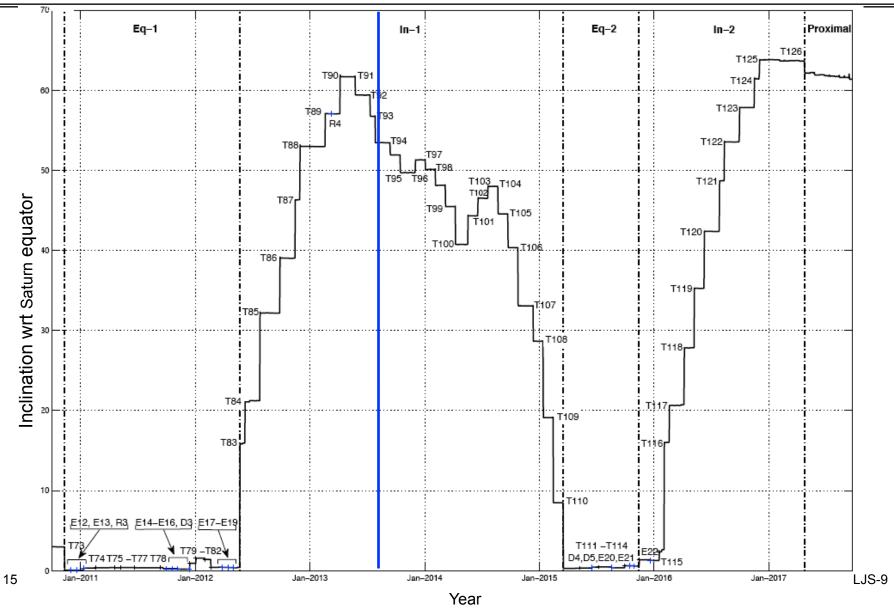




Cassíní Solstice Míssíon



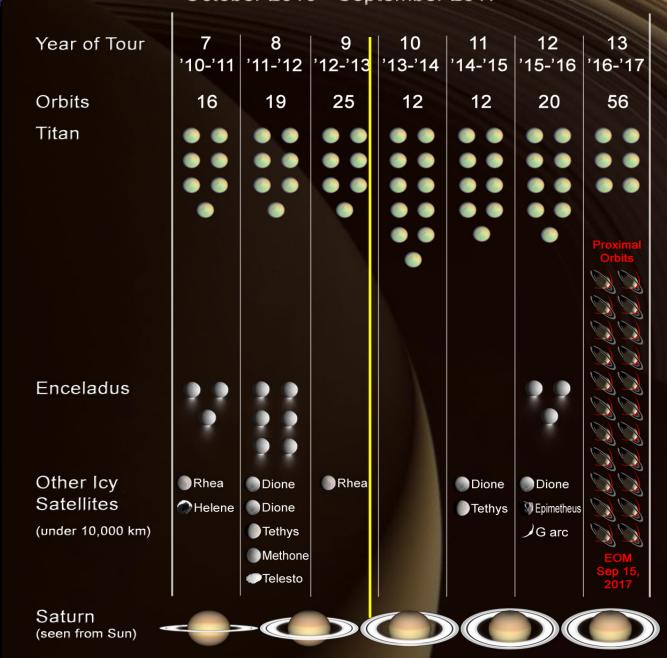
Solstice Mission Inclination Profile





Cassini Solstice Mission Overview October 2010 - September 2017



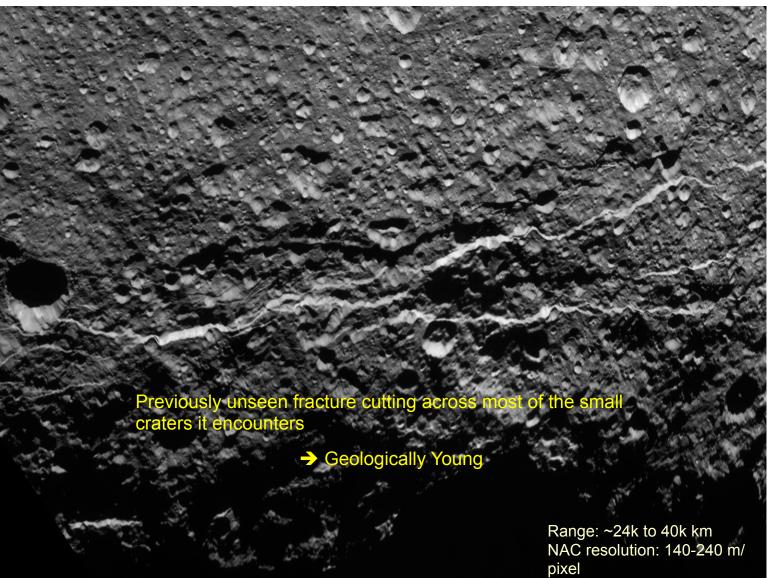




Cassini Solstice Mission



Last Rhea Flyby of the Cassini Mission!



15 July 20

Phase: ~40

LJS-11



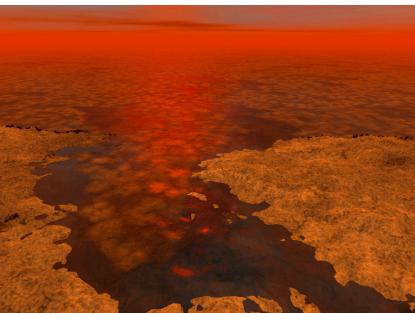
Cassini Solstice Mission

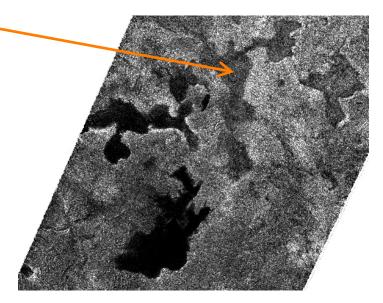
JPL

Titan Ice Floats

- Seasonal ice in Titan's lakes probably floats*.
- Methane ice exists if winter temperature is below freezing point of pure methane (-297°F).
- Ethane ice formed with 5 to 10% air will initially float but will sink if the temperature drops by just a few degrees.
- Radar data acquired in the winter shows a granular lake texture suggesting that ice may be making the lake surface bumpier in the winter than the flat, calm summer lakes.
- Unlike water ice in Earth's oceans, ice in Titan's lakes and seas may float, sink and rise again to the surface as the temperature changes.

* "Does Ice Float in Titan's Lakes and Seas?", Jason D. Hofgartner, Jonathan I. Lunine, *Icarus, In Press, Accepted Manuscript, Available online 30 November 2012*





Recent Titan Radar images show some of Titan's lakes. The top right lakes show a granular texture that may indicate floating ice.

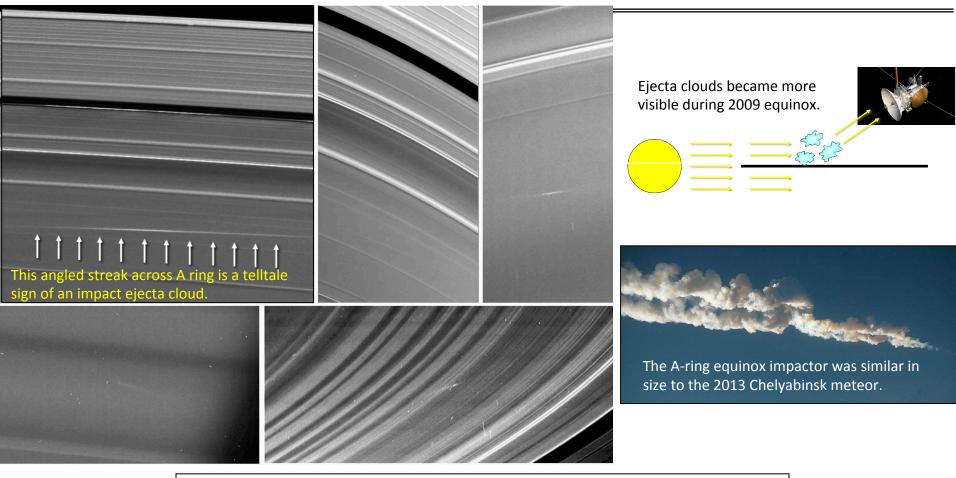
Cassini T92 ISS movie of Titan's south-polar vortex



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Meteor Strikes in Saturn's Rings



An orbiting cloud shears with time, as material closer to the planet orbits at a faster rate.

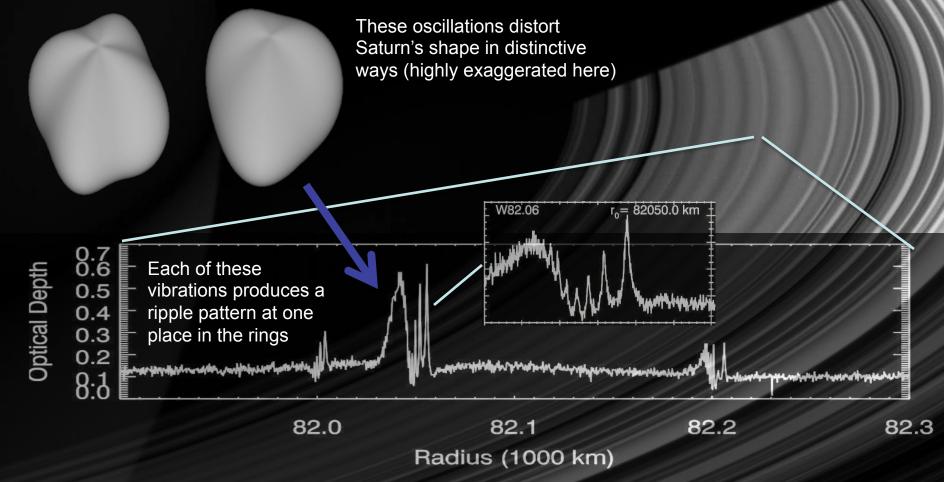






A Seismograph as Big as Saturn's Rings

Cassini scientists have discovered that Saturn's rings act as a seismograph that records large-scale oscillations, probably emanating from deep within the planet, that "ring" Saturn like a bell. In the same way that helioseismology tells us about activity inside the sun, "Kronoseismology" provides a completely new way to probe structure and activity in Saturn's interior.



* Kronoseismology: Using Density Waves in Saturn's C Ring to Prove the Planet's Interior, M.M. Hedman and P.D. Nicholson, CRSR, Cornell University Astrophysical Journal



Cassíní Solstice Míssíon



Eye Spied: Saturn's Behemoth Polar Hurricane

Stunning new views from NASA's Cassini spacecraft reveal the eye of an enormous hurricane locked in place at Saturn's north pole.





In this false color image, red indicates deep clouds, while green shows clouds¹that are higher in altitudes. The Sun is to the right in this image.



Cassini Solstice Mission

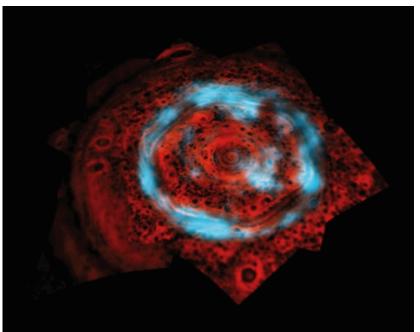


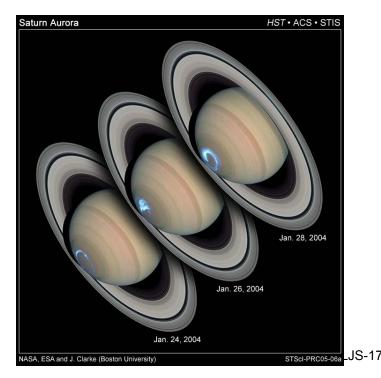
Saturn Auroral Campaign (April-May 2013)



Cassini auroral observations of Saturn's northern and southern aurora have been designed to coincide with Solar Maximum and with observations of Saturn's northern aurora by HST and ground-based IR

OPAG



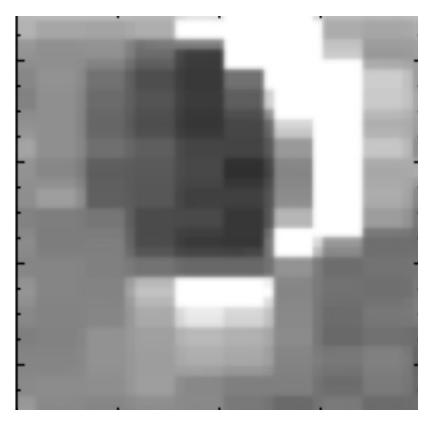


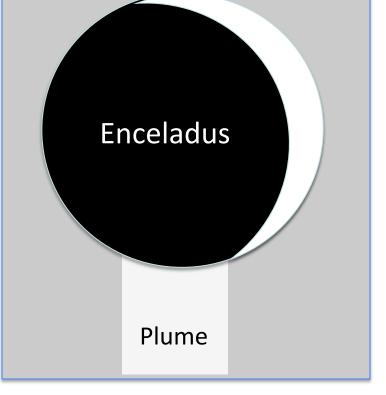


Cassini Solstice Mission



Variable Plume Activity on Enceladus





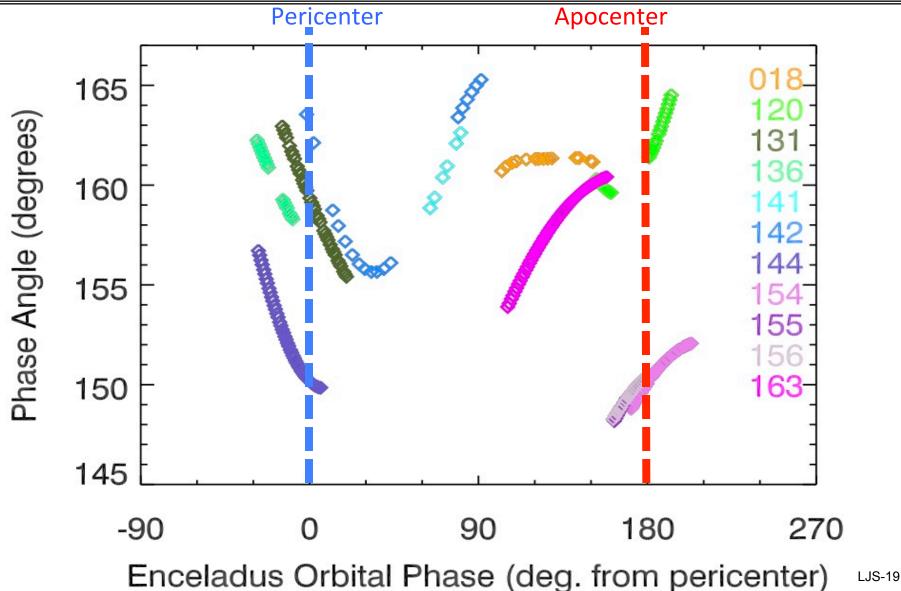
VIMS plume image (wavelength 0.9-1.1 microns) **Interpretive Drawing**



Cassini Solstice Mission



VIMS data cover a wide range of phase angles and orbital phases

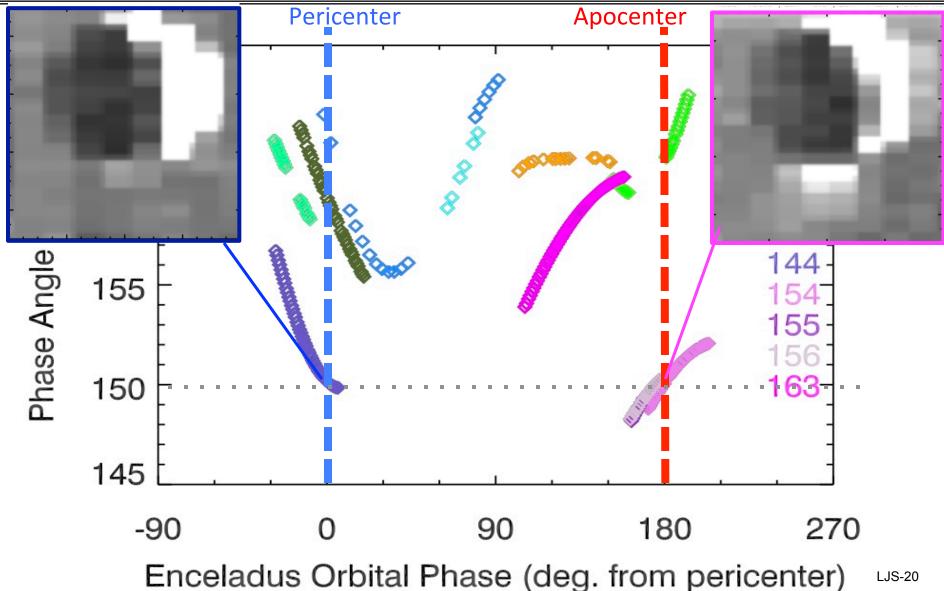


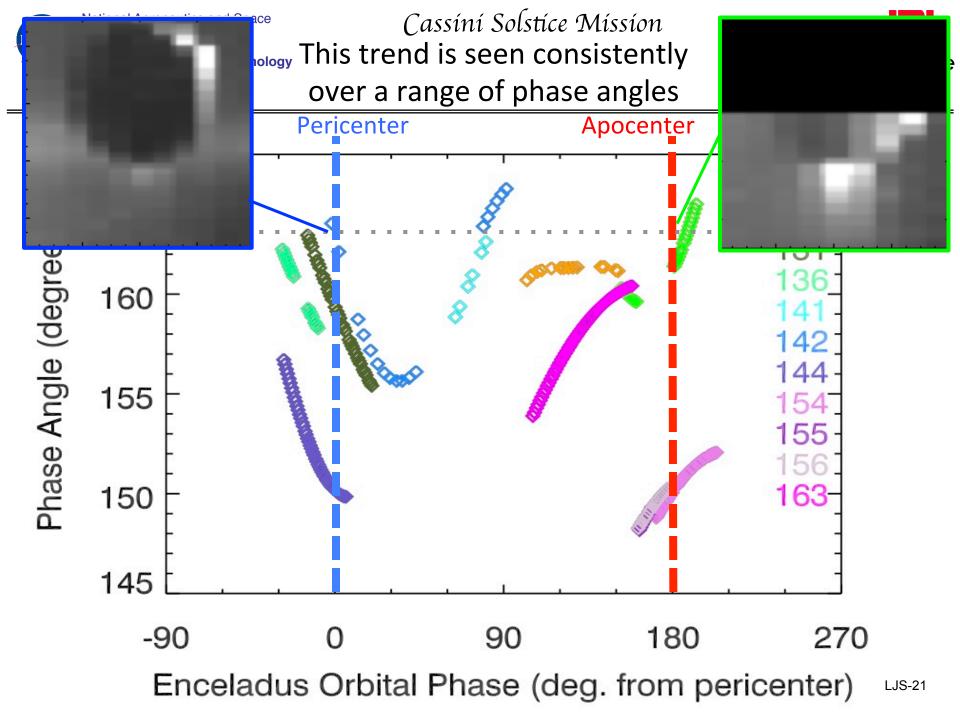


Cassini Solstice Mission



The plume is brighter when Enceladus is near its orbital apocenter



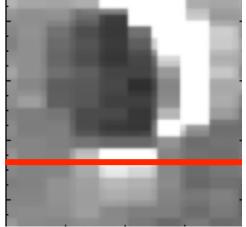




Cassini Solstice Mission



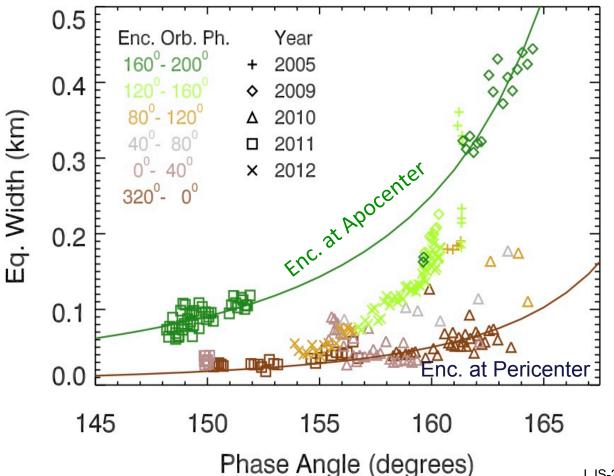
Enceladus Plume Variability



VIMS image Plume is consistently brighter when Enceladus is near apocenter of its orbit.

Hedman et al., Science

Wavelength=1.2 μ m, Altitude= 85 km (interpolated)



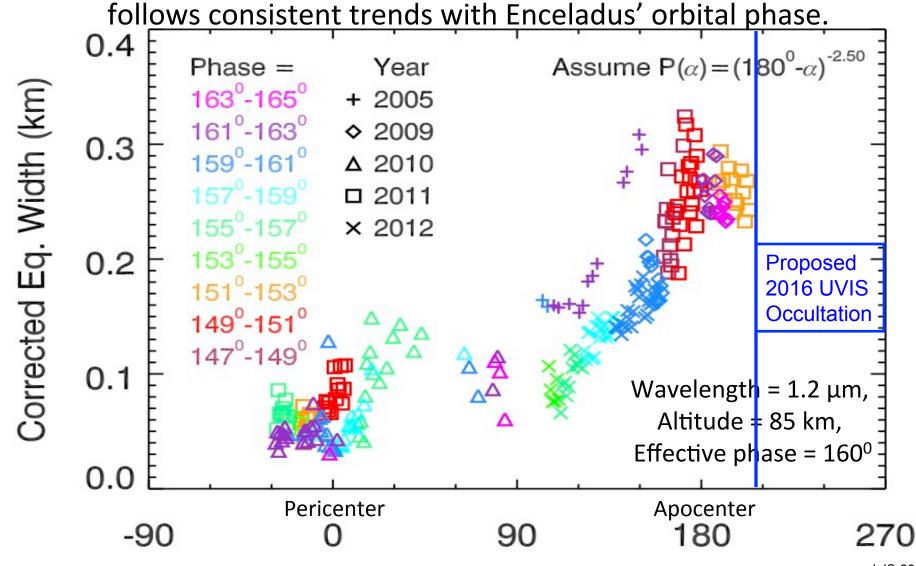
15 July 2013



Cassíní Solstice Mission



After correcting for phase angle variations, the plume's brightness



Enceladus Orbital Phase (deg. from pericenter)²³



Cassíní Solstice Míssíon



Consolidated Senior Review

- **Assume** similar format and schedule as past Planetary Science Division (PSD) Consolidated Senior Review (no guidelines released yet)
 - Senior review in 2014 (SR every 2 years) to provide best balanced science for scarce funding available
 - Missions: Cassini, LRO, MSL, MER, MEX, MRO, ODY
 - Science merits and performance will be evaluated
 - 35-page proposal to address FY15 FY16 extended mission
 - Cassini requesting addition of FY17 (final year of tour) to proposal as well
 - Two funding options: baseline and 85% option
- Cassini Proposal Submitted
- Questions from Panel to Project Offices
 - In addition, report on changes in Operations and Science
- Face to Face visit/oral presentation
- Senior Review Report submitted to PSD
- PSD Notification to Project Offices

Late June Mid-July FY15 start – 2 months

Late May

Mid-June





On July 19, 2013, at 5:27 pm EDT, look up and wave as NASA's Cassini spacecraft photographs Earth from Saturn. For more information see: http://saturn.jpl.nasa.gov/waveatsaturn

The View from Washington Saturn

On July 19, 2013, at 5:27 pm EDT, look up and wave as NASA's Cassini spacecraft photographs Earth from Saturn. For more information see: http://saturn.jpl.nasa.gov/waveatsaturn





Cassiní Solstice Mission



END



Cassíní Solstice Míssíon



Cassini Plasma Spectrometer (CAPS) Status

- CAPS powered off in June 2011 as a result of a high rail to chassis short
- CAPS turned on March 16, 2012 in time for E17 and E18 flybys
- CAPS functioned normally from March 16 to June 1, 2012
- CAPS solid state power switch (SSPS) tripped off on June 2, 2012
 - Multiple (at least 3) different hard and soft shorts occurred
- Extensive review conducted by NASA Engineering and Safety Center (NESC) team in conjunction with JPL, Cassini and CAPS teams
- NESC completed briefing to JPL and SwRI on April 26, 2013
- Final report TBD but expect no changes from briefing
- Summary of findings:
 - No root cause for shorts could be determined
 - If CAPS were to be be switched on again, the shorts will likely continue
 - High and low rail shorts are not benign to the power subsystem and represent a risk to the extended mission
- For the foreseeable future, CAPS will remain off

Baseline Allocation

Cassíní Solstíce Míssíon

- Cassini will be the only outer planet Flagship mission flying and collecting data through the mid- to late- portion of this decade
- Continued funding of Cassini will:
 - maintain the current scientifically rich tour
 - continue training of next generation of planetary scientists
 - continue healthy international collaboration, and
 - return multidisciplinary, synergistic science data as only a Flagship can do







Cassini Solstice Mission



Past SR: 17 Science Campaigns: 85% funding

Titan	Icy Satellites	Rings	Magneto- sphere	Saturn
(E.1.1) Titan Great Seas	(E.2.1) Enceladus	*(E.3.1) Changing Rings	(E.4.1) Periodicities	*(E.5.1) Aftermath of Giant Storm
(E.1.2) Titan Global Seasonal	(E.2.2) Rhea and Dione	(E.3.2) Composition, Origin, and Evolution of Rings	*(E.4.2) Aurora: Imaging the Magneto- sphere	(E.5.2) Seasonal and Temporal Changes
(E.1.3) Titan Interior Ocean	(E.2.3) Small Moons	(E.3.3) Rings, Protoplanets, and Exoplanets	(E.4.3) Magneto- spheric Interactions of Satellites	*(E.5.3) Polar Studies
Yellow: partially	narily accomplish accomplished risk of being pree	(E.4.4) Water Dominated Magneto- sphere	(E.5.4) Probe Saturn's Interior	

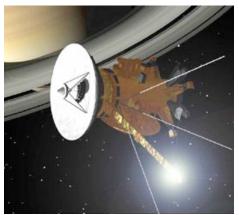


Cassíní Solstice Míssíon



Effects of 85% Allocation

- 85% allocation will result primarily in loss of jobs
 - Cassini spacecraft is built and launched, no major hardware purchases remain
 - Cassini is in orbit at Saturn, 7-year cruise is complete
- Loss of personnel would result in science loss and increased science risk in FY13-14



- Loss of operationally complex Ka-band science: Would reduce RSS operations and analysis costs
- Fewer people to implement science observations resulting in fewer and less complex science observations
- Slower response to instrument anomalies, would be best efforts basis
- Reduced calibration support, less well validated data to PDS
- Slower response time to new discoveries, additional opportunities might be missed
- Severely impact funding, educating and mentoring of the next generation of planetary scientists (team associates and postdocs)



Cassíní Solstice Míssíon



Cassini Solstice Science

- Cassini Solstice Mission enables unprecedented opportunities for unique, groundbreaking science
- Unique, compelling Juno-like end of mission science
- Direct relevance to the Planetary Decadal Survey and NASA's exploration program
- New Participating Scientist program actively involves broader science community

